AN ABSTRACT OF THE THESIS OF

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Title: <u>Petrogenesis of Compositionally Distinct Silicic Volcanoes in the Three Sisters Region of the</u> <u>Oregon Cascade Range: The Effects of Crustal Extension on the Development of Continental Arc</u> <u>Silicic Magmatism.</u>

Abstract Approve

Signature redacted for privacy.

Edward M. Taylor

The Three Sisters region of the Oregon High Cascades has developed three compositionally and petrogenetically distinct silicic (i.e., $SiO_2 \ge 58\%$) magma systems within the last 600 k.y. These silicic systems evolved from the same High Cascade mafic magma system and developed in the same 20x30 km area of the arc, but did not interact. The Broken Top system (BT) evolved to 71% SiO₂ through a combination of plag + px + Fe-Ti oxides ± ap (PPFA) fractionation and 20%-35% mixing of rhyolitic (74% SiO₂) crustal melts. In contrast, part of the Three Sisters system (3S) evolved to 66% SiO₂ through PPFA fractionation alone, while other parts evolved to 66% SiO₂ through PPFA fractionation coupled with $\ge 40\%$ mixing of rhyolitic ($\ge 72\%$ SiO₂) crustal melts.

The 3S system was intermittently active from ≤ 340 ka to 2 ka. The petrogenesis of intermediate composition rocks at Middle Sister (<340 ka, >100 ka) was controlled by PPFA fractionation to $\leq 66\%$ SiO₂. Rhyolite (72%-76% SiO₂) was first erupted in the 3S system at ≈ 100 ka, at the start of South Sister (SS) volcanism. Major and trace element abundances preclude derivation of 3S rhyolite through crystal fractionation, but are consistent with 20-30% dehydration melting of mafic amphibolite. The petrogenesis of intermediate composition rocks at SS was controlled by PPFA fractionation coupled with 30-40% rhyolitic magma mixing. However, the rhyolitic magma mixed into an essentially mafic system, which limited intermediate differentiation at SS to $\leq 66\%$ SiO₂.

The BT system was active from ≈ 600 ka to at least 200 ka. Major and trace element abundances preclude derivation of BT rhyolite (74% SiO₂) through crystal fractionation, but are consistent with $\approx 30\%$ dehydration melting of older tonalitic intrusions. BT petrogenesis was controlled by PPFA fractionation accompanied by 10-20% mixing of rhyolitic magmas to $\approx 63\%$ SiO₂, with $\approx 30\%$ rhyolite mixing from 63% to 71% SiO₂. In contrast to the 3S system, differentiation proceeded beyond 66% SiO₂ because rhyolitic magma was mixed into a more evolved ($\approx 60\%$ -65% SiO₂) system.

The observed temporal and spatial variations in petrogenesis were not controlled by regional changes in tectonic setting, crustal thickness or crustal composition. However, small-scale changes in the magnitude of crustal extension occurred in this area, and are thought to have controlled petrogenesis by localizing mid-crustal mafic magmatism and thus crustal heat flow.

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Petrogenesis of Compositionally Distinct Silicic Volcanoes in the Three Sisters Region of the Oregon Cascade Range: The Effects of Crustal Extension on the Development of Continental Arc Silicic Magmatism.

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PETROGENESIS OF COMPOSITIONALLY DISTINCT SILICIC VOLCANOES IN THE THREE SISTERS REGION OF THE OREGON CASCADE RANGE: THE EFFECTS OF CRUSTAL EXTENSION ON THE DEVELOPMENT OF CONTINENTAL ARC SILICIC MAGMATISM

1. INTRODUCTION

1.1 Overview of the Research Topic

This study was undertaken to examine the petrogenesis of silicic (i.e. $SiO_2 > 58\%$) rocks from the Three Sisters region of the Oregon Cascade Range, using modern analytical techniques to supplement detailed field studies of this geologically complex part of the High Cascades continental arc. Previous workers had shown that this area of the High Cascades was composed of at least 5 major volcanic centers, which had been active within the last ≈ 1 Ma and erupted a wide variety of volcanic rocks (e.g., Williams, 1944; Taylor, 1978, 1987; Hughes, 1983). Rhyolite was known to be unusually abundant in the Three Sisters area, compared to the rest of the Oregon High Cascades. Limited geochemical data from these studies indicated that there could be several distinct trends of silicic magma evolution, involving varying amounts of crystal fractionation and magma mixing. In addition, these trends occurred within a relatively small (600 km²) area of the Oregon Cascades.

Petrogenetic trends in continental arcs can be influenced by marked changes in the physical conditions of the arc. Arc-length variations in subduction rate (e.g., Carr et al., 1982; Nye and Reid, 1986) and subduction angle (e.g., Kay et al., 1980; Carr et al., 1990) have been shown to control mafic arc magmatism. Compositional heterogeneities in the upper mantle, which may be primary (e.g., Hughes, 1990) or the result of subduction-related metasomatism (e.g. Hawkesworth et al., 1984; Arculus and Powell, 1986) can produce compositionally distinct mafic magmas through partial melting. Changes in the thickness or composition of the crust under the arc (Tilton and Barreiro, 1980; Kay et al., 1980; Thorpe et al., 1982; Hildreth and Moorbath, 1988) can produce distinct assimilation trends in the mafic magma system and compositionally distinct silicic magmas through partial crustal melting. Given these factors, it is often unrealistic to construct an arclength petrogenetic model (e.g. Arculus and Johnson, 1978).

The Three Sisters area, however, is located in a relatively small area of the Oregon Cascades arc. It is extremely unlikely that any of the regional tectonic features listed above could control large variations in petrogenesis within a 20 km by 30 km segment of the Cascades arc. In addition, previous studies have shown that the mafic magma system in the Three Sisters area has the same degree of heterogeneity as in adjacent parts of the Oregon High Cascades, and that there are no spatial or temporal controls on mafic rocks within the Three Sisters area. If petrogenetically distinct silicic magma systems are present in the Three Sisters area, then these systems developed from and were influenced by the same mafic magma system. Regional mapping in central Oregon has shown that the Three Sisters region of the Oregon Cascades was located in a structurally complex area, where regional north-striking extensional features merged with northwest-striking features (Taylor, 1981; Macleod et al., 1982). The effect of crustal stress on igneous petrogenesis is poorly known. Empirical (Nakamura, 1977; Hildreth, 1981) and theoretical (Shaw, 1980) models generally conclude that extension favors the ascent of mafic magma in the crust, but these models have not been previously applied to continental arc petrogenesis. Although the observed structural variations apparently did not effect mafic petrogenesis in the Three Sisters area, it was hypothesized that changes in the orientation of crustal stress could have influenced silicic petrogenesis by controlling the flux of mafic magma into the lower(?) crust. Recent work by Hildreth and Moorbath (1988) has demonstrated that lower crustal processes, and especially crustal composition, can strongly affect the petrogenesis of mafic and silicic arc rocks. The Three Sisters area thus represents an ideal locality to study the effects of local crustal processes on calc-alkaline continental arc petrogenesis.

The role of crystal fractionation versus magma mixing in arc petrogenesis cannot be explained by a simple model. It has been repeatedly shown that both of these petrogenetic processes can operate in arc environments (e.g., Bowen, 1928; Larsen, 1938; Anderson, 1976; Gill, 1981; Sakuyama, 1984; and many others). Additional complexities introduced through magma recharge (O'Hara, 1977, 1980; Nielsen, 1988) and crustal assimilation (DePaolo, 1981) further obscure the geochemical signatures of fractionation and mixing. Recent advances in experimental petrology, and the availability of powerful desk-top computers, have allowed the development of complex petrogenetic models. These models are constrained by experimentally determined phase relationships (e.g., Grove et al., 1982) and trace element distribution coefficients (e.g., Irving, 1978; Nielsen, 1985), and allow for realistic simulations of a variety of petrogenetic hypotheses. Tests of crystal fractionation hypotheses are no longer limited to simple mass balance and Rayleigh fractionation constraints. Although these procedures still result in useful petrogenetic limits, any real test of the importance of crystal fractionation in arc petrogenesis must account for open system processes as well (i.e., recharge, eruption and assimilation). The models developed in this study constrain some of the effects of open system processes on the High Cascades magmatic system.

1.2 Location and Methods

The study area encompasses 600 km² of the Oregon central High Cascades, and includes the Three Sisters and Broken Top composite volcanoes (figure 1.1). In keeping with historic usage, the term Three Sisters area will be used to describe the area from \approx 5 km west of the Three Sisters peaks to \approx 15 km east of Broken Top volcano (figure 1.1). In addition, the terms Three Sister system and Broken Top system will be used extensively throughout this study. The



Figure 1.1: Location map of the Three Sisters region of the Oregon High Cascades. Major vents of the Three Sisters system: N: North Sister, M: Middle Sister, S: South Sister, O: Obsidian Cliffs. Major vents of the Broken Top system: BT: Broken Top, TMR: Tam MacArthur Rim, TLV: Todd Lake volcano, B: Mt. Bachelor. Prominent vents of the Tumalo volcanic center (Broken Top system): 3: Three Creek Bu., TH: Triangle Hill, E: Edison Bu. Three Sisters system refers to rocks associated with North, Middle and South Sister, including numerous silicic domes located within 5 km south and west of these composite volcanoes (figure 1.1). The Broken Top system refers to rocks associated with Broken Top, Tam MacArthur Rim, the Tumalo volcanic center (Hill, 1988), and Todd Lake volcano (Taylor, 1978)(figure 1.1).

Much of the study area is located in the Three Sisters Wilderness area of the Deschutes and Willamette National Forests; most parts of the area can be reached within a day's hike. Access to the rest of the area is provided by a myriad of logging roads. Field work was accomplished during the summers of 1986, 1987 and 1988, and consisted of selecting more than 200 samples for analysis, development of a stratigraphic section based on field relationships and published maps, and 1:62500 geologic mapping in the eastern Broken Top system. Sampling was guided by prior studies (Taylor, 1978, 1987, unpub. res.; Taylor et al., 1987; Wozniak, 1982; Clark, 1983; Hughes, 1983), and included all major silicic units within the study area.

A total of 127 High Cascade and 10 Deschutes Formation (Smith, 1986) rocks were selected for detailed geochemical analysis. Analytical methods are given in Appendix A. After completion of the geochemical analyses, 26 chemically-representative High Cascade units were selected for detailed mineralogical study with the O.S.U. Cameca SX-50 electron microprobe (Appendix A). Five K-Ar age determinations were also obtained to correlate and constrain the volcanic stratigraphy of the Three Sisters area (Chapter 3).

1.3 Summary of Previous Work

The Three Sisters area of the Oregon Cascade Range has been the subject of numerous reconnaissance geological studies since the late 1870's. A summary of early work, which focused on the geomorphology and glacial geology of this area, is presented in Hodge (1925). Mapping by Williams (1944, 1957) first characterized the major volcanos in the area, which were shown to be the products of discrete eruptions and not related to one "mega-volcano" (i.e., the Mt. Multnomah of Hodge, 1925). Williams also recognized that a large volume of pyroclastic material had been deposited east of the Three Sisters area prior to Quaternary volcanism. A 1:250000 reconnaissance map of the geology of Deschutes County by Peterson et al., 1976, included most of the study area. Detailed 1:62500 maps of the Three Sisters Wilderness area (Taylor et al., 1987), Mt. Bachelor area (Scott and Gardner, 1991), and the Broken Top area (Taylor, 1978,1987) cover much of the study area and were used extensively for sample selection.

There have been several petrologic studies on parts of the Three Sisters area. Anttonen (1972) characterized the major and trace element abundances of numerous rocks from the Three Sisters area, and modeled most of the variations through different amounts of partial melting in the descending oceanic plate. Taylor (1978,1987) mapped and analyzed rocks from the northwest and southwest parts of the Broken Top 15' quadrangle. Wozniak (1982) mapped the northern part

of South Sister and produced major element analyses for most map units. Clark's 1983 study of South Sister focused on Holocene activity, although older units were mapped and analyzed. Clark recognized the existence of a significant compositional gap at South Sister, which he concluded was produced through remobilization and mixing of previously evolved rhyolitic magmas into a fractionating magma system. Hughes (1983, 1990) and Hughes and Taylor (1986) have primarily examined the petrogenesis of 0-5 Ma mafic rocks in the central High Cascades. These studies conclude that much of the geochemical variations in the central Cascade mafic system is related to heterogeneities in the degree of mantle metasomatism or melting, and that mafic rocks erupted prior to \approx 5 Ma are at times petrogenetically distinct from <0.7 Ma rocks. The pyroclastic deposits around Bend, Oregon, have been examined in detail by Peterson et al. (1976), Taylor (1981), Mimura (1984), Hill (1985, 1987, 1988) and Hill and Taylor (1989). These studies generally conclude that these pyroclastic rocks are the products of 4 major eruptions from the High Cascades west of Bend.

2. REGIONAL GEOLOGY OF THE CENTRAL HIGH CASCADES AREA

2.1 Crustal Composition

The best known exposures of pre-Tertiary crust are located over 100 km northeast of the Three Sisters area, and are primarily composed of greenschist-to-blueschist facies metavolcanics and metasediments (Swanson, 1969). A lawsonite-bearing blueschist from this terrane has an 40 Ar- 39 Ar date of 223 ± 3 m.y., which is also the age of blueschist facies metavolcanic rocks in the east-central Klamath Mountains (Hotz et al., 1977). The occurrence in these two areas of blueschist facies rocks of the same age, with common structural trends and associated melange zones, supports continuation of Klamath Mountain-type rocks into central Oregon under the Tertiary volcanic cover (Hotz et al., 1977; Hietanen, 1981).

These greenschist facies rocks are likely part of a large, late Paleozoic-to-Jurassic suture system that extends from the northwestern Sierra Nevada through the Klamath Mountains, and which may possibly continue through Oregon and into the Idaho Batholith (Hietanen, 1981). Exposures in the Klamath Mountains of southwestern Oregon show that terranes associated with this suture zone are dominantly accreted island arc sequences of late Paleozoic to early Mesozoic age (e.g., Irwin, 1981).

A \approx 50 km wide, regional gravity anomaly (figure 2.1) is coincident with the apparent northern boundary of the Klamath Mountain-central Oregon suture zone (Thiruvathukal et al., 1970; Riddihough et al., 1986). The gravity anomaly, referred to as the Klamath-Blue Mountain Lineament (KBL), marks a change in crustal density from relatively lower density, older(?) crust to the southeast and higher density, younger(?) crust to the northwest (Thiruvathukal et al., 1970; Riddihough et al., 1986). The KBL may reflect changes in crustal thickness due to Basin & Range extension (Thiruvathukal et al., 1970), but regional seismic studies have failed to detect any significant changes in crustal thickness where the KBL crosses the High Cascades (Leaver et al., 1984). High heat flow associated with High Cascade volcanism (Blackwell et al., 1982, 1990) obscures the exact location of the KBL in the Oregon Cascades, but a linear extrapolation of the KBL intersects the High Cascade axis in the Three Sisters area (figure 2.1).

Limited isotopic data support the argument that the KBL marks a change in crustal composition. Although the Oregon Cascades are characterized by relatively primitive lead isotopic ratios (Church and Tilton, 1973; Church et al., 1986; Conrey, 1991), lead isotopes from sulfide minerals in the western Cascades (Church. 1986) are slightly less radiogenic north of the KBL than to the south. Strontium isotopic ratios for Quaternary rhyolitic rocks, which were likely produced through partial crustal melts, are lower north of the KBL at Mt. Jefferson (≈ 0.7033 ; Conrey, 1991) than south of the KBL at the Three Sisters (≈ 0.7036 ; this study) and Mt. Mazama (≈ 0.7037 ; Bacon



Figure 2.1: Regional Bouger gravity anomaly map of the Pacific Northwest, reproduced from Riddihough et al., 1986. Arrows indicate the Klamath-Blue Mountain gravity lineament (KBL) referred to in text. Small crosses at 122°W, 44°N represent North Sister, South Sister, and Broken Top volcanos. The KBL trends across the Oregon High Cascades in the Three Sisters area, although this relationship is obscured by the high heat flow (and thus a more negative Bouger anomaly) associated with High Cascade volcanism. It is possible that the Three Sisters area is located on, or very near to, the accretionary contact between older, Klamath Mtn.- type metamorphic rocks and younger meta-oceanic crust (Riddihough et al., 1986).

and Druitt, 1988). These isotopic data may indicate that slightly older, more isotopically evolved crust is located south of the KBL.

Metamorphic or older plutonic xenoliths have not been found in any rocks from the Three Sisters region, and thus the composition of the middle (\approx 15-30 km) to lower (\approx 30-45) crust under this part of the High Cascades remains unknown. Based on the above arguments, middle crustal lithologies under the Three Sisters area most likely consist of late Paleozoic accreted arc terranes. These terranes should be dominated by amphibolite-grade mafic metamorphic rocks, with relatively minor amounts of metamorphosed sedimentary and silicic igneous rocks.

2.2 Deep Crustal Structure

Magnetotelluric (Livelybrooks et al., 1989; Stanley et al., 1990) and seismic refraction (Leaver et al., 1984) studies have concluded that the crust beneath the central High Cascades consists of essentially four sub-horizontal layers (figure 2.2). Layer 1 is composed of ≤ 2 km of high resistivity Quaternary volcanic flows, which overlie 1-5 km of more conductive Tertiary volcanic rocks. Layer 2 extends to 11-12 km, and is composed of high resistivity, pre-Tertiary metamorphic rocks. The base of layer 2 is marked by a transition from brittle to ductile deformation, and represents the base of the upper crust (Stanley et al., 1990). Layer 3 (the middle crust) is highly conductive and is thought to represent amphibolite-facies rocks with small amounts of metamorphic fluid or partial melt (Stanley et al., 1990). Layer 3 extends to ≈ 30 km and is thought to be at temperatures of 500-600 °C (Stanley et al., 1990; Blackwell et al., 1990). Layer 4 is only resolved seismically (Leaver et al., 1984), and extends from ≈ 30 km to a depth of 44 km along the length of the Oregon High Cascades. The deep crust (i.e., layer 4) is thought to consist of mafic granulite at temperatures of ≤ 1100 °C, and mafic magma and intrusions (Stanley et al., 1990).

2.3 Regional tectonics

Most researchers in the last decade agree that the Oregon Cascade arc is the result of volcanism associated with subduction of the Juan de Fuca Plate (JDF) beneath the North American Plate (e.g., Priest et al., 1983; Taylor, 1990). Subduction of the JDF plate has likely occurred at a rate of ≈ 3 cm/yr since ≈ 5 Ma, although subduction rates as high as ≈ 8 cm/yr may have occurred between 5-35 Ma (Riddihough, 1984; Engebretson et al., 1986; Verplank and Duncan, 1987). The convergence direction between the JDF and North American plates has remained at about N50E since ≈ 10 Ma (Verplank and Duncan, 1987). Estimates for the dip of the subducting JDF plate beneath the Oregon Cascades are generally $65^\circ \pm 5^\circ$ (Rasmussen and Humphries, 1988; Keach et al., 1989). Variations in subduction rate, convergence angle and possibly subduction angle with time are thought to have controlled many of the features of Cascade volcanism throughout the last 35 Ma (Priest et al., 1983; Verplank and Duncan, 1987; Taylor, 1990;



Figure 2.2: Schematic model of crustal structure in the Oregon High Cascades, based on magnetotelluric, seismic, heat flow, and petrologic data, slightly modified from Stanley et al., 1990. Mafic magmatism is focused in an essentially anhydrous deep (≥30 km) crust. Magma segregation to middle crustal levels (15-30 km) permits differentiation to silicic compositions, and rhyolitic dehydration melts of amphibolitic crustal rocks. Magmas may pond at the brittle-ductile transition at ≈12 km (Stanley et al., 1990), and subsequently erupt through tectonically controlled magma conduits.

Priest, 1990). There is no evidence that any of these tectonic parameters have changed since 1 Ma, or would have resolvable variations along a 20 km segment of the High Cascades arc (i.e., within the Three Sisters area).

2.4 Regional Structure

The Oregon High Cascades are located in an area of active north-south compression and east-west horizontal extension of the crust, as shown (figure 2.3) by north-striking normal faults and vent alignments in this area (Nakamura, 1977; Zoback and Zoback, 1980). The orientation of regional horizontal extension has remained constant since ≈ 5 Ma, but was likely oriented to the northeast prior to ≈ 7 Ma (Sherrod and Pickthorn, 1989; Priest, 1990). Changes in the orientation (and probably the magnitude) of crustal extension in the Oregon Cascades have been linked to changes in the convergence angles between the Pacific and North American Plates (Atwater, 1970; Zoback and Zoback, 1980; Rogers, 1985; Priest, 1990), and to impingement of Basin and Range extension into the Cascades (Lawrence, 1976; Guffanti and Weaver, 1988; Wells, 1990).

The Brothers Fault Zone (BFZ) is expressed as a ≈ 200 km long zone of en-echelon normal faults that strike N40W to N50W (figure 2.3), and are thought to be the result of sub-surface dextral shear associated with Basin and Range extension (Lawrence, 1976). The BFZ may also represent a zone of transtensional shear associated with oblique subduction of the Juan de Fuca Plate (Gutmanis, 1989; Pezzopane and Weldon, 1990). Although the BFZ strikes towards the central part of the Oregon High Cascades, the BFZ is overlain by ≤ 0.7 to ≈ 3 Ma volcanic rocks near Newberry volcano (Peterson et al., 1976; MacLeod et al., 1982; Hawkins et al., 1988). The Tumalo Fault Zone (TFZ) (Peterson et al., 1976; Taylor, 1981) is a series of $\approx N20W$ -striking normal faults that offsets units as young as 0.3-0.4 Ma (Peterson et al., 1976; Taylor, 1981; Sarna-Wojcicki et al., 1987). The TFZ offsets Newberry lavas, and strikes into the northwest fissure zone of Newberry volcano (Peterson et al., 1976; MacLeod et al., 1982; MacLeod and Sherrod, 1988). The TFZ does not merge with the BFZ, as has been occasionally reported (Lawrence, 1976; Gutfanti and Weaver, 1988; Gutmanis, 1989), and has a markedly different orientation and age of deformation than the BFZ.

North-striking normal faulting at 5.4 to \approx 4 Ma near the latitude of Mt. Jefferson (figure 2.3) resulted in the formation of a discontinuous intra-arc graben (Taylor, 1981; Smith, 1986). North-striking normal faults, which are thought to represent the western boundary of this graben, are exposed between Mt. Jefferson and the Three Sisters (figure 2.3). These faults mark the contact between Quaternary late High Cascade lavas on the east, and older (4-7 Ma) early High Cascade lavas on the west (Priest et al., 1987, 1988; Black et al., 1987). However, the contact between early and late High Cascade units south of the Three Sisters area is generally depositional (Sherrod, 1986). The eastern boundary of the graben is well exposed at Green Ridge (figure 2.3),



Figure 2.3: Faults and vent alignments of central Oregon, modified from MacLeod and Sherrod, 1988. Approximate boundaries of the regional gravity gradient from Thiruvathukal et al., 1970. FZ = Fault Zone. The strike of vent alignments changes from north to northwest in the Broken Top area. Northwest-striking vent alignments are parallel with the strike of the Tumalo FZ, which continues south to Newberry volcano and eventually merges with the Walker Rim FZ.

where at least 1 km of stratigraphic offset has occurred since 5.4 Ma (Conrey, 1985; Smith, 1986; Hill et al., 1991). Green Ridge faults apparently merge with the northern part of the Tumalo Fault Zone (Peterson et al., 1976; Conrey, 1985), suggesting that initial motion along the TFZ occurred during graben formation (Smith, 1986). However, numerous graben-forming faults had continued activity to 2.4-1.7 Ma (Priest et al., 1988; Priest, 1990; Taylor, 1990), suggesting that the TFZ may be primarily a Quaternary tectonic feature. Although there is no concrete evidence that the High Cascade graben is continuous (i.e., Allen, 1966; cf. Smith, 1986), north-striking normal faults may occur beneath the Quaternary volcanic cover in the subsurface west of the TFZ (Smith, 1986; Livelybrooks et al., 1989; Taylor, 1990).

2.5 A Brief History of the Oregon Cascades

The Oregon Cascades have been a locus of subduction-related calc-alkaline magmatism since ≈ 37 Ma (McBirney et al., 1974; White and McBirney, 1978; Taylor, 1981, 1990; Priest et al., 1983; Priest, 1990). Following the nomenclature of Priest et al., (1983), units older than 7.4 Ma are generally associated with the Western Cascades (i.e., Callagan, 1933; Peck et al., 1964). Units from 7.4 to 4 Ma are referred to as early High Cascades, with younger units belonging to the late High Cascades. Although it is likely that the position, width, and petrogenesis of the Cascade volcanic arc have changed considerably from ≈ 37 Ma to ≈ 7 Ma (e.g., Priest, 1990; Wells, 1990), much of the early geological history of the Western Cascades remains unstudied. Summaries of Western Cascade petrogenesis are presented in Taylor (1990), Priest et al., (1983), and Priest (1990).

Stratigraphic records of early High Cascade volcanism are preserved ≈ 20 km west and east of the late High Cascade volcanic axis. Lavas and volcaniclastic sediments of the Deschutes Formation (Smith, 1986) were deposited from 7.4 to 5.4 Ma in a broad basin located $\approx 10(?)$ to ≈ 50 km east of the early High Cascade volcanic axis. Mafic flows and intrusions, which are temporally correlative with the Deschutes Formation, are also exposed from ≈ 20 to ≈ 40 km west of the present volcanic axis (Sherrod, 1986; Black et al., 1987; Priest et al., 1987, 1988). Based on major element and limited trace element analysis, early High Cascade mafic rocks appear to be petrogenetically distinct from late High Cascade mafic rocks (Hughes, 1983, 1990; Conrey, 1985; Smith, 1986; Hughes and Taylor, 1986; Conrey and Sherrod, 1988). Early High Cascade mafic rocks are commonly diktytaxitic, and probably evolved from a mantle source that was more depleted in incompatible elements than the source of the late High Cascades (Hughes, 1990).

The transition from early to late High Cascade volcanism is marked by formation of the High Cascade graben, and uplift of the Western Cascades (Taylor, 1981, 1990; Priest, 1990). Graben formation was accompanied by the syntectonic(?) eruption of anomalously large amounts of relatively nonevolved high aluminum olivine tholeiite, possibly in response to increased amounts

of crustal extension (Taylor, 1981; Hughes and Taylor, 1986; Hughes, 1990). The dominance of basalt gradually yielded to basaltic andesite by about 2 Ma (Hughes, 1990; Conrey and Sherrod, 1988). Volcanism since ≈ 2 Ma has been characterized by the eruption of mafic rocks that are more incompatible element-enriched than those associated with either the early High Cascades or graben formation (Hughes, 1983, 1990; Hughes and Taylor, 1986). Most of the present Oregon central High Cascade volcanos are younger than ≈ 0.7 Ma, and have been erupted onto the postgraben mafic platform (White and McBirney, 1978; Taylor, 1981, 1987, 1990). However, Mt. Hood (Wise, 1969; Priest and Vogt, 1982) and Mt. Jefferson (Conrey, 1991) overlie silicic rocks that are penecontemporaneous with the mafic platform.

Late High Cascade silicic volcanism has been restricted to large stratovolcanic centers, although isolated silicic domes have been occasionally erupted >20 km from these centers (e.g., Hayrick Butte, Benchmark Butte). The major silicic vents in the Oregon High Cascades are spaced ≈ 100 km apart, which is similar to the spacing of the larger silicic intrusions in the Western Cascades (Peck et al., 1964). This spacing may reflect the thermal limits of maintaining hot magmatic conduits through continental crust of this thickness and composition (cf., Marsh, 1984).

3. GEOLOGY OF THE THREE SISTERS AREA

3.1 Introduction

The geology of the Three Sisters area reflects the complexities associated with the eruption of at least 7 major silicic vent complexes, and numerous monogenetic vents, within a 600 km² area. Previously reported stratigraphic relationships (Taylor, 1978, 1981, 1987; Wozniak, 1982; Clark, 1983) have been constrained with new K-Ar age determinations in this study. These relationships are summarized in figure 3.1, and will be expanded upon below.

This section also describes the geology and age of a poorly-documented section of the Deschutes Formation, informally known as the Bull Springs Inlier (Taylor, 1978; Smith, 1986), which occurs ≈ 5 km east of Triangle Hill. In addition, new data on the petrology and source of the voluminous pyroclastic deposits that occur in the Bend area, and previously unreported strontium isotopic analyses for rocks of the Three Sisters area, will be presented in this section.

3.2 Deschutes Formation

The Deschutes Formation is an assemblage of lavas, pyroclastic deposits and volcanogenic sediments derived from 7.4 to ≈ 4 Ma early High Cascade volcanism in the Mt. Jefferson-Three Sisters area (Hales, 1975; Taylor, 1978, 1981; Conrey, 1985; Smith, 1986; Smith et al., 1987). The Deschutes Fm. is best exposed ≈ 30 km northeast of the Three Sisters area in the Deschutes Basin (Smith, 1986), where it forms a ≈ 250 m thick section dominated by volcanogenic sediments and pyroclastic deposits. Deschutes Fm. mafic lavas are exposed as far south as Bend (Smith, 1986; McDannel, 1989), and it is likely that many of the mafic lavas and vents ≤ 10 km east of the Tumalo Fault Zone are of Deschutes Formation age (Taylor, 1981; Smith, 1986; McDannel, 1989).

Although most Deschutes Fm. igneous units were erupted from the area around Mt. Jefferson (Hales, 1975; Conrey, 1985; Smith, 1986), numerous Deschutes Fm. ignimbrites were erupted from the Three Sisters area (Cannon, 1985; Smith, 1986; McDannel, 1989). Deschutes Fm. ignimbrites range in composition from andesite to rhyodacite, and commonly contain banded pumice (Smith, 1986). Several small exposures of Deschutes Fm. ignimbrite are preserved in the Bend area, and include a poorly welded, banded rhyodacite-andesite vitric lapilli tuff near the confluence of Tumalo Creek and the Deschutes River (samples 3S088M,S), a non-welded rhyodacitic tuff on the northwest margin of Awbrey Butte (3S080), and a highly devitrified, moderately welded tuff near Tumalo Dam (3S077). Samples of the Buckhorn Canyon (3S141) and Deep Canyon (3S142) ignimbrites, which were erupted from the Three Sisters area, were also analyzed.

Analyzed Deschutes Fm. rhyodacitic ignimbrites are compositionally similar to rhyodacites from the Broken Top system (figure 3.2). Broken Top rhyodacites were produced though the



Figure 3.1: Generalized stratigraphy of the Three Sisters area, constrained with new K-Ar dates from this study. Stratigraphic positions are relative, and do not represent absolute ages. Size of K-Ar date symbol represents range of 1σ error. Positions of the Tumalo Tuff and Desert Spring Tuff constrained by data in Sarna-Wojcicki et al., (1987, 1989). K-Ar data presented in Table 3.1.

mixing of rhyolitic crustal melts into a fractionating intermediate composition magma system (Chapter 8). It is likely that a similar magma system existed in the Three Sisters area at ≈ 5 Ma, and produced these Deschutes Fm. rhyodacite ignimbrites. Although rhyolitic crustal melts were produced in this area at ≈ 5 Ma, these melts did not coalesce into a discrete rhyolitic magma chamber. Rhyolitic volcanism in the Broken Top system at ≈ 0.4 Ma thus reflects a set of geologic conditions that were unique to this particular time.

3.2.1 The Bull Springs Inlier of the Deschutes Formation

A noteworthy outcrop of Deschutes Fm. lavas, informally known as the Bull Springs Inlier (Taylor, 1981; Smith, 1986; Hill, 1988a), occurs ≈ 5 km east of Triangle Hill (figure 3.3). This ≈ 4 km² outcrop of basaltic andesite-to-dacite lavas occurs ≈ 5 km west of the Tumalo Fault Zone, and is important to understanding the structural history of this part of the Oregon Cascades.

The Bull Springs Inlier (BSI) is primarily composed of a series of andesitic to dacitic lavas, which have normal paleomagnetic directions and are nearly horizontal (figure 3.3). These lavas are capped by an aphyric basaltic andesite lava, which has a reversed paleomagnetic direction. This basaltic andesite lava has a K-Ar date of 4.7 ± 0.1 Ma, which corresponds to an upper Deschutes Fm. age (table 3.1). This date is also within a 4.47-4.85 Ma reverse polarity event in the Gilbert polarity epoch (Mankinen and Dalrylmple, 1979); normal polarity rocks underneath this unit are thus older than 4.85 Ma. Quaternary ignimbrites and mafic lavas (≤ 0.7 Ma) are also deposited around, and on, the BSI (figure 3.3). The distribution of Quaternary units on the western margin of the BSI suggests that the western face of the BSI is a pre-Quaternary fault scarp.

The BSI represents an isolated flank of a Deschutes Fm. volcanic center, with the main vent complex presumedly located further to the west. Normal faulting on the western margin isolated the BSI in a manner analogous to the formation of Green Ridge, and resulted in the formation of a basin west of the BSI. Offsets in shallow (≤ 5 km) crustal conductors observed ≈ 7 km northwest of the BSI (Livelybrooks et al., 1989) are consistent with this hypothesis. In addition, rotary drilling by UNOCAL Geothermal 6.5 km north-northwest of the BSI (Hole #79-9, T16S R10E, SWNE27) encountered >100 m of volcanic sediments underlying ≈ 100 m of Quaternary lavas and ignimbrites, further indicating that the area west to northwest of the BSI was a basin.

Formation of the High Cascade graben in the Green Ridge area began at ≈ 5.4 Ma, and likely continued until ≈ 5.3 Ma (Smith, 1986; Smith et al., 1987). If the BSI was created through the continuation of normal faulting south of Green Ridge, then this faulting did not reach the BSI area until <4.7 Ma. It is also unlikely that the >1 km of vertical displacement associated with the formation of Green Ridge (Conrey, 1985; Hill et al., 1991) was accommodated entirely by the Tumalo Fault Zone east of the BSI (Smith, 1986), or else the BSI would not have been preserved at the surface unless it had a pre-faulting elevation >2280 m ($\approx 7500^{\circ}$). These relationships



Figure 3.2: Chondrite-normalized (Anders and Ebihara, 1982) rare earth element diagram for Deschutes Formation (DF) rhyodacitic ignimbrites analyzed in this study. Broken Top (BT) rhyodacite 3S098 included for comparison. The similarity between DF and BT rhyodacites suggests that these units had a similar petrogenesis; rhyolitic magma was available for mixing during DF magmatism, but did not accumulate and form discrete rhyolitic magma chambers.





Base: NW 1/4, Shevlin Park 7.5' quadrangle Deschutes County, Oregon.

Contour interval = 100'



Geology of the Bull Springs Inlier

B. Hill, 1986, 1987

Description of Map Units Qal: Quaternary alluvium Qc: Colluvium Qba: Quaternary High Cascade basaltic andesite lavas. Qsp: Shevlin Park Tuff Qds: Desert Springs Tuff Tba: Pliocene basaltic andesite Ta: Pliocene-Miocene andesite Td: Pliocene-Miocene dacite

> Contact, located ±50' Contact, located ±200'

Numbers refer to analyzed samples (3S00-).

| Series ¹ | Sample # | Sample wt (g) | K (wt %) ² | ⁴⁰ Ar _{rad} X10 ⁻¹² mol/g | ⁴⁰ Ar _{rad} (%) | Date (± 1σ) | Comments |
|---------------------|--------------------|------------------|--------------------------|---|--|----------------|--|
| DF | 3S002 | 4.223 | 0.92 | 7.462 | 45.7 | 4.7 ± 0.1 Ma | Top of the Bull Springs Inlier, p-mag reversed |
| BT³ | 3S015 ³ | n/a | 0.984 | n/a | 6.6 | 0.63 ± 0.09 Ma | Basaltic andesite cone, 1.5 km west of Triangle Hill |
| ЗT | 3S067 | 3.034 | 1.79 | 1.418 | 17.2 | 460 ± 30 ka | Todd Lake volcano, fresh dacite flow. |
| ЗТ | 3S010 | 4.115 | 1.05 | 0.619 | 13.8 | 340 ± 20 ka | Triangle Hill, non-vesiculated andesite bomb |
| ЗТ | 3S046 | 2.499 | 2.33 | 0.859 | 13.8 | 213 ± 9 ka | Tam MacArthur Rim, glassy rhyodacite flow |
| S | 3S159 | 4.021 | 0.59 | 0.373 | 1.8 | 0.4 ± 0.2 Ma | Middle Sister, vesicular porphyritic basalt |
| S | 3S139 | 2.524 | 2.77 | 0.455 | 6.4 | 95 ± 10 ka | Obsidian Cliffs, fresh rhvolite obsidian |
| S | 3S162 | 3.027 | 1.65 | 0.267 | 7.7 | 93 ± 11 ka | South Sister, glassy andesite flow 3S162. |

Notes:

'Series: DF = Deschutes Fm.; BT = Broken Top system; 3S = Three Sisters system.

²K determined by Atomic Absorption Spectrometry at the University of Oregon, C. McBirney, analyst.

³Analysis at the University of Arizona Isotope Geochemistry Lab, 4/88, on feldspar concentrate from 3S015. Analysis courtesy of G. Priest, Or. Dept. Geol. Min. Ind., 1988. n/a = data not reported.

- Isotopic ratios determined at Oregon State University, R. Duncan, principal investigator.

- K-Ar ages calculated using the constants of Steiger and Jager (1977): $\lambda_{e} = 0.581 \times 10^{-10} \text{ yr}^{-1}$, $\lambda_{B} = 4.962 \times 10^{-10} \text{ yr}^{-1}$, and ${}^{40}\text{K/K}_{\text{total}} = 1.167 \times 10^{-4} \text{ mol/mol}$.

Table 3.1: Whole rock K-Ar age determinations for rocks from the Three Sisters area of the Oregon High Cascades. The date of Middle Sister basalt 3S159 is probably at the younger limits of 1σ error, based on stratigraphic relationships; empirical analysis of reported High Cascade K-Ar dates indicates that analyses with <2% radiogenic Ar are generally inaccurate as well as imprecise. Other dates correspond to the stratigraphic positions

indicate that normal faults of significant vertical displacement are probably buried under Quaternary units between the BSI and Triangle Hill (figure 3.3). These faults probably strike sub-parallel to the Tumalo Fault Zone; the BSI is located ≈ 10 km eastward of the projected strike of the Green Ridge faults, and pre-Quaternary rocks are not exposed in the area between the BSI and Triangle Hill.

3.3 Tumalo Volcanic Center

One of the more prominent topographic features of the Three Sisters area is a broad highland that extends east of Broken Top volcano (figure 1.1). This area was referred to as the "silicic highland" by Taylor (1978, 1981), but a more regionally descriptive name was needed to describe this highland and associated vents. The name "Tumalo volcanic center" was selected by Hill and Taylor (1988b) to describe both the "silicic highland", and a series of cogenetic andesitic cinder cones and rhyolitic domes that extend from Melvin Butte ≈ 25 km south to Edison Butte (figure 3.4).

Most of the TVC is covered by younger mafic lavas erupted from the Tam MacArthur Rim-Broken Top area, and by glacial deposits (Taylor, 1978, 1981). The most prominent feature of the TVC is a roughly circular arrangement of mafic cinder cones and silicic domes, centered on Triangle Hill (figure 3.4). This ≈3 km diameter feature is also coincident with a prominent negative Bouger gravity anomaly of 10 mgal (Couch et al., 1982), indicating that low density material underlies the Triangle Hill vent complex. Shallow crustal xenoliths form $\approx 1\%$ of the ejecta of Triangle Hill, and range in size from <1 mm to ≈ 1 m in diameter. The most abundant xenoliths consist of fresh to hydrothermally(?) altered (mixed layer clays ± zeolite(?)) silicic lavas, which commonly do not resemble rocks exposed at the surface. Xenoliths of rhyolitic vent breccia are also common, as are mafic to intermediate lavas. Xenoliths of an intensely welded, crystal-rich silicic ignimbrite have also been erupted at Triangle Hill. These ignimbrite xenoliths are petrographically identical to samples of an intensely welded ignimbrite exposed ≈ 8 km south in the valley of Tumalo Creek (figure 3.4). This ignimbrite is probably associated with early volcanic activity at the TVC, and projects horizontally to ≈300 m beneath Triangle Hill. The abundance and composition of silicic xenoliths, combined with the negative gravity anomaly, indicates that the Triangle Hill vent complex overlies a major silicic vent.

All rocks exposed in the TVC area have normal paleomagnetic directions, and are likely younger than 720 ka. Based on limited stratigraphic information and the relatively steep slope of the cone, Triangle Hill appears to be the youngest vent in the Triangle Hill complex. Triangle Hill has a K-Ar date of 340 ± 20 ka, which probably represents the age of the last eruption of the TVC (Table 3.1). An additional K-Ar date of 0.63 ± 0.09 Ma was obtained (G. Priest, written comm.,

Figure 3.4: Geologic map of the Tumalo volcanic center and associated pyroclastic deposits. Data from Taylor (1981, 1987; unpub. res. 1974-1981), Hill (1985), Scott and Gardner (1991), and this study. Geology of the Bull Springs Inlier (BSI) shown in detail in figure BSI. Tumalo Tuff incorporates exposures of the Bend Pumice, Tumalo Tuff, and Lava Island Tuff (Taylor, 1981).

Location symbols: MB: Melvin Butte; TH: Triangle Hill; BB: Bearwallow Butte; BSI: Bull Springs inlier; TMR: Tam MacArthur Rim; BT: Broken Top; TCk: Tumalo Creek valley; EB: Edison Butte; B: Mt. Bachelor.



Figure 3.4: Geologic map of the TVC.

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1988) on a plagioclase separate from an eroded basaltic andesite cone (3S015) 1.5 km northwest of Triangle Hill (Table 3.1). This date probably represents the age of early TVC volcanic activity, and is supported in the next section by evidence that the oldest TVC pyroclastic eruption occurred around 0.63 Ma.

3.3.1 Pyroclastic Deposits of the TVC

A sequence of four major pyroclastic deposits are exposed in the area extending east of the TVC to the town of Bend (figure 3.4). These deposits represent the eruption of $\geq 14 \text{ km}^3$ (dense rock equivalent, DRE) of andesitic-to-rhyolitic magma from the Tumalo volcanic center, from ≈ 600 ka to ≤ 320 ka. Although these deposits have been studied in some detail (Taylor, 1981; Mimura, 1984; Hill, 1985; Hill and Taylor, 1989), confusion remains as to the source and age of these rocks. Data presented in this section will show that these rocks were clearly related to late Quaternary eruptions from the Tumalo volcanic center.

3.3.1a Desert Spring Tuff

The Desert Spring Tuff (Taylor, 1981) probably is the oldest of the Pleistocene High Cascade pyroclastic deposits. Where the basal contact is exposed, the Desert Spring Tuff overlies rocks of the Deschutes Formation. The Desert Spring Tuff contains at least two distinct flow lobes that form one cooling unit. Although complete sections are not exposed, the Desert Spring Tuff has an average thickness of about 12 meters, with one preserved section 30 meters thick. Although preservation is generally poor and outcrops are widely scattered, the Desert Spring Tuff probably represents the eruption of $\approx 2 \text{ km}^3$ (DRE) of rhyodacitic magma.

The Desert Spring Tuff is characterized by black rhyodacitic pumice up to 0.5 m in diameter, which contains up to 15 wt. % phenocrysts of plagioclase (An_{45-40}) , orthopyroxene $(Wo_3En_{55}Fs_{42})$, augite $(Wo_{40}En_{38}Fs_{22})$, magnetite, ilmenite, and apatite. Pyroxene phenocrysts are characterized by abundant inclusions of apatite. Although outcrops of the Desert Spring Tuff are scattered, the distribution of outcrops and the general westward increase in extent of welding indicates that this unit was erupted from the TVC area (Taylor, 1981; Mimura, 1984). In addition, the mineralogy and chemistry of the Desert Spring Tuff is very similar to several TVC domes such as Bearwallow Butte (table 3.2), which further supports a TVC source for this unit.

Glass shards from the Desert Spring Tuff have been correlated on the basis of major element geochemistry with the Rye Patch Dam ash bed (Davis, 1978), which is a widespread ash found in central Nevada and in core from Tulelake, northern California (Sarna-Wojcicki et al., 1989). The Rye Patch Dam ash has an age of ≈ 0.63 Ma, based on stratigraphic position in the Tulelake core (Sarna-Wojcicki et al., 1989). Paleomagnetic directions of the Desert Spring Tuff and Rye Patch Dam ash are also similar, further supporting a correlation between these two units

| Sample | 3 s005 | $\begin{array}{r} \textbf{35009} \\ 68.2 \\ 0.61 \\ 16.1 \\ 3.44 \\ 0.09 \\ 0.87 \\ 2.35 \\ 5.56 \\ 2.58 \\ 0.16 \\ 101.16 \end{array}$ | 35XBP | Avg Do | 3 s079 | 3 s028 | 35003 | | |
|---|---------------|---|-------|-------------|---------------|---|--------------|--|--|
| SiO2 | 68.0 | | 74.2 | 74.6 | 67.0 | 58.5 | 61.5 | | |
| TiO2 | 0.72 | | 0.16 | 0.15 | 0.81 | 1.51 | 1.28 | | |
| A12O3 | 15.6 | | 14.2 | 13.9 | 17.1 | 17.4 | 17.6 | | |
| FeO* | 4.24 | | 1.82 | 1.63 | 4.12 | 7.88 | 6.53 | | |
| MnO | 0.09 | | 0.07 | 0.06 | 0.13 | 0.16 | 0.22 | | |
| MgO | 0.45 | | 0.08 | 0.05 | 0.83 | 2.26 | 1.50 | | |
| CaO | 2.12 | | 0.81 | 0.69 | 2.39 | 5.36 | 3.83 | | |
| Na2O | 4.64 | | 4.46 | 5.18 | 4.33 | 4.78 | 5.16 | | |
| K2O | 3.25 | | 3.47 | 3.49 | 2.16 | 1.02 | 1.65 | | |
| P2O5 | <u>0.19</u> | | 0.03 | <u>0.02</u> | <u>0.16</u> | <u>0.37</u> | <u>0.47</u> | | |
| TOTAL | 99.86 | | 99.63 | 99.77 | 98.01 | 100.09 | 99.92 | | |
| Rb | 77 | 52 | 84 | 84 | 44 | 16 | 26 | | |
| Ba | 755 | 732 | 794 | 796 | 616 | 503 | 519 | | |
| Sr | 209 | 237 | 65 | 50 | 278 | 498 | 396 | | |
| Cs | 1.40 | 0.95 | 3.10 | 1.37 | 1.60 | 0.80 | 1.20 | | |
| SC | 9.52 | 8.72 | 4.29 | 4.30 | 11.05 | 21.82 | 18.33 | | |
| V | 25 | 23 | 0 | 3 | 36 | 159 | 105 | | |
| Co | 5.87 | 3.72 | 0.50 | 0.66 | 2.12 | 15.38 | 11.60 | | |
| Ni | 12 | 8 | 11 | 11 | 7 | 3 | 7 | | |
| Cr | 3 | 0 | 1 | 1 | 3 | 10 | 4 | | |
| Zn | 66 | 65 | 105 | 79 | 87 | 66 | 108 | | |
| La | 25.2 | 25.9 | 28.7 | 23.2 | 19.8 | $ 19.3 \\ 39.9 \\ 24.1 \\ 6.24 \\ 1.92 \\ 0.96 \\ 3.2 \\ 0.49 \\ 1.95 \\ 51.02 \\ 0.93 $ | 20.8 | | |
| Ce | 57.2 | 57.7 | 59.7 | 52.5 | 39.9 | | 49.8 | | |
| Nd | 27.1 | 26.0 | 26.5 | 20.4 | 22.1 | | 23.9 | | |
| Sm | 6.27 | 6.02 | 5.95 | 4.42 | 6.06 | | 6.13 | | |
| Eu | 1.49 | 1.40 | 0.83 | 0.61 | 1.37 | | 1.71 | | |
| Tb | 0.97 | 0.88 | 0.94 | 0.73 | 0.87 | | 0.88 | | |
| Yb | 3.7 | 3.7 | 4.2 | 3.3 | 3.4 | | 2.9 | | |
| Lu | 0.51 | 0.53 | 0.58 | 0.50 | 0.50 | | 0.43 | | |
| La/Sm | 2.54 | 2.72 | 3.05 | 3.32 | 2.06 | | 2.14 | | |
| EHREE | 54.09 | 52.75 | 57.77 | 46.72 | 50.23 | | 45.99 | | |
| Eu/Eu* | 0.72 | 0.71 | 0.42 | 0.41 | 0.70 | | 0.86 | | |
| Zr | 293 | 277 | 231 | 217 | 231 | 176 | 203 | | |
| Hf | 7.9 | 7.5 | 6.7 | 6.6 | 5.5 | 4.6 | 5.2 | | |
| Ga | 24 | 20 | 20 | 20 | 21 | 24 | 20 | | |
| Y | 34 | 33 | 36 | 28 | 36 | 32 | 30 | | |
| Nb | 18.6 | 16.1 | 16.9 | 17.4 | 14.3 | 9.5 | 12.9 | | |
| Ta | 1.01 | 0.98 | 1.20 | 1.02 | 0.66 | 0.61 | 0.72 | | |
| Th | 7.5 | 5.8 | 8.1 | 8.1 | 3.8 | 3.0 | 3.3 | | |
| U | 2.6 | 1.9 | 2.7 | 2.5 | 1.4 | 0.9 | 1.2 | | |
| <pre>3S005: Desert Spring Tuff on eastern flank of Bull Springs Inlier. 3S009: Bearwallow Butte rhyodacite dome. 3SXBP: Bulk Bend Pumice, quarry ≈1mi SW Bend. Avg Do: Average of 8 TVC rhyolite domes. 3S079: Columbia Canyon lapilli-fall tuff.</pre> | | | | | | | | | |

3S028: Century Drive Tuff, 2 mi NW. of Wanoga Butte. 3S003: Shevlin Park Tuff at west margin of Bull Springs Inlier.

Table 3.2: Compositions of pyroclastic units associated with the Tumalo volcanic center, and representative TVC silicic domes. Note that the composition of the Bend Pumice is nearly identical to the average composition of TVC rhyolite domes, and that the Desert Spring Tuff is compositionally similar to TVC rhyodacite domes.
(Gardner et al., 1992). If the Desert Spring Tuff represents the proximal equivalent of the 0.62 Ma Rye Patch Dam ash, then early TVC volcanism likely predated eruption of the Todd Lake volcano.

3.3.1b Tumalo Tuff

Previous workers (Taylor, 1981; Mimura, 1984; Hill, 1985; Hill and Taylor, 1989) have described a rhyolitic eruptive sequence consisting of a lapilli-fall tuff (Bend Pumice) directly overlain by a poorly welded ignimbrite (Tumalo Tuff), which is in turn overlain by a highly devitrified ignimbrite (Lava Island Tuff). Data presented in this section will support previous arguments (Hill, 1985; Hill and Taylor, 1989) that these units are the products of one eruptive event, and that the "Bend Pumice" (after Taylor, 1981) should be referred to as a member of the Tumalo Tuff.

The Bend Pumice member of the Tumalo Tuff (BP) is a rhyodacitic vitric lapilli-fall tuff that is best exposed along the roads leading to Tumalo State Park. The basal zone of the BP consists of pumice lapilli, ash, and perlitic obsidian, which have been locally reworked with alluvial gravels and sands. The basal zone ranges to 2 meters in thickness, and is thought to represent the preliminary stage of a climactic eruption (Hill, 1985). The basal zone is overlain by 3-13 meters of air-fall lapilli and ash, which progressively increases in average grain size up section except for a small reversal near the top of the deposit (Hill, 1985).

The median diameter of pumice-fall deposits generally shows an exponential decrease with distance from the vent (Fisher, 1964; Walker, 1971). The significant variations in median grain size and degree of sorting observed within the BP section require that samples used for granulometric analysis must be collected from a constant stratigraphic horizon in the deposit. Six ≈ 2 kg BP samples from 0.25 m below the contact with the overlying Tumalo Tuff were collected and manually sieved using standard sediment sieves (figure 3.5). These data were combined with grainsize data from Hill (1985) in order to better constrain the source of the Bend Pumice. The observed grain size distributions in the BP (figure 3.5) support the hypothesis that this unit was erupted from west of Bend (Hill, 1985). Comparison of the measured median grain sizes with distribution data from other pumice-fall deposits (Fisher, 1964; Walker, 1971) also suggests that the source of the BP was ≈ 10 to ≈ 20 km from these deposits (Hill, 1985).

A roughly 7x6x6-cm basaltic ($\rho \approx 2$ g/cm³) rock fragment was collected near the top of a Bend Pumice outcrop 9 km east-northeast of Triangle Hill (figure 3.4). Because rock fragments this size are generally transported ballistically, rather than convectively, from the vent, ballistic trajectories were calculated for this clast using the equations of Wilson (1972). The maximum ballistic transport distance for this clast is ≈ 2 km, which is unreasonably low. It is likely that this clast had components of both convective and ballistic transport (i.e., Walker, 1981), because the nearest reasonable vent locality was 9 km to the west at Triangle Hill. In essence, the occurrence



| | D | N | G | H | С | E | H | L | K | T |
|-----------------|------|------|------|------|------|------|------|------|------|------|
| -64 | 10.8 | 0.0 | 20.7 | 0.0 | 0.0 | 0.0 | 0.0 | 2.9 | 25 | 21 |
| -54 | 17.1 | 4.0 | 6.6 | 10.0 | 3.2 | 3.0 | 3.8 | 3.7 | 4.8 | 4.0 |
| -4Φ | 9.9 | 4.7 | 11.2 | 13.2 | 10.4 | 8.4 | 13.1 | 13.8 | 12 0 | 11 3 |
| - 34 | 12.5 | 14.3 | 16.7 | 18.9 | 17.6 | 15.2 | 19.6 | 19.8 | 21 1 | 20.2 |
| -2 | 11.7 | 20.3 | 15.3 | 18.7 | 19.7 | 18.2 | 20.2 | 18.6 | 21.9 | 20.2 |
| -1≢ | 13.5 | 24.6 | 13.6 | 16.8 | 20.6 | 21.5 | 19.6 | 18.4 | 20.7 | 20.4 |
| 0₽ | 8.9 | 17.3 | 6.9 | 8.9 | 12.5 | 13.3 | 11.9 | 10.6 | 9.5 | 10.8 |
| 1 ē | 7.5 | 9.1 | 4.5 | 7.0 | 9.3 | 10.4 | 8.3 | 8.5 | 5.1 | 6.4 |
| >1 | 8.1 | 5.6 | 4.4 | 6.4 | 6.7 | 10.0 | 3.4 | 3.6 | 2.4 | 4.2 |
| | | | _ | | | | | | | |
| ⁹ 50 | 3.0 | 1.7 | 3.3 | 2.6 | 2.2 | 1.8 | 2.4 | 2.5 | 2.6 | 2.4 |

Figure 3.5: Isopleth map of the median pumice diameter from the upper 0.25 m of the Bend Pumice member of the Tumalo Tuff. Sites A, B, F, and J from Hill (1985). Precision of these data is about ± 0.2¢, based on duplicate analyses from adjacent sites. Note the overall west-to-east decrease in median diameter, which suggests a source near Triangle Hill.

of this relatively large basaltic clast in the BP at this location clearly indicates proximity to the source vent.

Westward increases in average grain size, unit thickness (Hill, 1985), and size of volcanic rock fragments all indicate the BP was erupted from west of Bend, and most likely from the TVC area. In addition, the composition of the BP is nearly identical to the composition of TVC rhyolite domes (table 3.2). This rhyolite composition only occurs in the central High Cascades of Oregon at the TVC. The source of the BP (and cogenetic Tumalo and Lava Island Tuffs) was most likely the large silicic vent complex buried beneath Triangle Hill.

The Tumalo Tuff is a pink-to-tan rhyodacitic vitric ignimbrite that directly overlies the Bend Pumice. The absence of a normally graded top to the Bend Pumice, and the non-erosive basal contact of the Tumalo Tuff with the BP, indicates that the Tumalo Tuff was produced through collapse of a plinean eruption column (Hill, 1985). In addition, the compositions of the Bend Pumice and Tumalo Tuff are identical (Hill, 1985). It is thus misleading to refer to the Bend Pumice and Tumalo Tuff as separate units in the stratigraphy of the Bend, Oregon area. Eruption of the Tumalo Tuff was preceded by the deposition of a lapilli-fall tuff, which is herein referred to as the Bend Pumice member of the Tumalo Tuff.

The basal ≈ 1 m of the Tumalo Tuff consists of a coarse pumice-depleted zone (layer 2a) and occasional surge (layer 1p) deposits. The main nonsorted part of the Tumalo Tuff (layer 2b) is as thick as 22 meters, and was deposited as a single flow unit. Welding ranges from nonwelded to slightly welded (small amounts of pumice collapse). With the exception of the basal 1 meter and the interior of large pumice bombs, the Tumalo Tuff has undergone low-grade vapor phase alteration (Hill, 1985).

Imbrication of pumice clasts in the Tumalo Tuff indicates a northeast direction of flow (Mimura, 1984), which has been used to contend that the Tumalo Tuff was erupted from vents located 30-45 km southwest of Bend (Mimura, 1984). However, flow directions do not always indicate the direction to the source vent, because most ignimbrites are channeled by the preexisting topography of the area (e.g., Suzuki and Ui, 1983). The direct association of the Tumalo Tuff with the Bend Pumice indicates that the Tumalo Tuff was erupted from the TVC, but was channeled by northeast-trending drainages as it flowed off of the TVC highland (Hill, 1985). The Tumalo Tuff represents the eruption of a minimum of 10 km³ of nonzoned rhyodacitic magma (Hill, 1985).

The Tumalo Tuff has a distinct mineral assemblage: Plagioclase $(An_{20}) >$ ferrohypersthene $(Wo_3En_{30}Fs_{57}) \ge pargasitic hornblende \ge magnetite \ge ilmenite \ge apatite \ge augite \ge zircon$ (Appendix B). It is noteworthy that the Tumalo Tuff contains the most iron-rich orthopyroxenes that have been observed in the central High Cascades of Oregon. Banded pumice clasts, which

represent the mingling of rhyodacitic and unrelated dacitic magmas (Hill, 1985), are also found in proximal (western) exposures of the Tumalo Tuff ignimbrite.

The Tumalo Tuff eruption has been correlated on the basis of composition and paleomagnetic direction (Gardner et al., 1992) with the Loleta Ash, which has an estimated age of slightly less than 0.40 Ma (Sarna-Wojcicki et al., 1987). K-Ar age determinations for 4 plagioclase separates from the Tumalo Tuff yielded an average date of 0.29 \pm 0.12 Ma (Sarna-Wojcicki et al., 1987). Fragments of rhyolite obsidian from the basal zone of the Bend Pumice have K-Ar dates of 0.39 \pm 0.01 Ma and 0.41 \pm 0.02 Ma (Sarna-Wojcicki et al., 1989), further indicating that the Tumalo Tuff is younger than 0.4 Ma. The age of the Tumalo Tuff is proabably 0.35-0.38 Ma, based on the stratigraphic position of the Loleta ash (Sarna-Wojcicki et al., 1987), age of similar composition rhyolite obsidian clasts in the basal BP (Sarna-Wojcicki et al., 1989), and age of overlying deposits (this study). A \approx 0.35 Ma age for the Tumalo Tuff is at least 0.5 m.y. younger than previously reported ages (Armstrong et al., 1975).

3.3.1c Lava Island Tuff

The Lava Island Tuff is a purple-to-grey, intensely devitrified ignimbrite that in places overlies the Tumalo Tuff. The basal contact of the Lava Island Tuff with the Tumalo Tuff is sharp and erosive, with no intervening deposits. In contrast, the basal contact of the Tumalo Tuff with the unconsolidated Bend Pumice is non-erosive. The Lava Island Tuff has relative abundances of pumice, rock fragments and ash that appear similar to welded sections of the Tumalo Tuff. However, the Lava Island Tuff lacks a basal coarse-depleted 2a layer, which is present in the Tumalo Tuff. The intense devitrification and turbulent emplacement of the Lava Island Tuff indicates that this ignimbrite was emplaced with a higher gas content than the Tumalo Tuff. The Lava Island Tuff has a mineralogy that is similar to the Tumalo Tuff, except that the ferrohypersthene is rimmed with iron-rich augite ($Wo_{40}En_{20}Fs_{40}$) and both pyroxenes contain abundant apatite inclusions. The Lava Island Tuff probably represents a flow lobe of the Tumalo Tuff, which was derived from a deeper, less degassed part of the magma chamber during the waning stages of the Tumalo Tuff eruption (Taylor, 1981; Hill and Taylor, 1989).

3.3.1d Pumice of Columbia Canyon

This informally named (Hill and Taylor, 1989) dacitic, vitric lapilli-fall tuff occurs underneath the Shevlin Park Tuff in a steep canyon along the Columbia Canal (T17S R11E S17), and overlies several mafic flows north of the TVC. It has an average grain size (about 1 cm) that is similar to the Bend Pumice at the Columbia Canyon outcrop, and contains abundant angular fragments of black obsidian up to 4 cm in diameter, both of which suggest that this unit was erupted from a nearby TVC vent. The pumice contains phenocrysts of plagioclase (An₄₀) > augite $(Wo_{40}En_{43}Fs_{35}) \approx$ hypersthene $(Wo_3En_{63}Fs_{34}) >$ amphibole \geq magnetite \geq ilmenite > apatite (Appendix B). The moderately hydrated pumice of Columbia Canyon (sample 3S079) cannot be directly correlated with any other High Cascade dacite (table 3.2). Glass shards from the pumice of Columbia Canyon have a major element composition that is very similar to Summer Lake ash bed NN (Davis, 1985), which suggests that these two units are related (Sarna Wojcicki et al., 1989).

3.3.1e Century Drive Tuff

The Century Drive Tuff (Taylor, 1981) is a variably welded vitric ignimbrite that contains both rhyodacitic(?) and andesitic pumice. Andesitic pumice (table 3.2) contains phenocrysts of plagioclase (An₃₅) > augite (Wo₄₁En₄₂Fs₁₇) > olivine (Fo₇₁) > hypersthene (Wo₂En₆₇Fs₃₁) \geq magnetite \geq ilmenite (Appendix B). Although the Century Drive Tuff is restricted to scattered outcrops in the area south and west of Bend (figure 3.4), it appears to be more densely welded in exposures closer to the TVC, suggesting eruption from a TVC vent. Outcrops of Century Drive Tuff also occur at elevations of 5600-5400' \approx 25 km south of Triangle Hill, which further supports eruption from a TVC vent (figure 3.4). Eruption of the Century Drive Tuff probably involved \leq 0.5 km³ (DRE) of andesitic magma.

3.3.1f Shevlin Park Tuff

The Shevlin Park Tuff (Taylor, 1981) is an andesitic (table 3.2) ignimbrite that is distributed around a 180° sector east of and centered on Triangle Hill (Taylor, 1987; Hill and Taylor, 1989; Figure 3.4). The Shevlin Park Tuff forms a single cooling unit, and may contain two flow lobes and occasional base-surge deposits in more proximal exposures. The Shevlin Park Tuff contains plagioclase (An₃₆) > augite (Wo₄₂En₄₃Fs₁₅) \geq hypersthene (Wo₃En₆₅Fs₃₂) > olivine (Fo₇₁) > titanomagnetite (Appendix B). The distribution of the Shevlin Park Tuff, along with observed increases in degree of welding and average pumice size, clearly indicates this unit was erupted from the TVC. Imbrication of pumice lapilli in the Shevlin Park Tuff also indicates a radial direction of flow away from the Triangle Hill area (Mimura, 1984). Eruption of the Shevlin Park Tuff probably involved ≥ 1 km³ (DRE) of andesitic magma.

A noteworthy Shevlin Park Tuff outcrop occurs in the upper reaches of the North Fork of Squaw Creek (T16S R9E S29-30). At this location the Shevlin Park Tuff is overlain by a normal polarity basaltic andesite flow, which is in turn overlain by the oldest basaltic andesites clearly associated with North Sister (Taylor, 1987). Because North Sister is the oldest stratovolcano of the Three Sisters, this exposure clearly demonstrates that pyroclastic volcanism associated with the TVC predated construction of the Three Sisters.

3.4 Todd Lake Volcano

Some of the oldest Quaternary-age rocks in the Broken Top system are associated with the Todd Lake volcano (Williams, 1944; Taylor, 1978; Figure 3SLOC). Todd Lake volcano (TLV) is a highly eroded composite volcano, which consists primarily of dacitic flows and intrusions; rhyodacite and basaltic andesite lavas were also erupted from the TLV (Taylor, 1978). A dacite flow from the middle section of TLV has a K-Ar date of 460 ± 30 ka (table 3.1), and all rocks associated with the TLV have normal paleomagnetic directions. TLV is overlain by mafic flows associated with Broken Top-Tam MacArthur Rim volcanism, and its stratigraphic relationship to the Tumalo volcanic center is not known.

3.5 Tam MacArthur Rim and Broken Top

Tam MacArthur Rim (TMR) represents a distinct volcanic center in the Broken Top system. Numerous basaltic andesite to rhyodacite dikes and vent breccias exposed by glaciation at TMR clearly demonstrate that this area was a locus of volcanic activity, even though these eruptions did not result in the construction of a stratovolcano (Taylor, 1978). Although stratigraphic relationships between TMR and the TVC are obscured by Pleistocene glacial deposits, it is likely that some mafic lavas from TMR partially cover the older TVC highland (Taylor, 1978).

A K-Ar date of 213 ± 9 ka was obtained on a glassy rhyodacite flow located at the top of the TMR section (table 3.1). This rhyodacite is also interbedded with a series of basaltic andesite flows that comprise the base of Broken Top volcano (Taylor, 1978). Although Broken Top volcano has been highly dissected by Pleistocene glaciation, this relatively large degree of erosion has developed in less than 200 k.y. In addition, a deposit of unconsolidated pumice lapilli from a small rhyodacite dome eruption on the northwest flank of Broken Top is preserved at an elevation of 7800'. A similar relationship is observed near North Sister, where lapilli from the ≈ 100 ka eruption of Obsidian Cliffs are preserved on the western flank of Little Brother at an elevation of 7400' (E.M. Taylor, pers. comm., 1986). This relationship suggests that Broken Top rhyodacitic volcanism may have continued until after early Wisconsin glaciation at ≈ 110 ka (Jouzel et al., 1987), assuming that early Wisconson glaciation covered these high-standing ridges.

3.6 The Three Sisters

North Sister is the oldest of the Three Sisters stratovolcanos (Taylor, 1981), and stratigraphically overlies the Shevlin Park Tuff (Taylor, 1987). North Sister is primarily constructed of basaltic andesite, although several small dikes of andesite are exposed on the northeast flank of the stratovolcano (E.M. Taylor, pers. comm., 1988). Uppermost North Sister lavas can be traced into the lower parts of the Middle Sister stratigraphic section. Mafic volcanism at North Sister has not been studied in detail, although reconnaissance work by Hughes (1983) indicates that North Sister is a representative part of the Oregon High Cascades mafic magma system.

The lower parts of Middle Sister are primarily composed of mafic lavas, and andesitic to dacitic lavas are restricted to the upper parts of the Middle Sister section (Taylor, 1981; Taylor et al., 1987). Near the saddle between Middle and South Sister, a porphyritic basalt lava relatively high in the Middle Sister section (3S159) overlies dacite lava from the basal section of South Sister. A K-Ar date of 0.4 ± 0.2 Ma on this basalt has low precision and accuracy due to large amounts of atmospheric Ar (table 3.1). In addition, the youngest dacite from Middle Sister bifurcated around an early South Sister rhyolite dome in the saddle between Middle and South Sister. These stratigraphic relationships indicate that late Middle Sister and early South Sister volcanism was contemporaneous.

South Sister is the youngest of the Three Sisters stratovolcanos, and has erupted $\ge 1 \text{ km}^3$ of rhyolite as recently as ≈ 2 ka (Scott, 1987; Taylor, 1978). Geothermal drilling on the southern flanks of South Sister at Devils Lake encountered 0.15 ± 0.05 Ma basalt at 1819', near the base of the hole (Priest et al., 1989). Stratigraphic relations in the drill core indicate that this basalt is overlain by Devils Hill rhyolites, which are in turn overlain by South Sister andesites. In addition, a K-Ar date of 93 ± 11 ka was measured on an andesite flow near the base of the South Sister section. These data indicate that the bulk of South Sister is <100 ka.

3.7 Strontium Isotopes

There are few published strontium isotopic analyses for rocks from the Three Sisters area (table 3.3). A previously unreported 87 Sr/ 86 Sr analysis on the Tumalo Tuff (0.70375 ± 5) was provided by J. Dasch for this study (table 3.3). In addition, six 87 Sr/ 86 Sr analyses obtained by UNOCAL Geothermal as part of a regional geochemical study were released for this study (table SR76). These limited data suggest that mafic rocks in the Three Sisters area may have slightly higher 87 Sr/ 86 Sr (\approx 0.7035) than most Quaternary mafic rocks at Mt Jefferson (0.7028-0.7036; Conrey, 1991). Rhyolites from the Three Sisters and Broken Top systems have identical 87 Sr/ 86 Sr (\approx 0.7037). Petrogenetic models in chapter 9 show that these rhyolites were produced through partial crustal melts, indicating that the crust beneath this part of the Oregon Cascades is isotopically primitive. Rhyodacites at Mt. Jefferson, which were also inferred to have been produced through partial crustal melting (Conrey, 1991), have lower 87 Sr/ 86 Sr (\approx 0.7033). This relationship may indicate that Mt. Jefferson is located over relatively younger mafic crust than found under the Three Sisters area (Chapter 2).

Three Sisters area

Published ⁸⁷Sr/⁸⁶Sr analyses:

| <u>Location</u> | SiO2 | ppm Sr | $\frac{8}{Sr} = 1\sigma$ | Source |
|-----------------|---------|--------|--------------------------|-----------------------|
| Three Sisters | 58% | 355 | 0.7036 ± 6 | Church & Tilton, 1970 |
| Three Sisters | 54% | 330 | 0.704 ± 1 | Church & Tilton, 1970 |
| South Sister | <53->64 | n/a | 0.7034-0.7039 | Leeman, 1982 |

New ⁸⁷Sr/⁸⁶Sr analyses:

| Location | Sample# | SiO, | ppm Sr | ppm Rb | $\frac{8}{Sr}/\frac{80}{Sr} \pm 1\sigma$ | Source |
|-----------------|---------------------|------|--------|--------|--|--------------|
| Tumalo Tuff | , 3245 | 74 | 68.3 | 79.2 | 0.70375 ± 5 | Dasch, 1984 |
| Three Ck. Bu. | 35032 | 74 | 52 | 86 | 0.70368 ± 2 | UNOCAL, 1989 |
| Obsidian Cliffs | 3S139 | 76 | 119 | 79 | 0.70363 ± 3 | UNOCAL, 1989 |
| Sister Sp. | 3S136 | 64 | 392 | 28 | 0.70370 ± 3 | UNOCAL, 1989 |
| Yapoah lava | 2TFJ-437 | 56 | 650 | 21 | 0.70346 ± 1 | UNOCAL, 1989 |
| E. N. Sister | ² TS-692 | 53 | 530 | 4 | 0.70351 ± 2 | UNOCAL, 1989 |
| Older Belknap | C-1 | 53 | n/a | n/a | 0.70345 ± 2 | UNOCAL, 1989 |

- Dasch: Analysis at Johnson Space Center, 1984.

- UNOCAL: Analyses at U.C.S.B. by J. Mattinson, 1985. Relative to NBS 987 Sr standard: 0.710254.

1: Sample numbers correspond to equivalent samples from this study.

2: Sample numbers correspond to equivalent samples Hughes, 1983.

Table 3.3: Compilation of published and new ⁸⁷Sr/⁸⁶Sr isotopic analyses for rocks from the Three Sisters area. Tumalo Tuff sample from Hill (1985), analysis courtesy of J. Dasch, written comm., 1984. UNOCAL samples are previously unreported data from UNOCAL Geothermal, Santa Rosa, California, and published with permission of UNOCAL (D. Carrier, written comm., 1989). Samples are correlated with equivalent whole-rock analyses, as other geochemical data are not available. These data indicate that mafic samples contain less radiogenic Sr than silicic samples, but still have enriched ⁸⁷Sr/⁸⁶Sr relative to most samples from Mt. Jefferson (Conrey, 1991), and some samples from the flanks of Newberry volcano (Goles and Lambert, 1990). Rhyolite analyses have identical ⁸⁷Sr/⁸⁶Sr at 1σ, which indicates that these units did not sample isotopically distinct or highly evolved crustal compositions.

4. MINERALOGY

Rocks from the study area display a wide variety of textures and mineral abundances, which systematically with rock composition. Phenocryst abundances range from trace amounts in some rhyolites to greater than 35% in some andesites and dacites (figure 4.1), and all rock compositions have samples that vary from nearly aphyric (< 1%) to moderately phyric (> 20%). The phenocryst assemblage is dominated by plagioclase for all rock compositions, with subordinate amounts of clinopyroxene, orthopyroxene, olivine, Fe-Ti oxides, and amphibole in relative proportions that are compositionally controlled. The mineralogy of Three Sisters and Broken Top systems can be characterized, however, as an anhydrous, plagioclase - two pyroxene assemblage; hydrous phenocrysts are restricted to trace amounts of amphibole in some rhyodacites and rhyolites. Disequilibrium crystallization features, such as complex plagioclase zoning, abundant glass inclusions in plagioclase phenocrysts, reaction rims around mafic phases, and multiple phenocryst populations within single units, are also common in rocks from the study area. These textural features are important characteristics of both the Three Sisters and Broken Top volcanic systems, and are evidence of the dynamic, open-system differentiation that produced these rocks.

Thin sections from 115 samples were studied in detail, and the petrographic data are summarized in appendix B. Of these samples, 21 were selected for detailed microprobe analysis. Most of these samples represented typical compositions from the study area, based on geochemical abundances and petrographic characteristics, while other samples had unusual petrographic features (i.e. complex zoning, multiple phenocryst populations) that warranted additional study. In addition, the compositions of minerals separated from 5 pyroclastic units in the Bend area were determined with the microprobe.

4.1 Feldspar

Phenocrysts and microlites of plagioclase are ubiquitous in the studied units, and range in abundance from 25% in some dacites and basaltic andesites to trace amounts in nearly aphyric rhyolites (Appendix B). Plagioclase usually occurs as 1-2 mm-long euhedral to subhedral crystals, but ranges in length up to 4 mm. Anorthoclase has been recognized through microprobe analysis as a rare groundmass phase in a Tam MacArthur Rim dacite (3S036), and as a 10 μ m rim on several andesine phenocrysts in a South Sister dacite (3S123) (table 4.1).

Variations in plagioclase composition determined through microprobe analysis are plotted in figures 4.2a-c for each rock type. Rhyolites from the Three Sisters system (figure 4.2a) contain plagioclase that is consistently more calcic (An_{50-25}) than found in Broken Top rhyolites (An_{25-15}). This compositional difference is probably controlled by bulk composition and not by temperature, as the Three Sisters rhyolites are consistently more calcic than Broken Top rhyolites: molar



Figure 4.1: Visual estimates of phenocryst modal abundance versus silica content of the host rock. Note that Three Sisters rhyolites are more phyric than Broken Top rhyolites.

| Sample | SiO2 | A1203 | FeO | MgO | CaO | Na 20 | K20 | BaO | Total | Ab | Or | An |
|-------------------|-------|-------|------|------|-------|-------|------|------|--------|-------|-------|-------|
| 3S003 rim (2) | 57.75 | 25.99 | 0.43 | 0.02 | 7.83 | 6.96 | 0.28 | 0.07 | 99.33 | 60.60 | 1.62 | 37.79 |
| 3S005 rim (2) | 56.02 | 25.78 | 0.34 | 0.01 | 8.26 | 6.71 | 0.39 | 0.04 | 97.53 | 58.21 | 2.22 | 39.58 |
| 3S009 rim (2) | 60.11 | 24.01 | 0.24 | 0.02 | 5.57 | 7.71 | 0.69 | 0.08 | 98.41 | 68.63 | 4.00 | 27.38 |
| 3S017 GM (3) | 53.09 | 28.39 | 1.01 | 0.09 | 11.42 | 4.63 | 0.18 | 0.01 | 98.82 | 41.88 | 1.06 | 57.06 |
| 3S028b rim | 57.39 | 25.65 | 0.34 | 0.06 | 7.37 | 7.49 | 0.30 | 0.05 | 98.65 | 63.67 | 1.69 | 34.64 |
| 3S030 GM & R (4) | 54.57 | 27.86 | 1.00 | 0.02 | 10.62 | 5.01 | 0.31 | 0.02 | 99.40 | 45.21 | 1.82 | 52.97 |
| 3S030 #3 rim | 59.95 | 24.35 | 0.74 | 0.00 | 6.30 | 7.16 | 0.70 | 0.06 | 99.26 | 64.49 | 4.15 | 31.36 |
| 3S036 HiK GM (2) | 64.90 | 21.14 | 0.58 | 0.02 | 4.04 | 6.66 | 2.01 | 0.06 | 99.41 | 65.23 | 12.91 | 21.85 |
| 3S036 GM (2) | 58.76 | 25.29 | 0.78 | 0.00 | 7.26 | 6.60 | 0.67 | 0.06 | 99.42 | 59.69 | 3.99 | 36.33 |
| 3S042 GM (4) | 64.68 | 21.15 | 0.33 | 0.00 | 2.26 | 8.83 | 1.23 | 0.25 | 98.73 | 81.07 | 7.43 | 11.49 |
| 3S045 GM & R (5) | 55.62 | 27.25 | 1.14 | 0.04 | 9.60 | 5.57 | 0.36 | 0.01 | 99.60 | 50.13 | 2.14 | 47.73 |
| 3S049 GM & R (3) | 52.78 | 29.09 | 0.92 | 0.07 | 12.04 | 4.45 | 0.33 | 0.02 | 99.70 | 39.33 | 1.91 | 58.76 |
| 3S058 GM & R (3) | 59.40 | 25.23 | 0.24 | 0.00 | 6.64 | 7.19 | 0.44 | 0.06 | 99.19 | 64.49 | 2.57 | 32.94 |
| 3s061 GM & R (2) | 55.97 | 27.01 | 0.57 | 0.01 | 9.38 | 5.73 | 0.39 | 0.03 | 99.09 | 51.32 | 2.29 | 46.40 |
| 3S069 GM (2) | 50.73 | 30.36 | 1.02 | 0.01 | 13.55 | 3.70 | 0.21 | 0.00 | 99.57 | 32.68 | 1.22 | 66.10 |
| 3S074 GM & R (3) | 53.46 | 28.11 | 1.18 | 0.06 | 11.30 | 4.45 | 0.43 | 0.01 | 99.01 | 40.53 | 2.59 | 56.88 |
| 3S079 rim (3) | 58.70 | 25.46 | 0.42 | 0.02 | 7.72 | 7.04 | 0.25 | 0.04 | 99.63 | 61.38 | 1.41 | 37.22 |
| 3S090 GM (5) | 53.66 | 28.71 | 0.82 | 0.05 | 11.24 | 4.82 | 0.21 | 0.01 | 99.53 | 43.15 | 1.24 | 55.62 |
| 3S098 GM & R (4) | 62.10 | 23.72 | 0.36 | 0.01 | 4.82 | 8.18 | 0.53 | 0.09 | 99.80 | 73.09 | 3.12 | 23.78 |
| 3S112 GM & R (3) | 57.13 | 26.47 | 0.87 | 0.01 | 8.56 | 6.09 | 0.45 | 0.04 | 99.62 | 54.78 | 2.68 | 42.54 |
| 3S114 GM & R (3) | 57.27 | 26.50 | 0.29 | 0.02 | 8.43 | 6.40 | 0.34 | 0.03 | 99.27 | 56.73 | 1.97 | 41.30 |
| 3S115 GM low An | 59.69 | 25.09 | 0.88 | 0.03 | 6.48 | 7.15 | 0.53 | 0.06 | 99.91 | 64.54 | 3.13 | 32.33 |
| 3S115 GM high An | 55.15 | 28.45 | 0.69 | 0.03 | 10.40 | 5.11 | 0.22 | 0.01 | 100.06 | 46.45 | 1.33 | 52.22 |
| 3S122 GM & R (5) | 60.72 | 24.70 | 0.27 | 0.01 | 6.36 | 7.10 | 0.58 | 0.08 | 99.82 | 64.53 | 3.50 | 31.97 |
| 3S123 GM | 59.20 | 24.94 | 0.76 | 0.01 | 6.79 | 7.17 | 0.68 | 0.08 | 99.62 | 63.04 | 3.95 | 33.01 |
| 3S123z rim (2) | 62.60 | 22.01 | 0.69 | 0.01 | 4.07 | 7.16 | 2.16 | 0.04 | 98.73 | 66.13 | 13.02 | 20.85 |
| 35131 GM & #3 rim | 61.45 | 23.12 | 0.50 | 0.01 | 4.68 | 7.97 | 0.79 | 0.09 | 98.59 | 71.95 | 4.65 | 23.40 |
| 35131 GM & R (4) | 54.51 | 27.33 | 0.96 | 0.07 | 10.13 | 5.37 | 0.26 | 0.03 | 98.66 | 48.18 | 1.56 | 50.26 |
| 3S134 GM (2) | 60.08 | 24.34 | 0.57 | 0.00 | 5.78 | 7.57 | 0.51 | 0.08 | 98.93 | 68.22 | 3.01 | 28.78 |
| 38137 GM (3) | 58.66 | 25.62 | 0.51 | 0.01 | 1.23 | 6.99 | 0.39 | 0.06 | 99.47 | 62.20 | 2.26 | 35.54 |
| JEADP FIM (J) | 62.4/ | 22.33 | 0.24 | 0.02 | 4.05 | 8.49 | 0.76 | 0.10 | 98.46 | 75.60 | 4.48 | 19.92 |
| LI EIM (2) | 03.52 | 23.08 | 0.13 | 0.00 | 4.32 | 0.40 | 0.57 | 0.10 | 100.27 | /5.39 | دد.د | 21.28 |

Table 4.1: Average compositions of groundmass feldspars (GM) and plagioclase rim compositions (R), determined through microprobe analysis. Number of samples averaged shown in parenthesis. All averages have standard deviations ≤ 7%.



Figure 4.2a: Compositional range of plagioclase phenocrysts and microlites for individual rhyolite and rhyodacite samples, plotted on part of the Ab-An-Or feldspar ternary. + Broken Top system, o Three Sisters system. Rhyolite and Rhyodacite are combined plots of all samples.



Figure 4.2b: Compositional range of plagioclase phenocrysts and microlites for individual dacite samples, plotted on part of the Ab-An-Or feldspar ternary. Dacite is a plot of all dacitic samples, symbols the same as in figure 4.2a.



Figure 4.2c: Compositional range of plagioclase phenocrysts and microlites for individual andesite, basaltic andesite, and basalt samples, plotted on part of the Ab-An-Or feldspar ternary. Symbols as in figure 4.2a. Andesite and BA & Bs are combined plots of all samples.

 $CaO/Na_2O = 0.52$ vs. 0.25 at 72% SiO₂, and 0.24 vs. 0.15 at 76% SiO₂, respectively. The largest range in plagioclase composition tends to occur in basaltic andesites, although each rock type contains samples with small (≤ 10 An) and samples with large (10-50 An) compositional variations. Plagioclase phenocrysts with the largest compositional range generally contain an anomalously calcic core, strongly zoned sodic rims, and tend to be the larger crystals in a seriate distribution.

Average phenocryst rim and groundmass plagioclase compositions (table 4.1), which should represent compositions that are relatively close to equilibrium with the magma at the time of eruption, are compared with rock SiO₂ in figure 4.3a. The expected trend (cf. Gill, 1982; Ewart, 1982) of more albitic plagioclase with increasing rock SiO₂ is present, along with variations of \pm 10 An for similar composition rocks. As expected, most mafic units have groundmass plagioclase more calcic than An₅₀. The groundmass plagioclase in basaltic andesites 3S030 and 3S131, however, is unusually albitic; both of these units have a coarsely crystalline, sub-ophitic groundmass, which indicates that these units cooled slowly at the surface and that the groundmass plagioclase probably crystallized at a temperature that was significantly lower than in the pre-eruptive magma.

Most of the studied plagioclase phenocrysts display complex oscillatory zoning, which is characteristic of most calc-alkaline systems (Ewart, 1982). Both oscillatory and simply zoned plagioclase tend to be normally zoned, with rim An 10-60% less than core An (figure 4.3a). Reverse zoning is uncommon, but is observed both optically and analytically in several units (figure 4.3b). The most strongly zoned plagioclase phenocrysts also tend to have at least one zoning reversal (i.e. normal-reverse-normal) and complex oscillatory zoning. Both simple and oscillatory normal zoning is the expected result of equilibrium fractionation of a slowly cooled magma, and is supported by experimental studies (Lofgren, 1980), numerical simulations (Loomis, 1982), and by its common occurrence in most calc-alkaline volcanic systems (Ewart, 1982; Gill, 1982).

Strong compositional zonation is also a characteristic of plagioclase phenocrysts that contain abundant glass inclusions. While a variety of terms such as "sieved", "dusty", "fritted" or "spongy" have been applied to plagioclase with glass inclusions, two genetically distinct types of plagioclase glass inclusions are generally recognized. The first type, herein referred to as *fritted*, consists of irregularly-shaped sodic cores mantled by a zone of fine (μ m-scale) glass inclusions in a more calcic rim (figure 4.4a). Fritted plagioclase has been reported in a variety of igneous rocks (e.g. Larsen et al., 1938; MacDonald & Katsura, 1965; Koyaguchi, 1986; Glazner et al., 1990), and is generally thought to represent dissolution of albitic plagioclase that has been incorporated into a more anorthitic melt with subsequent crystallization of a reversely zoned rim (do.; Sakuyama, 1981; Tsuchiyama & Takahashi, 1983; Donaldson, 1985; Tsuchiyama, 1985). Fritted plagioclase, however, is uncommon and poorly developed in this part of the High Cascades; only 4% of 114 examined rocks have more than 50% of the plagioclase with some evidence of weak fritting, and between 5%



Figure 4.3: A) Average composition of groundmass plagioclase and rim of phenocrysts, plotted against silica content of the host rock. Compositions listed in table 4.1. B) Percent change in core-to-rim An content of plagioclase phenocrysts, versus silica content of the host rock. Plagioclase compositions determined through microprobe analysis. Negative values indicate reversed zoning.

B



Figure 4.4: A) Poorly developed fritted rim (An₆₁) on An₄₀ plagioclase phenocryst, basaltic andesite 3S074. This sample is one of few in the study area that contains fritted or resorbed plagioclase. B) Typical sieved plagioclase phenocryst, with nonzoned An₇₄ core and normally zoned, glass-free rim (An₇₃₋₅₈). Sample from basaltic andesite 3S090.

and 50% of the plagioclase displays some evidence of fritting in 18% of these rocks (Appendix B). In addition, half of the samples that are identified as fritted contain highly rounded sodic cores and glass-free, reversely zoned calcic rims. These samples clearly indicate dissolution of albitic plagioclase in an anorthitic melt (Donaldson, 1985), but lack the fine glass inclusions that are characteristic of a fritted texture.

The second type of plagioclase is referred to as sieved, which consists of calcic cores that are nonzoned to weak normal zoned and contain abundant, large (generally $\geq 10\mu$ m) glass inclusions, mantled by a normally zoned, glass-free, more sodic rim (figure 4.4b). Unlike fritted plagioclase, sieved plagioclase is a common petrographic feature in the studied rocks; between 5% and 50% of the plagioclase phenocrysts are significantly sieved in 64% of these rocks, and $\geq 50\%$ of the plagioclase is sieved in 11% of the rocks. Sieved plagioclase has also been reported in a large variety of igneous rocks (do., Luhr & Carmichael, 1980), and is most likely produced through disequilibrium crystallization produced by undercooling of the melt (Lofgren, 1980; Swanson, 1977; Corrigan, 1982; Glazner et al., 1990). Lofgren (1974) observed that a transition from tabular to skeletal (i.e. sieved) plagioclase morphology generally occurred between 40°C and 140°C undercooling, and that larger amounts of undercooling produced acicular, dendritic plagioclase. Isothermal cooling experiments by Lofgren (1973) also produced sieved plagioclase when cooling rates were roughly 2°C/hr, but an increase in cooling rate to 32°C/hr produced acicular, dendritic plagioclase. Growth rates of about 10⁻⁷ cm/sec for skeletal plagioclase in undercooled tholeiitic melts were determined by Corrigan (1982). While this rate will presumedly decrease with decreasing temperature and increasing viscosity of the melt, it is likely that the 1-2 mm sieved plagioclase commonly observed in the studied rocks crystallized on the order of tens of days.

Magma mixing is the most common mechanism that is thought to produce the large degrees of undercooling necessary for the crystallization of sieved plagioclase (e.g. Larsen et al., 1938; Dungan & Rhodes, 1978; Hibbard, 1981; Ussler and Glazner, 1989). Evidence for some degree of magma mixing occurs in most arc volcanic systems (e.g. Eichelberger, 1975; Gill, 1982), and there is little reason to doubt that magma mixing is capable of producing the degree of cooling necessary to crystalize sieved plagioclase. Because at least 5% of the plagioclase is sieved in 75% of the studied rocks, understanding its petrogenetic connotations is obviously important. Large amounts (\geq 50%) of sieved plagioclase, however, are only observed in some basaltic andesites and in Holocene South Sister rhyolites. Previous studies by Hughes (1983) and Clark (1983) have indicated that basaltic andesites in this part of the High Cascades can be generated through a variety of open-system magmatic processes. This model will also be supported in the following chapter. It is thus not surprising that basaltic andesites in the study area contain abundant sieved plagioclase, which was likely produced through magma mixing.

However, roughly 90% of the plagioclase in Holocene South Sister rhyolites is sieved, but less than 20% of the plagioclase in all other Three Sisters rhyolites is sieved (figure 4.4C-D). This is significant because Holocene South Sister rhyolites have nearly the same composition as other Three Sisters rhyolites, and minor compositional differences between these rhyolites types cannot be attributed to mixing of compositionally distinct magmas (Chapter 9). It is possible that chemically similar but thermally distinct rhyolitic magmas were mixed to produce the necessary undercooling for sieved plagioclase crystallization in the Holocene units, as temperature drops of only about 100°C have been shown to produce sieved plagioclase (Lofgren, 1974; Swanson, 1977). Although work by Nelson and Montana (1989) has shown that plagioclase resorption (i.e., fritted) in anhydrous basaltic melts can be induced by drops in pressure of 2-6 kbar, the effects of decompression on plagioclase morphology in hydrous silicic systems have not been studied. However, the relatively lower position of the plagioclase liquidus under hydrous conditions (Kushiro, 1979; Sekine et al., 1979) may result in the formation of sieved plagioclase with isothermal decompression. Sieved plagioclase in the Holocene South Sister rhyolites thus may have recorded the recharge of a Holocene magma chamber by a fresh batch of cooler but compositionally similar rhyolitic magma, or may indicate that an episode of decompression occurred in the Holocene magma chamber.

4.2 Pyroxenes

Pyroxenes are the most abundant ferromagnesian silicate minerals in the central High Cascades, and an anhydrous, two-pyroxene assemblage characterizes the mineralogy in the study area for rocks with 55% to 72% SiO₂. Abundances of pyroxene phenocrysts, which were determined through visual estimation of thin sections, typically range from 1-5% for mafic rocks, 3-10% for intermediate rocks, and 1-3% for silicic rocks. Petrographic data for the occurrence of pyroxenes are summarized in figure 4.5. Clinopyroxene (augite) is the most abundant pyroxene in rocks having 53% to \approx 58% SiO₂, with the exception of a Middle Sister basaltic andesite (3S030). Augite is essentially a trace phase by 72% SiO₂, which is where the ferromagnesian assemblage changes to orthopyroxene \pm amphibole. Pyroxenes from 27 representative mafic to silicic rocks from the study area were analyzed with the electron microprobe (Appendix C-2), and their compositional range is shown in figure 4.6.

4.2.1 Clinopyroxene

Phenocrysts of clinopyroxene (augite) are commonly coarser-grained (1-2 mm) and more irregularly formed (subhedral to anhedral) than orthopyroxene. Augite compositions are plotted in



Figure 4.4: C) Non-polarized and D) plane-polarized photomicrographs of Holocene South Sister rhyolite 3S122. Plagioclase phenocrysts in Holocene Three Sisters rhyolites have complex normal oscillatory zoning ($\approx An_{40-30}$) and contain abundant coarse glass inclusions. However, Pleistocene Three Sisters rhyolites typically contain <10% sieved plagioclase but display the same complex normal oscillatory zoning ($\approx An_{40-30}$) as the Holocene rhyolites.



Figure 4.5: Visual estimates of clinopyroxene and orthopyroxene modal abundances versus silica content of the host rock. The high content of opx in basaltic andesite 3S030 is due to abundant 1-2mm sub-ophitic opx in the sample.





Figure 4.6: Compositional range of clinopyroxene and orthopyroxene for all samples, plotted on part of the Wo-En-Fs ternary. Compositions in mol % Wo and Fs. Samples grouped according to range in SiO₂ of the host rock.

\$

figure 4.6, and representative augite analyses are given in table 4.2. Augite compositions tend to be less magnesian with increasing SiO₂ in the host rock, and range from $\approx En_{50}$ at 53% SiO₂ (3S049) to $\approx En_{34}$ at 72% SiO₂ (3S098). Visible sector zoning and concentric zoning is scarce in augites from mafic rocks. Core-to-rim zonation in En content of more than $\approx 10\%$ was not found in any of the augites (figure 4.7; augites $> \approx 50\%$ En). Nonetheless, small degrees of zonation are probably significant, owing to the limited amount of variability that can occur in iron and magnesium abundances in the pyroxene crystal lattice while calcium content remains relatively constant (i.e. 16-19%). Augite rim compositions that differ by at least 5% from core compositions probably represent non-equilibrium crystallization and may be evidence of open system differentiation, although no apparent correlation between degree of zoning and either rock composition or augite composition was observed.

Some of the augites in numerous intermediate composition rocks (Appendix B) have thin overgrowth rims of a finely crystalline, low birefringence mineral that is probably a low-Ca pyroxene. Microprobe analysis of rimmed augites in dacites from South Sister (3S061) and Broken Top (3S115) determined that these overgrowths are pigeonite. A South Sister andesite (3S112) also had several phenocrysts of hypersthene with thin rims of pigeonite. The transition from augiteorthopyroxene to augite-pigeonite generally occurs at either temperatures greater than ≈1000°C for these composition rocks (Lindsley, 1983), or in the more iron-enriched (calcium-poor) members of a comagmatic suite (cf. Huebner, 1980). The pigeonite rims probably formed on pyroxene xenocrysts that were introduced into a more silicic (i.e. calcium-poor) melt through magma mixing. As rimmed pyroxenes do not normally develop during closed-system fractionation (e.g. Huebner, 1980), their occurrence is another petrographic indicator of open-system magmatic processes.

4.2.2 Orthopyroxene

Subhedral to euhedral phenocrysts of orthopyroxene generally range up to 2 mm in length, and are the dominant ferromagnesian mineral in most rocks with $SiO_2 \ge 64\%$ (figure 4.5). Orthopyroxene compositions are also plotted in figure 4.6, and representative orthopyroxene compositions are in table 4.2. Analyzed orthopyroxenes show an overall decrease in En content with increasing rock SiO_2 , and range from $\approx En_{74}$ at 58% (3S003) to $\approx En_{39}$ at 74% SiO_2 (BPb). Unlike augites, orthopyroxene compositions in rhyolites are distinct for the Broken Top and Three Sisters magmatic systems (figure 4.6). The orthopyroxenes in Broken Top rhyolites are ferrohypersthenes with Wo contents that are typical for High Cascade orthopyroxenes (Wo₃), while Three Sisters rhyolites contain hypersthenes with consistently low Wo contents (Wo₂). It is not surprising that ferrohypersthenes are found in Broken Top rhyolites, as these rocks are highly iron enriched: Fe' [= molar Fe/(Fe + Mg)] ≥ 0.9 ; Three Sisters rhyolites have Fe' ≤ 0.9 and thus contain hypersthene. Lindsley (1983) observed that with constant rock composition and decreasing

| | | | Cl: | inopyrox | enes | | |
|--|---|--|---|--|---|---|---|
| Sample Rx SiO2 | 3S049b 53.0 | 3S045a 56.0 | 3S028b 58.5 | 3S112b 60.0 | 3S115b 62.9 | 3S134c 65.7 | 35098c 71.1 |
| SiO2 TiO2 Al2O3 FeO Fe2O3 MgO MnO CaO Na2O Cr2O3 TOTAL | 51.94 0.48 2.96 6.00 0.71 17.44 0.21 18.34 0.38 <u>0.25</u> 98.71 | $51.23 \\ 0.82 \\ 3.06 \\ 10.54 \\ 0.00 \\ 14.47 \\ 0.30 \\ 17.45 \\ 0.47 \\ 0.03 \\ 98.36 $ | 50.83 0.64 1.93 9.26 1.84 14.23 0.64 19.00 0.38 0.01 98.75 | 52.88 0.75 3.12 13.08 0.00 12.73 0.39 16.77 0.78 0.03 100.53 | $52.44 \\ 0.64 \\ 1.65 \\ 9.71 \\ 0.74 \\ 15.09 \\ 0.33 \\ 19.29 \\ 0.35 \\ <0.01 \\ 100.22$ | 51.150.411.3713.280.9512.380.7818.600.340.0199.26 | 51.950.260.6214.410.3511.790.7119.330.31 $<0.0199.73$ |
| Wo En Fs | 38.7 51.1 10.2 | 37.9 43.7 18.4 | 40.8 42.6 16.6 | 37.3 39.4 23.4 | 40.1 43.6 16.3 | 39.7 36.8 23.5 | 40.7 34.5 24.8 |
| Sample Rx SiO2 | 3S028c 58.5 | 35112c 60.0 | Or 35115a 62.9 | thopyrox 3S134b 65.7 | enes 3S098b 71.1 | BPb 74.0 | 3S058a 73.7 |
| SiO2 TiO2 Al2O3 FeO Fe2O3 MgO MnO CaO Na2O Cr2O3 TOTAL | 51.97 0.27 0.69 14.64 5.01 25.12 0.90 1.45 0.04 <u>0.02</u> 100.12 | $52.14 \\ 0.54 \\ 0.53 \\ 23.38 \\ 0.44 \\ 20.25 \\ 0.63 \\ 1.92 \\ 0.06 \\ < 0.01 \\ 99.89 \\ \end{cases}$ | $53.94 \\ 0.32 \\ 0.98 \\ 18.36 \\ 0.34 \\ 24.68 \\ 0.52 \\ 1.37 \\ 0.04 \\ 0.04 \\ 100.58 $ | $51.550.210.6726.720.0617.731.221.610.05\leq 0.0199.81$ | 51.02 0.20 0.40 28.56 0.30 16.11 1.40 1.85 0.03 < <u>0.01</u> 99.88 | $\begin{array}{c} 49.07\\ 0.10\\ 0.22\\ 31.78\\ 2.24\\ 12.97\\ 1.98\\ 1.29\\ 0.05\\ \underline{<0.01}\\ 99.70\end{array}$ | 52.43 0.12 0.39 24.43 0.00 20.01 1.17 0.92 0.01 <u>0.02</u> 99.50 |
| Wo En Fs | 2.99 72.00 25.01 | 3.9 57.7 38.4 | 2.7 68.1 29.2 | 3.3 51.3 45.4 | 3.9 47.0 49.1 | 2.8 39.5 57.7 | 1.9 57.1 41.0 |
| | 35131 (3) | 3S061b Core | 3SO61b Rim | Pigeonit 3S112d Core | es 3S112d Rim | 3S115c Core | 3S115c Rim |
| SiO2 TiO2 Al2O3 FeO Fe2O3 MgO MnO CaO Na2O Cr2O3 TOTAL | $52.30 \\ 0.45 \\ 0.66 \\ 21.14 \\ 0.27 \\ 19.82 \\ 0.58 \\ 4.38 \\ 0.06 \\ 0.02 \\ 99.68 \\ \end{array}$ | $52.30 \\ 0.52 \\ 1.67 \\ 7.34 \\ 1.62 \\ 15.47 \\ 0.37 \\ 20.42 \\ 0.34 \\ \underline{< 0.01} \\ 100.05 \\ \end{array}$ | 51.67 0.40 0.37 23.66 0.36 17.95 1.06 3.93 0.08 < <u>0.01</u> 99.48 | 53.880.300.8617.041.4825.040.511.890.020.02101.03 | 52.22 0.62 0.85 20.15 0.00 18.11 0.59 7.14 0.12 <u>0.03</u> 99.82 | 52.510.661.819.540.8815.140.4519.380.340.340.03100.73 | 52.240.360.4023.260.0018.260.993.990.100.0499.62 |
| Wo En Fs | 9.0 56.4 34.7 | 42.6 44.9 12.6 | 8.2 51.8 40.1 | 3.8 69.1 27.2 | 14.7 51.9 33.4 | 40.2 43.7 16.2 | 8.2 52.6 39.2 |

Table 4.2: Representative pyroxene microprobe analyses. Rx SiO2 refers to SiO2 in host rock,
mol% Wo, En and Fs. Pigeonite sample 3S131 is an average of 3 30-50 μ groundmass
crystals, others samples are single-crystal rim compositions unless otherwise stated.



Figure 4.7: Range of core-to-rim zoning measured in pyroxene phenocrysts. Zoning is expressed as the percent change from core to rim En (mol %) content, with negative numbers indicating reversed zoning. Clinopyroxenes are generally < En_{so}, and orthopyroxenes are generally > En_{so}.

temperature, the Ca content of orthopyroxene will decrease while the Ca content of augite will increase. The anomalously low Wo contents of hypersthene in Three Sisters rhyolites suggests that these minerals may have crystallized at lower temperatures than Broken Top system ferrohypersthenes.

In general, analyzed orthopyroxenes are more strongly zoned with respect to Fe & Mg than augites, and have core-to-rim changes in En content of up to 25% (figure 4.7). There is also a general tendency, which was not observed for augites, for the strongest zonation in Fe & Mg to occur in more magnesian orthopyroxenes. This is probably due to the greater range in iron and magnesium contents that can be accommodated in orthopyroxene, relative to augite, for the same series of rocks (i.e., En_{74-39} vs. En_{52-34}).

4.2.3 Pigeonite

Pigeonite (i.e. Wo_5 to Wo_{15}) has been recognized through microprobe analysis as both a stable groundmass mineral in basaltic andesite 3S131, and as occasional 5-15 μ rims around other pyroxenes, as stated above. Representative pigeonite analyses are in table 4.2. Basaltic andesite 3S131 is slightly iron enriched (Fe' = 0.58) relative to other High Cascade basaltic andesites (Fe' \approx 0.4-0.5) (Hughes, 1983; this study). As pigeonite is generally the stable low-calcium pyroxene phase for mafic magmas with Fe' greater than \approx 0.45 (Huebner, 1980), it is not surprising that the pyroxene assemblage in 3S131 is augite + pigeonite.

4.3 Olivine

Phenocrysts of olivine range in abundance from 5% to 10% in basalts, from 2% to 5% in most basaltic andesites, and $\leq 1\%$ in some andesites (Appendix B). Olivine phenocrysts are generally subhedral and $\leq 2mm$ in diameter, and euhedral phenocrysts are only observed in some basalts. Xenocrysts of olivine are found in several dacites and rhyodacites, and are always embayed and rimmed with orthopyroxene.

Opacitic olivine was found in some vesicular mafic rocks. The formation of opacitic olivine is a deuteric process involving the oxidation of olivine to magnetite \pm enstatite at low pressures (Haggerty & Baker, 1967), and does not indicate that olivine was out of equilibrium with the magma. Microprobe analysis of olivine in basalt 3S049 (table 4.3) shows that fresh phenocrysts of \approx Fo₈₀ olivine alter to Fo₉₅ olivine (3S049b) and a symplectite of magnetite + orthopyroxene (3S049b opacite). Lesser degrees of deuteric alteration (\pm hydration) results in the formation of iddingsite on olivine rims and fractures, which is commonly observed in mafic rocks.

| | SiO2 | FeO | MgO | MnO | CaO | NiO | Total | Fo | Fa | Mo | Li | Те | FM |
|--------------------------|-------|-------|-------|------|------|------|--------|-------|-------|------|------|------|------|
| 3S003a core | 37.15 | 25.12 | 36.24 | 0.46 | 0.13 | 0.00 | 99.09 | 71.76 | 27.91 | 0.19 | 0.00 | 0.52 | 0.28 |
| 3SOO3a rim | 37.92 | 24.99 | 35.55 | 0.42 | 0.16 | 0.01 | 99.05 | 71.54 | 28.21 | 0.23 | 0.01 | 0.48 | 0.29 |
| 3S003b core | 37.34 | 25.53 | 35.98 | 0.42 | 0.15 | 0.00 | 99.41 | 71.34 | 28.40 | 0.21 | 0.00 | 0.47 | 0.29 |
| 3S003b rim | 38.08 | 24.78 | 37.77 | 0.30 | 0.19 | 0.02 | 101.14 | 73.05 | 26.88 | 0.26 | 0.02 | 0.33 | 0.27 |
| $3S017a$ $150\mu m$ core | 37.81 | 21.81 | 39.01 | 0.35 | 0.16 | 0.10 | 99.24 | 75.99 | 23.84 | 0.22 | 0.10 | 0.39 | 0.24 |
| 3S017a rim | 37.09 | 25.57 | 36.08 | 0.38 | 0.15 | 0.14 | 99.39 | 71.40 | 28.39 | 0.21 | 0.14 | 0.42 | 0.29 |
| 35028 core | 37.33 | 25.49 | 35.81 | 0.45 | 0.17 | 0.04 | 99.28 | 71.26 | 28.46 | 0.24 | 0.04 | 0.51 | 0.29 |
| 35028 rim | 36.87 | 27.17 | 37.34 | 0.37 | 0.16 | 0.00 | 101.91 | 70.88 | 28.94 | 0.22 | 0.00 | 0.40 | 0.29 |
| 3S028 rim 2nd point | 38.52 | 24.88 | 36.57 | 0.37 | 0.21 | 0.04 | 100.59 | 72.29 | 27.59 | 0.30 | 0.04 | 0 42 | 0.29 |
| $3S045$ $120\mu m$ euhed | 36.84 | 26.44 | 35.01 | 0.54 | 0.16 | 0.08 | 99.06 | 69.96 | 29.65 | 0.22 | 0.09 | 0.62 | 0.20 |
| $3S045 60 \mu m$ subh | 36.88 | 25.74 | 35.46 | 0.57 | 0.19 | 0.04 | 98.88 | 70.79 | 28.84 | 0.27 | 0.05 | 0.65 | 0.29 |
| 35045a 250µm core | 37.73 | 22.04 | 38.73 | 0.44 | 0.22 | 0.06 | 99.22 | 75.67 | 24.16 | 0.31 | 0.06 | 0.49 | 0.25 |
| 35049 30µm GM | 35.95 | 31.72 | 30.47 | 0.53 | 0.24 | 0.00 | 98.91 | 62.96 | 36.77 | 0.35 | 0 00 | 0.62 | 0.25 |
| 35049 50µm GM | 35.77 | 33.33 | 29.82 | 0.67 | 0.19 | 0.01 | 99.79 | 61.15 | 38.35 | 0.28 | 0.01 | 0.78 | 0.37 |
| 3SO49a core | 38.62 | 16.99 | 42.67 | 0.31 | 0.16 | 0.19 | 98.94 | 81.64 | 18.24 | 0.22 | 0.20 | 0.34 | 0.19 |
| 35049a rim | 36.47 | 27.92 | 33.36 | 0.45 | 0.17 | 0.05 | 98.41 | 67.85 | 31.87 | 0.24 | 0.05 | 0.53 | 0.32 |
| 3SO49b By opacite | 41.40 | 5.24 | 53.14 | 0.19 | 0.07 | 0.12 | 100.16 | 94.66 | 5.24 | 0.09 | 0.11 | 0.20 | 0.05 |
| 3S049b Opacite | 37.61 | 32.80 | 26.81 | 0.33 | 0.62 | 0.18 | 98.35 | 59.64 | 40.94 | 0.99 | 0.22 | 0.42 | 0.41 |
| 3S049c core | 38.79 | 17.29 | 42.88 | 0.23 | 0.17 | 0.22 | 99.59 | 81.55 | 18.44 | 0.24 | 0.23 | 0.25 | 0.19 |
| 3S049c rim | 38.65 | 18.76 | 40.84 | 0.33 | 0.16 | 0.11 | 98.85 | 79.41 | 20.46 | 0.23 | 0.11 | 0.37 | 0.21 |
| 3SO69a 200 μ m core | 39.91 | 14.93 | 45.13 | 0.29 | 0.18 | 0.15 | 100.60 | 84.29 | 15.65 | 0.24 | 0.16 | 0.31 | 0.16 |
| 3S069a rim | 39.46 | 15.49 | 44.56 | 0.38 | 0.10 | 0.12 | 100.10 | 83.45 | 16.28 | 0.13 | 0.12 | 0.40 | 0.17 |
| 3S069b 200µm core | 40.22 | 12.68 | 46.89 | 0.24 | 0.14 | 0.34 | 100.50 | 86.77 | 13.16 | 0.18 | 0.34 | 0.25 | 0.13 |
| 3S069b rim | 39.52 | 15.76 | 44.10 | 0.33 | 0.13 | 0.17 | 100.01 | 83.14 | 16.67 | 0.17 | 0.17 | 0.36 | 0.17 |
| 3S074 30µm opacitic rim | 37.70 | 23.20 | 37.68 | 0.55 | 0.10 | 0.08 | 99.31 | 73.97 | 25.55 | 0.14 | 0.08 | 0.61 | 0.26 |
| 3S074a core | 39.41 | 17.11 | 42.75 | 0.35 | 0.15 | 0.27 | 100.03 | 81.52 | 18.30 | 0.20 | 0.28 | 0.38 | 0.19 |
| 3S074a rim | 39.17 | 16.55 | 42.28 | 0.29 | 0.18 | 0.16 | 98.63 | 81.93 | 18.00 | 0.25 | 0.17 | 0.32 | 0.18 |
| 35090a 120µm core | 37.50 | 23.86 | 36.77 | 0.40 | 0.16 | 0.07 | 98.75 | 73.14 | 26.63 | 0.22 | 0.07 | 0.46 | 0.27 |
| 35090a rim | 35.74 | 33.97 | 28.92 | 0.56 | 0.23 | 0.08 | 99.50 | 60.08 | 39.61 | 0.35 | 0.09 | 0.66 | 0.40 |
| 35090b 300µm core | 37.17 | 25.65 | 35.90 | 0.44 | 0.15 | 0.10 | 99.41 | 71.18 | 28.54 | 0.21 | 0.11 | 0.49 | 0.29 |
| 35090b rim | 35.98 | 30.27 | 32.11 | 0.52 | 0.20 | 0.03 | 99.11 | 65.20 | 34.49 | 0.29 | 0.03 | 0.60 | 0.35 |
| 3S131 150µm core | 35.05 | 39.89 | 24.84 | 0.61 | 0.25 | 0.10 | 100.72 | 52.42 | 47.23 | 0.38 | 0.11 | 0.73 | 0.48 |
| 3S131a 100µm core | 35.75 | 35.87 | 28.24 | 0.51 | 0.24 | 0.00 | 100.61 | 58.25 | 41.51 | 0.36 | 0.00 | 0.59 | 0 42 |
| 3S131a rim | 35.24 | 39.05 | 25.56 | 0.68 | 0.22 | 0.05 | 100.81 | 53.59 | 45.94 | 0.33 | 0.06 | 0.81 | 0.47 |
| 3S131b 15M GM | 34.16 | 44.49 | 20.15 | 0.71 | 0.31 | 0.03 | 99.85 | 44.49 | 55.11 | 0.49 | 0.04 | 0.89 | 0.56 |
| 3S131b rim | 33.91 | 45.61 | 20.02 | 0.80 | 0.31 | 0.02 | 100.66 | 43.67 | 55.82 | 0.48 | 0.02 | 0.99 | 0.57 |
| | | | | | | | | | | | | | / |

Table 4.3: Olivine compositions determined through microprobe analysis.FM = Molar Fe / (Fe + Mg + Mn); mol %'s Fo, Fa, Mo, Li and Te.Opacite refers to a symplectic intergrowth of magnetite ± enstatite produced through deuteric alteration.

Olivine phenocrysts in basaltic andesites and andesites commonly have embayed rims, which are either opacitic or rimmed by $a \le 0.1$ mm zone of orthopyroxene. Embayed rims in olivine do not indicate resorption, but are probably a growth feature characteristic of an increased cooling rate in the magma (Donaldson, 1976). Orthopyroxene rims formed subsequent to embayment, due to the reaction of a silica-undersaturated phase with a more silicic melt. Embayed and rimmed olivines are thus one of the expected products of magma mixing.

Olivine compositions determined through microprobe analysis ranged from Fo_{87} chrysolite in basalt 3S069 to $=Fo_{45}$ hortonolite in basaltic andesite 3S131 (table 4.3). Forsterite content generally decreases with increasing silica in the host rock (figure 4.8), although unusually Fe-rich olivines were found in basaltic andesites 3S049 and 3S131; Olivine in most calc-alkaline basaltic andesites is generally between Fo_{85} and Fo_{70} (Ewart, 1982). The high Fe content of olivine (Fo₄₅) in basaltic andesite 3S131 reflects the anomalously high Fe content of the rock (Fe' = 0.58) compared to average High Cascade basaltic andesites (Fe' \leq 0.5; Hughes, 1983).

4.4 Iron-Titanium Oxides

Phenocrysts to microphenocrysts of opaque minerals from the system FeO-Fe₂O₃-TiO₂, which will be referred to as Fe-Ti oxides, occur in all of the studied units. Fe-Ti oxide abundances typically range from 2-5% in mafic rocks, 1-3% in intermediate units, and $\leq 1\%$ in silicic rocks. The Fe-Ti oxides commonly occur as inclusions in all other major phases, and phenocryst sizes typically range to 0.5 mm. Analyses of representative Fe-Ti oxides from the study area are in table 4.4. Microprobe analysis of 27 representative mafic to silicic rocks from the study area (Appendices C-4 and C-5) confirmed the ubiquitous occurrence of titanomagnetite (magnetite-ülvospinel solid solution) when opaque minerals are present. The occurrence of ilmenite (ilmenite-hematite solid solution) was restricted to samples with SiO₂ \geq 59%.

Exsolution lamellae are present in titanomagnetites from ignimbrites 3S005 and the Lava Island Tuff, rhyolite dome 3S042, and in ilmenites from rhyodacite dome 3S009. All of the units with exsolved Fe-Ti oxides also have light to moderate deuteric vapor-phase(?) alteration, indicating that the exsolution lamellae were produced through oxidation and not through slow cooling (Buddington & Lindsley, 1964). Because oxidation-induced exsolution can alter the Ti/Fe ratio in iron-titanium oxides (Furuta et al., 1985), these oxides with exsolution lamellae do not represent magmatic compositions.

Coexisting titanomagnetite-ilmenite pairs from most units have Mg/Mn ratios that plot within the 2σ limits of equilibrium distribution of Bacon & Hirschmann (1988) (figure 4.10). This equilibrium relationship has been calibrated with high-silica rhyolite data, but the compositions and temperatures of formation of the Fe-Ti oxides in this study are comparable to the data used by



Figure 4.8: Compositional range of olivine phenocrysts and microlites versus silica content of the host rock. Note that high Fe olivine (hortonolite) is associated with unusually Fe-rich basaltic andesite 3S131. Fe-rich olivines in 3S049 are groundmass microlites.

| Sample # Rx SiO2 | 35049 53.0 | 3S045 56.0 | 3S028 58.5 | 3S123 63.5 | 3S098 71.1 | 3S058 73.7 | BP-b 74.0 |
|---------------------|---------------|---------------|---------------|---------------|---------------|---------------|---------------|
| SiO2 TiO2 | 0.07 | 0.10 | 0.06 13.66 | 0.05 14.49 | 0.08 17.32 | 0.09 8.49 | 0.11 14.42 |
| A12O3 | 0.77 | 0.57 | 2.24 | 2.17 | 1.30 | 1.82 | 1.35 |
| FeO | 35.35 | 44.04 | 40.23 | 42.41 | 45.38 | 37.58 | 43.24 |
| Fe2O3 | 48.22 | 36.84 | 41.27 | 39.61 | 35.73 | 51.29 | 40.31 |
| MnO | 0.28 | 0.34 | 0.60 | 0.47 | 0.72 | 0.54 | 0.93 |
| MgO | 2.57 | 0.30 | 2.18 | 1.42 | 1.15 | 0.86 | 0.50 |
| Cr203 | 0.02 | _0.02 | 0.03 | 0.00 | 0.01 | 0.07 | 0.11 |
| TOTAL | 97.28 | 97.73 | 100.27 | 100.63 | 101.68 | 100.74 | 100.97 |
| XÜsp | 29.2 | 45.7 | 38.2 | 40.6 | 48.3 | 24.3 | 41.0 |

Fe-Ti-Cr spinels

| Sample # Rx SiO2 | 35030 54.3 | 3S069 51.8 | 3S069 51.8 | 3S112 60.0 | 3S045 56.0 |
|---|---|---|--|---|--|
| SiO2 TiO2 Al2O3 FeO Fe2O3 MnO MgO Cr2O3 TOTAL | $\begin{array}{c} 0.05 \\ 18.39 \\ 1.40 \\ 44.13 \\ 31.01 \\ 0.44 \\ 2.18 \\ 0.96 \\ 98.56 \end{array}$ | $\begin{array}{c} 0.07 \\ 1.78 \\ 3.51 \\ 27.85 \\ 50.79 \\ 0.43 \\ 3.19 \\ \underline{11.52} \\ 99.14 \end{array}$ | 0.04 1.44 2.26 27.06 61.95 0.51 3.27 <u>2.84</u> 99.37 | $\begin{array}{c} 0.07 \\ 20.46 \\ 1.53 \\ 47.45 \\ 24.66 \\ 0.38 \\ 1.20 \\ \underline{2.15} \\ 97.89 \end{array}$ | 0.07 2.03 20.60 21.50 20.02 0.32 9.38 <u>24.04</u> 97.95 |
| XÜsp | 52.1 | 5.2 | 4.2 | 58.7 | 5.2 |
| | | | | | |
| | | | I | lmenites | |
| Sample # Rx SiO2 | 3S028 58.5 | 3S123 63.5 | I 35098 71.1 | 1menites 3S058 73.7 | BPb-a 74.0 |

16.63 14.18 8.39 18.44

XHem

Table 4.4: Representative Fe-Ti oxide compositions, and compositions of unusually Cr-rich (Cr₂O₃ ≥1 wt%) Fe-Ti oxides, determined by microprobe analysis. RxSiO2 refers to SiO₂ in host rock. Ferric-ferrous iron partitioning, mole fractions ülvospinel (XÜsp) and hematite (XHem) calculated according to Carmichael (1967). Sample locations in appendix D-1.

6.12



Figure 4.9: Molar Mg/Mn ratios for coexisting magnetite-ilmenite pairs. Equilibrium between magnetite and ilmenite is indicated by Mg/Mn ratios that plot within the 2 σ error limits of Bacon and Hirschmann (1988); samples 3S003 and 3S115 are clearly out of equilibrium.

Bacon & Hirschmann (1988). Although a slight compositional dependence on Mg/Mn partitioning has been suggested by Grunder & Mahood (1988), the distribution of most of the analyses from this study within the 2σ equilibrium limits strongly indicates that these titanomagnetite-ilmenite pairs are in equilibrium. Fe-Ti oxides from dacite 3S115 and andesite 3S003, however, are strongly out of equilibrium with respect to Mg/Mn abundances, which suggests that one or both of these phases are xenocrysts.

Chromian titanomagnetites (cf., Haggerty, 1976) were analyzed in basalt 3S069, basaltic andesite 3S030, and andesite 3S112 (table 4.4). Chromian titanomagnetite phenocrysts are commonly observed in relatively non-evolved (i.e. high Cr, Ni, Co) basalts (Haggerty, 1976). Basalt 3S069 has a very high Cr abundance (415 ppm), which supports petrographic evidence that chromian titanomagnetite was a phenocryst in this unit. Bulk rock samples 3S030 and 3S112 both have Cr ≤ 100 ppm, indicating a more evolved composition than 3S069. Other analyzed Cascade mafic units with Cr ≤ 100 ppm contain titanomagnetite as the main Fe-Ti oxide, indicating that the chromian titanomagnetites in 3S030 and 3S112 are likely xenocrysts from less evolved basaltic melts.

An inclusion of titanian chromite (table 4.4) was also found in a plagioclase glomerocryst (An_{85}) from basaltic andesite 3S045. Both the unusually high calcium content of the plagioclase and the titanian chromite inclusion strongly indicate that this glomerocryst was derived from a composition that is much more primitive than observed in exposed rocks in this part of the High Cascades.

4.5 Amphibole

Trace amounts of amphibole phenocrysts are found in about half of the studied rhyolites and occasionally in some rhyodacites (Appendix B). Amphibole phenocrysts range in size up to 5 mm, are generally euhedral, and have brown to brownish-green pleochroism. Glomerocrysts containing partially to wholly opacitized amphibole, and fine-grained glomerocrysts of orthopyroxene + opaque minerals pseudomorphically replacing amphibole (i.e. ghosts), are also found in three basaltic andesites and a dacite (Appendix B). As amphibole phenocrysts have only been observed in rhyolites, discrete amphibole "ghosts" are probably direct evidence of magma mixing with rhyolitic melts.

Amphiboles from 9 units (55 points) were analyzed on the microprobe, and representative analyses are summarized in table 4.5. Cation distributions were calculated on the basis of 23 oxygens according to the method of Spear and Kimball (1984), using the average of the highest and lowest acceptable values for Fe^{3+} to calculate the amphibole formula. Following the terminology of Leake (1978), all amphiboles in the study area are calcic and have a general composition of Magnesian Hastingsite, but range in composition to Edenite and Pargasitic Hornblende (table 4.5). There is a surprising degree of variation in amphibole composition; amphiboles from individual

| Label | 3 S053-A core 53A | 3SO53-A rim 53a | 3S053-B core 53B | 3S053-B rim 53b | BPT-C core BC | BPT-C rim Bc | 3 5060-A core 60A | 35060-A rim 60a |
|--|---|--|---|--|--|--|---|--|
| SiO2 TiO2 A12O3 FeO Fe2O3 FeO* MgO MnO CaO Na2O K2O H2O F- C1- Total | $\begin{array}{r} 44.18\\ 1.47\\ 9.94\\ 12.74\\ 3.32\\ 15.73\\ 13.10\\ 0.33\\ 10.77\\ 1.57\\ 0.34\\ 1.74\\ 0.54\\ \hline 0.09\\ 100.14 \end{array}$ | $\begin{array}{r} 43.45\\ 1.81\\ 9.28\\ 10.19\\ 6.10\\ 15.68\\ 13.71\\ 0.20\\ 10.94\\ 1.81\\ 0.43\\ 1.88\\ 0.21\\ 0.08\\ 100.10\\ \end{array}$ | $\begin{array}{r} 47.80\\ 1.30\\ 5.57\\ 14.74\\ 2.45\\ 16.95\\ 13.48\\ 0.48\\ 10.87\\ 0.98\\ 0.74\\ 1.96\\ 0.04\\ 0.19\\ 100.59\end{array}$ | 47.29 1.98 8.03 15.77 0.00 15.77 13.12 0.28 10.71 1.56 0.41 1.92 0.24 <u>0.09</u> 101.41 | $\begin{array}{r} 42.15\\ 3.29\\ 10.00\\ 15.92\\ 3.62\\ 19.18\\ 10.81\\ 0.52\\ 10.17\\ 2.22\\ 0.29\\ 1.88\\ 0.21\\ 0.04\\ 101.10\end{array}$ | $\begin{array}{r} 43.20\\ 3.13\\ 10.05\\ 19.14\\ 0.00\\ 19.14\\ 10.06\\ 0.52\\ 10.06\\ 2.07\\ 0.28\\ 1.77\\ 0.28\\ 1.77\\ 0.46\\ \underline{0.05}\\ 100.79\end{array}$ | 42.68 3.17 11.77 11.61 0.97 12.48 13.81 0.16 11.24 1.89 0.43 1.92 0.24 0.02 99.90 | 41.41 3.78 11.51 8.73 5.70 13.86 13.28 0.19 11.20 2.01 0.36 1.93 0.17 <u>0.01</u> 100.29 |
| | Cat | ions on | the basi | s of 23 | Oxygen, | average | Fe3+ | |
| Si Al(TV) | 6.53 | 6.42 | 7.04 | 6.89 | 6.29 | 6.48 | 6.27 | 6.08 |
| m (1 v) | 1.4/ | 1.00 | 0.90 | 1.12 | ±•/± | T. 72 | 1.75 | 1.92 |
| Ti Al(VI) Fe2+ Fe3+ Mg Mn Ca Na | 0.16 0.26 1.57 0.37 2.89 0.04 1.71 0.00 | 0.20 0.03 1.27 0.67 3.02 0.03 1.73 0.06 | 0.14 0.01 1.81 0.28 2.96 0.06 1.72 0.02 | 0.22 0.26 1.92 0.00 2.85 0.04 1.67 0.05 | 0.37 0.05 1.99 0.41 2.10 0.07 1.63 0.09 | 0.35 0.26 2.40 0.00 2.25 0.07 1.62 0.05 | 0.35 0.31 1.43 0.11 3.02 0.02 1.77 0.00 | 0.42 0.07 1.07 0.64 2.90 0.02 1.76 0.13 |
| Ti Al(VI) Fe2+ Fe3+ Mg Mn Ca Na Na Na(IV) K | 0.16 0.26 1.57 0.37 2.89 0.04 1.71 0.00 0.45 0.06 | 0.20 0.03 1.27 0.67 3.02 0.03 1.73 0.06 0.46 0.08 | 0.14 0.01 1.81 0.28 2.96 0.06 1.72 0.02 0.26 0.14 | 0.22 0.26 1.92 0.00 2.85 0.04 1.67 0.05 0.39 0.08 | 0.37 0.05 1.99 0.41 2.10 0.07 1.63 0.09 0.55 0.06 | 0.35 0.26 2.40 0.00 2.25 0.07 1.62 0.05 0.55 0.05 | 0.35 0.31 1.43 0.11 3.02 0.02 1.77 0.00 0.54 0.08 | 0.42 0.07 1.07 0.64 2.90 0.02 1.76 0.13 0.45 0.07 |
| Ti Al(VI) Fe2+ Fe3+ Mg Mn Ca Na Na(IV) K OH Cl F | 0.16 0.26 1.57 0.37 2.89 0.04 1.71 0.00 0.45 0.06 1.73 0.02 0.25 | 0.20 0.03 1.27 0.67 3.02 0.03 1.73 0.06 0.46 0.08 1.88 0.02 0.10 | 0.14 0.01 1.81 0.28 2.96 0.06 1.72 0.02 0.26 0.14 1.93 0.05 0.02 | 0.22 0.26 1.92 0.00 2.85 0.04 1.67 0.05 0.39 0.08 1.87 0.02 0.11 | 0.37 0.05 1.99 0.41 2.10 0.07 1.63 0.09 0.55 0.06 1.89 0.01 0.10 | 0.35 0.26 2.40 0.00 2.25 0.07 1.62 0.05 0.55 0.05 1.77 0.01 0.22 | 0.35 0.31 1.43 0.11 3.02 0.02 1.77 0.00 0.54 0.08 1.88 0.01 0.11 | 1.92 0.42 0.07 1.07 0.64 2.90 0.02 1.76 0.13 0.45 0.07 1.92 0.00 0.08 |

Amphibole types

3S053-A Edenitic Hornblende, Holocene Devils Hill pumice. 3S053-A Magnesian Hastingsitic Hornblende. 3S053-B Magnesio-Hornblende. 3S053-B Edenite. BPT-C Magnesian Hastingsitic Hornblende, top of Bend Pumice. BPT-C Edenitic Hornblende. 3S060-A Pargasitic Hornblende, Holocene Rock Mesa pumice. 3S060-A Magnesio-Hastingsite.

Table 4.5: Representative microprobe analyses of amphiboles from rhyolites. Recalculations using the method of Spear & Kimball (1984), using average values of Fe³⁺ on a 23-O basis; water calculated on a 15 cation, 2(OH, F, Cl) basis. Amphibole types from Leake (1978).

Three Sisters units generally have a larger compositional variation than found in Broken Top units, and amphiboles from each system plot in discrete to partially overlapping fields (e.g. figure 4.10a-d). All amphiboles, with the exception of 3S090 and 3S009, were analyzed during the same microprobe run and no significant drift in probe calibration occurred; the compositional variations are not due to analytical error.

Increases in the Al content of igneous amphiboles can be produced through increases in temperature (Blundy & Holland, 1990), or through increases in total pressure at a relatively fixed temperature (Hollister et al., 1987; Hammarstron & Zen, 1986; Johnson & Rutherford, 1989). While most of the analyzed amphiboles have a consistent linear decrease in total Al with increasing Si (figure 4.10b), 7 points have significantly higher total Al. Excess Al will be in octahedral coordination, and will generally substitute for Fe^{3+} , Ti or $(Fe^{2+} + Mg)$ on the M2 site (Hawthorne, 1981). If the observed increases in total Al are a function of pressure or temperature, there should be a positive correlation between Al(IV) and another M2 cation. No correlation between Mg, Fe^{2+} , or Ti and Al is observed (e.g. figure 4.10c), although a slight negative correlation is observed with Fe^{3+} (figure 4.10d). This strongly suggests that the observed variation in Al content of amphiboles in the same rock was not produced through changes in magmatic pressure or temperature. In addition, all of the analyzed amphiboles have very similar anion abundances (Appendix C-3). Consequently, it is difficult to appeal to variations in magmatic water content that shifted the compositional field into a more aluminous range.

It is unclear why there is such a wide range of amphibole composition, both within individual rocks and between rocks of very similar composition. The Bend Pumice has no significant compositional zonation from the base to the top of the unit (Hill, 1985), yet the Bend Pumice amphiboles show a large amount of compositional variation. The amphibole variations are not systematic, and do not reflect progressive core to rim differences in composition. Given the high degree of ordering in the amphibole structure (Hawthorne, 1981; Blundy & Holland, 1990), it is speculated that the observed compositional variations may have been produced through variations in chemical diffusion rates, much in the same way that oscillatory zoning can be formed in plagioclase (Loomis, 1982; Grove et al., 1985).

4.6 Accessory Minerals

Apatite occurs as small ($\leq 0.1 \text{ mm}$) inclusions in plagioclase and pyroxene phenocrysts in most units with SiO2 $\geq 57\%$ (Appendix B), and can be recognized as a groundmass mineral in many dacites and rhyodacites. Based on experimental studies on a Cascade-Aleutian rock suite, Green and Watson (1982) concluded that apatite saturation could occur in this system at $\geq 57\%$ SiO2, but that the temperatures necessary for apatite crystallization in more mafic magmas were unrealistically low (≈ 800 °C at 7.5 kb). Green and Watson suggested that the apatite inclusions



Figure 4.10: A) Compositional range of amphibole for all samples. Analyses in appendix C-3, normalized to 23-O and average Fe³⁺ (Spear and Kimball, 1984) and nomenclature from Leake (1978). Note the wider range of scatter for Three Sisters samples (53 & 60) than for Broken Top samples. B) Total Al versus Si variations for analyzed amphiboles. Note the two distinct fields of total Al at constant Si.



Figure 4.10: C) & D) Cation abundances for analyzed amphiboles. Samples with excess Al(VI) do not show a correlation with Ti, $Mg + Fe^{2+}$, or Fe^{3+} abundance. If the excess Al(VI) contents of these amphiboles was a function of temperature or pressure, then there should be a positive correlation between AL(VI) and another M2 cation (Ti, $Mg + Fe^{2+}$, or Fe^{3+}) (cf., Blundy and Holland, 1990).
were produced through local saturation of P_2O_5 at the phenocryst-melt interface, and that the common occurrence of apatite inclusions in most mafic rock minerals does not indicate that the bulk melt was saturated in P_2O_5 . Experimental studies by Harrison and Watson (1984) confirmed that the relatively slow rate of phosphorus diffusion in these melts would result in local saturation of apatite at the phenocryst-melt boundary. The magmas in the study area were probably not saturated with apatite until about 62% SiO2 (Green & Watson, 1982), although local saturation in apatite probably occurred by 57% SiO2. Petrogenic models for the interval 57% to 62% SiO2 thus need to account for the effect of apatite fractionation, even though apatite is not saturated in the melt.

Zircon was recognized in the groundmass of several rhyodacites and rhyolites as a very small (< 0.1 mm), highly birefringent crystal (Appendix B). It is likely that most other rhyolites contained zircon, but it was not present in the thin section due to its very low abundance in the rock.

Small ($\leq 0.2 \text{ mm}$) xenocrysts of biotite, associated with plagioclase glomerocrysts, were occasionally found in thin sections of Three Sisters rhyolites. The biotite xenocrysts generally had faint opacitic rims, and only rarely occurred as a discrete phase. A 30µm shred of biotite was also analyzed in dacite 3S036 (table 4.6). Single xenocrysts of quartz (0.3 mm) and sanidine (0.5 mm) were found in dacite 3S036.

Rounded inclusions of pyrrhotite ($\approx Fe_{0.89}S$) were found in an orthopyroxene from rhyodacite 3S005, and in an ilmenite from rhyolite 3S114 (table 4.6). The pyrrhotite inclusion were probably trapped as immiscible sulfide blebs, as they are rounded and wholly enclosed by the host mineral. It is likely that sulfide blebs in silicic rock minerals represent local saturation of sulfur at the crystal-melt interface (Bacon, 1989).

Phlogopite was recognized as a vapor phase mineral in two rocks that had undergone small amounts of deuteric alteration: rhyodacite 3S009 and basaltic andesite 3S074. Vapor phase phlogopite is characterized by high F contents (R. Nielsen, pers. comm., 1990), which are found in the analyzed phlogopites (table 4.6). Although many other vapor phase minerals undoubtably occur in some rocks in the study area, these deuterically altered rocks were generally not sampled.

MICAS

| | 3 500 9 | 35074 | 35036 |
|--------|----------------|--------|--------|
| SiO2 | 46.30 | 41.80 | 53.87 |
| TiO2 | 0.14 | 0.85 | 1.20 |
| A12O3 | 11.02 | 10.79 | 15.39 |
| FeO | 1.80 | 4.99 | 8.30 |
| MgO | 24.12 | 26.29 | 9.65 |
| MnO | 0.18 | 0.05 | 0.06 |
| CaO | 0.03 | 0.00 | 0.70 |
| Na2O | 0.95 | 0.74 | 3.35 |
| K20 | 8.05 | 8.83 | 7.56 |
| Cr203 | 0.00 | 0.00 | 0.03 |
| H2O | 0.11 | 0.18 | 3.35 |
| F- | 8.71 | 8.39 | 2.43 |
| C1- | 0.05 | 0.10 | 0.18 |
| Total | 101.46 | 103.01 | 106.07 |
| O=F,Cl | 3.69 | 3.56 | 1.06 |
| TOTAL | 97.77 | 99.45 | 105.01 |

3S009 15 μ m clot of vapor phase phlogopite. 3S074 30 μ m vapor phase(?) phlogopite by plagioclase glomerocryst. 3S036 30 μ m biotite shred in groundmass.

SULFIDES

| | Weig | ght % | Ato | mic % |
|---------------------------------|--|--|--|--|
| | 35005 | 35114 | 35005 | 35114 |
| S Fe Cu Ni Co Zn | 38.61 58.56 0.18 0.40 0.13 0.03 | 37.97 59.61 0.00 0.15 0.04 0.00 | 53.17 46.29 0.12 0.30 0.10 0.02 | 52.52 47.34 0.00 0.11 0.03 0.00 |
| Total | 97.904 | 97.767 | | |

 $35005\ 10\mu m$ rounded pyrrhotite inclusion in opx. $35114\ 15\mu m$ rounded pyrrhotite inclusion in ilmenite.

Table 4.6: Microprobe analyses of micas and sulfides found in rocks from the study area. High F contents of the micas indicates a deuteric origin. Sulfide inclusions probably represent the effects of local saturation of the melt with sulfur (Bacon, 1989).

5. ESTIMATES OF INTENSIVE VARIABLES

5.1 Fe-Ti Oxide Temperatures and Oxygen Fugacities

Temperatures and oxygen fugacities for equilibrium pairs of titanomagnetite-ilmenite were derived according to the method of Buddington and Lindsley (1964), using the recalculation procedure of Carmichael (1967) to account for the distribution of minor components (table 5.1). Temperatures and oxygen fugacities were also calculated from the solution model of Ghiorso and Sack (1991). The Ghiorso and Sack solution model accounts for the significant variations in thermodynamic parameters that are produced by nonideal mixing of minor components in the spinel structure, which have not been accounted for in previous solution models for Fe-Ti oxide geothermometry (e.g. Anderson and Lindsley, 1988).

Temperatures derived from Buddington and Lindsley (1964) are plotted against temperatures from Ghiorso and Sack (1991) in figure 5.1. Most of the temperatures derived from Ghiorso and Sack (1991) are within 30°C of the temperatures derived from Buddington and Lindsley (1964), which corresponds to the 1σ analytical uncertainty reported by Buddington and Lindsley (1964). Calculated differences in temperatures of >30°C are restricted to some Broken Top series pyroclastic units, which also have titanomagnetite-ilmenite Mg/Mn ratios that are close to the 2σ limits of equilibrium (Bacon and Hirschmann, 1988). However, the temperatures calculated for these two units using the method of Ghiorso and Sack are unreasonably low (<750°C), based on the lack of low temperature phases (i.e., sanidine and quartz) in these units. This may indicate that temperature-fO₂ equilibria relationships in Ghiorso and Sack's formulations are more sensitive to disequilibriation effects at low temperatures. Fe-Ti oxide temperatures and oxygen fugacities determined from Buddington and Lindsley's calibration curves will be used in this study, as these temperatures correspond more closely to the observed modal mineralogies in silicic rocks from the study area.

Temperatures and oxygen fugacities for equilibrium titanomagnetite-ilmenite pairs are plotted in figure 5.2. Equilibrium data are limited in this study, but several important trends can be recognized. Most of the samples plot within the field of orthopyroxene-bearing silicic rocks recognized by Carmichael and others (1974), which correlates with the 2-pyroxene mineralogy found in most of the studied rocks. Broken Top rhyolites that plot within the orthopyroxene field, however, contain phenocrysts of amphibole along with orthopyroxene. These amphibole-bearing rhyolites either had oxygen fugacities that were lower, or temperatures that were higher, than found in the orthopyroxene-amphibole field of Carmichael and others (chapter 6, 1974). In addition, Three Sisters rhyolites had oxygen fugacities that were significantly higher than Broken Top rhyolites (figure 5.2), and plot near the amphibole-biotite field or in the extrapolated orthopyroxene-amphibole field.

| | | | | | | | | | | Budd | ington L | Ghi | orso & |
|-------------------------|---|-------|-------|-------|------|------|------|------|--------------|--------|-----------|------|--------|
| | | Mols | Mols | Mols | Molt | | Mols | Mols | Mol% | Lindsl | ey (1964) | Sack | (1991) |
| Titanomagnetite | | Usp | Mgt | X +++ | Usp | Ilmn | Ilmn | Hem | R2O 3 | T(°C |) f0- | T(°C | f0. |
| 3S005* Avg (2) | R | 0.154 | 0.286 | 0.289 | 35.0 | (4) | 0.62 | 0.05 | 7.04 | 780 | -14.9 | 743 | -15.92 |
| 3S028* Avg (6) | A | 0.179 | 0.267 | 0.269 | 40.1 | (3) | 0.55 | 0.11 | 16.97 | 960 | -10.4 | 1012 | -9.47 |
| $3S058$ 50μ , euhed | h | 0.108 | 0.336 | 0.339 | 24.3 | (1) | 0.53 | 0.12 | 18.44 | 825 | -12.3 | 864 | -11.46 |
| 3S079* Avg (2) | D | 0.180 | 0.274 | 0.277 | 39.7 | (2) | 0.56 | 0.11 | 16.49 | 965 | -10.4 | 1016 | -9.37 |
| 3S098* Avg (3) | R | 0.220 | 0.233 | 0.236 | 48.6 | (1) | 0.61 | 0.06 | 8.39 | 885 | -12.7 | 862 | -13.35 |
| $3S114$ 50μ fresh | h | 0.122 | 0.327 | 0.330 | 27.1 | (1) | 0.53 | 0.13 | 19.68 | 900 | -11.1 | 917 | -10.47 |
| 3S123* Avg (2) | d | 0.185 | 0.265 | 0.268 | 41.1 | (2) | 0.57 | 0.09 | 13.86 | 925 | -11.4 | 964 | -10.54 |
| 3S134* Avg (2) | d | 0.232 | 0.204 | 0.208 | 53.2 | (1) | 0.61 | 0.05 | 7.58 | 910 | -12.4 | 872 | -13.41 |
| 3S137* Avg (2) | d | 0.251 | 0.197 | 0.201 | 56.0 | (2) | 0.60 | 0.07 | 10.51 | 1005 | -10.6 | 988 | -10.90 |
| BP* Avg (7) | Н | 0.183 | 0.259 | 0.263 | 41.4 | (2) | 0.62 | 0.04 | 6.12 | 820 | -14.2 | 738 | -16.59 |
| Lava Is* Avg (2) | H | 0.142 | 0.299 | 0.303 | 32.1 | (1) | 0.58 | 0.08 | 12.03 | 830 | -13.0 | 833 | -13.11 |

 Table 5.1: Average compositions of Fe-Ti oxides used in calculations. End-member compositions calculated using the procedure of Carmichael (1967), and the number of samples averaged for each calculation shown in parantheses. Temperatures and oxygen fugacities for equilibrium Fe-Ti oxide pairs calculated from the curves of Buddington & Lindsley (1964), and the solution model of Ghiorso & Sack (1991).



Figure 5.1: Temperatures for titanomagnetite-ilmenite pairs. Temperatures determined through the solution model of Ghiorso and Sack (1991) plotted against temperatures determined from the experimental curves of Buddington and Lindsley (1964). Error field of ± 30°C corresponds to reported 1 σ error for each method.



Figure 5.2: Temperatures and oxygen fugacities for titanomagnetite-ilmenite pairs. T and fO₂ data derived from the curves of Buddington and Lindsley (1964). Fields of orthopyroxene, opx-amphibole and biotite ± amphibole stability from Carmichael et al., 1974.

5.2 2-Pyroxene Geothermometry

Many of the analyzed units with $SiO_2 > 55\%$ contain phenocrysts of both clinopyroxene and orthopyroxene. Pyroxene phenocrysts in these units are commonly subhedral to euhedral with fresh, nonreacted rims, indicating equilibrium crystallization between the two pyroxenes. Rocks with visibly equilibrated pyroxenes are appropriate for use in pyroxene geothermometry.

Pyroxene thermometry is based on the relationship that with increasing temperature, the Ca content of clinopyroxenes (i.e. augite) generally decreases while the Ca content of orthopyroxene or pigeonite increases (Lindsley, 1983). Relatively simple graphical and solution models (e.g. Wood and Banno, 1973) have been developed into experimentally constrained thermodynamic solution models (Davidson and Lindsley, 1985). The Davidson and Lindsley (1985) solution model is preferred as it accounts for the effect of pressure on calculated temperatures ($\approx 8^{\circ}C/kbar$), and for the effects of minor components (esp. Ti and Al) on the activities of Wollastonite (Wo), Enstatite (En), and Ferrosilite (Fs).

Major element pyroxene analyses are recalculated into mol percent Wo, En, and Fs using the projection of Lindsley and Anderson (1983), which accounts for the distribution of the minor components Ti, Al, Cr, Mn, and Na. Lindsley (1983) observed that only pyroxenes with Wo + En + Fs \geq 90% are suitable for use in 2-pyroxene geothermometry. Numerous pyroxenes analyzed for this study that had Al₂O₃ \geq 2.5 wt.% or TiO₂ \geq 0.8% (Appendix C-2) had Wo + En + Fs \leq 90% and were thus not suitable for pyroxene thermometry calculations.

Lindsley (1983) observed an analytical + experimental error of $\pm 50^{\circ}$ C for pyroxenes with $\leq 2 \mod \%$ minor components, with an increase in error of + 5°C for each mol percent increase in minor components. These errors also apply to the solution model of Davidson and Lindsley (1985). There is also a pressure dependence of $\leq 8^{\circ}$ C/kbar for the Davidson and Lindsley (1985) pyroxene geothermometer. Whereas total pressure is a poorly constrained intensive variable for the systems in this study, it seems reasonable to assume that most of the phenocrysts in the studied rocks crystallized between 5 kbars (rhyolites?) to 10 kbars (andesites?) total pressure. Temperatures calculated from the Davidson and Lindsley (1985) solution model, using estimated pressures of 5 kbars and 10 kbars, are shown in table 5.2, along with appropriate estimated uncertainties and pyroxene compositions. An increase of 5 kbars in estimated pressure generally results in an increase of $\approx 30^{\circ}$ C in temperature, which is always less than the $\pm 50^{\circ}$ C analytical error (table 5.2). Reasonable (i.e. ± 5 kbars) variations in total pressure will not significantly change the temperatures calculated from this model.

The Davidson and Lindsley (1985) solution model uses the measured augite composition to iteratively calculate the temperature and low-Ca pyroxene composition that should be in equilibrium with the analyzed augite. Relative equilibration between augite and coexisting orthopyroxene or pigeonite can thus be evaluated by comparing the Wo and En abundances calculated by the solution

| | | | Aug | ite (1 | nol%) | C | рх | (mol%) | | |
|--|--|--|---|--|---|---|--|---|---------------------------------------|--|
| Sample | Pa | ir | En | WO | Others | En | | WO O | the | ers |
| 35003 avg & c | Срх | -Opx | 48.1 | 38.2 | 9.0 | 73. | 1 | 3.1 | 5. | . 7 |
| 35005 | Срх | -Opx | 39.3 | 39.1 | 6.7 | 53. | 6 | 3.2 | 3. | . 1 |
| 35009 | Cpx | -Opx | 40.6 | 40.6 | 5.6 | 57. | 2 | 3.0 | 3. | . 3 |
| 3S028 avg & c-core | Срх | -Opx | 47.4 | 38.6 | 9.7 | 68. | 3 | 2.9 | 6. | .0 |
| 35030 | Cpx | -Opx | 46.4 | 36.6 | 7.4 | 69. | 2 | 4.3 | 4. | . 3 |
| 35036 | Cpx | -Opx | 47.5 | 39.2 | 9.1 | 69. | 5 | 3.5 | 5. | . 4 |
| 35045 | Cpx | -Opx | 45.5 | 35.7 | 9.3 | 71. | 4 | 4.1 | 4. | .1 |
| 3S061 orthopyroxene | Cpx | -Opx | 46.7 | 39.2 | 8.4 | 71. | 2 | 2.8 | 4 | .9 |
| 3S061 pigeonite | Cpx | -Piq | 46.7 | 39.2 | 8.4 | 51. | 2 | 11.0 | 3. | 9 |
| 3S079 c rim & avgs | Cpx | -Opx | 49.0 | 37.6 | 5.4 | 68. | 9 | 3.0 | 5. | 8 |
| 3S090c w/ a-core | Cpx | -Opx | 45.1 | 41.2 | 7.2 | 73. | 1 | 3.5 | 8 | .3 |
| 35090c w/ a-rim | Cpx | -opx | 45.1 | 41.2 | 7.2 | 69. | 6 | 3.7 | 5 | 0 |
| 3S098 glomero opx | Cox | xq0- | 35.4 | 40.1 | 5.0 | 51. | 3 | 3.6 | 3 | 9 |
| 3S098 phenocrysts | Cox | -Opx | 35.4 | 40.1 | 5.0 | 48. | 7 | 3.6 | 3 | . 7 |
| 3S112 Opx-C rim | Cox | -0x | 45.5 | 35.5 | 8.9 | 58. | 3 | 4.0 | 3 | 8 |
| 3S112 phenocrysts | Cox | -07X | 45.5 | 35.5 | 8.9 | 69. | 7 | 3.7 | 4 | 8 |
| 3S112 pigeonite | Cox | -Pia | 45.5 | 35.5 | 8.9 | 50. | 9 . | 17.3 | 4 | 2 |
| 3S115 phenocrysts | Cox | -Opx | 45.2 | 38.6 | 7.7 | 65. | 8 | 3.2 | Δ | 1 |
| 3S115 pigeonite | Cox | -Pia | 45.2 | 38.6 | 7.7 | 51. | 8 1 | 11 1 | 3 | 4 |
| 3S123 glomerocryst | Cox- | -0px | 45.0 | 40.0 | 8.2 | 68 | 4 | 3 2 | Δ. | 7 |
| 3S123 phenocrysts | Cox | -Opx | 46.5 | 39.3 | 8.9 | 67. | 6 | 3 4 | 4 | Δ |
| 3S131 pigeonite | Cox- | -Pig | 42.9 | 35.6 | 6.7 | 55 | 7 - | 11 0 | 2 | 7 |
| 35134 | Cnx- | -00x | 38 8 | 38 1 | 7 2 | 54 | <u>.</u> | 3 / | 3. | 1 |
| 35137 | Cox | -Onx | 44 3 | 30.1 | 9 7 | 62 | 5 | 3.5 | 6 | 0 |
| 3SXPB | Cnx- | -Onx | 44.2 | 38 9 | 8.2 | 41 | 2 | 2.0 | 7 | 5 |
| 3SXBPa rim & avg | Cox | -Opx | 45.8 | 38 2 | 8.8 | 41 | 2 | 2.9 | 7 | 5 |
| Lava Island | Cpx- | -Opy | 12 9 | 30.2 | 9.9 | 41. | 1 | 2.7 | 6 | 0 |
| Leve ibidid | opr | Opr | 76.7 | 57.5 | 2.0 | 40. | 4 | J. T | 0. | • • |
| | | 5 kba | rs | 1 | 0 kbars | म | rroi | r B | CT. | |
| Sample | En | Wo | T(°C) | En | Wo T | (°C) | +10 | יש בי | (0 (| 1 |
| 35003 avg & c | | | - (- / | | | (~/ | | - | · · | -) |
| | 73.9 | 3.4 | 1098 | 74.3 | 3.6 | 1129 | 69 | | | |
| 35005 | 73.9 | 3.4 | 1098 960 | 74.3 | 3.6 3.1 | 1129 988 | 69 55 | 780 | + | 30 |
| 35005 35005 35009 | 73.9 57.2 60.2 | 3.4 3.0 2.6 | 1098 960 922 | 74.3 57.8 60.9 | 3.6 3.1 2.6 | 1129 988 949 | 69 55 57 | 780 | ± | 30 |
| 35005 35005 35009 35028 avg & c-core | 73.9 57.2 60.2 73.0 | 3.4 3.0 2.6 3.3 | 1098 960 922 1078 | 74.3 57.8 60.9 | 3.6 3.1 2.6 3.4 | 1129 988 949 1109 | 69 55 57 70 | 780 | ± + | 30 |
| 3S005 3S009 3S028 avg & c-core 3S030 | 73.9 57.2 60.2 73.0 69.3 | 3.4 3.0 2.6 3.3 3.8 | 1098 960 922 1078 1116 | 74.3 57.8 60.9 73.4 | 3 3.6 3 3.1 9 2.6 4 3.4 5 4.0 | 1129 988 949 1109 1149 | 69 55 57 70 61 | 780 960 | ± ± | 30 30 |
| 35005 35009 35028 avg & c-core 35030 35036 | 73.9 57.2 60.2 73.0 69.3 73.8 | 3.4 3.0 2.6 3.3 3.8 3.1 | 1098 960 922 1078 1116 1062 | 74.3 57.8 60.9 73.4 69.6 74.2 | 3 3.6 3 3.1 9 2.6 4 3.4 5 4.0 9 3.2 | 1129 988 949 1109 1149 1092 | 69 55 57 70 61 67 | 780 960 | ± + | 30 30 |
| 3S005 3S009 3S028 avg & c-core 3S030 3S036 3S045 | 73.9 57.2 60.2 73.0 69.3 73.8 66.9 | 3.4 3.0 2.6 3.3 3.8 3.1 4.0 | 1098 960 922 1078 1116 1062 1125 | 74.3 57.8 60.9 73.4 69.6 74.2 | 3 3.6 3 3.1 2 6 4 3.4 5 4.0 2 3.2 2 4.2 | 1129 988 949 1109 1149 1092 1159 | 69 55 57 70 61 67 61 | 780 960 | ± ± | 30 30 |
| 35005 35009 35028 avg & c-core 35030 35036 35045 35061 orthopyroxene | 73.9 57.2 60.2 73.0 69.3 73.8 66.9 72.2 | 3.4 3.0 2.6 3.3 3.8 3.1 4.0 3.1 | 1098 960 922 1078 1116 1062 1125 1051 | 74.3 57.8 60.9 73.4 69.6 74.2 67.2 | 3 3.6 3 3.1 9 2.6 4 3.4 5 4.0 2 3.2 2 4.2 7 3.2 | 1129 988 949 1109 1149 1092 1159 1081 | 69 55 57 70 61 67 61 64 | 780 960 | ± ± | 30 30 |
| 3S005 3S009 3S028 avg & c-core 3S030 3S036 3S045 3S061 orthopyroxene *3S061 pigeopite | 73.9 57.2 60.2 73.0 69.3 73.8 66.9 72.2 68.1 | 3.4 3.0 2.6 3.3 3.8 3.1 4.0 3.1 6.9 | 1098 960 922 1078 1116 1062 1125 1051 1013 | 74.3 57.8 60.9 73.4 69.6 74.2 67.2 72.7 68.4 | 3 3.6 3 3.1 9 2.6 4 3.4 5 4.0 2 3.2 2 4.2 7 3.2 4 7.1 | 1129 988 949 1109 1149 1092 1159 1081 | 69 55 57 70 61 67 61 64 | 780 960 | ± ± | 30 30 |
| 35005 35009 35028 avg & c-core 35030 35036 35045 35061 orthopyroxene *35061 pigeonite 35079 c-rim & avgs | 73.9 57.2 60.2 73.0 69.3 73.8 66.9 72.2 68.1 74.9 | 3.4 3.0 2.6 3.3 3.8 3.1 4.0 3.1 6.9 3.6 | 1098 960 922 1078 1116 1062 1125 1051 1013 1125 | 74.3 57.8 60.9 73.4 69.6 74.2 67.2 72.7 68.4 75.2 | 3 3.6 3 3.1 2 2.6 3.4 3.4 5 4.0 2 3.2 2 4.2 7 3.2 4 7.1 3 8 | 1129 988 949 1109 1149 1092 1159 1081 1026 | 69 55 57 70 61 67 61 64 60 | 780 960 | ± ± + | 30 30 |
| 3S005 3S009 3S028 avg & c-core 3S030 3S036 3S045 3S061 orthopyroxene *3S061 pigeonite 3S079 c-rim & avgs *3S090c w/ a-core | 73.9 57.2 60.2 73.0 69.3 73.8 66.9 72.2 68.1 74.9 70.3 | 3.4 3.0 2.6 3.3 3.8 3.1 4.0 3.1 6.9 3.6 2.4 | 1098 960 922 1078 1116 1062 1125 1051 1013 1125 959 | 74.3 57.8 60.9 73.4 69.6 74.2 67.2 72.7 68.4 70.2 70.2 | 3 3.6 3 3.1 2 6 3 4 5 4.0 2 3.2 4 2 4 3.2 4 2 7 3.2 4 7.1 2 3.8 2 5 | 1129 988 949 1109 1149 1092 1159 1081 1026 1156 987 | 69 55 57 70 61 67 61 64 60 69 | 780 960 965 | + + + + | 30 30 30 |
| 3S005 3S009 3S028 avg & c-core 3S030 3S036 3S045 3S061 orthopyroxene *3S061 pigeonite 3S079 c-rim & avgs *3S090c w/ a-core | 73.9 57.2 60.2 73.0 69.3 73.8 66.9 72.2 68.1 74.9 70.3 70.3 | 3.4 3.0 2.6 3.3 3.8 3.1 4.0 3.1 6.9 3.6 2.4 2.4 | 1098 960 922 1078 1116 1062 1125 1051 1013 1125 959 959 | 74.3 57.8 60.9 73.4 69.6 74.2 67.2 72.7 68.4 75.2 70.9 | 3 3.6 3 3.1 2 6 3 4 5 4.0 2 3.2 4 2 4 2 4 3.2 4.2 3.2 4.2 3.2 4.2 3.2 4.2 3.2 4.2 3.2 4.2 3.2 4.2 3.2 4.2 3.2 5.5 2.5 6 2.5 | 1129 988 949 1109 1149 1092 1159 1081 1026 1156 987 | 69 55 70 61 67 61 64 60 81 | 780 960 965 | +1 +1 +1 | 30 30 30 |
| 3S005 3S009 3S028 avg & c-core 3S030 3S036 3S045 3S061 orthopyroxene *3S061 pigeonite 3S079 c-rim & avgs *3S090c w/ a-core *3S090c w/ a-rim | 73.9 57.2 60.2 73.0 69.3 73.8 66.9 72.2 68.1 74.9 70.3 70.3 49 4 | 3.4 3.0 2.6 3.3 3.8 3.1 4.0 3.1 6.9 3.6 2.4 2.4 2.7 | 1098 960 922 1078 1116 1062 1125 1051 1013 1125 959 959 881 | 74.3 57.8 60.9 73.4 69.6 74.2 67.2 72.7 68.4 75.2 70.9 70.9 | 3.6 3.1 2.6 3.4 5.4 5.4 5.4 5.4 5.4 5.4 3.2 4.2 7.1 2.5 2.5 2.5 2.5 2.5 | 1129 988 949 1109 1149 1092 1159 1081 1026 1156 987 987 | 69 55 57 70 61 67 61 64 60 69 81 65 | 780 960 965 | ± ± ± | 30 30 30 |
| 3S005 3S009 3S028 avg & c-core 3S030 3S036 3S045 3S061 orthopyroxene *3S061 pigeonite 3S079 c-rim & avgs *3S090c w/ a-core *3S090c w/ a-rim *3S098 glomero opx *3S098 phenocrysts | 73.9 57.2 60.2 73.0 69.3 73.8 66.9 72.2 68.1 74.9 70.3 70.3 49.4 49.4 | 3.4 3.0 2.6 3.3 3.8 3.1 4.0 3.1 6.9 3.6 2.4 2.4 2.7 2.7 | 1098 960 922 1078 1116 1062 1125 1051 1013 1125 959 959 881 881 | 74.3 57.8 60.9 73.4 69.6 74.2 67.2 72.7 68.4 75.2 70.9 70.9 50.0 | 3 3.6 3 1 9 2.6 3.4 3.4 5 4.0 2 3.2 4.2 3.2 4.2 3.2 4.2 3.2 4.2 3.2 4.2 3.2 4.2 3.2 4.2 3.2 4.2 3.2 4.2 3.2 5.2 2.5 2.5 2.5 2.5 2.7 2.7 2.7 | 1129 988 949 1109 1149 1092 1159 1081 1026 1156 987 987 906 | 69 55 70 61 64 69 81 65 59 | 780 960 965 | ± ± + | 30 30 30 |
| 3S005 3S009 3S028 avg & c-core 3S030 3S036 3S045 3S061 orthopyroxene *3S061 pigeonite 3S079 c-rim & avgs *3S090c w/ a-core *3S090c w/ a-core *3S098 glomero opx *3S098 phenocrysts 3S112 opy-c rim | 73.9 57.2 60.2 73.0 69.3 73.8 66.9 72.2 68.1 74.9 70.3 70.3 49.4 49.4 66.7 | 3.4 3.0 2.6 3.3 3.8 3.1 4.0 3.1 6.9 3.6 2.4 2.4 2.7 2.7 4.1 | 1098 960 922 1078 1116 1062 1125 1051 1013 1125 959 959 881 881 | 74.3 57.8 60.9 73.4 69.6 74.2 67.2 72.7 68.4 75.2 70.9 50.0 50.0 50.0 | 3.6 3.1 2.6 3.4 4.0 2.3.2 4.2 3.2 4.2 3.2 4.2 3.2 4.2 3.2 4.2 3.2 4.2 3.8 2.5 2.7 <td< td=""><td>1129 988 949 1109 1149 1092 1159 1081 1026 1156 987 987 906 906</td><td>69 55 70 61 60 69 85 59 59</td><td>780 960 965 885</td><td>+ + + +</td><td>30 30 30 30</td></td<> | 1129 988 949 1109 1149 1092 1159 1081 1026 1156 987 987 906 906 | 69 55 70 61 60 69 85 59 59 | 780 960 965 885 | + + + + | 30 30 30 30 |
| 3S005 3S009 3S028 avg & c-core 3S030 3S036 3S045 3S061 orthopyroxene *3S061 pigeonite 3S079 c-rim & avgs *3S090c w/ a-core *3S090c w/ a-core *3S098 glomero opx *3S098 phenocrysts 3S112 opx-c rim 3S112 phenocrysts | 73.9 57.2 60.2 73.0 69.3 73.8 66.9 72.2 68.1 74.9 70.3 70.3 49.4 49.4 66.7 66.7 | 3.4 3.0 2.6 3.3 3.8 3.1 4.0 3.1 6.9 3.6 2.4 2.4 2.7 2.7 4.1 | 1098 960 922 1078 1116 1062 1125 1051 1013 1125 959 881 881 1129 1129 | 74.3 57.8 60.9 73.4 69.6 74.2 67.2 72.7 68.4 75.2 70.9 50.0 50.0 50.0 67.0 | 3 3.6 3 1 2 3.1 2 3.4 4 0 2 3.2 4.2 3.2 4.2 3.2 4.2 3.2 4.2 3.2 4.2 3.2 4.2 3.2 4.2 3.8 2.5 2.5 2.7 2.7 4.2 4.2 | 1129 988 949 1109 1149 1092 1159 1081 1026 1156 987 987 906 1163 | 69 55 70 61 67 61 60 85 59 55 59 55 55 55 55 55 55 55 55 55 55 | 780 960 965 885 | ± ± ± ± | 30 30 30 30 |
| 3S005 3S009 3S028 avg & c-core 3S030 3S036 3S045 3S061 orthopyroxene *3S061 pigeonite 3S079 c-rim & avgs *3S090c w/ a-core *3S090c w/ a-core *3S098 glomero opx *3S098 phenocrysts 3S112 phenocrysts *3S112 pigeonite | 73.9 57.2 60.2 73.0 69.3 73.8 66.9 72.2 68.1 74.9 70.3 70.3 49.4 49.4 66.7 66.7 61.9 | 3.4 3.0 2.6 3.3 3.8 3.1 4.0 3.1 6.9 3.6 2.4 2.4 2.7 2.7 4.1 4.1 | 1098 960 922 1078 1116 1062 1125 1051 1013 1125 959 959 881 881 1129 1129 | 74.3 57.8 60.9 73.4 69.6 74.2 72.7 68.4 75.2 70.9 50.0 50.0 67.0 67.0 | 3 3 6 3 1 2 6 3 4 0 2 3 4 0 2 3 2 4 0 2 3 2 4 2 3 2 4 2 4 7 1 2 3 8 9 2 5 2 5 2 5 9 2 5 2 7 1 2 3 8 9 2 5 2 5 2 7 1 2 3 8 2 5 2 5 2 7 1 2 3 2 5 2 5 2 7 1 2 3 2 7 1 2 5 2 7 1 2 3 2 7 1 2 3 2 7 1 2 3 2 7 1 2 3 2 7 1 2 3 | 1129 988 949 1109 1149 1092 1159 1081 1026 1156 987 987 906 906 1163 1163 | 69 55 70 61 67 61 60 85 59 59 59 59 61 | 780 960 965 885 | + + + + | 30 30 30 30 |
| 3S005 3S009 3S028 avg & c-core 3S030 3S036 3S045 3S061 orthopyroxene *3S061 pigeonite 3S079 c-rim & avgs *3S090c w/ a-core *3S090c w/ a-core *3S098 glomero opx *3S098 glomero opx *3S098 phenocrysts 3S112 phenocrysts *3S12 pigeonite 3S115 phenocrysts | 73.9 57.2 60.2 73.0 69.3 73.8 66.9 72.2 68.1 74.9 70.3 70.3 49.4 49.4 66.7 66.7 61.9 | 3.4 3.0 2.6 3.3 3.8 3.1 4.0 3.1 6.9 3.6 2.4 2.4 2.7 2.7 4.1 9.8 2.2 | 1098 960 922 1078 1116 1062 1125 1051 1013 1125 959 959 881 881 1129 1129 1049 | 74.3 57.8 60.9 73.4 69.6 74.2 72.7 68.4 75.2 70.9 50.0 67.0 67.0 67.0 67.0 | 3.6 3.1 2.6 3.4 4.0 2.3.2 4.2 3.2 4.2 3.2 4.2 3.2 4.2 3.2 4.2 3.8 2.5 2.5 2.7 4.2 <td< td=""><td>1129 988 949 1109 1149 1092 1159 1081 1026 1156 987 906 906 1163 1163 1163</td><td>6957017667 664609155996610 669155996610</td><td>780 960 965 885</td><td>± ± ± ±</td><td>30 30 30 30</td></td<> | 1129 988 949 1109 1149 1092 1159 1081 1026 1156 987 906 906 1163 1163 1163 | 6957017667 664609155996610 669155996610 | 780 960 965 885 | ± ± ± ± | 30 30 30 30 |
| 3S005 3S009 3S028 avg & c-core 3S030 3S036 3S045 3S061 orthopyroxene *3S061 pigeonite 3S079 c-rim & avgs *3S090c w/ a-core *3S090c w/ a-core *3S098 glomero opx *3S098 glomero opx *3S098 phenocrysts 3S112 phenocrysts *3S115 phenocrysts | 73.9 57.2 60.2 73.0 69.3 73.8 66.9 72.2 68.1 74.9 70.3 70.3 49.4 49.4 66.7 66.7 61.9 68.8 64.7 | 3.4 3.0 2.6 3.3 3.8 3.1 4.0 3.1 6.9 3.6 2.4 2.4 2.7 2.7 4.1 9.8 3.2 7 | 1098 960 922 1078 1116 1062 1125 1051 1013 1125 959 959 881 881 1129 1129 1095 1049 | 74.3 57.8 60.9 73.4 69.6 74.2 72.7 68.4 75.2 70.9 50.0 67.0 67.0 67.0 67.0 62.0 67.0 62.0 | 3.6 3.1 2.6 3.4 4.0 3.2 4.2 3.2 4.2 3.2 4.2 3.2 4.2 3.2 4.2 3.8 2.5 2.7 4.2 3.8 2.5 2.7 4.2 3.3 3.3 3.3 3.3 | 1129 988 949 1109 1149 1092 1159 1081 1026 1156 987 906 906 1163 1163 1163 1107 1080 | 6957017667 66469159994607 6609155994607 | 780 960 965 885 | + + + + | 30 30 30 30 |
| 3S005 3S009 3S028 avg & c-core 3S030 3S036 3S045 3S061 orthopyroxene *3S061 pigeonite 3S079 c-rim & avgs *3S090c w/ a-core *3S098 glomero opx *3S098 glomero opx *3S098 phenocrysts 3S112 phenocrysts *3S112 pigeonite 3S115 pigeonite 3S115 pigeonite | 73.9 57.2 60.2 73.0 69.3 73.8 66.9 72.2 68.1 74.9 70.3 49.4 49.4 66.7 66.7 61.9 68.8 64.7 68.8 64.7 68.7 | 3.4 3.0 2.6 3.3 3.8 3.1 4.0 3.1 6.9 3.6 2.4 2.4 2.7 2.7 4.1 9.8 3.2 7.3 | 1098 960 922 1078 1116 1062 1125 1051 1013 1125 959 959 881 881 1129 1129 1095 1049 1049 | 74.3 57.8 60.9 73.4 69.6 74.2 72.7 68.4 75.2 70.9 50.0 67.0 67.0 67.0 69.3 67.0 69.3 67.0 69.3 | 3.6 3.1 2.6 3.4 4.0 3.2 4.2 3.2 4.2 3.2 4.2 3.2 4.2 3.2 4.2 3.8 2.5 2.7 4.2 3.8 2.5 2.7 4.2 3.3 3.3 3.3 3.3 3.3 | 1129 988 949 1109 1149 1092 1159 1081 1026 1156 987 906 906 1163 1163 1163 1163 1107 1080 1026 | 695701766666865999410074 | 780 960 965 885 | +1 +1 +1 +1 | 30 30 30 30 |
| 3S005 3S009 3S028 avg & c-core 3S030 3S036 3S045 3S061 orthopyroxene *3S061 pigeonite 3S079 c-rim & avgs *3S090c w/ a-core *3S090c w/ a-core *3S098 glomero opx *3S098 glomero opx *3S098 phenocrysts 3S112 phenocrysts *3S112 pigeonite 3S115 pigeonite 3S123 glomerocryst 3S123 phenocrysts | 73.9 57.2 60.2 73.0 69.3 73.8 66.9 72.2 68.1 74.9 70.3 49.4 49.4 66.7 61.9 68.8 64.7 69.4 71.9 | 3.4 3.0 2.6 3.3 3.8 3.1 4.0 3.1 6.9 3.6 2.4 2.7 2.7 4.1 9.8 3.2 7.3 2.8 | 1098 960 922 1078 1116 1062 1125 1051 1013 1125 959 959 881 881 1129 1129 1095 1049 1014 1002 | 74.3 57.8 60.9 73.4 69.6 74.2 72.7 68.4 75.2 70.9 50.0 67.0 67.0 67.0 69.3 64.9 70.0 70.0 70.0 70.0 70.0 70.0 70.0 70 | 3.6 3.1 2.6 3.4 4.0 3.2 4.2 3.2 4.2 3.2 4.2 3.2 4.2 3.2 4.2 3.8 2.5 2.7 3.8 2.5 2.7 4.2 3.3 7.5 3.3 7.5 2.2 | 1129 988 949 1109 1149 1092 1159 1081 1026 1156 987 906 906 1163 1163 1163 1163 1107 1080 1026 1031 | 695701766666859994100742 | 780 960 965 885 | ± ± ± ± | 30 30 30 30 |
| 3S005 3S009 3S028 avg & c-core 3S030 3S036 3S045 3S061 orthopyroxene *3S061 pigeonite 3S079 c-rim & avgs *3S090c w/ a-core *3S090c w/ a-core *3S098 glomero opx *3S098 glomero opx *3S098 phenocrysts 3S112 phenocrysts *3S115 pigeonite 3S115 pigeonite 3S123 glomerocrysts *3S121 pigeonite | 73.9 57.2 60.2 73.0 69.3 73.8 66.9 72.2 68.1 74.9 70.3 49.4 49.4 66.7 61.9 68.8 64.7 69.4 71.9 68.4 71.9 68.8 64.7 69.4 71.9 67.2 72.2 72.2 72.2 72.2 72.2 72.2 72.2 72.2 72.2 72.2 72.2 72.2 72.2 72.2 74.9 70.3 70.3 49.4 66.7 61.9 68.8 64.7 69.4 71.9 72.7 72.2 | 3.4 3.0 2.6 3.3 3.8 3.1 4.0 3.1 6.9 3.6 2.4 2.7 2.7 4.1 9.8 3.2 7.3 2.8 3.6 | 1098 960 922 1078 1116 1062 1125 1051 1013 1125 959 959 881 881 1129 1029 1049 1014 1002 1046 | 74.3 57.8 60.9 73.4 69.6 74.2 72.7 68.4 75.2 70.9 50.0 67.0 67.0 67.0 69.3 64.9 70.4 50.0 67.0 67.0 67.0 67.0 67.0 67.0 67.0 6 | 3.6 3.1 2.6 3.4 4.0 3.2 4.2 3.2 4.2 3.2 4.2 3.2 4.2 3.2 4.2 3.8 2.5 2.7 2.5 2.7 4.2 3.3 7.5 3.2 3.2 | 1129 988 949 1109 1149 1092 1159 1081 1026 1156 987 906 906 1163 1163 1163 1163 1107 1080 1026 1031 | 69570176666686599941007420 | 780 960 965 885 925 | * * * * * | 30 30 30 30 30 |
| 3S005 3S009 3S028 avg & c-core 3S030 3S036 3S045 3S061 orthopyroxene *3S061 pigeonite 3S079 c-rim & avgs *3S090c w/ a-core *3S098 glomero opx *3S098 glomero opx *3S098 glomero opx *3S098 phenocrysts 3S112 phenocrysts *3S115 pigeonite 3S115 pigeonite 3S123 glomerocrysts *3S131 pigeonite 3S131 pigeonite | 73.9 57.2 60.2 73.0 69.3 73.8 66.9 72.2 68.1 74.9 70.3 49.4 49.4 66.7 61.9 68.8 64.7 69.4 71.9 57.7 55.7 | 3.4 3.0 2.6 3.3 3.1 4.0 3.1 6.9 3.6 2.4 2.7 2.7 4.1 9.8 2.7 2.7 4.1 9.8 3.1 9.2 7.3 2.8 3.1 9.2 | 1098 960 922 1078 1116 1062 1125 1051 1013 1125 959 959 881 881 1129 1029 1049 1014 1002 1046 1065 | 74.3 57.8 60.9 73.4 69.6 74.2 72.7 68.4 75.2 70.9 50.0 50.0 67.0 67.0 69.3 64.9 72.4 56.5 70.9 56.5 70.0 57.0 57.0 57.0 57.0 57.0 57.0 57 | 3.6 3.1 2.6 3.4 4.0 3.2 4.2 3.2 4.2 3.2 4.2 3.2 4.2 3.2 4.2 3.8 2.5 2.5 2.7 4.2 3.3 7.5 3.2 3.3 7.5 3.2 3.2 3.3 7.5 3.2 3.2 3.3 3.2 3.3 3.2 3.3 3.2 3.2 3.2 3.2 3.2 3.3 3.4 3.2 3.3 3.4 3.5 3.4 3.5 3.6 3.7 3.8 3.9 </td <td>1129 988 949 1109 1149 1092 1159 1081 1026 1156 987 906 906 1163 1163 1163 1163 1163 1107 1080 1026 1031</td> <td>65701714091599941074280</td> <td>780 960 965 885 925</td> <td>± ± ± ± ±</td> <td>30 30 30 30 30</td> | 1129 988 949 1109 1149 1092 1159 1081 1026 1156 987 906 906 1163 1163 1163 1163 1163 1107 1080 1026 1031 | 65701714091599941074280 | 780 960 965 885 925 | ± ± ± ± ± | 30 30 30 30 30 |
| 3S005 3S009 3S028 avg & c-core 3S030 3S036 3S045 3S061 orthopyroxene *3S061 pigeonite 3S079 c-rim & avgs *3S090c w/ a-core *3S090c w/ a-core *3S098 glomero opx *3S098 glomero opx *3S098 phenocrysts 3S112 opx-c rim 3S112 phenocrysts *3S115 phenocrysts *3S115 pigeonite 3S123 glomerocryst *3S131 pigeonite 3S134 *3S137 | 73.9 57.2 60.2 73.0 69.3 73.8 66.9 72.2 68.1 74.9 70.3 49.4 49.4 66.7 61.9 68.8 64.7 69.4 71.9 57.7 55.9 57.7 55.9 57.7 57.7 55.9 57.7 | 3.4 3.0 2.6 3.3 3.1 4.0 3.1 6.9 3.6 2.4 2.7 2.7 4.1 9.8 2.7 3.1 9.6 3.1 9.6 3.2 | 1098 960 922 1078 1116 1062 1125 1051 1013 1125 959 959 881 881 1129 1029 1049 1014 1002 1046 1065 984 | 74.3 57.8 60.9 73.4 69.6 74.2 72.7 68.4 75.2 70.9 50.0 50.0 67.0 62.0 69.3 64.9 72.4 57.5 50.0 50.0 50.0 50.0 50.0 50.0 50.0 | 3.6 3.1 2.6 3.4 4.0 3.2 4.2 3.2 4.2 3.2 4.2 3.2 4.2 3.2 4.2 3.8 2.5 2.7 3.8 2.5 2.7 4.2 9.9 3.3 7.5 3.2 9.8 3.2 9.8 3.2 9.8 3.2 9.8 3.2 9.8 3.2 9.8 3.2 9.8 3.4 9.8 3.4 9.8 3.4 9.8 9.8 9.8 9.8 3.4 9.8 9.8 3.4 | 1129 988 949 1109 1149 1092 1159 1081 1026 1156 987 906 906 1163 1163 1163 1163 1107 1080 1026 1031 1075 1076 | 657017140915999410742800 | 780 960 965 885 925 910 | * * * * * * | 30 30 30 30 30 30 |
| 3S005 3S009 3S028 avg & c-core 3S030 3S036 3S045 3S061 orthopyroxene *3S061 pigeonite 3S079 c-rim & avgs *3S090c w/ a-core *3S090c w/ a-core *3S090c w/ a-rim *3S098 glomero opx *3S098 glomero opx *3S098 phenocrysts 3S112 phenocrysts *3S115 pigeonite 3S123 glomerocryst *3S131 pigeonite 3S123 phenocrysts *3S131 pigeonite 3S134 3S137 | 73.9 57.2 60.2 73.0 69.3 73.8 66.9 72.2 68.1 74.9 70.3 49.4 49.4 66.7 61.9 68.8 64.7 69.4 71.9 57.7 55.9 67.5 7.5 67.5 | 3.4 3.0 2.6 3.3 3.1 4.0 3.1 6.9 3.6 2.4 2.4 2.7 2.7 4.1 9.8 3.2 7.3 2.8 3.1 9.6 3.3 1 9.6 3.3 | 1098 960 922 1078 1116 1062 1125 1051 1013 1125 959 959 881 881 1129 1029 1049 1014 1002 1046 1065 984 1027 | 74.3 57.8 60.9 73.4 69.6 74.2 72.7 68.4 70.9 50.0 50.0 50.0 67.0 62.0 50.0 67.0 69.3 64.9 72.4 556.1 556.1 | 3.6 3.6 3.1 2.6 3.4 4.0 3.2 4.2 3.2 4.2 3.2 4.2 3.2 4.2 3.8 2.5 2.7 2.5 2.7 4.2 9.9 3.3 7.5 3.2 9.8 3.4 3.2 9.8 3.4 3.2 9.8 3.4 3.2 9.8 3.4 3.2 9.8 3.4 3.1 3.2 9.8 3.4 3.4 3.4 3.4 3.4 3.4 3.4 3.4 3.4 3.4 < | 1129 988 949 1109 1149 1092 1159 1081 1026 1156 987 906 906 1163 1163 1163 1163 1163 1163 1163 11 | 657017166666865999410074280000 | 780 960 965 885 925 925 910 | + + + + + + + + + | 30 30 30 30 30 30 30 |
| 3S005 3S009 3S028 avg & c-core 3S030 3S036 3S045 3S061 orthopyroxene *3S061 pigeonite 3S079 c-rim & avgs *3S090c w/ a-core *3S090c w/ a-core *3S090c w/ a-rim *3S098 glomero opx *3S098 glomero opx *3S098 phenocrysts 3S112 phenocrysts *3S112 pigeonite 3S115 phenocrysts *3S115 pigeonite 3S123 glomerocryst *3S131 pigeonite 3S134 3S137 *3SXBP | 73.9 57.2 60.2 73.0 69.3 73.8 66.9 72.2 68.1 74.9 70.3 49.4 66.7 61.9 68.8 64.7 69.4 71.9 57.7 55.9 67.5 67.5 67.6 | 3.4 3.0 2.6 3.3 3.1 4.0 3.1 6.9 3.6 2.4 2.7 2.7 4.1 9.8 3.2 7.3 2.8 3.1 9.6 3.3 3.1 9.6 3.3 1 9.6 3.3 | 1098 960 922 1078 1116 1062 1125 1051 1013 1125 959 959 881 881 1129 1029 1049 1014 1002 1046 1065 984 1016 1065 | 74.3 57.8 60.9 73.4 69.2 74.2 72.7 68.4 70.9 50.0 50.0 50.0 50.0 67.0 62.0 50.0 67.0 50.0 67.0 50.0 50.0 50.0 50.0 50.0 50.0 50.0 5 | 3.6 3.6 3.1 2.6 3.4 4.0 3.2 4.2 3.2 4.2 3.2 4.2 3.2 4.2 3.8 2.5 2.7 2.5 2.7 4.2 9.9 3.3 7.5 3.2 9.8 3.1 3.2 | 1129 988 949 1109 1149 1092 1159 1081 1026 1156 987 906 906 1163 1163 1163 1163 1163 1163 1163 11 | 65701716666686599941074280080 | 780 960 965 885 925 925 910 1005 820 | + + + + + + + + + + + + + + + + + + + | 30 30 30 30 30 30 30 30 |
| 3S005 3S009 3S028 avg & c-core 3S030 3S036 3S045 3S061 orthopyroxene *3S061 pigeonite 3S079 c-rim & avgs *3S090c w/ a-core *3S090c w/ a-core *3S090c w/ a-rim *3S098 glomero opx *3S098 glomero opx *3S098 phenocrysts 3S112 phenocrysts *3S115 pigeonite 3S115 phenocrysts *3S115 pigeonite 3S123 glomerocryst *3S131 pigeonite 3S134 3S137 *3SXBP *3SXBPa rim & avgs | 73.9 57.2 60.2 73.0 69.3 73.8 66.9 72.2 68.1 74.9 70.3 49.4 49.4 66.7 61.9 68.8 64.7 69.4 71.9 57.7 55.9 67.5 67.6 64.6 | 3.4 3.0 2.6 3.3 3.1 4.0 3.1 6.9 3.6 2.4 2.7 2.7 4.1 9.8 2.7 2.7 4.1 9.8 3.2 3.1 9.6 3.0 1 3.0 3.1 9.6 3.0 2.4 | 1098 960 922 1078 1116 1062 1125 1051 1013 1125 959 959 881 881 1129 1095 1049 1014 1002 1046 1065 984 1016 1027 1069 | 74.3 57.8 60.9 73.4 69.2 74.2 72.7 68.4 70.9 50.0 50.0 50.0 50.0 50.0 67.0 62.3 67.0 50.0 67.0 50.0 50.0 50.0 50.0 50.0 50.0 50.0 5 | 3.6 3.6 3.1 2.6 3.4 4.0 3.2 4.2 3.2 4.2 3.2 4.2 3.2 4.2 3.2 4.2 3.8 2.5 2.7 4.2 3.8 7.5 3.3 7.5 3.3 3.4 3.2 3.3 3.3 3.3 3.3 3.3 3.3 3.3 3.3 3.3 3.3 3.4 3.2 3.4 3.5 | 1129 988 949 1109 1149 1092 1159 1081 1026 1156 987 906 906 1163 1163 1163 1163 1163 1163 1163 11 | 65701714091599941074280088 667788 67788 67788 67788 | 780 960 965 885 925 925 910 1005 820 820 | + + + + + + + + + + + + + + + + + + + | 30 30 30 30 30 30 30 30 30 30 30 |

Table 5.2: 2-pyroxene thermometry (Davidson and Lindsley, 1985). *Sample = pyroxene pairs out of equilibrium, B&L = temperatures using curves of Buddington and Lindsley (1964). See text for additional explanations of data.

model for the low Ca pyroxene with the measured Wo and En abundance. Most of the calculated compositions are within ± 0.5 mol% Wo and ± 5 mol% En of the measured Wo and En abundances (table 5.2), which corresponds to the qualitative limits of "equilibrium" of the Davidson and Lindsley (1985) model. Pyroxene pairs that are apparently out of equilibrium fall into two groups: calculated En contents that are higher than measured En (9113, 9113a, MTT, LI, 65p, 112p, 115p), or calculated Wo contents that are lower than measured Wo (all pigeonite samples and 90c, 90r, 98, 98g). These samples do not yield accurate temperatures, and are not used in the following discussions. In addition, naturally occurring pigeonites with the same En contents as the studied pigeonites (En₅₁ to En₅₆) are stable at temperatures greater than ≈ 1000 °C (Lindsley, 1983), which can thus be used as a minimum temperature for pigeonite-bearing rocks in this study.

5.3 Combined Estimates of Temperature

Equilibrium temperatures calculated from the 2-pyroxene thermometer of Davidson and Lindsley (1985) at 5 kbars pressure are equivalent within 1σ error limits to the corresponding Fe-Ti oxide temperatures of Buddington and Lindsley (1964) (column BL&, table 5.2) for most samples. Temperatures calculated from either the Fe-Ti oxide equilibration curves of Buddington and Lindsley or 2-pyroxene thermometer of Davidson and Lindsley are plotted against rock SiO₂ in figure 5.3; Fe-Ti oxide temperatures are preferred, as these temperatures are generally more precise than 2-pyroxene temperatures. The Fe-Ti oxide temperature for rhyodacite 3S005, however, is unreasonably low (780°); low temperature minerals such as sanidine and quartz, which are normative for this composition, are not observed. The 2-pyroxene temperature (960°C) is thus used for 3S005. Based on these limited data, there are no significant differences in temperature between the Broken Top and Three Sisters systems for rocks of similar composition. Temperatures are apparently continuous from ≈1150°C at 55% SiO₂ to ≈850°C at 74% SiO₂, although there is significant variation in temperature (± 50-75°C) for rocks of similar composition (figure 5.3). These temperatures should be viewed as good approximations, as many of the analyzed mineral pairs were out of equilibrium and additional microprobe analysis is needed before temperature relationships can be adequately described.

5.4 Estimates of Pressure

There are no mineralogical relationships in any of the studied rocks that are known to be uniquely pressure dependent. The Al content of calcic amphiboles has been proposed as an empirical geobarometer for silica saturated igneous rocks (e.g. Hammarstrom and Zen, 1986). This geobarometer may be reasonably accurate at temperatures of ≈ 750 °C for calcic amphiboles in equilibrium with melt, fluid, biotite, quartz, sanidine, plagioclase, sphene and titanomagnetite (i.e.



Figure 5.3: Temperatures calculated from Fe-Ti oxide or 2-pyroxene data plotted against silica content of the host rock. Approximate range of pigeonite stability from Lindsley (1983).

Johnson and Rutherford, 1989). High Cascade rhyolites lack many of these phases, indicating that the Al content of amphibole may not reflect an equilibrated, pressure-dependent condition. The Al-in-amphibole geobarometer is also strongly temperature dependent (Blundy and Holland, 1990), and has an associated error of ± 4 kbars when used with well-constrained temperature data.

The recent development of a thermodynamically constrained solution model of Al siting in amphibole by Blundy and Holland (1990) removes many of the mineralogical constraints imposed by the amphibole geobarometer of Hammarstrom and Zen (1986). This solution model is valid for plagioclase-bearing, silica saturated rocks, and is best applied as a geothermometer. Although the pressure dependence of the model is poorly constrained, independently determined temperatures may be applied to the model to derive estimates of total pressure (± 4 kbars) (Blundy and Holland, 1990). The solution model was applied to amphibole-bearing silicic rocks in the study area, in an attempt to provide some limits on total pressure. Estimated temperatures from table 5.2, which have uncertainties of at least \pm 50°C, were used in the model, along with the upper and lower limits of Si found in the amphiboles of each unit (table 5.3). Calculated pressures ranged from -2 \pm 4 kbars to 6 \pm 4 kbars (figure 5.4), and serve only to indicate that the magmas probably equilibrated last at upper to middle crustal levels.

Pressures calculated for amphiboles with excess Al are no greater than those calculated for amphiboles from the same unit that lack excess Al (figure 5.4). This relationship supports the previous hypothesis (Chapter 4.5) that the excess Al in some amphiboles was not produced through higher pressure or higher temperature crystallization; increasing the estimated crystallization temperature of the excess Al amphiboles would have the effect of lowering the calculated pressure, and it is not intuitive why potentially higher pressure crystallization should have occurred at lower temperature (i.e. 750°C).

Amphiboles from Broken Top system rhyolites cluster at 2 ± 2 kbars (figure 5.4), whereas Three Sisters amphiboles show a much wider range of calculated pressure (i.e. 2 ± 4 kbars). Three Sisters amphiboles that are magnesio-hornblendes also have consistently lower pressures than other amphiboles (table 5.3). The magnesio-hornblendes may have crystallized at a temperature lower than 850 °C; when the estimated temperature of these amphiboles is lowered to 750 °C, their calculated pressures are comparable to other Three Sisters amphiboles. Whereas the excess Al in some amphiboles may be an artifact of the method used to calculate Fe³⁺/Fe²⁺ (Spear and Kimball, 1984), the magnesio-hornblendes may represent a lower temperature stage of crystallization in some Three Sisters rhyolites.

Rocks from the study area can be compared to other calc-alkaline systems that have experimentally constrained phase relationships. If a magma system is evolving primarily through crystal fractionation, then the fractionating assemblage will reflect the temperature, pressure and composition of the system. The effects of pressure and water content on phase relationships of

| | Si | ALvi | Т | Xab | Ρ | Label | Amphibole type |
|-----------------------|------|------|------|------|-------|-------|-----------------------------------|
| | | | °C | | kbars | | |
| 3S009 core | 6.46 | 0.09 | 900 | 0.69 | 0.6 | 9 | Magnesian Hastingsitic Hornblende |
| 3S053-A 250µm core | 6.53 | 0.26 | 850 | 0.65 | 2.3 | 53A | Edenitic Hornblende |
| 38053-A rim | 6.42 | 0.03 | 850 | 0.65 | 3.2 | 53a | Magnesian Hastingsitic Hornblende |
| 3S053-B 100µm | 7.04 | 0.01 | 850 | 0.65 | -2.4 | 53B | Magnesio-Hornblende |
| 3S053-B rim | 6.89 | 0.26 | 850 | 0.65 | -0.9 | 53b | Edenite |
| 38053-С 200µm соге | 6.62 | 0.27 | 850 | 0.65 | 1.6 | 53C | Edenitic Hornblende |
| 3S053-C rim | 6.54 | 0.26 | 850 | 0.65 | 2.2 | 53c | Edenitic Hornblende |
| 3S053-D 100µm | 6.74 | 0.08 | 850 | 0.65 | 0.5 | 53* | Magnesio-Hornblende |
| 3S053-Е 50µm FRAG rim | 6.69 | 0.07 | 850 | 0.65 | 1.0 | 53 | Magnesio-Hornblende |
| 3S060-A 250µm core | 6.27 | 0.31 | 850 | 0.65 | 4.4 | 60A | Pargasitic Hornblende |
| 35060-A rim | 6.08 | 0.07 | 850 | 0.65 | 5.9 | 60a | Magnesio-Hastingsite |
| 3S060-B 200µm FRAG | 6.29 | 0.07 | 850 | 0.65 | 4.2 | 60* | Magnesio-Hastingsitic Hornblende |
| 38060-С 250µm соге | 6.22 | 0.05 | 850 | 0.65 | 4.8 | 60C | Magnesio-Hastingsite |
| 35060-C rim | 6.24 | 0.07 | 850 | 0.65 | 4.6 | 60c | Magnesio-Hastingsite |
| 3S060-D >250µm core | 7.01 | 0.04 | 850 | 0.65 | -2.0 | 60D | Magnesio-Hornblende |
| 35060-D rim | 7.02 | 0.05 | 850 | 0.65 | -2.1 | 60d | Magnesio-Hornblende |
| 3S060-E 200µm FRAG | 6.88 | 0.08 | 850 | 0.65 | -0.8 | 60 | Magnesio-Hornblende |
| 3S090 glomero | 6.16 | 0.06 | 1000 | 0.5 | 2.5 | 90 | Magnesio-Hastingsite |
| 3S090a 2ND POINT | 6.29 | 0.17 | 1000 | 0.5 | 1.4 | 90a | Pargasitic Hornblende |
| 3S090a glomero | 6.18 | 0.08 | 1000 | 0.5 | 2.3 | 90a | Magnesio-Hastingsite |
| BP LRZ-A >250µm core | 6.41 | 0.07 | 850 | 0.76 | 2.1 | LA | Magnesian Hastingsitic Hornblende |
| BP LRZ-A rim | 6.47 | 0.06 | 850 | 0.76 | 1.6 | la | Magnesian Hastingsitic Hornblende |
| BP LRZ-B >250µm core | 6.48 | 0.03 | 850 | 0.76 | 1.5 | lΒ | Magnesian Hastingsitic Hornblende |
| BP LRZ-B rim | 6.45 | 0.05 | 850 | 0.76 | 1.8 | lЬ | Magnesian Hastingsitic Hornblende |
| BP LRZ-C core | 6.46 | 0.04 | 850 | 0.76 | 1.7 | lC | Magnesian Hastingsitic Hornblende |
| BP LRZ-C rim | 6.26 | 0.09 | 850 | 0.76 | 3.3 | lc | Magnesian Hastingsitic Hornblende |
| BPB-A core | 6.23 | 0.08 | 850 | 0.76 | 3.5 | bA | Magnesian Hastingsitic Hornblende |
| BPB-A rim | 6.31 | 0.08 | 850 | 0.76 | 2.9 | ba | Magnesian Hastingsitic Hornblende |
| BPB-B 200µm core | 6.34 | 0.08 | 850 | 0.76 | 2.7 | bB | Magnesian Hastingsitic Hornblende |
| BPB-B rim | 6.45 | 0.06 | 850 | 0.76 | 1.8 | bb | Magnesian Hastingsitic Hornblende |
| BPB-C 250µm core | 6.39 | 0.09 | 850 | 0.76 | 2.3 | ЬС | Magnesian Hastingsitic Hornblende |
| BPB-C rim | 6.29 | 0.06 | 850 | 0.76 | 3.0 | bc | Magnesian Hastingsitic Hornblende |
| BPB-D >250µm core | 6.24 | 0.06 | 850 | 0.76 | 3.4 | b0 | Magnesian Hastingsitic Hornblende |
| BPB-D rim | 6.29 | 0.09 | 850 | 0.76 | 3.0 | bd | Magnesian Hastingsitic Hornblende |
| BPB-E core | 6.36 | 0.06 | 850 | 0.76 | 2.5 | b | Magnesian Hastingsitic Hornblende |
| BPT-A core | 6.52 | 0.04 | 850 | 0.76 | 1.2 | BA | Magnesian Hastingsitic Hornblende |
| BPT-A rim | 6.45 | 0.07 | 850 | 0.76 | 1.8 | Ba | Magnesian Hastingsitic Hornblende |
| BPT-B >250µm core | 6.37 | 0.05 | 850 | 0.76 | 2.4 | BB | Magnesian Hastingsitic Hornblende |
| BPT-B rim | 6.35 | 0.05 | 850 | 0.76 | 2.6 | Bb | Magnesian Hastingsitic Hornblende |
| BPT-C core | 6.29 | 0.05 | 850 | 0.76 | 3.0 | BC | Magnesian Hastingsitic Hornblende |
| BPT-C rim | 6.48 | 0.26 | 850 | 0.76 | 1.5 | Bc | Edenitic Hornblende |
| BPT-D core | 6.31 | 0.05 | 850 | 0.76 | 2.9 | BD | Magnesian Hastingsitic Hornblende |
| BPT-D rim | 6.33 | 0.08 | 850 | 0.76 | 2.7 | Bd | Magnesian Hastingsitic Hornblende |
| LI-A 200µm core | 6.46 | 0.02 | 850 | 0.75 | 1.8 | LA | Magnesian Hastingsitic Hornblende |
| LI-A rim | 6.45 | 0.10 | 850 | 0.75 | 1.9 | La | Magnesian Hastingsitic Hornblende |
| LI-B >250µm core | 6.39 | 0.04 | 850 | 0.75 | 2.4 | LB | Magnesian Hastingsitic Hornblende |
| LI-B rim | 6.49 | 0.24 | 850 | 0.75 | 1.6 | Lb | Edenitic Hornblende |
| LI-C >200µm core | 6.34 | 0.08 | 850 | 0.75 | 2.8 | LC | Magnesian Hastingsitic Hornblende |
| LI-C rim | 6.29 | 0.09 | 850 | 0.75 | 3.1 | Lc | Magnesian Hastingsitic Hornblende |

Table 5.3: Estimated pressures of crystallization of amphibole-bearing rocks, using the geobarometer of Blundy and Holland (equation 5, 1990). All pressures have a minimum uncertainty of ± 4 kbars. Si and Al(VI) cation abundances calculated from Spear and Kimball (1984) (Appendix C-5), estimated temperatures from table 5.2, mol fraction albite from table 4.1, and amphibole types from Leake (1978).



Figure 5.4: Estimated pressures calculated from the solution model of Blundy and Holland (1990) for the Al content of amphibole. Amphibole data described in text and given in table 5.3. Al(VI) calculated from the method of Spear and Kimball (1984).

calc-alkaline rocks from the oceanic part of the Aleutian arc have been investigated by Baker and Eggler (1983, 1987). Because the mafic systems in the High Cascades and Aleutians have similar compositions (calc-alkaline, parental high alumina basalt) and range of temperatures, the pressure-dependent phase relationships determined on the Aleutian system should be applicable to the High Cascades system. However, the following chapter on mafic petrogenesis will show that differentia-tion of the High Cascades mafic system involved significant amounts of silicic magma mixing, which could potentially alter the phase relationships between the High Cascades and Aleutian systems.

Using the pseudo-ternary phase projections from Baker and Eggler (1983), mafic rocks from this part of the High Cascades plot (figure 5.5) close to the anhydrous 8 kbar phase boundaries of Baker and Eggler (1983, 1987). Andesites and dacites plot between the 8 kbar and 5 kbar-2% H_2O boundaries. Extrapolation of the 5 kbar-2% H_2O boundary to mafic compositions (cf., Baker, 1987) also indicates that the High Cascade mafic magmas may have contained <2% H_2O , although it is possible that higher water contents (>2% H_2O) combined with higher pressures (>8 kbars) might result in similar phase relationships. These data suggest that the evolution of High Cascade mafic magmas probably occurred at deep crustal levels (30-40 km) and relatively low water contents ($\leq 2\%$ H_2O).

5.5 Estimates of Water Content

Although water content is an extensive variable, it is rarely measured directly and must be estimated by comparison with experimental systems. This part of the High Cascades is characterized by an anhydrous mineral assemblage from basalt to rhyodacite, with trace amounts of amphibole in most rhyolites (chapter 4). In contrast, andesites, dacites and rhyodacites at Mt. Mazama contain significant amounts of amphibole, which probably indicates a relatively higher H_2O content (Bacon and Druitt, 1988).

5.5.1 Amphibole equilibria

Magmas that were in presumed equilibrium with plag + opx + cpx can be used to infer H_2O contents by comparison with the experimental data of Merzbacher and Eggler (1984) for Mt. St. Helens andesites and dacites. These experiments were run between 1 atm. and 4 kbars, and show that amphibole saturation occurs at higher water contents at both higher pressures and temperatures. Andesites, dacites and rhyodacites from the study area generally plot (figure 5.6) in the area of $\geq 4\%$ H₂O, but at water contents below the amphibole stability field. The lack of amphibole phenocrysts in rocks that plot close to or within the amphibole stability field of Merzbacher and Eggler (1983) suggests that crystallization might have occurred at pressures > 4 kbars, which would shift the amphibole stability field away from the opx apex. In contrast, rocks



Figure 5.5: Pseudoternary projection from plag + mgt of all High Cascade rocks from the study area, after Baker and Eggler (figure 1, 1983). Phase boundaries determined from melting experiments of Aleutian calc-alkaline arc rocks at 8 kbars anhydrous (solid line), 5 kbars, 2 wt% H_2O (dashed line), and 2 kbars, 2 wt% H_2O (dotted line), from Baker and Eggler (1983,1987). High Cascade mafic rocks plot near the 8 kbars anhydrous boundaries. Note that the mafic data plot above an extrapolated 5 kbars hydrous boundary; increasing pressure at constant water content would effectively raise this boundary towards the diopside apex. Mafic rocks may have thus evolved at pressures >8 kbars, if estimated water contents of 1-2 wt% are correct. Although the trend of the silicic data is toward lower pressures or higher water contents, more evolved compositions may reflect mixing between rhyolites and mafic rocks.



Figure 5.6: Intermediate and silicic rocks from the study area projected onto part of the plag-opxqtz+or pseudoternary of Merzbacher and Eggler (figure 3, 1984). Projection is from diopside + mgt, and is modified from Baker and Eggler (1983) to include opx instead of ol. Contours = position of the opx-plag cotectic at a given melt H_2O content. Field of amphibole stability is for intermediate to silicic calc-alkaline rocks to 4 kbars pressure (Merzbacher and Eggler, 1984). Note that the trend of amphibole-bearing rocks from Mt. Mazama (Bacon and Druitt, 1988) corresponds to the amphibole stability field, but that rocks from the study area generally plot below the Mazama data and lack amphibole. Andesitic and dacitic magmas from Middle Sister (filled circles) had apparently higher water contents than similar rocks from South Sister (open circles). from Mt. Mazama generally plot within the amphibole stability field (figure 5.6), contain significant amounts of amphibole, and are characterized by differentiation involving amphibole fractionation (Bacon and Druitt, 1988). Climactic rhyodacites from Mt. Mazama also plot at higher water contents than rhyolites from the study area, even though both groups contain amphibole phenocrysts (figure 5.6). Lower temperatures in the rhyolites would stabilize amphibole at lower water contents than in Mazama rhyodacites, although differences in pressure and composition should also control the stability of amphibole.

The composition of amphiboles from the study area is very similar to the composition of Mazama amphiboles (Druitt and Bacon, 1989) and amphiboles produced experimentally from Mt. St. Helens dacite (Rutherford and Devine, 1988). However, some of the amphiboles in the study area are more Mg-depleted (Mg' = Mg/Mg + Fe²⁺ \approx 0.5-0.7) than Mazama (Mg' \approx 0.6-0.7) or Mt. St. Helens (Mg' \approx 0.6-0.7) amphiboles. This relationship is consistent with the restriction of amphibole in the study area to silicic rhyodacites and rhyolites, and indicates that these magmatic systems are comparable.

5.5.2 Plagioclase equilibria

Plagioclase-melt equilibrium is a function of temperature, composition, and water content of the melt (Kudo and Weill, 1970; Housh and Luhr, 1991). Synthetic, 1 atm. relationships (Kudo and Weill, 1970) have been modified by experimental data on natural systems by Housh and Luhr (1991), who have developed albite and anorthite exchange reactions between plagioclase and melt. These exchange reactions can be used to estimate the water content of the melt, if the temperature, pressure and plagioclase composition are known. The minimum uncertainties in estimated water contents derived from the exchange reactions are ± 0.54 wt% for albite and ± 0.33 wt% for anorthite.

The plagioclase exchange reactions were applied to representative samples from the study area, using the BASIC program TWATER from Housh and Luhr (1991). Variations in total pressure between 10 kbar and 3 kbar had no significant effect (±0.1 wt%) on calculated water contents, and estimated pressures of 8 kbars were used in all calculations. Uncertainties in temperature of ± 30°C, however, resulted in changes of ± ≈1 wt% in calculated water content. Water contents for South Sister rhyolite 3S114 (An₄₁, T=900°C) were (Ab) 2.8 ± 0.6 and (An) 2.9 ± 0.4 wt%, which corresponds to the water content derived from figure 5.6 (≈3%). Broken Top rhyolite 3SXBP (An₂₄, T=820°C) water contents were (Ab) 4 ± 1 and (An) 5 ± 1 wt%. Both Broken Top and Three Sisters rhyolites contain trace amounts of amphibole and project onto the same area in figure 5.6, and should have similar water contents (≈4%). Broken Top rhyodacite 3S009 (An₃₈, T=950°C), which contains a trace of amphibole, had calculated water contents of (Ab) 3 ± 1 and (An) 4 ± 1 wt%, and dacite 3S115 (An₅₄) had lower water contents: (Ab) 1.6 ± 0.9 , (An) 2.7 ± 0.8 wt%. South Sister dacite 3S123 (An₅₀, T=925°C) had higher calculated water contents (Ab: 4.0 ± 0.7 , An: 5.4 ± 0.6 wt%) than either Broken Top dacite or Middle Sister dacite 3S137 (An₄₅, T=1000°C), (Ab) 2.5 ± 0.6 , (An) 3.6 ± 0.5 wt%. However, Middle Sister dacites and andesites plot in an area of higher water content (figure 5.6) than South Sister andesites and dacites. The lower estimated temperature of SS dacite 3S123 may be in part responsible for the higher calculated water content in this unit. Application of the Housh and Luhr (1991) plagioclase exchange model to rocks in the study area is limited by the relatively large uncertainties in estimated temperature associated with these rocks.

Although andesites, dacites and rhyodacites from the study area probably contained $\geq 4\%$ H₂O, amphibole was not stable in these melts and was not a fractionating phase. If these rocks evolved primarily through crystal fractionation, the parental mafic rocks must have been hydrous, and probably contained 1-2% H₂O (assuming H₂O is incompatible and >50% crystallization to derive andesite from high alumina basalt). Rhyolitic magmas probably contained 3-4% H₂O, and crystallized trace amounts of amphibole. Petrogenetic models developed in the following chapters must consider the effects of elevated pressure and water content, in relation to commonly used anhydrous, 1 atmosphere petrogenetic models.

6. GEOCHEMISTRY

6.1 Introduction

One hundred and twenty-seven High Cascade samples were analyzed with standard X-ray fluorescence and instrumental neutron activation analysis techniques for this study. Details of these analytical methods are in Appendix A, and geochemical analyses are listed in Appendix D. Analyzed elements and associated analytical errors are summarized in table 6.1. Analytical precision is primarily a function of the concentration of an element in the sample; higher concentrations give larger signals above background "noise", and thus result in more precise analyses. The range of the 1σ errors in table 6.1 combines errors due to the precision of the analytical technique with errors of accuracy in reproducing standard values.

The analyzed High Cascade samples are all younger than 720 ka, and were all produced under the same regional tectonic conditions. Ten rocks from the Deschutes Formation were also analyzed. These samples were discussed in Chapter 3, but are not included in the following geochemical models of High Cascade petrogenesis because they probably formed under different tectonic conditions than the late High Cascades (e.g., Priest et al., 1983). An additional 28 analyses of late High Cascade mafic to intermediate composition rocks from the Three Sisters area, which are from Hughes (1983, 1990), are also included in the data set used for geochemical modeling.

6.2 Terminology

The general classification of aphanitic calc-alkaline rocks is at times confusing and arbitrary, but is generally based on silica content and significant changes in modal mineralogy. The classification scheme used in this study (figure 6.1a) is a modification of the total alkali-silica variation diagram of Le Maitre (1984), which generally conforms to historical use of volcanic rock names in the Cascades and preserves petrologically significant distinctions between silicic rock types. The major difference between the classifications in this study (figure 6.1a) and Le Maitre (1984) is the addition of the term "rhyodacite" for rocks with 68-72% SiO₂. Rhyodacite is usually the most silicic rock at other Oregon High Cascade volcanos (Wise, 1969; Conrey, 1991; Bacon and Druitt, 1988), but higher silica rocks are common in the Three Sisters area. It is thus appropriate to reserve the term "rhyolite" for aphanitic calc-alkaline rocks with $\geq 72\%$ SiO₂. In this study, mafic refers to rocks with $\leq 58\%$ SiO₂ and silicic to rocks with $\geq 58\%$ SiO₂.

6.3 General Major Element Trends

Quaternary rocks of the Three Sisters area form a typical calc-alkaline assemblage, following the widely accepted definitions of Peacock (1931) (figure 6.1b), and Irving and Barager (1971)

| Element | Method | Range of 1ø <u>analytical error</u> ¹ |
|--------------------------------|--------|---|
| SiO ₂ | XRF | ± 0.5 wt% |
| TiO ₂ | XRF | ± 0.05 wt% |
| Al ₂ Ó ₂ | XRF | ± 0.5 wt% |
| Feð | XRF | ± 0.3 wt% |
| MnO | XRF | ± 0.01 wt% |
| MgO | XRF | ± 0.05 wt% |
| CaO | XRF | ± 0.1 wt% |
| Na ₂ O | XRF | ± 0.1 wt% |
| К ₂ Ó | XRF | ± 0.05 wt% |
| P205 | XRF | ± 0.01 wt% |
| Rb | XRF | 3% - 3% |
| Ba | XRF | 78 - 38 |
| Sr | XRF | 38 - 38 |
| Cs | INAA | 15% - 3% |
| Sc | INAA | 3% - 3% |
| V | XRF | 3% - 3% |
| Co | INAA | 3% - 3% |
| Ni | XRF | 5% - 10% |
| Cr | INAA | 3% - 15% |
| Zn | INAA | 10% - 6% |
| La | INAA | 3% - 3% |
| Ce | INAA | 48 - 38 |
| Nd | INAA | 78 - 58 |
| Sm | INAA | 3% - 3% |
| Eu | INAA | 3% - 3% |
| Tb | INAA | 48 - 38 |
| Yb | INAA | 48 - 48 |
| Lu | INAA | 3% - 3% |
| Zr | XRF | 3% - 3% |
| Hf | INAA | 48 - 38 |
| Ga | XRF | 78 - 58 |
| Y | XRF | 6% - 4% |
| Nb | XRF | 10% - 5% |
| Та | INAA | 6% - 3% |
| Th | INAA | 5% - 3% |
| U | INAA | 6% - ≥20% |

- - - - -

¹ Range of analytical error for mafic to silicic samples. Error incorporates counting statistics and accuracy in reproduction of standards during analytical runs. XRF accuracy determined through analyses of standard samples by W.S.U., U.S.G.S., and other labs.

Table 6.1: Summary of analytical methods and associated 1σ errors for mafic-to-silicic samples. XRF = X-ray fluorescence, INAA = instrumental neutron activation analysis. Details of analytical procedures are explained in Appendix A.



Figure 6.1: A) Total alkalies versus silicia classification diagram, modified from Le Maitre (1984). Dashed diagional line represents original dacite-rhyolite boundary, which has been subdivided into a rhyodacite field. B) Peacock diagram showing variations in total alkalies and lime that intersect in the calc-alkaline field (dashed lines).

(figure 6.2a). Rocks from the study area are primarily metaluminous (molar Al/Ca+Na+K \leq 1)(Shand, 1947), but some rhyolites are slightly peraluminous (molar Al/Ca+Na+K = 1-1.2) (figure 6.2b). Eight analyses of rhyodacitic to dacitic pumice from unconsolidated surficial deposits (3S001, 3S007, 3S014, 3S018, 3S024, 3S041, 3S044, 3S079) appeared strongly peraluminous but had low major element totals (98%-96%) and were anomalously low in alkalies due to hydration (Appendix D). These hydrated samples are not used in the following major element discussions, but will be included in models utilizing relatively immobile trace elements.

Major element variation diagrams in figure 6.3 provide useful first-order constraints on the petrogenesis of the Broken Top and Three Sisters magmatic systems, and define the key petrologic problems in the study area. Many inter-element variations are relatively well constrained from 76% to $\approx 58\%$ SiO₂, but significantly less constrained from $\approx 58\%$ to 48% SiO₂. In the most general sense, the mafic system appears to be more open (i.e. greater amounts of recharge, assimilation and mixing) than the silicic system.

Most of the 48% to 54% SiO_2 samples are from the Three Sisters system, as basalts are relatively uncommon in the Broken Top system. Studies of mafic magmatism by Hughes (1983, 1990), and Hughes and Taylor (1986) have not identified any spatial distinctions among Quaternary basalts from the Three Sisters to Mt. Jefferson area. Basalts and basaltic andesites from the Broken Top system plot within the same compositional range as mafic rocks from the Three Sisters system; there is no evidence to suggest that either magmatic system has compositionally distinct mafic rocks.

There is, however, a significant compositional gap in the Three Sisters system between 66% and 72% SiO₂, but the same interval in the Broken Top system shows continuous elemental variations (figure 6.3). A similar compositional gap was also recognized by Clark (1983) at South Sister volcano. There is also a small but significant compositional gap in the Broken Top system between 72% and 74% SiO₂. This compositional gap is not only expressed by a 2 wt% gap in SiO₂, but other major and trace element variations have a similar compositional gap between rhyodacites and rhyolites in the Broken Top system. I must emphasize that sampling throughout the study area focused on silicic rocks, and was guided by the detailed studies of Wozniak (1982), Clark (1983), Hughes (1983) and Taylor (1978, 1987, unpub. res.). It is thus extremely unlikely that the compositional gaps represent inadequate sampling of the silicic rock suite exposed at the surface.

Broken Top rhyolites often plot off of the trend defined by $\approx 65\%$ to 72% SiO₂ Broken Top rocks. It will be shown that the 72% to 74% SiO₂ compositional gap in this system is petrologically significant. In contrast, Three Sisters rhyolites generally form linear trends from 72% to 76% SiO₂, which always project to more poorly defined trends at $\leq 65\%$ SiO₂.

These observations of major element distributions define the main petrologic problems that will be addressed in this study:



Figure 6.2: A) AFM diagram defining the boundary between tholeiitic and calc-alkaline series rocks, from Irving and Barager (1971). A = Na₂O+K₂O, F = FeO, M = MgO, all in weight percent. B) Alumina saturation diagram, after Shand (1947). The rocks that appear to be slightly peraluminous (i.e., corundum normative) are rhyolites.



Figure 6.3: Major element-silica variation diagrams.



Figure 6.3: Major element-silica variation diagrams (continued).



Figure 6.3: Major element-silica variation diagrams (continued).

- 1) Why are Broken Top rhyolites compositionally distinct from Three Sisters rhyolites?
- 2) What is the contribution (if any) of each rhyolite type to the geochemical evolution of their associated systems? Did rhyolites from one system interact with the other system?
- 3) Why are rhyolites relatively abundant in this part of the Oregon Cascades, but rarely found at other Oregon Cascade volcanos?
- 4) Why does the Three Sisters system have a compositional gap from 66% to 72% SiO₂, but this interval in the Broken Top system represents continuous elemental variations? What is the significance of the compositional gap in the Broken Top system between 72% and 74% SiO₂?
- 5) Why does the amount of inter-element variation apparently decrease with increasing silica content? Is this simply a matter of closure, that is, does increasing the percentage of silica result in a decrease in the abundance and possible variations of other elements (cf., Chayes, 1964). Or does this trend reflect other magmatic processes?

All of these questions must be evaluated in light of the close spatial and temporal associations of these two magmatic systems. The primary goal of this study is to explain the physical processes that led to the development of two distinct silicic magma systems in this part of the Oregon Cascades. Silicic magmatism cannot be studied, however, without first considering the petrogenesis of associated mafic rocks that are likely progenitors to the silicic series.

7. MAFIC PETROGENESIS

7.1 Introduction

The petrogenesis of mafic rocks in the central part of the Oregon Cascades has been examined in detail by Hughes (1983, 1990) and Hughes and Taylor (1986). These studies have examined late Miocene to Holocene basalts, basaltic andesites, and andesites that have been generated in sometimes different tectonic regimes, and erupted at spatially and compositionally distinct volcanic complexes. The fundamental conclusions of their research can be summarized as follows: High Cascade basalts erupted since the late Pliocene were generated from an ocean island basalt-like mantle that was variably enriched in large ion lithophile elements. Variable (but undefined) amounts of mantle metasomatism are required to generate the chemically distinct types of basaltic andesite they recognized in the High Cascades, along with lesser amounts of crystal fractionation and crustal assimilation.

This study was designed to examine silicic magmatism in detail, and the data base is not representative enough of the mafic system to allow for rigorous evaluation of the model presented above. The following sections will, however, impose significant constraints on the general magmatic processes invoked by Hughes (1983, 1990), and Hughes and Taylor (1986).

7.2 Crystal Fractionation

The simplest model of mafic petrogenesis to test is whether or not a series of rocks are related through crystal fractionation. Crystallization is probably the single most important petrogenetic process in the compositional evolution of a magma, as it is the inescapable result of the loss of heat from a magmatic system. Assimilation of wall rocks and the mixing of magmas must also result in some crystallization of the parental magma, as crystallization supplies the heat required to raise an assimilant to its solidus and produce melting (Bowen, 1928). It will be shown that even in a dynamic, open magmatic system, the crystallizing mineral assemblage will exert the most control on the major element evolution of the system (cf., Defant and Nielsen, 1990).

The mineralogy, and inferred crystallization sequence, of mafic rocks in the study area is olivine (ol) \pm plagioclase (plag) to \approx 53% SiO₂, clinopyroxene (cpx) + ol + plag for 53%-56% SiO₂, cpx + ol + plag \pm orthopyroxene (opx) for 56%-58% SiO₂, and cpx + opx + plag (\pm trace ol) to 63% SiO₂. Phenocrysts of Fe-Ti oxides are also present in all mafic rocks, and apatite occurs as a trace phenocryst in rocks with \geq 57% SiO₂. The modal proportions of these phases in a rock, however, do not necessarily reflect the fractionating mineral assemblage. Crystal sorting by size and density in a convecting magma body (Marsh and Maxey, 1985) can readily change the phenocryst proportions observed in a rock. In-situ crystallization (Langmuir, 1989; McBirney and Noyes, 1979) can also involve the apparent nonmodal fractionation of minerals. In addition, some basalts from the study area and many other calc-alkaline continental arcs (cf. Gill, 1981; Thorpe, 1981) show evidence of clinopyroxene crystallization, yet lack that mineral in the mode.

Recent experimental work on the phase relationships of different types of basaltic melts, under variable pressures and water contents, can be used to constrain crystal fractionation models more accurately than with the modal mineralogy. Much of the work on natural systems has focused on anhydrous melts at 1 atmosphere pressure (Walker et al, 1979; Grove et al., 1982; Grove and Bryan, 1983), which are used to define the cotectics in figure 7.1. The general effects of increasing pressure under anhydrous conditions on some basaltic phase relationships has been studied by Thompson (1974, 1975), Stolper (1980), Takahashi and Kushiro (1983), and Elthon and Scarfe (1984). These studies show that increasing pressure will generally expand the orthopyroxene and clinopyroxene phase volumes at the expense of olivine \pm plagioclase. Low pressure (≤ 2.5 kbar) experiments on water-undersaturated basalts by Spulber and Rutherford (1983) indicate the olivine phase volume is expanded while the plagioclase volume contracts. Increasing water content will also increase the anorthite content of the liquidus plagioclase (Eggler, 1972; Sekine et al., 1979) and stabilize magnetite as an early crystallizing phase (Grove and Sisson, 1991).

Rocks from the study area and 1 atmosphere cotectics from Grove et al. (1982) are plotted in figure 7.1. The samples plot well off of the 1 atmosphere cotectics, but note that the low-Ol end of the data is sub-parallel to these cotectics, especially for higher silica compositions. This may be due to the effect of closure at higher silica compositions, but similar offsets at Medicine Lake volcano were interpreted by Grove and Baker (1984) to represent high pressure (5-10 kbars?) fractionation at water undersaturated conditions. Under these conditions olivine should be the liquidus phase, with early crystallization of magnetite, augite and calcic plagioclase. Olivine and augite dominate the crystallizing assemblage, with lesser amounts of plagioclase and magnetite (Grove and Baker, 1984; Grove and Kinzler, 1986; Grove and Sisson, 1991), leading to a calcalkaline mafic fractionation trend.

Anhydrous, high pressure experiments on high magnesium (MgO ≥ 9 wt%) basalt (Elthon and Scarfe, 1984; Gust and Perfit, 1987) and high aluminum (Al₂O₃ ≥ 17 wt%) basalt (Baker and Eggler, 1983, 1987; Johnston, 1986) can also provide useful constraints on the effects of pressure on basalt crystallization, and generally confirm the above mineralogical relationships. It is not possible, however, to rigorously apply the phase relationships from these studies to the mafic rocks in the study area, as differences in composition will change the phase relationships. For example, higher MgO contents will greatly expand olivine stability for P ≤ 20 kbars (figure 7.2). Compositional differences in studied high aluminum basalts also results in changes in liquidus temperatures and garnet stability fields (figure 7.3). While liquidus temperatures and exact orders of crystallization cannot be determined for High Cascade rocks from these data, the experimental data confirm that higher pressures will result in early crystallization of augite in anhydrous basaltic melts.



Figure 7.1: Low pressure phase relationships in the system olivine (Ol)- diopside (Di)- quartz (Qtz)- plagioclase, projected from plagioclase following the method of Grove et al., 1982:

Wt % oxide is converted to mole %, and cation normalized. Mineral components:

$$\begin{split} \text{SUM} &= \text{SiO}_2 - \text{CaO} - 2(\text{KO}_{0.5} + \text{NaO}_{0.5}) + \text{Cr}_2\text{O}_3 + \text{TiO}_2\\ \text{Qtz} &= (\text{SiO}_2 - 0.5(\text{FeO} + \text{MgO}) - 1.5\text{CaO} - 0.25\text{AlO}_{1.5} - 2.75(\text{KO}_{0.5} + \text{NaO}_{0.5})\\ &+ \text{Cr}_2\text{O}_3 + 0.5\text{TiO}_2)/\text{SUM}\\ \text{Plag} &= 0.5(\text{AlO}_{1.5} - \text{KO}_{0.5} + \text{NaO}_{0.5})/\text{SUM}\\ \text{Oliv} &= 0.5(\text{FeO} + \text{MgO} + 0.5(\text{AlO}_{1.5} - \text{KO}_{0.5} - \text{NaO}_{0.5}) - \text{CaO} - 2\text{TiO}_2 - \text{Cr}_2\text{O}_3)/\text{SUM}\\ \text{Diop} &= \text{CaO} - 0.5\text{AlO}_{1.5} + 0.5(\text{KO}_{0.5} + \text{NaO}_{0.5})/\text{SUM}\\ \text{Or} &= \text{KO}_{0.5}/\text{SUM}\\ \text{Sp} &= (\text{Cr}_2\text{O}_3 + \text{TiO}_2)/\text{SUM}\\ \text{Normalize} \text{Qtz} + \text{Plg} + \text{Oliv} + \text{Diop} = 100\%, \text{ then normalize} \text{Qtz} + \text{Oliv} + \text{Diop} = 100\% \text{ for a plagioclase projection.} \end{split}$$

+ = Broken Top system rocks, o = Three Sisters system rocks.



Figure 7.2: Comparison of high pressure phase relationships for high-Mg basalts NT-23 (A) (Elthon and Scarfe, 1984) and MK-15 (B) (Gust and Perfit, 1987). Significant differences in MgO, CaO, Na₂O and K₂O (C) result in different phase relationships. High-Mg basalts in the High Cascades are significantly different from these compositions, which precludes direct application of these high pressure phase relationships to the Cascade mafic system.



Figure 7.3: Comparison of high pressure phase relationships for high-Al basalts AT-1 (A) (Baker and Eggler, 1983) and SSS.1.4 (B) (Johnston, 1986). Significant differences in composition (C) result in different phase relationships for these rocks. High-Al basalts in the High Cascades are also significantly different from these compositions, which again precludes direct application of these high pressure phase relationships to the Cascade mafic system.

Pearce element ratios (PER) (Pearce, 1968) are commonly used to test if a series of rocks are cogenetic (cf. Russell and Nicholls, 1988). In order to quantitatively model with PERs, however, the extensive properties of one of the constituents in the system must not change; assimilation or mixing in the magma system will change the extensive properties of all the components in the system, and remove the constraints necessary for robust PER modeling. Disequilibrium mineral features, such as sieved plagioclase and rimmed pyroxenes (Chapter 4), indicate that the mafic magma system undoubtedly evolved with a significant component of mixing or assimilation. Most continental arc mafic systems are open magmatic systems (cf. Hildreth et al., 1986; Grove and Kinzler, 1986), and that rigorous application of PER models to these systems is not justified. PER models can, however, provide useful limits on the types and relative proportions of minerals that may have fractionated in the mafic magma system.

PER discrimination diagrams show the effects of fractionating different minerals on major element abundances. The hypothesis that clinopyroxene is a significant fractionating phase in the mafic system, along with olivine, plagioclase, and magnetite, was tested using numerator coefficients from Stanley and Russell (1989) and assuming that potassium was a conserved element. Plagioclase and olivine + oxide (figure 7.4a) fractionate in roughly equal proportions (\approx 1:1), olivine + oxide and clinopyroxene (figure 7.4b) fractionate at roughly 1.5:1, and plagioclase fractionates at roughly 2:1 with clinopyroxene (figure 7.4c). A combination of these trends results in a fractionating assemblage of plag:ol+ox:cpx of \approx 40:35:25, which is very similar to the assemblages used by Grove and Baker (1984) to generate numerous mafic calc-alkaline trends (i.e. plag > ol+ox ≥ cpx with ol+ox+cpx > plag).

Simple mass balance calculations, however, (table 7.1) show that fractionating these mineral proportions from a representative parental high MgO Cascade basalt (TFJ-438) fails to reproduce depletions in FeO* and TiO₂ observed in Cascade rocks (figure 7.5a). Mass balance models using these mineral proportions and more iron-rich olivine (Fo₇₀) and augite (En₄₀) also fail to produce the observed FeO* and TiO₂ depletions. However, simple mass balance models using Ol:Cpx:-Plag:Mgt of 25:22:40:13 produce the best fit to the observed trends for all major elements except Na₂O, which is relatively enriched in the model (figure 7.5b); elevated Na₂O was expected because in the model a plagioclase composition is fixed at An₈₅.

The role of magnetite fractionation in the petrogenesis of calc-alkaline arc rocks is at times controversial. A common problem involves the necessity of having 10%-15% magnetite in the fractionating assemblage in order to prevent titanium enrichment and create a silica enrichment trend (Gill, 1981; Grove and Baker, 1984; Grove and Kinzler, 1986). Cumulate gabbroic xenoliths from island arc volcanos, however, rarely contain $\geq 5\%$ magnetite (Beard and Day, 1988). It is often assumed that if $\geq 5\%$ of the fractionating assemblage were magnetite, then the cumulate xenoliths should also contain $\geq 5\%$ magnetite (cf., Conrey, 1990). This argument against $\geq 5\%$



Figure 7.4: Pearce element discrimination diagrams, using numerator coefficents from Stanley and Russell (1989) and assuming K is a conserved element. Slope of the trends gives fractionation ratio for tested minerals. A) Plag-Ol + Oxide fractionation occurred at ≈1:1. B) Ol + Oxide - Cpx fractionation at ≈1.5:1. C) Plag - Cpx fractionation at ≈2:1. Combination of these fractionation relationships gives an assemblage of Plag:Ol+Ox:Cpx of ≈40:35:25.

| Minera1 | | SiO2 | TiO2 | A1203 | Fe0* | MnO | MgO | CaO | Na20 | K20 | P205 |
|-----------|------|------|------|-------|------|-----|------|------|------|------------|------|
| Olivine | F085 | 39.9 | | | 14.9 | 0.3 | 45.1 | 0.2 | | | |
| Augite | En51 | 51.9 | 0.5 | 3.0 | 6.6 | 0.2 | 17.4 | 18.3 | 0.4 | | |
| Plag | An85 | 46.6 | | 34.0 | 0.5 | | 0.1 | 17.1 | 1.6 | 0.0 | |
| Magnetite | | 0.1 | 15.5 | 0.6 | 77.2 | 0.3 | 0.3 | | | | |

Starting composition

| | SiO2 | TiO2 | A1203 | FeO* | MnO | MgO | CaO | Na20 K | 20 F | 205 |
|---------|------|------|--------|------|-----|-----|-----|--------|------|-----|
| TFJ-438 | 49. | 8 1. | 5 16.5 | 9.8 | 0.1 | 9.5 | 8.6 | 3.3 | 0.7 | 0.2 |

Ending Compositions

| 0 | 1:Cpx:Plq:Mqt | <u></u> | | | | | | | | | | <u> </u> |
|----|---------------|---------|-----|------|-----|-----|-----|-----|-----|-----|-----|----------|
| 1) | 30:25:40:05 | 54.0 | 1.9 | 18.0 | 9.5 | 0.1 | 4.0 | 6.7 | 5.0 | 1.2 | 0.3 | 0.50 |
| 2) | 25:22:40:13 | 59.0 | 0.9 | 18.5 | 4.5 | 0.1 | 4.0 | 6.4 | 5.5 | 1.4 | 0.4 | 0.66 |

Observed range in composition at 4% MgO

| SiO2 | TiO2 | A1203 | FeO* | MnO | MgO | CaO | Na 20 | K20 | P205 |
|------|------|-------|------|-------|-----|------|--------|--------|------|
| 58 | 1. | 0 18 | 6 | 0.13 | 4 | 6.5 | 5 4 | 1.3 | 0.4 |
| ±3 | ±0. | 2 ±1 | ±1 | ±0.02 | ±1 | ±0.5 | 5 ±0.5 | 5 ±0.4 | ±0.1 |

Table 7.1: Mineral compositions used in mafic mass balance calculations, determined through microprobe analysis of phenocrysts in basalts and basaltic andesites (magnetite). F: Fraction crystallized. Simple mass balance calculations used a 2% crystallization interval, in which fixed proportions of the above minerals were removed from the parental composition. The derivative composition was re-normalized to 100%, and crystallization was modeled until MgO reached ≈4%, which is the point orthopyroxene appears as a phenocryst phase. Ending compositions for (1) calc-alkaline fractionation bounded by Pearce element plots, and (2) best fit to observed trends.



Figure 7.5: Results of simple mass balance models (symbols) plotted with best-fit trends for High Cascade mafic rock (lines). Right y-axis corresponds to Al₂O₃, other oxides on left y-axis.
A) Test of model bounded by Pearce element calculations, using an Ol:Cpx:Plag:Mgt of 30:25:40:05. Note the poor fit of the model to observed FeO and TiO₂ trends. B) Best-fit mass balance model using TFJ-438 as a starting composition and a Ol:Cpx:Plag:Mgt at 25:22:40:13. Calculations discussed in table 7.1.
magnetite fractionation assumes that the gabbroic xenoliths actually represent cumulates from the studied fractionation interval. Mafic magmas, however, have low viscosity ($\approx 10^2 \cdot 10^3$ poise, Murase and McBirney, 1973) and thus have a limit on the density of a xenolith that can be entrained and erupted. With all other factors being equal (i.e., size, composition, temperature, ascent rate etc.), a gabbroic xenolith with an Ol:Cpx:Plg:Mgt of 25:21:40:14 ($\rho \approx 3.4$) will have about a 33% greater settling velocity in a basaltic magma ($\rho \approx 2.7$) that a xenolith with an Ol:Cpx:Plg:Mgt of 30:25:40:5 ($\rho \approx 3.2$). The lack of gabbroic cumulate xenoliths with $\geq 5\%$ magnetite in many arc rocks may thus reflect the inability of basaltic magma to transport a relatively high density xenolith to the surface, and the absence of these xenoliths does not necessarily invalidate fractionation models involving $\geq 5\%$ magnetite.

The major element variation diagrams (figure 6.3) have a large degree of scatter at mafic compositions, which indicates that the mafic series is not the product of a uniform petrologic process. Simple mass balance calculations, however, show that the overall **trends** in the mafic system are consistent with a general petrogenetic model **dominated** by fractionation of calcic plagioclase, sub-equal amounts of olivine and augite, and $\geq 10\%$ magnetite. Similar degrees of major element scatter in fractionating mafic systems are often explained by invoking loosely constrained "open system processes", such as magma mixing, recharge, and assimilation (e.g., Hildreth et al., 1986; Hildreth and Moorbath, 1988; Hughes, 1990). However, limits can be imposed on the effect of open system processes using numerical mixing, recharge and assimilation models (e.g., O'Hara, 1980; DePaolo, 1981; Nielsen 1985, 1988, 1990; Langmuir, 1989).

7.3 Open System Models

The composition and proportion of minerals fractionating from an evolving magma are controlled primarily by temperature, pressure and the composition of the melt. Experimental data for the temperature and compositional dependence on phase equilibria can be used to model liquid lines of descent in anhydrous magma systems at low pressure (Nielsen and Dungan, 1983; Nielsen, 1985, 1988, 1990; Defant and Nielsen, 1990). The essence of this approach is to use experimental data to construct temperature-composition expressions that relate the composition of the studied system to a calculated liquidus mineralogy, and to crystallize the system under various conditions of recharge, eruption and mixing.

Trace element partitioning is strongly dependent on the composition of the system and temperature, and to much lesser extent on pressure (Irving, 1978; Mysen and Virgo, 1980; Nielsen, 1985). In order to account for compositional effects on trace element partitioning, the Nielsen (1985,1988) model assumes that the melt consists of two components: network formers (Si, NaAl, and KAl) and network modifiers (Al in excess of Na+K, all other elements). The use of a two

lattice melt model permits estimation of the activity of trace elements in the melt from their concentrations. The partitioning of the element into the fractionating mineral is then related to mineral component reactions, which account for substitution and siting effects (e.g., Gallahan and Nielsen, in press). This approach relates the activity of the trace element in the melt to it's activity in the fractionating phase, which results in compositionally compensated distribution coefficients.

The advantage of this model (referred to as CHAOS by Nielsen, 1990) is that the behavior of all components in the system is allowed to vary as the system evolves, as opposed to remaining fixed for relatively long intervals of differentiation (i.e. mass balance calculations or Rayleigh fractionation). The distribution coefficients used in the model are calculated for the phases that are in equilibrium with the evolving melt, and are thus more accurate than the general mineral/matrix distribution coefficients commonly used in trace element modeling. This approach permits detailed calculation of major and trace element liquid lines of descent, and allows quantitative assessment of magma recharge, mixing, assimilation and eruption effects on the evolving system.

The experimental data that constrain CHAOS models, however, are from anhydrous melts at 1 atmosphere. Previous arguments have indicated that the mafic system in this part of the Oregon Cascades evolved at elevated pressures (\approx 10 kbars), and that the system was not anhydrous but water undersaturated. CHAOS models will thus be dominated by the plagioclase + olivine cotectic, with later crystallization of augite and magnetite; the Oregon Cascades mafic system should have early crystallization of all four of these phases, and the plagioclase should be more calcic than modeled with CHAOS. For this reason, the following CHAOS models are not presented as rigorous simulations of Oregon Cascade mafic petrogenesis, but rather as a useful test of what **degrees** of elemental variation can be expected from different open system processes.

7.3.1 Crystal fractionation

Differences between liquid lines of descent modeled with CHAOS and those observed in the studied mafic system are represented in figure 7.6. The closed system CHAOS model with TFJ-438 (Hughes, 1983) as the starting composition used a 2% crystallization interval and was produced with no recharge, assimilation or eruption, which is referred to as a 0-0-0 model. Early crystallization of plag + ol results in little silica enrichment until the appearance of augite on the liquidus at $\approx 60\%$ crystallization. A 0-0-0 model was also calculated using basalt 3S069 as the starting composition; augite fractionation occurred at only $\approx 40\%$ crystallization in this model. It is apparent that simple mass balance models more accurately account for the observed major element variations in this system than 0-0-0 CHAOS models using either TFJ-438 or 3S069 as parental compositions (figure 7.6).

It is also apparent from figure 7.6 that there are numerous "parental" compositions that can be used as starting compositions in these models. Primitive basalts (i.e. close to equilibrium with



Figure 7.6: Closed system (0-0-0) models projected from plagioclase (Grove et al., 1982). Note that an unreasonably large amount of crystallization (≈60%) is needed to reach the olivine-augite cotectic using TFJ-438 as the starting composition. If 3S069 is used as the starting composition, only ≈40% crystallization is necessary to reach the olivine-augite cotectic. The simple mass balance model, however, more closely corresponds to the High Cascade mafic trend than either 0-0-0 model.

mantle peridotite) are rarely found in volcanic arcs (Nye and Reid, 1987), and have not been reported in the Oregon Cascades. High Cascade basalts that are high in MgO, Ni and Cr have higher incompatible element abundances than some High Cascade basaltic andesites (Hughes, 1983), which further obscures parental relationships. The geochemical diversity observed in this area (e.g., figure 6.3) shows that there are many basalts in this part of the High Cascades that can serve as parental compositions. The purpose of the following models is not to rigorously describe the liquid line of descent for one parent to a specific daughter, but rather to understand the geological processes that are acting on the magma system as a whole. Given the current uncertainties in the composition of parental melts, the effects of pressure and water content on the liquid line of descent, trace element distribution coefficients, and the composition of possible crustal assimilants, rigorous geochemical modeling of the High Cascade mafic system is not justified. For the purposes of this study, a relatively primitive basalt was used as a parental composition in the petrogenetic models; basalt 3S069 has relatively high MgO (8.8%), Ni (148 ppm) and Cr (451 ppm), low incompatible element abundances, and phenocryst modes of 10% olivine but only 0.1% plagio-clase.

7.3.2 Magma recharge and eruption

Although the absolute elemental abundances produced by CHAOS models will not accurately simulate the High Cascade magmatic system, due to the low pressure data used to construct the models, the relative abundances generated by CHAOS yield valuable information on the geochemical effects of various open system processes. These CHAOS models provide the means to test various open system hypotheses that are frequently invoked, but rarely constrained, in most mafic petrogenetic studies.

Fractionation paired with magma recharge and eruption is a relatively simple open-system model that can be simulated with CHAOS. A steady state model can be constructed by setting the recharge rate to twice the crystallization rate and the rate of eruption equal to the rate of crystallization, which is referred to as a 2-0-1 model. Randomized recharge events were set to occur after 20% of the system had been crystallized and erupted, using 5% crystallization increments. The modeled system was randomly sampled after every 15% interval of crystallization; variations in the periodicity of recharge (10%-40%), crystallization increment (1%-10%) and sampling (10%-40%) only produced small deviations from the above model.

A 2-0-1 model for 3S069 (figure 7.7) illustrates the main features of a recharging mafic magma system. The most significant feature is the inability of a recharging mafic system to routinely fractionate to compositions more silicic than basaltic andesite. Recharge rates that permit a steady state (i.e., long lived) magma chamber continuously drive the bulk composition of the system back towards a parental composition, and constrain the system to higher temperature



Figure 7.7: 2-0-1 recharge model using 3S069 as the starting composition. Note that the overall trend of the model parallels the 1 atmosphere cotectics (cf., figure 7.6), with significant deviations produced through the mixing of parental and evolved magmas. Increased pressure should expand the augite stability field at the expense of olivine, and lower the augite-olivine cotectic into the High Cascade data field (arrows). Much of the observed variation in the High Cascade mafic series can be accounted for by the 2-0-1 recharge model.

cotectics (O'Hara and Mathews, 1981; Nielsen, 1988). Thus the only way to generate silicic melts from a basaltic parent through crystal fractionation (i.e., without assimilation) is to close the system to significant amounts of recharge, which greatly limits the amount of evolved magma that can be produced (e.g., >50-60% crystallization necessary for andesite).

The liquid line of descent in the 2-0-1 model, however, is still controlled by the position of the 1 atm. cotectics, with only minor deviations off the cotectics produced through mixing (figure 7.7). This is in part an artifact of the anomalously large 1 atm. olivine-plagioclase stability field used in the models, which limits the fractionation of augite to intervals of >40% crystallization. Previous discussions have shown the necessity of augite and magnetite crystallization early in the fractionation sequence. Early fractionation of augite would result in more apparent scatter in the modeled 2-0-1 trends because the augite-olivine cotectic would be reached at much lower degrees of crystallization. Increasing the sampling, mixing interval and simulation length could also produce a little more apparent scatter off of the cotectics in modeled systems (Nielsen, 1988).

Most models of continental arc magmatism contain a large, long lived mafic magma reservoir near the base (40-60 km) of the crust (e.g., Gill, 1981; Marsh, 1984; Hildreth and Moorbath, 1988; Bergantz, 1990). Magmatic differentiation in this reservoir is probably dominated by recharge, with lesser amounts of wall rock assimilation due to a large volume-to-surface ratio of the magma reservoir and the difficulty in mixing silicic partial melts down into the mafic reservoir (Huppert and Sparks, 1988). The relative importance of fractionation versus assimilation in mafic differentiation cannot be rigorously tested until additional experimental work on high pressure (~10 kbars) phase relationships is completed. With that caveat in mind, some of the major (figure 7.7) and trace (figure 7.8a) element variations observed in High Cascade rocks from 49% to about 58% SiO₂ are reproduced by a 2-0-1 recharge and eruption model, once the effects of higher pressure fractionation are qualitatively accounted for. One effect of recharge is to maintain compatible element abundances at higher levels than observed in closed-system fractionation models (Defant and Nielsen, 1990). Some High Cascade mafic rocks have compatible element abundances (e.g., Ni, Cr, V, Co) that are consistent with a 2-0-1 model (figure 7.8a). Although magma recharge and eruption must exert a significant degree of control on the mafic differentiation trends in this part of the High Cascades, these processes do not explain all of the variations observed within the High Cascade mafic system (e.g., figure 7.8a-b). Further complexities must be introduced by recharge of different primary melt compositions (i.e., Hughes, 1983, 1990; Hughes and Taylor, 1986), or by crustal assimilation, in order to account for the observed geochemical variations.

Pearce (1983) recognized that incompatible elements with low ionic potentials (Cs, Rb, Sr, Ba, large ion lithophiles (LIL) and light rare earths (LRE)) are characteristically enriched in subduction zone magmas, relative to heavy rare earth (HRE) and high field strength (HFS) elements. The low ionic potentials of the enriched elements favor aqueous mobility, which has lead



Figure 7.8: A) Representative compatible-incompatible trace element diagram for High Cascade mafic samples. Note that modeled element distributions are very similar to some observed compositions; most trace elements that are compatible in olivine (± augite) follow this pattern. Although modeled abundances will at times deviate from observed abundances, due to the 1 atm. crystallization path in the model, the modeled patterns generally reflect the observed distribution patterns. B) Representative compatible-incompatible element diagram for modeled elements that do not follow observed trends. The large amount of observed Sr variation is not explained by the modeled degrees of magma recharge and eruption, and must have been produced through another open system processes. The 2-0-1 model also over estimates the amount of plagioclase fractionation, due to the large stability field of plagioclase in 1 atm. anhydrous mafic systems.

Pearce (1983) and others (e.g., Hawkesworth and Ellam, 1989; Hughes, 1990) to speculate that heterogeneities in primary melt composition may be produced by variable degrees of fluid enrichment in the mantle wedge. This hypothesis, however, is difficult to test, as there are few constraints placed on the degree and types of enrichments that are possible in the mantle wedge. The effects of fluid enrichment in subduction zone magmatism is a fundamental problem in igneous petrology, and one that cannot be adequately addressed by the data in this study. The elements that are most strongly enriched in a metasomatized mantle source, however, are also the elements that are incompatible in the major rock forming silicate minerals. Incompatible element enrichments can also be produced through mixing silicic partial melts into the mafic system. The next section will show that this hypothesis is valid for the High Cascades mafic magma system.

7.3.3 Assimilation

Experimental work has shown that silicic melts can be produced through partial melting of basaltic and metabasaltic rocks under a wide range of temperatures, pressures, and water contents (e.g. Helz, 1976; Spulber and Rutherford, 1983; Ellis and Thompson, 1986; Rutter and Wyllie, 1988; Beard and Lofgren, 1989). Low degrees ($\approx 10-30\%$) of melting generally produces rhyolitic partial melts, which are broadly similar in composition to rhyolites in the study area (chapter 9) and resemble granitic minimum melts (cf., Tuttle and Bowen, 1958).

Rhyolitic partial melts may be abundant near the base of the crust, where the ponding and crystallization of mafic magmas will generate sufficient heat for crustal fusion (e.g., Patchett, 1980; Huppert and Sparks, 1988; Bergantz, 1989). Assuming a geothermal gradient of 20°C/km and the base of the crust at \approx 40 km, lower crustal rocks should be at \approx 800°C. Metabasalt should have a specific heat of \approx 0.3 cal/g °C, a heat of fusion (or latent heat of crystallization) of \approx 100 cals/g and a solidus at \approx 1000°C (after Grove et al., 1982). A melt-to-wallrock ratio of \approx 1.5:1 should thus result in partial melts of the metabasaltic wallrock; if the melts convect, this ratio will be lower (cf. Huppert and Sparks, 1988).

It is likely that the temperature of these rhyolitic melts was close to the liquidus. Petrographic study has shown that plagioclase phenocrysts with resorbed, low An cores and reverselyzoned rims are rarely found in mafic rocks from the study area. Such reversely-zoned phenocrysts are most likely xenocrysts from the mixing or assimilation of partially crystalline silicic magmas or wall rocks. For example, Conrey (1991) has shown that mixing a silicic magma into the Mt. Jefferson mafic system usually resulted in abundant resorbed and reversely-zoned plagioclase phenocrysts. The absence of abundant reversely-zoned phenocrysts does not prove that the mixed silicic magma was aphyric, however, because the lower temperature phenocrysts may have been completely resorbed during homogenization. The repeated absence of these phenocrysts, however, may indicate that mixing involved sparsely phyric silicic magmas. Sparsely phyric rhyolites from the

study area have phenocryst equilibrium temperatures of around 850°C (Chapter 5), and may thus have had a liquidus temperature in excess of 900°C.

The experimental and thermal data support the hypothesis that rhyolitic partial melts should be associated with any large, long-lived mafic magma system ponded near the base of the crust. The effects of assimilation/mixing of these rhyolitic melts into the mafic magma system can be modeled using CHAOS. Because the rhyolitic melts are assumed to be cooler ($\approx 900 \,^\circ$ C) than the mafic melts ($\geq 1100 \,^\circ$ C), mixing will result in crystallization of the evolving magma. In order to produce a steady-state simulation, recharge was set to 2, assimilation (i.e., mixing of rhyolite) to 0.3, and eruption at 1.3, with mixing events randomized at 35% intervals of fractionation. Larger amounts of assimilation resulted in very short simulations, presumably due to the quenching effect of the rhyolitic assimilant. Small rates of assimilation (10-20%) generally resembled 2-0-1 models, and rarely evolved to andesitic compositions. Ranges of assimilation-to-crystallization for AFC-type models in the literature range up to 1:1 (DePaolo, 1981), although 0.3-0.5:1 is probably more realistic (Taylor, 1980; Hildreth and Moorbath, 1988; Grunder and Mahood, 1988).

The main results of a 2-0.3-1.3 model using rhyolite 3S032 as the assimilant are shown in figures 7.9-7.11. The primary effect of assimilation is to drive the evolving system off of the cotectics to a much greater extent than observed for 2-0-1 recharge models (figure 7.9). The assimilating system will still be bounded by the cotectics, especially the cpx-opx-plag cotectic, but the higher temperature cotectics should be less well defined. The degree of major element variation for the assimilation model more closely corresponds to the observed variations than found in 2-0-1 models (figure 7.9), although this may be in part due to the limited mafic data set. More importantly, this model allows the mafic system to routinely evolve beyond 58% SiO₂, which was not possible with a 2-0-1 recharge model.

Trace elements that are compatible in olivine and pyroxene, as represented by Ni in figure 7.10a, are generally enriched in the assimilation model relative to 2-0-1 models. Ni bulk distribution coefficients will not be strongly affected by changes in pressure, because Ni remains highly compatible in high and low pressure mafic assemblages due to the early crystallization of olivine. The assimilation model can directly account for numerous samples that lie off of the main recharge-fractionation trend (figure 7.10a). However, Sc-Rb variations (figure 7.10b) are not duplicated by CHAOS models, because augite does not fractionate early in anhydrous, 1 atmosphere mafic systems. The lack of a Sc enrichment trend in High Cascade basalts indicates that augite is an early fractionating phase in this system, which is consistent with differentiation at pressures of \approx 5-10 kbars (cf., Elthon and Scarfe, 1984). Note that although the absolute abundances of Sc and Rb from CHAOS models in figure 7.10b are higher than observed in the High Cascade rocks, the degree of scatter in the 2-0.3-1.3 model is very similar to the scatter in the mafic High Cascade rocks.



Figure 7.9: 2-0.3-1.3 assimilation model using 3S069 as the starting composition, and rhyolite 3S032 as the assimilant. Note that the evolving system is much more offset from the cotectics than observed in the 2-0-1 recharge model. The degree of variation in the assimilation model closely corresponds to observed variations in the mafic system. The mixing of rhyolitic melts into the mafic system is necessary for the mafic system to evolve beyond 58% SiO₂, and to explain some of the observed variations that cannot be accounted for through recharge or closed system fractionation.



Figure 7.10: Representative compatible-incompatible trace element diagram for High Cascade mafic samples. A) Note that the assimilation model accounts for numerous samples that lie above the main recharge (2-0-1) trend, and the greater degree of variation shown by the assimilation model. Trace elements that are compatible in olivine follow this pattern.
B) Assimilation is also required to explain many of the distributions of trace elements that are compatible in augite. Although the absolute abundances of Sc and Rb do not match observed abundances due to the late crystallization of cpx in 1 atm. models, the overall patterns in the 2-0.3-1.3 model closely match much of the observed distribution.

The distribution of other incompatible to moderately compatible trace elements (figure 7.11), which are poorly constrained by 2-0-1 models, can be accounted for by a 2-0.3-1.3 assimilation model. Ba/La and La/Sm ratios will not be strongly affected by the fractionation of olivine, pyroxene, magnetite or plagioclase. The large degree of variation in these elemental ratios (figure 7.11) also indicates that assimilation must be occurring with fractionation and recharge.

However, neither recharge or assimilation CHAOS models adequately account for the relatively high observed abundances of Sr and Al_2O_3 (figure 7.12). This may be in part due to the unusually low abundance of Sr (440 ppm) in 3S069 relative to other analyzed basalts (550-600 ppm). Simple mass balance models discussed above required $\approx 40\%$ plagioclase in the fractionating assemblage to match the observed Al_2O_3 distributions. The observed enrichments in Sr (figure 7.12) are consistent with 50% crystallization of an assemblage containing 40% plagioclase (Sr distribution coefficient of 1.5 to 2), assuming the initial concentration of Sr was 550 ppm. CHAOS models consistently over estimate the degree of plagioclase fractionation ($\approx 60\%$ plagioclase in the fractionating assemblage), due to the relatively large stability field of plagioclase at 1 atmosphere, anhydrous conditions.

7.4 Summary of Mafic Petrogenesis

The High Cascade mafic magma system cannot be described by any single petrogenetic model. Profound uncertainties in mineralogic relationships at 10 kbars pressure and water undersaturated conditions precludes a rigorous evaluation of the role of crystal fractionation in the evolution of mafic rocks. The possibility of metasomatic enrichments in mantle-derived basaltic melts and heterogeneities in crustal assimilants further obfuscates general paradigms for mafic petrogenesis. There are, however, certain limits that can be placed on the evolution of mafic rocks in the study area.

The models presented above have shown the versatility of crystal fractionation and recharge in explaining some of the observed geochemical variations, even though these models are limited by the lack of high pressure mineralogical constraints. The High Cascade mafic system undoubtedly has numerous parental compositions (cf. Hughes, 1990), each of which can evolve a separate "micro-system" or blend into the larger magma system. To simply choose a basalt and differentiate it through Rayleigh fractionation is not a realistic test of the crystal fractionation hypothesis. The effects of magma recharge and eruption, which must occur in any open long-lived system, must be accounted for before the crystal fractionation hypothesis can be fully tested.

Silicic magmas must also at times be mixed into the mafic system, in order to account for many of the observed trace element variations from 49% to $\approx 58\%$ SiO₂, and to differentiate above 58% SiO₂. Thermal and experimental data show that silicic magmas can be generated through partial melts of essentially mafic lower crust. Indeed, it is difficult to envision how a large, long-



Figure 7.11: Representative incompatible trace element diagams. A) The degree of variation between elements with similar geochemical behavior can be successfully described with the 2-0.3-1.3 assimilation model. Note that the 2-0-1 recharge model cannot significantly fractionate Ba from La, because these elements have very similar mineral/melt distribution coefficients (≈0.1-0.2) in the fractionating mafic system. The assimilation model, however, results in significant Ba/La fractionation, due to the mixing of a rhyolitic melt with a different Ba/La than the starting composition. B) Fractionation of plag + px + mgt cannot significantly change the La/Sm ratio. Mixing of a rhyolitic melt with a different La/Sm is required to produce many of the observed REE distributions.



Figure 7.12: Sr-Rb (A) and Al₂O₃-MgO (B) variations cannot be accounted for by either recharge or assimilation CHAOS models. CHAOS models over estimate the amount of plagioclase fractionation at 60%; simple mass balance and Rayleigh fractionation are consistent with ≈40% plagioclase fractionation. Over fractionation of plagioclase results in anomalously low values of Sr and Al₂O₃, and limits the degree of Sr-Rb variation.

lived mafic magma system could be located in crustal rocks and not produce partial melts! The absence of restitic xenoliths or xenocrysts in presumedly mixed magmas indicates that either the interval of time after mixing was long enough to permit complete resorption of the incorporated material, or that the incorporated silicic magma was aphyric. Incompatible element enrichments, which are often explained as metasomatic enrichments in the mantle source, also can be produced in mafic magmas through reasonable amounts of silicic magma mixing paired with recharge. Metasomatic enrichments in the mantle are thus not a prerequisite for mafic magmatism in this part of the High Cascades.

The observed decrease in geochemical variation with increasing silica is not simply a matter of closure, but reflects fundamental changes in the processes that control magma evolution. A large mafic system that is open to recharge by parental basaltic melts cannot differentiate past $\approx 58\%$ SiO₂; recharge rates that permit a long-lived system also drive that system continuously back towards a parental composition. Routine differentiation to compositions more silicic than andesite must require either stopping the input of primary basaltic melts into the system, or increased rates of silicic magma mixing/crustal assimilation. Magma evolution above $\approx 58\%$ SiO₂ must proceed without recharge, which removes an important "randomizing" factor from the magma system. The relatively lower degrees of geochemical variation above $\approx 58\%$ SiO₂ thus reflects the lower number of "randomizing" magmatic processes. Crystal fractionation may occur in a system that is essentially closed to mafic input, and the mixing of silicic partial melts may become a more important petrogenetic process. The next chapter will test these hypothesis, and start to define more narrow evolutionary paths towards rhyolitic magmatism.

Finally, it is important to note that the Quaternary High Cascades mafic magma system does not display any apparent spatial or temporal controls on petrogenesis. For example, rocks that were apparently produced with relatively large amounts of assimilation (i.e., $\approx 30\%$) are not restricted to any specific location or time; they occur throughout the Quaternary High Cascades section. The petrogenetically distinct silicic magma systems in this part of the High Cascades (chapters 8 & 9) thus evolved from the same mafic magma system.

8. PETROGENESIS OF ANDESITES, DACITES AND RHYODACITES

8.1 Introduction

Models of mafic petrogenesis in the preceding chapter have demonstrated the importance of open system magmatic processes in generating the observed major and trace element variations for mafic rocks in this part of the Oregon High Cascades. Although the dynamics of magma recharge and assimilation of crustal melts produces distinct compositional trends, neither of these processes appears to be spatially or temporally controlled for mafic rocks in the Quaternary Oregon Cascades. A lack of spatial or temporal control supports the argument that the evolution of mafic rocks in this part of the High Cascades occurred in a large, long-lived magma system ponded at the base of the lower crust.

Compositional trends associated with more silicic rocks, however, often show spatial and temporal controls. This chapter will examine the petrogenesis of andesites, dacites and rhyodacites, and show that there are at times distinct spatial and temporal controls on the distribution of these rock types within the study area. For convenience, andesites, dacites, and rhyodacites will be referred to as the intermediate series throughout this chapter.

Unlike basalts and basaltic andesites, the andesites, dacites and rhyodacites were almost exclusively erupted from large vent complexes in the study area (stratovolcanos and the TVC). Exceptions are limited to small dacitic domes like Kokostick Butte (figure 1.1), which are as much as 5-10 km away from a central vent complex. This relationship is true for other volcanic centers in the High Cascades (cf., Conrey, 1991; Bacon and Druitt, 1988; Leeman et al., 1990), and for most other continental arcs (e.g. Gill, 1981; Thorpe, 1982; Grove and Kinzler, 1986).

There is a distinct temporal progression to the intermediate series in the Three Sisters system. The oldest stratovolcano, North Sister, consists primarily of basalt and basaltic andesite, with only a few small andesitic dikes emplaced in the upper part of the northwestern flank (Taylor, 1981; pers. comm., 1987). Middle Sister contains relatively abundant andesite and dacite, but basalt and basaltic andesite comprise the bulk of the volcano (Taylor, 1981). The youngest stratovolcano, South Sister, consists primarily of andesite and dacite, with lesser amounts of basaltic andesite and rhyolite (Taylor, 1981; Wozniak, 1982; Clark, 1983); rhyodacite (i.e. 68-72% SiO₂) is not found at any vents in the Three Sisters system. In addition, andesites and dacites associated with Middle Sister (MS) are also compositionally and texturally distinct from South Sister (SS) andesites and dacites. The Three Sisters thus show increasing amounts of different types of intermediate magmatism with both time and southward progression of volcanism. In contrast, there is no apparent spatial or temporal correlation among intermediate series rocks, including rhyodacite, were erupted throughout the history of each of these two vent complexes. Petrogenetic models for intermediate composition rocks will ultimately address the following points:

- 1) Did the same petrogenetic processes found in the High Cascade mafic system act in the intermediate system, or did these processes change?
- 2) Why did the intermediate series in the Three Sisters system only evolve to ≈66% SiO₂, but evolved to 72% SiO₂ in the Broken Top system?
- 3) Why are Middle Sister intermediate rocks compositionally distinct from South Sister intermediate rocks?

This chapter will begin to examine the compositional differences between the Three Sisters and Broken Top magmatic systems, and present models to explain these differences. The preceding chapter has identified the two dominant petrogenetic processes that control magma evolution within the High Cascades: The mixing of silicic partial melts into the evolving magma system, and crystal fractionation associated with mafic recharge.

Useful boundaries can be placed on the relative importance of open system fractionation versus silicic magma mixing through semi-qualitative modeling. The strength of these models lies not in the absolute values derived, but rather in the relative changes that are observed both within and between the studied magmatic systems. However, the geochemical behavior of trace elements in andesites and dacites is only moderately well constrained by experimental and empirical data. Large variations in many reported mineral/matrix distribution coefficients for intermediate composition rocks (Appendix E) precludes detailed trace element modeling. Phase relationships in intermediate composition rocks are also poorly constrained, as are the effects of water content and pressure. With these uncertainties, no useful purpose is served by presenting rigorously quantitative (i.e., grossly oversimplified) petrogenetic models.

8.2 Three Sisters System

Andesites and dacites associated with Middle Sister (MS) are compositionally and texturally distinct from South Sister (SS) andesites and dacites. Less than 5% of the plagioclase phenocrysts in MS intermediate rocks contain abundant glass inclusions, but \geq 50% of the plagioclase phenocrysts in SS intermediate rocks are usually sieved (Appendix B). Numerous SS dacites contain \approx 0.1-0.5% olivine phenocrysts, but MS dacites lack olivine (Appendix B). Intermediate series rocks from MS also tend to be less phyric (10-15% total phenocrysts) than SS rocks (10-35% total phenocrysts). These mineralogical features at SS suggest disequilibrium crystallization associated with magma mixing, and MS textures indicate equilibrium crystallization.

Although dacites from both MS and SS range to 66% SiO₂, MS dacites are characterized by generally lower compatible and incompatible trace elements (table 8.1). The textural and geochemical differences between these two systems may indicate that the MS intermediate series was produced dominantly through crystal fractionation, and the SS intermediate series through mixing between mafic and rhyolitic melts.

If crystal fractionation controlled the petrogenesis of intermediate composition rocks, then the crystallizing phases should be represented by the minerals observed in these rocks. Intermediate composition rocks from both the Three Sisters and Broken Top systems have contain phenocrysts of plag + cpx + opx + mgt ± ol from $\approx 56\%$ to $\approx 63\%$ SiO₂, with trace amounts of olivine observed in some SS dacites to 66% SiO₂. Augite and orthopyroxene occur in nearly equal proportions in the intermediate series rocks. Amphibole phenocrysts are only observed in intermediate rocks as a trace phase (1 crystal per thin section) in two TVC rhyodacite (68% SiO₂) domes. Trace amounts of highly resorbed, opacitic amphibole xenocrysts were found in TVC andesite 3S090 and SS dacite 3S150. There is, however, no mineralogical evidence to support resolvable amounts of amphibole fractionation in the intermediate series. Apatite occurs as inclusions in phenocrysts for most units with $\geq 57\%$ SiO₂, which is interpreted to represent local saturation in P₂O₅ at the mineral-melt interface (Chapter 4). Small (≤ 0.5 mm) groundmass crystals of apatite are often observed in rocks with SiO₂ $\geq 62-63\%$. This observation, and the experimental work of Green and Watson (1982), indicates that apatite fractionation did not occur in the intermediate series until $\approx 62\%$ SiO₂.

Simple mass balance calculations (table 8.2) using the above anhydrous phenocrysts show that any fractionation hypothesis must involve a significant amount of plagioclase (>50%) to deplete the system in Al₂O₃ and CaO. Magnetite must also be present (\geq 5%) to prevent TiO₂ enrichment and, in conjunction with olivine, give a silica enrichment trend. The best-fit to the observed intermediate major element trends is obtained with a fractionating assemblage of 8% ol, 15% cpx, 10% opx, 60% plag (An₇₅), 7% mgt and 0.4% apatite (table 8.2). Although small variations in either the fractionating proportions or in mineral compositions can have a significant effect on the dacite generated through simple mass balance models (table 8.2), these models support a fractionation hypothesis for the MS intermediate series.

Trace element variations also support a fractionation-dominated origin for MS dacites, but indicate that the fractionating system was probably open to some amounts of recharge and assimilation. Observed depletions in Sc relative to Rb enrichment (figure 8.1a) are consistent with $\approx 40\%$ equilibrium fractionation of an assemblage having a bulk distribution coefficient for Sc (D_{Sc}) ≈ 2 , with D_{Rb} = 0.01. Experimentally determined Sc distribution coefficients for pyroxenes in intermediate composition melts can be calculated from Nielsen et al., (in press), giving a

| | <u>Middle</u> | Sister | S Si | lster | |
|-------------|---------------|--------------|---------------|----------------|-----------|
| | 35132 | 3\$134 | 3\$130 | 35119 | |
| SiO2 | 61.9 | 65.7 | 62.0 | 65.8 | |
| TiO2 | 1,10 | 0.73 | 0,75 | 0,91 | |
| A1203 | 17 2 | 16 2 | 16 5 | 16 0 | |
| FeO* | 5 01 | 5 03 | 5 25 | 1 10 | |
| reo. | 0.14 | 5.03 | 5.25 | 4.49 | |
| MnO | 0.14 | 0.14 | 0.10 | 0.10 | |
| MgO | 1.86 | 0.99 | 3.85 | 1.35 | |
| CaO | 4.59 | 2.72 | 5.72 | 3.52 | |
| Na2O | 5.52 | 6.26 | 4.08 | 5.30 | |
| K20 | 1.47 | 1.89 | 1.64 | 2.34 | |
| P2O5 | 0.33 | 0.25 | 0.12 | 0.26 | |
| Rb | 24 | 31 | 35 | 46 | |
| Ba | 508 | 607 | 456 | 653 | |
| Sr | 483 | 272 | 352 | 329 | |
| Ca | 0.68 | 0.85 | 0.69 | 1.27 | |
| 55 | 0.00 | 0.00 | 0.05 | 1.4.4.1 | |
| Sc | 16.01 | 12.81 | 15.68 | 11.45 | |
| v | 58 | 12 | 109 | 67 | |
| Co | 9.76 | 3.56 | 20.66 | 7.73 | |
| Ni | <40 | 18 | 49 | <40 | |
| Cr | 3 | <6 | 84 | 5 | |
| Zn | 152 | 208 | ≈60 | 140 | |
| C 11 | 175 | 200 | | 140 | |
| La | 18.6 | 21.3 | 14.1 | 20.6 | |
| Ce | 44.1 | 53.8 | 28.5 | 47.1 | |
| Nd | 23.1 | 28.0 | 15.0 | 20.4 | |
| Sm | 5.29 | 6.75 | 3.15 | 4.95 | |
| Eu | 1.61 | 1.80 | 0.90 | 1.28 | |
| Tb | 1.04 | 1.27 | 0.57 | 0.90 | |
| Yb | 2.7 | 4.2 | 2.0 | 2.6 | |
| Lu | 0.41 | 0.62 | 0.28 | 0.41 | |
| La/Sm | 2 22 | 1 99 | 2 83 | 2 63 | |
| | 17 69 | 66 22 | 2.03 | 2.03 | |
| LIRES | 4/.00 | 00.22 | 0.02 | 44.22 | |
| Eu/Eu* | 0.80 | 0.// | 0.83 | 0./5 | |
| Zr | 176 | 241 | 110 | 206 | |
| Hf | 4.6 | 7.0 | 2.9 | 5.6 | |
| Ga | 20 | 21 | 13 | 22 | |
| Y | 27 | 38 | 19 | 26 | |
| Nb | 10.2 | 12.7 | 7.4 | 10.9 | |
| Ta | 0.54 | 0.62 | 0.45 | 0.74 | |
| Th | 2.3 | 2.9 | 3.0 | 4.7 | |
| U | 0.9 | 1.3 | 1.1 | 1.7 | |
| | | | | | |
| 35132 | Andesite | lava 1 mi | east of Linto | on Spring, ele | -vation = |
| 35134 | Lano Meg | a dacitic fl | ow dome east | margin 6460 |)' |
| 20120 | Lane Mes | flow 0 E mi | SW of Obeing | lian Falle | |
| 20110 | Anuesite | 100 U.5 M1 | . SW OF UDS10 | atan Falls. | |
| 32113 | Dacite f | IOW(7) in SS | E FLANK S. SI | lster, /320'. | |

Table 8.1: Comparision of representative Middle Sister and South Sister andesites and dacites. Note
the significant differences in compatible (V, Co, Cr) element and REE abundances between
MS and SS types. Σ HREE = chondrite-normalized Tb + Yb + Lu.

| | | Mineral compositions | | | | | | | | | |
|--------|------|----------------------|------|-------|--------|------|--------|------|------|-----|------|
| | | SiO2 | TiO2 | A1203 | FeO* | MnO | MgO | CaO | Na2O | K2O | P205 |
| Oliv | F085 | 39.9 | | | 14.9 | 0.3 | 45.1 | 0.2 | | | |
| Срх | En51 | 51.9 | 0.5 | 3.0 | 6.6 | 0.2 | 17.4 | 18.3 | 0.4 | | |
| Plag | An75 | 47.8 | | 33.1 | 0.6 | | 0.1 | 16.2 | 2.2 | 0.1 | |
| Орх | En72 | 52.0 | 0.3 | 0.7 | 19.2 | 0.9 | 25.1 | 1.5 | 0.0 | | |
| Mgt | BA45 | 0.1 | 15.5 | 0.6 | 77.2 | 0.3 | 0.3 | | | | |
| Ap | | | | | | | | 56.0 | | | 43.0 |
| | | | | | | | | | | | |
| | | | | St | arting | comp | ositio | n | | | |
| | | SiO2 | TiO2 | A1203 | FeO* | MnO | MgO | CaO | Na2O | K20 | P205 |
| 3\$106 | 5 | 55.9 | 1.0 | 18.3 | 7.1 | 0.1 | 4.6 | 7.8 | 4.0 | 0.9 | 0.2 |

BEST FIT

| Mod | el 1: | 84 01 | ., 15% | Cpx, | 104 Obx | , 60% | Plag, | 78 M | gt, | 0.4% Ap |
|----------|-------|-------|--------|------|---------|-------|-------|------|-----|---------|
| | SiO2 | TiO2 | A1203 | FeO* | MnO | MgO | CaO | Na2O | к20 | P205 |
| F=0.40 | 65.6 | 0.8 | 16.3 | 4.7 | 0.1 | 0.9 | 3.6 | 6.3 | 1.7 | 0.3 |
| MS 3S134 | 65.7 | 0.7 | 16.2 | 5.0 | 0.1 | 1.0 | 2.7 | 6.3 | 1.9 | 0.3 |

| | Model 2: | 10% 01, | 13% C | Срх, 5% | Opx, 62% | Plag, | 10% 1 | Mgt, | 0% Ap |
|-------------------|-------------------------|-----------------------------|-----------------------|------------------------------|-----------------------|-------------------|--------------------|-------------------|--------------------|
| F=0.44 MS 3S13 | SiO2 65.8 34 65.7 | TiO2 A1 0.7 16 0.7 16 | 203 F .2 5 .2 5 | FeO* Mr. 5.0 0. 5.0 0. | MgO 1 1.9 1 1.0 | CaO 4.3 2.7 | Na2O 5.9 6.3 | K2O 1.6 1.9 | P2O5 0.4 0.3 |

Model 3: Same mineral %'s as model 1, 8% Ol, 15% Cpx, 10% Opx, 60% Plag, 7% Mgt, 0.4% Ap But more evolved compositions:

| Oliv Fo70 Cpx En39 Plag An60 Opx En68 | SiO2 36.8 52.9 52.7 53.9 | TiO2 0.8 0.3 | A12O3 3.1 29.1 1.0 | FeO* 26.4 13.1 0.9 18.7 | MnO 0.5 0.4 0.5 | MgO 35.0 12.7 0.1 24.7 | CaO 0.2 16.8 12.1 1.4 | Na2O 0.8 4.5 0.0 | к20 0.3 | P205 |
|--|--------------------------------------|--------------------|-----------------------------|-------------------------------------|--------------------------|------------------------------------|-----------------------------------|---------------------------|-------------------|--------------------|
| F=0.26 MS 3S134 | SiO2 65.5 65.7 | TiO2 0.7 0.7 | A12O3 17.9 16.2 | FeO* 1.8 5.0 | MnO 0.1 0.1 | MgO 1.6 1.0 | CaO 4.9 2.7 | Na2O 5.6 6.3 | K2O 1.8 1.9 | P2O5 0.3 0.3 |

Table 8.2: Simple mass balance models used to test the crystal fractionation hypothesis for Middle Sister dacites, using the mass balance method described in table 7.1. Starting composition is a representative High Cascade basaltic andesite, and mineral compositions are from analyzed samples in this study. Model 1 represents best fit to observed MS compositions. Models 2 and 3 show how small changes in the fractionating proportions (2) or mineral compositions (3) can significantly change the modeled dacite composition.



Figure 8.1: Representative compatible-incompatible element plots for Three Sisters rocks. A) Depletions in Sc relative to Rb enrichment in the MS intermediate series are consistent with open system fractionation of a typical High Cascade basaltic andesite, involving ≈20% pyroxene. Note the relatively low depletions in Sc relative to Rb enrichment for the SS intermediate series. B) Depletions in V relative to Rb enrichment in the MS intermediate series also are consistent with open system fractionation of a typical High Cascade basaltic andesite, involving ≈7% magnetite. Note the relatively low depletions in V relative to Rb enrichment for the SS intermediate series also are consistent with open system fractionation of a typical High Cascade basaltic andesite, involving ≈7% magnetite. Note the relatively low depletions in V relative to Rb enrichment for the SS intermediate series.

cpx $K_{dSc} = 5$ and opx $K_{dSc} \approx 2$. The resulting $D_{Sc} \approx 1$ (using mass balance %'s) is too low to have produced the observed depletions in Sc relative to Rb enrichment in figure 8.1a, using simple equilibrium fractionation.

Steady-state recharge of the fractionating system by basaltic andesite melt, however, will produce larger incompatible element enrichments relative to compatible element depletions than are possible through simple equilibrium fractionation (Nielsen, 1988; Defant and Nielsen, 1990). This relationship is shown in figure 8.2 for Sc and V, two elements that are strongly compatible in px + mgt. Opening the evolving system to recharge (2-0-1) by parental basaltic andesite melts results in mixed compositions that are lower in compatible elements, relative to a closed system (0-0) model. Combining recharge with modest amounts of silicic magma mixing ($\approx 20\%$) also results in incompatible element enrichments, relative to both closed system fractionation and recharge. Thus, if only 5 ppm of the observed Rb enrichments were due to simple fractionation effects and the other 5 ppm were due to open system recharge and eruption, the resulting depletions in Sc for the MS intermediate series could be produced with an effective $D_{Sc}\approx 1$.

The same arguments can be applied to V depletions in figure 8.1b. To produce the observed depletions in V through simple equilibrium fractionation would necessitate a $D_V \approx 7$ and a magnetite $K_{dV} \approx 80$, which is twice as high as generally measured in andesitic rocks (Appendix E). Open system processes indicated by Rb enrichments, however, would lower the necessary magnetite K_{dV} to ≈ 40 , which corresponds to published mineral/matrix distribution coefficients.

The distribution of Zr versus Rb also clearly shows that the MS intermediate series was produced primarily through fractionation and recharge of a basaltic andesite to dacite system, but that a significant amount of silicic magma mixing is required to account for the observed SS patterns (figure 8.3). Zr remains an incompatible element during fractionation until zircon saturation is reached, which is a function of both temperature and composition of the melt (Watson and Harrison, 1983). Application of Watson and Harrison's experimental results shows that zircon saturation for Three Sisters (and BT) intermediate compositions requires >500 ppm Zr in the melt for temperatures >930°C, but ≤ 250 ppm Zr at temperatures of $\leq 860°$ C. Zr should thus behave incompatibly during fractionation of the intermediate series, as temperatures are likely $\geq 930°$ C (Chapter 5). The Zr/Rb of the MS series is very close to the Zr/Rb produced by CHAOS 0-0-0 and 2-0-1 models, indicating that fractionation paired with basaltic andesite recharge can produce the MS intermediate series (figure 8.3). The buffering of Zr with respect to increasing Rb in the SS series cannot be produced through zircon fractionation, and must represent a significant amount ($\geq 20\%$) of rhyolite mixing into the evolving intermediate system.

Rare earth element (REE) trends are also consistent with a fractionation hypothesis for MS intermediate series rocks. Fractionation of cpx + opx + ol + plag + mgt will have little effect on La/Sm, due to the low distribution coefficients for these elements (Nielsen, 1990; Nielsen et al., in



Figure 8.2: Results of intermediate system CHAOS models. Models used a typical High Cascade basaltic andesite (3S106) under closed system fractionation (0-0-0, F=0.24), steady-state fractionation with randomized recharge events (2-0-1) and fractionation paired with recharge and mixing of Three Sisters rhyolite magmas (2-0.2-1.2). Both open system models were run until F=800. Relative to the High Cascades, the models are enriched in compatible elements due to the late crystallization of augite (Sc) and magnetite (V) in CHAOS simulations. These models show that incompatible elements are relatively enriched when the system is open to rhyolite mixing, and that magma recharge will result in lower compatible element abundances relative to simple fractionation.



Figure 8.3: Results of CHAOS models showing variations in Zr vs. Rb. A) Closed system fractionation (0-0-0) or recharge (2-0-1) models account for the observed variation between Zr and Rb in the MS system, as zircon is not a fractionating phase for these compositions. Open system (2-0.2-1.2) models using rhyolite 3S139 are necessary to successfully model Zr vs. Rb variations in the SS system. B) CHAOS trends superimposed on Three Sisters data. Note that the 2-0.2-1.2 model is relatively enriched in Zr, indicating that higher assimilation rates are probably necessary to account for the SS data. Higher assimilation rates, however, quench the system during CHAOS simulations.

press). The incompatible heavy REE's Tb, Yb, and Lu (cpx $K_d \approx 0.8$) will also be enriched in the melt, with negligible fractionation between these elements (Nielsen et al., in press). The fractionation of apatite, however, can potentially alter La/Sm, due to the relatively high compatibility of the MREE's (Appendix E). However, apatite fractionation is limited to differentiation above 62% SiO₂, and constrained by mass balance calculations to involve $\leq 0.5\%$ of the fractionating assemblage. REE trends observed for MS intermediate rocks are consistent with the fractionation hypothesis, with little MREE depletion or La/Sm fractionation occurring with differentiation from a basaltic andesite parent (figure 8.4). The trends observed for SS rocks, however, are not consistent with a fractionation hypothesis; the large increases in La/Sm and depletions in HREE abundances necessitate $\geq 5\%$ apatite in the fractionating assemblage, which is not supported by either mass balance or experimental models (Green and Watson, 1982).

The intermediate rocks at Middle Sister are thus the products of open-system differentiation from a periodically recharged basaltic andesite to dacite magma system. Although small amounts of rhyolite magma mixing may have contributed to some of the geochemical variations observed in the MS intermediate series, this series was essentially produced through crystal fractionation. In addition, this system must have been closed to large amounts of recharge by parental **basaltic** melts or else it would not have been able to differentiate into intermediate compositions. Because mafic volcanism has occurred throughout the history of this area, it is highly unlikely that closure of the evolving MS intermediate system to mafic recharge was due to a cessation of mafic magmatism at depth. Segregation of the MS intermediate magmatic system to shallower crustal levels (20-30 km??), however, may have served to effectively isolate the system from basaltic recharge.

In contrast to Middle Sister, crystal fractionation is not capable of generating the South Sister intermediate series. Fractionation of pyroxene, plagioclase, olivine and magnetite cannot generate the observed La/Sm enrichments, which are associated with little increase in HREE abundance (figure 8.4). Fractionation of $\geq 5\%$ apatite might produce some of the observed HREE depletions, but P₂O₅ generally increases with differentiation in the SS intermediate series (table 8.1), precluding any significant amount of apatite fractionation; simple mass balance calculations show that crystallization of only 0.5% apatite will buffer the melt in P₂O₅. The observed decreases in Zr/Rb (figure 8.3) could have been produced through zircon fractionation, but the SS intermediate series was too hot ($\geq 950^{\circ}$ C) to have been saturated in zircon at only ≈ 200 ppm Zr (Watson and Harrison, 1983).

South Sister intermediate rocks generally plot along mixing lines between basaltic andesite and Three Sisters rhyolite (figures 8.1 and 8.3). Linear mixing of $\approx 50\%$ basaltic andesite (e.g., 3S106) with $\approx 50\%$ rhyolite (e.g., 3S140) reproduces many of the general major and trace element abundances of SS intermediate rocks, including REE patterns (figure 8.5a). Crystallization must have accompanied mixing, as the rhyolite was probably at a significantly lower temperature





Figure 8.4: Rare earth element plots for Three Sisters rocks. A) Chondrite-normalized (Anders and Ebihara, 1982) REE diagram for a typical High Cascade basaltic andesite (3S106), MS dacite (MS 134) and SS dacite (SS 119). Note that the REE trends for the MS dacite and basaltic andesite are parallel, but the SS dacite has distinct HREE and LREE slopes.
B) Chondrite-normalized La/Sm (i.e. LREE slope) versus normalized Tb+Yb+Lu (HREE abundance) for Three Sisters rocks. Fractionation will produce little change in La/Sm with increasing HREE abundance (MS dacites). Mixing with Three Sisters rhyolites must have occurred for SS dacites in order to produce the observed increases in La/Sm with little change in HREE abundance.

Δ

B



Figure 8.5: Mixing diagrams for South Sister dacite petrogenesis. A) Chondrite-normalized REE diagram for a 50% mixture of basaltic andesite 3S106 with 50% Three Sisters rhyolite 3S140. The resulting mixture (140-106) has a REE pattern that is parallel (very slight MREE depletion) to SS dacite 3S119. Fractionation of plag + px + mgt ± ap will enrich all REEs to SS dacite levels. B) Enrichment-depletion diagram for Three Sisters (140-106) and Broken Top (32-106) rhyolite mixes with basaltic andesite 3S106, plotted relative to SS dacite 3S119. Although both mixtures have similar major element characteristics, a mix with Broken Top rhyolite results in anomalously high incompatible element abundances (Ta, Rb, Th, HREE), which precludes developing the observed compatible element depletions in 3S119 through crystal fractionation. A mixture with a Three Sisters rhyolite permits subsequent fractionation of trace elements to observed abundances.

123

Α

B

(\approx 900°C) than the basaltic andesite (\approx 1100°C). As shown by CHAOS models, the effect of crystallization will be to shift the mixing vectors towards MS fractionation trends, which is observed for SS intermediate rocks in figures 8.1 and 8.3. Mixing paired with \approx 20-30% fractionation of pyroxene, plagioclase and magnetite will also increase total REE abundances without REE fractionation, and produce relative depletions in Eu (cf., Middle Sister dacites). This type of fractionation will increase the REEs in the rhyolite-basaltic andesite mixture to the levels observed in SS dacites (figure 8.5a).

The mixing hypothesis for the SS intermediate series was also tested using rhyolite 3S032 from the Broken Top system as the silicic component. Although the overall patterns obtained with a BT rhyolite mix were similar to the 3S rhyolite mix (figure 8.5b), incompatible element abundances es were generally too high (especially HREE) and compatible element abundances too low to allow for any significant amount of fractionation subsequent to mixing; the depletions in Mg, Ca, Sr and transition metals observed in SS dacites require that mixing is accompanied by $\approx 20-30\%$ fractionation of pyroxene and plagioclase (figure 8.5b).

Evolution of the South Sister intermediate series must have been dominated by mixing subequal amounts of basaltic andesite and rhyolite magmas, with subordinate amounts of crystal fractionation. A detailed study of mafic-silicic magma mixing by Frost and Mahood (1987) in the Lamarck pluton, and the theoretical models of Sparks and Marshall (1986), concluded that homogenization of similarly mixed magmas would only occur if the mass fraction of mafic magma was ≥ 0.5 ; smaller mass fractions result in quenching of the mafic magma. Mixing lesser amounts of mafic magma into a rhyolitic magma body should thus result in mafic enclaves (cf., Bacon, 1986; Linneman and Meyers, 1990) and not a homogenous dacite or rhyodacite. Trace amounts of small (≤1 mm) mafic clots were observed in the most silicic SS dacites (3S119 and 3S150), but were absent from other studied intermediate rocks. The occurrence of these mafic clots in only the most silicic SS dacites further indicates that these compositions represent the upper silica limit for mafic mixing/homogenization. Restriction of the SS intermediate system to $\leq 66\%$ SiO₂ is not an arbitrary boundary, but most likely reflects the lowest percentages of mafic magma that could be homogenized in a Three Sisters rhyolitic melt. In addition, ≈50:50 mixing of a more silicic magma (i.e. 61% SiO₂ andesite) into the rhyolitic magma would result in a mixture that was more silicic than 66% SiO₂. Pyroxene and plagioclase fractionation coupled with mixing would further increase the SiO₂ content of the mixture, resulting in the formation of a rhyodacitic magma (i.e., ≥68% SiO_2). The absence of any rocks at South Sister with $SiO_2 \ge 66\%$ indicates that the lower silica component of mixing was mafic and not intermediate in composition.

The lack of distinct mixing trends between MS intermediate rocks and Three Sisters rhyolites may indicate that rhyolitic melts were not sufficiently abundant during MS evolution to generate intermediate compositions through extensive mafic-rhyolitic magma mixing. The previous discussion

of mafic magmatism has shown the necessity of $\approx 20-30\%$ silicic magma mixing/assimilation to produce many of the observed open-system trends. These silicic partial melts, however, were likely generated in the lower crust adjacent to a large, long-lived mafic system, and subsequently homogenized into that system. The effects of silicic magma mixing are not restricted to mafic rocks associated with silicic volcanic centers, but are present in mafic rocks throughout the Oregon High Cascades (cf., Conrey and Sherrod, 1988; Conrey, 1991; Hughes, 1990; this study). This observation suggests that silicic mixing is a regional (i.e., deep crustal) feature, and can occur in areas that lack intermediate to silicic volcanos.

Geochemical modeling also indicates that it is unlikely that the regional High Cascades mafic system could differentiate to intermediate compositions, due to the large amounts of basaltic recharge necessary to maintain thermal balance. The MS intermediate system must have segregated from the large mafic system near the base of the crust, and probably resided at some intermediate crustal level (30-20 km?). The lack of abundant rhyolite mixing could indicate that the rhyolitic magmas resided at a different crustal depth, or that there were no effective connections between the rhyolite and MS intermediate system. It is also possible that rhyolitic magmas were not present at mid-crustal levels during MS evolution. Stratigraphic correlations and K-Ar dates (Chapter 3) strongly indicate that rhyolitic volcanism in the Three Sisters is a relatively young (\approx 90 ka) feature, and probably post-dates most Middle Sister activity. The lack of petrologic evidence for rhyolite mixing, which is so abundant at South Sister, further supports the argument that rhyolitic magmatism was not appreciably developed during the evolution of Middle Sister.

Although it cannot be proven with the present data, it is likely that Middle Sister volcanism was contemporaneous (\pm 50 ka) with Broken Top volcanism. There is, however, no evidence to suggest that the MS intermediate system interacted in any way with the Broken Top magmatic system. In addition, there is no evidence to suggest that the Broken Top and South Sister magmatic systems interacted to generate intermediate composition rocks, supporting the argument that South Sister is probably younger that the Broken Top system (≤ 100 ka vs. $\approx 100-200$ ka).

8.3 Broken Top Intermediate System

The mineralogy of the Broken Top system intermediate series is the same as observed in the Three Sisters system: plag + cpx + opx + mgt \pm ol from $\approx 56\%$ to $\approx 63\%$ SiO₂, with no ol and subequal amounts of cpx and opx from $\approx 63\%$ to 72% SiO₂. Apatite is a groundmass phase for SiO₂ $\geq 62\%$, and amphibole is only observed as a rare trace phase in two rhyodacites. Unlike the Three Sisters system, there are no textural or mineralogical distinctions among similar composition rocks from different vents within the Broken Top system, i.e., rocks from the TVC show the same phases and textures as rocks from Tam MacArthur Rim or Broken Top vents.

Major elements usually vary linearly with increasing SiO₂ for Broken Top system basaltic andesites through rhyodacites (cf., figure 6.3), with the exceptions of Al₂O₃, Na₂O, and P₂O₅. These elements show relatively large degrees of scatter to 65% SiO₂, but more linear distributions from 65% to 72% SiO₂. The small degree of scatter observed for most elements above 65% SiO₂ could possibly represent a liquid line of descent along a well-defined cotectic, or a mixing trend between rhyolite and less silicic magmas. Representative analyses of Broken Top intermediate series rocks are given in table 8.3.

The fractionation hypothesis can first be tested through simple mass balance models (table 8.4). Differentiation from 63% to 65% SiO₂ always results in increasing Na₂O when the fractionated plagioclase is at least as calcic as andesine. A fractionating assemblage of 54% An₆₀ plag, 20% cpx, 14% opx, 9% mgt, 1.5% ilmn and 1.5% ap (model #1) duplicates the composition of those Broken Top dacites that have unusually high Na₂O (>5.5 wt.%). High Na₂O produced through crystal fractionation is also a characteristic of Middle Sister dacites (figure 8.6).

In contrast, differentiation for most of the Broken Top intermediate series (lower Na₂O) cannot be modeled through crystal fractionation. Above 65% SiO₂, fractionation models must account for decreasing Na₂O with increasing SiO₂, which necessitates \geq 75% An₂₅ plagioclase in the crystallizing assemblage (models 2-4, table 8.4). Phenocrysts more sodic than \approx An₃₀ are rarely found in Broken Top rhyodacites, and it is highly unlikely that relatively hydrous (3-4% H₂O; chapter 5) dacitic to rhyodacitic magmas would have crystallized oligoclase rather than andesine. Even when \geq 75% oligoclase is used in mass balance models, fractionation of opx ± hb ± cpx cannot produce the observed variations in CaO, MnO and MgO (table 8.4) above 65% SiO₂.

Assimilation of rhyolitic crust (i.e. mixing with rhyolitic magma) paired with crystal fractionation (cf., DePaolo, 1981) is necessary to produce most of the Broken Top intermediate series. Models were constructed (table 8.5) using Broken Top rhyolite 3S032 as the assimilant, which has higher SiO₂ and K₂O than all other Broken Top intermediate series rocks. Although crystal fractionation of basaltic andesite can produce Broken Top low silica ($\approx 63\%$ SiO₂) dacites (table 8.4), smaller proportions of rhyolite assimilation (R = mass assimilated / mass fractionated = 0.1 to 0.2) with plag + ol + cpx + opx + mgt + ilmn + ap fractionation can also produce many Broken Top series low-silica dacites (table 8.5). However, the low silica dacites did not evolve to high Na₂O dacite compositions with any significant amounts of rhyolite mixing, because mixing these compositions would result in lower Na₂O abundances.

The main Broken Top intermediate series above 63% SiO₂ is characterized by decreasing Na₂O with increasing SiO₂, which cannot be produced through realistic crystal fractionation models (table 8.4). However, assimilation of rhyolitic magma at R = 0.3 paired with fractionation can reproduce the main Broken Top intermediate trend to rhyodacitic compositions (table 8.5, model #2). A second, higher Na₂O trend ($\geq 65\%$ SiO₂) can also be generated from the high-Na₂O dacites

| | | <u> </u> | Na | Main Series | | | |
|--------------|----------------|----------|-------|-------------|-------|--|--|
| Sample | 3 S10 7 | 35100 | 35098 | 35036 | 35037 | | |
| SiO2 | 63.0 | 64.9 | 71.1 | 64.9 | 69.6 | | |
| TiO2 | 1.10 | 0.97 | 0.40 | 1.02 | 0.54 | | |
| A1203 | 16.2 | 16.3 | 14.9 | 16.0 | 15.4 | | |
| FeO* | 5.84 | 4.85 | 2.74 | 4.94 | 3.31 | | |
| MnO | 0.13 | 0.14 | 0.08 | 0.11 | 0.07 | | |
| Mao | 1 84 | 1 22 | 0.00 | 1 33 | 0.36 | | |
| CaO | 1 23 | 3 26 | 1 40 | 2 / 3 | 1 64 | | |
| Nalo | 4.2J | 5.20 | 6 12 | 5.45 | 1.04 | | |
| Nazu | 3.45 | 1 00 | 0.13 | 2.24 | 2.30 | | |
| R20 | 1.0/ | 1.99 | 2.11 | 2.23 | 2.93 | | |
| P205 | 0.33 | 0.30 | 0.07 | 0.33 | 0.11 | | |
| Rb | 37 | 39 | 58 | 41 | 59 | | |
| Ba | 558 | 548 | 721 | 649 | 736 | | |
| Sr | 375 | 300 | 135 | 363 | 181 | | |
| 51 | 1 10 | 1 20 | 2 04 | 0 50 | 0 00 | | |
| CB | 1.10 | 1.29 | 2.04 | 0.50 | 0.90 | | |
| Sc | 16.27 | 14.07 | 9.25 | 13.60 | 8.73 | | |
| v | 98 | 24 | 1 | 58 | 15 | | |
| Co | 9.63 | 3.84 | 1.59 | 7.14 | 2.50 | | |
| Ni | 4 | 6 | 10 | 9 | 10 | | |
| Cr | 8 | 4 | 2 | <5 | <5 | | |
| Zn | 135 | 207 | 135 | 80 | 54 | | |
| 011 | 100 | 207 | 100 | 00 | 54 | | |
| La | 21.8 | 24.6 | 24.4 | 24.1 | 22.2 | | |
| Ce | 51.5 | 62.6 | 61.3 | 50.4 | 53.7 | | |
| Nd | 28.6 | 34.0 | 30.8 | 27.9 | 24.0 | | |
| Sm | 6.50 | 7.84 | 6.46 | 6.59 | 5.99 | | |
| Eu | 1.58 | 2.04 | 1.28 | 1.65 | 1.28 | | |
| Th | 1.00 | 1.45 | 1.28 | 0.95 | 0.84 | | |
| Yb | 3 6 | 4 8 | 4.2 | 3 4 | 3.5 | | |
| Tu | 0 51 | 0 77 | 0 66 | 0 51 | 0 50 | | |
| Lu La /Sm | 2 12 | 1 00 | 2 20 | 2 21 | 2.24 | | |
| | 2.12 | 1.70 | 2.30 | 2.JL | 2.34 | | |
| 2HREE | 54.21 | //.55 | 07.07 | 52.10 | 50.02 | | |
| Eu/Eu* | 0.74 | 0.75 | 0.56 | 0.77 | 0.66 | | |
| Zr | 222 | 272 | 274 | 250 | 308 | | |
| Hf | 5.8 | 7.7 | 8.3 | 6.3 | 8.4 | | |
| Ga | 21 | 19 | 20 | 18 | 19 | | |
| Y | 34 | 42 | 27 | 34 | 30 | | |
| Nb | 14 4 | 20 0 | 15 5 | 15 7 | 16.8 | | |
| Ta | 0 71 | 1 03 | 1 90 | 0 80 | 1 00 | | |
| Ta | 2 2 | 1.02 | 5.50 | 4.0 | 5.4 | | |
| 111 | 3.3 | 3.0 | 5.5 | 4.0 | 2.4 | | |
| 0 | 1.5 | 1.3 | 6.6 | T.0 | 2.2 | | |
| | | | | | | | |

3S107: Representative High Cascade low silica dacite, Broken Top. 3S100: High Na₂O dacite, S. Fork Squaw Ck., 6600'. 3S098: High Na₂O rhyodacite, older dome at Squaw Ck. Falls. 3S036: Main series dacite, east Tam MacArthur Rim. 3S037: Main series rhyodacite, east Tam MacArthur Rim.

Table 8.3: Representative Broken Top system intermediate series rocks, with examples of high Na₂O dacite and rhyodacite. Differentiation to high Na₂O is a feature of crystal fractionation, as assimilation of relatively low Na₂O (≈5%) rhyolite will lower Na₂O in the evolving magmas. Note that evolution to high Na₂O dacite results in increasing HREE's, but that main series differentiation is characterized by relatively constant HREE abundances.



Figure 8.6: Variations in Na₂O and SiO₂ for High Cascades rocks in this study. Crystal fractionation with plagioclase as sodic An₃₀ will produce a Na₂O enrichment trend; significant amounts of rhyolite mixing ($\approx 20-30\%$) and fractionation of sodic plagiocase (An₂₅₋₃₀) is necessary to buffer or deplete Na₂O in the intermediate series. Note that several Broken Top dacites have the same elemental abundances as Middle Sister dacites, and that the South Sister and Broken Top series follow different mixing paths due to the different rhyolite compositions in each system.

| | | SiO2 | TiO2 | A1203 | FeO* | MnO | MgO | CaO | Na2O | K20 | P205 | |
|--------------------------------|-------|--------------------|--------------------|-----------|---------|---------|--------|--------|------------------|------|------|--|
| Cpx E | Sn37 | 51.2 | 0.4 | 1.4 | 13.5 | 0.8 | 12.4 | 18.6 | 0.3 | | | |
| Plag A | An60 | 52.7 | | 29.1 | 0.9 | | 0.1 | 12.1 | 4.5 | 0.3 | | |
| Plag A | An35 | 58.7 | | 25.6 | 0.5 | | 0.0 | 7.2 | 7.0 | 0.4 | | |
| Plag A | an20 | 62.5 | | 22.3 | 0.2 | | 0.0 | 4.1 | 8.5 | 0.8 | | |
| Ony F | 3n68 | 53.9 | 0.3 | 1.0 | 18.7 | 0.5 | 24.7 | 1.4 | 0.0 | | | |
| Mat | 11100 | 0 1 | 14 5 | 2.2 | 78 1 | 0.7 | 1 4 | | 0.0 | | | |
| Timn | | 0.1 | 12.2 | 2.2 | 51 5 | 0.6 | 2 5 | | | | | |
| Apt | | 0.0 | 76.6 | | 51.5 | 0.0 | 2.5 | 56.0 | | | 43.0 | |
| upc upc | | 42 1 | 3 8 | 10 4 | 14 7 | 03 | 12 9 | 10.7 | 2.3 | 0.3 | 40.0 | |
| 116 | | 14.14 | 5.0 | 10.4 | ± 1 + 1 | 0.5 | 11.17 | 10.1 | 2.0 | 0.0 | | |
| Model #1: Starting composition | | | | | | | | | | | | |
| | | SiO2 | TiO2 | A1203 | FeO* | MnO | Mao | CaO | Na20 | K20 | P205 | |
| 35107 | | 63 0 | 1 10 | 16 2 | 5.84 | 0.13 | 1.84 | 4.23 | 5.45 | 1.87 | 0.33 | |
| 55107 | | 03.0 | 1.10 | 10.2 | 3.04 | 0.10 | 1.04 | 1125 | 0.10 | 1.07 | 0.00 | |
| | 5 | 4% An. | . 20% | Cpx, | 14% 0 | px, 9% | Mqt, | 1.5% 1 | lmn, | 1.5% | Ap | |
| | | 0 | j , | - · | · · · | | | | ŕ | | | |
| F=0.88 | В | 65.1 | 0.97 | 16.1 | 4.85 | 0.12 | 1.28 | 3.37 | 5.82 | 2.09 | 0.29 | |
| 3S100 | | 64.9 | 0.97 | 16.3 | 4.85 | 0.14 | 1.22 | 3.26 | 6.22 | 1.99 | 0.30 | |
| | | | | | | | | | | | | |
| | | | Mod | iels #: | 2-#4: | Starti | ng con | posit: | ion | | | |
| | | SiO2 | TiO2 | A1203 | FeO* | MnO | MgO | CaO | Na20 | K20 | P205 | |
| 3S100 | | 64.9 | 0.97 | 16.3 | 4.85 | 0.14 | 1.22 | 3.26 | 6.22 | 1.99 | 0.30 | |
| | | | | | | | | | | - | | |
| | | 76 | * An ₂₅ | 5, 14% | Hb, 7 | .3% Mg1 | t, 18 | Ilmn, | 1.78 | Ар | | |
| F= 0.5 | 56 | 71.8 | 0.39 | 14.6 | 2.80 | 0.16 | 0.81 | 1.59 | 6.10 | 2.79 | 0.06 | |
| 35098 | | 71.1 | 0.40 | 14.9 | 2.74 | 0.08 | 0.31 | 1.40 | 6.13 | 2.77 | 0.07 | |
| | | | | | | | | | | | | |
| | | | | M | odel # | 3: 351 | 00 wit | :h | | | | |
| | | 76 | 58 An ₂ | 5,78 | Hb, 7. | 3% Mgt | , 18 : | Ilmn, | 1.7% | Ар | | |
| F - 0 1 | 56 | 71 / | 0 53 | 15 0 | 2 40 | 0 13 | 0 53 | 1 97 | 6 19 | 2 80 | 0 06 | |
| 20000 | 50 | 71 1 | 0.33 | 1/ 0 | 2.40 | 0.13 | 0.33 | 1 40 | 6 13 | 2.00 | 0.00 | |
| 33090 | | / 1 • 1 | 0.40 | 14+2 | 2.14 | 0.00 | 0.51 | 1.40 | 0.13 | 2.11 | 0.07 | |
| | | | | М | odel # | 4: 351 | 00 wit | :h | | | | |
| | | 779 | Anar | , 13% | Opx, 5 | 5.3% Ma | t, 3% | Ilmn, | 1.7% | Ap | | |
| | | | 0 | | | | | · | | - | | |
| F = 0.! | 58 | 70.6 | 0.40 | 15.3 | 2.58 | 0.11 | 0.39 | 2.39 | 6.23 | 2.77 | 0.07 | |
| 35098 | | 71.1 | 0.40 | 14.9 | 2.74 | 0.08 | 0.31 | 1.40 | 6.13 | 2.77 | 0.07 | |
| | | | | | | | | | | | | |
| | | | Me | odel # | 5, mai | n tren | d: 351 | 107 wi | th | | | |
| | 7 | 1% An ₃ | so ^{, 12} | ξ Cpx, | 98 OI | px, 3.5 | 8 Mgt | , 3% I | lmn, | 1.5% | Ар | |
| | | SiO2 | Ti02 | A1203 | FeO* | MnO | ΜαΩ | CaO | Na ₂₀ | K20 | P205 | |
| $\mathbf{F} = 0$ | | 60 7 | 0 52 | 15 / | 4 22 | | | 2 06 | E 24 | 2 06 | 0.00 | |
| | 44 | 07.4 | U. 7 7 | 1.21 - 44 | 4 . 1/ | 0.09 | 0.32 | Z.Un | - 3 4 | 2.90 | 0.09 | |
| 35037 | 44 | 69.6 | 0.53 | 15.4 | 4.32 | 0.09 | 0.32 | 2.00 | 5.38 | 2.90 | 0.09 | |

Table 8.4: Simple mass balance models for Broken Top dacite to rhyodacite differentiaion. Model #1 shows that crystal fractionation from 63% to 65% SiO₂ produces a high Na₂O trend. Continued fractionation above 65% SiO₂ fails to produce the Broken Top main intermediate series compositions (Models 2-4); note that An₂₅ plag is required to buffer Na₂O, and the poor correspondence between observed and modeled CaO, MnO and MgO. Anhydrous assemblages (model #4) produce the best fit for most elements, but fail to account for CaO abundances. Using a lower Na₂O parent still requires An₃₀ plag (model #5) to buffer Na₂O in the main trend; plag phenocrysts <An₃₀ are rarely found in rhyodacites. Magma differentiation above 65% SiO₂ in the Broken Top system was not controlled by crystal fractionation.

| | | SiO 2 | TiO2 | A1203 | FeO* | MnO | MgO | CaO | Na2O | K20 | P205 |
|------|------|--------------|------|-------|------|------|------|------|------|------|------|
| 01v | Fo70 | 36.8 | | | 26.4 | 0.5 | 35.0 | 0.2 | | | |
| Срх | En39 | 52.9 | 0.8 | 3.1 | 13.1 | 0.4 | 12.7 | 16.8 | 0.8 | | |
| Plg | An60 | 52.7 | | 29.1 | 0.9 | | 0.1 | 12.1 | 4.5 | 0.3 | |
| P1g | An40 | 59.2 | | 26.1 | 0.4 | | 0.0 | 8.1 | 6.2 | 0.5 | |
| Plg | An35 | 58.7 | | 25.6 | 0.5 | | 0.0 | 7.2 | 7.0 | 0.4 | |
| P1g | An20 | 62.5 | | 22.3 | 0.2 | | 0.0 | 4.1 | 8.5 | 0.8 | |
| Opx1 | En68 | 53.9 | 0.3 | 1.0 | 18.7 | 0.5 | 24.7 | 1.4 | 0.0 | | |
| Opx2 | En55 | 52.4 | 0.1 | 0.4 | 24.8 | 1.1 | 20.2 | 0.9 | 0.1 | | |
| Mgt | | 0.1 | 14.5 | 2.2 | 78.1 | 0.7 | 1.4 | | | | |
| Ilmn | | 0.0 | 42.2 | | 51.5 | 0.6 | 2.5 | | | | |
| Apt | | | | | | | | 56.0 | | | 43.0 |
| Hb | | 42.1 | 3.8 | 10.4 | 14.7 | 0.3 | 12.9 | 10.7 | 2.3 | 0.3 | |
| 3503 | 2 | 74.5 | 0.15 | 13.7 | 1.70 | 0.06 | 0.01 | 0.71 | 5.26 | 3.48 | 0.02 |
| | - | | | | | | | | | | |

Starting composition

 SiO2 TiO2 Al203 FeO*
 MnO
 MgO
 CaO
 Na2O
 K2O
 P2O5

 3S093
 57.5
 1.25
 17.4
 7.60
 0.14
 3.37
 6.73
 4.63
 1.17
 0.27

AFC Model #1 R = 0.1

65% Ans7 Plag, 5% Ol, 15% Cpx, 8% Opx1, 5.6% Mgt, 1% Ilmn, 0.4% Ap

| 35107 | 63.0 1.10 | 16.2 | 5.84 | 0.13 | 1.84 | 4.23 | 5.45 | 1.87 | 0.33 |
|----------|-----------|------|------|------|------|------|------|------|------|
| F = 0.54 | 63.4 1.08 | 16.1 | 5.79 | 0.13 | 1.82 | 4.29 | 5.49 | 1.85 | 0.31 |

AFC Model #2 main series, R = 0.3 69% An₃₀ Plag, 14% Cpx, 9% Opx2, 3.5% Mgt, 3% Ilmn, 1.5% Ap

Starting Composition

CaO Na20 K20 SiO2 TiO2 A1203 FeO* MnO MgO P205 5.84 1.84 4.23 5.45 1.87 0.33 3S107 63.0 1.10 16.2 0.13 F = 0.50 70.6 0.48 15.43.40 0.05 0.39 1.76 5.40 3.03 0.09 35037 69.6 0.54 15.4 3.31 0.07 0.36 1.64 5.38 2.93 0.11

AFC Model #3, R = 0.35 75% An₂₅ Plag, 8% Cpx, 8% Opx2, 4.5% Mgt, 3% Ilmn, 1.5% Ap

Starting composition SiO2 TiO2 A1203 FeO* MnO MqO CaO Na2O K2O P205 0.97 16.3 4.85 0.14 1.22 3.26 6.22 1.99 0.30 3S100 64.9 F = 0.62 71.02.84 0.10 0.37 1.58 6.16 2.86 0.09 0.42 15.2 71.1 0.40 14.9 2.74 0.08 0.31 1.40 6.13 2.77 0.07 35098

Table 8.5: Mass balance models of assimilation with crystal fractionation for Broken Top intermediate series magmatism. Assimilation of rhyolite 3S032 was modeled by addition of R * F (Ratio of mass assimilated/mass fractionated * 2% Fractionation increment) of 3S032 to each fractionation increment. Model #1 shows that small amounts (R = 0.1-0.2) of crustal assimilation paired with fractionation of basaltic andesite can produce Broken Top low silicia dacites. Main series rhyodacites (Model #2) can be produced through AFC of low silica dacites, with assimilation rates of 0.3. Model #3 shows that AFC involving higher assimilation rates and fractionation of more sodic plagioclase may generate high Na₂O rhyodacites from high Na₂O dacites. through slightly higher assimilation proportions (R = 0.35) and increased fractionation of more sodic plagioclase (table 8.5, model #3).

Compatible element variations (e.g., Sc and V, figure 8.7) in the main Broken Top intermediate series also do not support a simple fractionation hypothesis. Using the arguments developed for the Middle Sister series, incompatible element abundances in the main BT intermediate series are too high (relative to compatible element depletions) to have originated through simple crystal fractionation of the mass-balance constrained mineral assemblage. The observed variations are also not consistent with the effects of recharge by andesitic melts, which will maintain essentially constant compatible element abundances but increase incompatible element abundances. The compatible element variations are consistent, however, with a model similar to the one developed for SS, which requires 20-30% of silicic magma mixing to account for the observed variations. The amount of mixing should be less than at South Sister, as much of the main BT series is intermediate between the MS (fractionation dominated) and SS (mixing dominated) fields (figure 8.7). Broken Top units that plot closest to the MS fields are part of the high Na₂O trend, which was apparently produced through crystal fractionation.

The distribution of Zr versus Rb also clearly shows that mixing of rhyolitic magma is required to account for the observed variations (figure 8.8). It has been shown that Zr remained an essentially incompatible element during intermediate series fractionation, due to the elevated temperatures (\geq 930°C) and low concentrations of Zr (\leq 350 ppm) that prevented zircon saturation in the BT melts. CHAOS models show that neither simple fractionation (0-0-0) nor recharge (2-0-1) can account for the observed Zr-Rb distributions, but that a model involving \approx 20% mixing of rhyolite (2-0.2-1.2) reproduces the observed scatter and relative increases in Zr and Rb to 72% SiO₂ (figure 8.8). Note that mixing of >20% rhyolite with the differentiating basaltic andesite system would probably result in lower Zr/Rb than observed in the Broken Top system.

Variations in REE's also show that rhyolite mixing is required to produce the BT intermediate series. The REE's form essentially sub-parallel patterns, increase in overall abundance, and develop small negative Eu anomalies from basaltic andesite to $\approx 63\%$ SiO₂ (figure 8.9a), which is characteristic of plagioclase-pyroxene fractionation with small (R ≈ 0.1) rhyolite assimilation rates. However, two distinct REE trends develop above 63% SiO₂ in the Broken Top system. In the high Na₂O trend, the REE's continue to enrich until 65% SiO₂ (figure 8.9a), which is consistent with plagioclase-pyroxene fractionation and results in high Na₂O dacites. Differentiation in the high Na₂O trend from 65% to 72% SiO₂ results in small MREE and HREE depletions, which was produced by apatite fractionation (model 3, table 8.5) paired with magma mixing (figure 8.9a). In contrast, the main trend has essentially constant REE abundances from 63% to $\approx 65\%$ SiO₂, and only minor LREE and HREE enrichments developed by $\approx 70\%$ SiO₂ (figure 8.9b).



Figure 8.7: Compatible-incompatible element trends for Broken Top and Three Sisters rocks. Variations in Sc (A) and V (B) indicate that the main BT intermediate series contains components of fractionation (i.e., Middle Sister) and rhyolite mixing (i.e., South Sister). Observed compatible element depletions are consistent with a model involving fractionation paired with ≈20-30% rhyolite mixing in an open basaltic andesite system. Note that intermediate samples from the TVC are indistinguishable from BT-TMR samples. Samples with ≤40 ppm Rb that plot close to the MS field are part of the high Na₂O series.

Δ


Figure 8.8: Results of CHAOS models showing variations in Zr vs. Rb. A) Closed system fractionation (0-0-0) or recharge (2-0-1) models cannot account for the observed fractionation between Zr and Rb, as zircon is not a fractionating phase for these compositions. Open system (2-0.2-1.2) models using rhyolite 3S032 successfully model Zr vs. Rb variations. B) CHAOS trends superimposed on BT system data. Note that samples with significant disequilibrium mineral textures do not form distinct groups.



La Ce Pr Nd Sm Eu Gd Tb Dy Ho Er Tm Yb Lu

Figure 8.9: Rare earth element diagrams for Broken Top system petrogenesis. A) Broken Top intermediate series differentiation to 63% SiO₂ is characterized by increasing REE content with little REE fractionation, and development of small negative Eu anomalies. Above 63% SiO₂, the high Na₂O series is characterized by increasing REE to 65% SiO₂, which is thought to represent differentiation without rhyolite mixing. Depletions in the HREE and increases in La/Sm above 65% SiO₂ are the product of $\approx 35\%$ rhyolite mixing paired with apatite fractionation. The main series (B) is characterized by constant MREE to $\approx 70\%$ SiO₂, with small enrichments in LREE and HREE. This feature indicates small amounts of apatite fractionation paired with $\approx 30\%$ rhyolite mixing.

B

The increase in La/Sm and depletions in the HREE's and MREE's in the high Na₂O trend might be a feature of apatite fractionation. Large amounts (10-20%) of amphibole fractionation could produce similar REE features (cf. Grove and Donnelly-Nolan, 1986; Bacon and Druitt, 1988), but none of the BT intermediate series rocks contain more than a rare trace of amphibole. Augite REE K_d's are only \approx 0.1-0.5 (Nielsen et al., in press), and 20% augite fractionation will not significantly affect the REE's. Apatite distribution coefficients for the REE's in silicic compositions are in the range of 10 for La and 20 to 30 for the MREE's (Watson and Green, 1981; Fujimaki, 1986; Appendix E), with the MREE's having the highest K_d's. The observed depletions in the REE's from 65% to 71% SiO₂ would require \approx 30% fractionation (determined by Rb and Th enrichment) with bulk distribution coefficients of D_{La}=1, D_{Sm}=1.5, D_{Eu}=2.3, D_{Tb}=1.4, and D_{Yb}=1.4, and thus 5-10% apatite in the fractionating assemblage. Mass balance constraints limit apatite to \approx 1.5% of the fractionating assemblage, which can not produce the observed REE patterns unless apatite distribution coefficients are much larger than reported. Magma evolution above 65% SiO₂ in the high Na₂O requires significant amounts of mixing with rhyolitic magmas (R = 0.35, table 8.5) in order to generate the observed REE patterns.

In contrast to the first trend, the increases in La/Sm and buffering of the MREE's in the main trend could have been produced by relatively lesser amounts of magma mixing along with apatite fractionation. The observed depletions in the REE's from 63% to 70% SiO₂ would require $\approx 40\%$ fractionation with bulk distribution coefficients of D_{La}=0.7, D_{Sm}=1.1, D_{Eu}=1.3, D_{Tb}=1.0, and D_{Yb}=0.8, and thus $\approx 5-3\%$ apatite in the fractionating assemblage; note that although the interval of fractionation has increased relative to the first trend, the bulk distribution coefficients have decreased. Mass balance calculations constrain apatite to $\approx 1.5\%$, and the fractionation of apatite \pm pyroxene (+ plag) is incapable of producing the REE variations observed in the second series. Mixing of rhyolite magma at R = 0.3 paired with apatite fractionation (cf., table 8.5) is capable of buffering the REE's in the main intermediate series. CHAOS (2-0.2-1.2) models involving 20% mixing of rhyolite 3S032 into an open basaltic andesite system (figure 8.10) confirm that the mixing of rhyolite with a parental basaltic andesite magma system, and subsequent fractionation of pyroxene \pm apatite, is capable of producing the observed REE features of the main BT intermediate series to 72% SiO₂.

Intermediate series dacites and rhyodacites with elevated REE's and Na₂O are not temporally or spatially distinct from intermediate rocks of the main trend. Rocks from both trends have the same range in mineralogy, texture, and degree of crystallinity. Unlike the Three Sisters system, the Broken Top system does not show any spatial or temporal constraints on intermediate series petrogenesis. There are, however, only two high Na₂O dacites in the sample set (3S100 and 3S008) for comparison. Both of these samples lack disequilibrium mineral features (Appendix B), which supports a crystal fractionation hypothesis.



Figure 8.10: Results of a CHAOS 2-0.2-1.2 model involving periodic mixing of 20% rhyolite 3S032. Fractionation of apatite associated with mixing generally buffers the intermediate series HREE's at rhyolite abundances. High HREE, high Na₂O dacites and rhyodacites are the result of crystal fractionation without rhyolite mixing to 65% SiO₂, and higher rates of rhyolite mixing and fractionation above 65% SiO₂. Although all the rhyolites are associated with the older Tumalo volcanic center (TVC), younger Broken Top & Tam MacArthur Rim magmatism involved mixing with this type of rhyolite.

8.4 Summary of Intermediate Series Petrogenesis

In summary, magma evolution to 63% SiO₂ in the Broken Top series was characterized by open system crystal fractionation with small amounts ($R \approx 0.1$) of rhyolite mixing. Above 63%SiO₂, two chemically distinct trends developed in the Broken Top system. Magma evolution from 63% to 72% SiO₂ in the main Broken Top trend was characterized by crystal fractionation paired with $\approx 30\%$ rhyolite mixing, which maintained fairly constant levels of Na₂O and the REE's during differentiation. A second, high Na₂O and high HREE trend was produced through crystal fractionation from 63% to 65% SiO₂, and higher rates of rhyolite mixing (R = 0.35) paired with fractionation from 65% to 72% SiO₂. Unlike the Three Sisters intermediate system, rocks from both Broken Top intermediate trends are spatially, temporally and mineralogically identical.

It is important to note that crystal fractionation as a dominant evolutionary process is limited in occurrence to ≤65% SiO₂ in both the Middle Sister and Broken Top systems. Magma evolution above 65% SiO₂ in the Middle Sister system probably did not occur because rhyolitic magmas were not available for mixing into the fractionated system. In contrast, intermediate series differentiation probably continued above 65% SiO2 in the Broken Top system because rhyolitic magmas were available for mixing into an evolved, intermediate composition magma system. Arguments developed for South Sister petrogenesis showed that rhyolite is probably limited to $\leq 50\%$ of any magma mixture, or else the mafic component would be thermally quenched (i.e., Frost and Mahood, 1987). South Sister intermediate compositions were thus limited to ≤65% SiO₂, which represents the limit of 50% rhyolite - 50% basaltic andesite mixing. In the Broken Top system, fractionation continued beyond basaltic andesite (58% SiO₂) to 65% SiO₂, thus increasing the SiO₂ content of potentially mixing system 7 wt.% SiO₂ above the South Sister system. The fact that the Broken Top system evolved through mixing + fractionation to 72% SiO₂ but the same process in the South Sister system was limited to 65% SiO2 is probably a function of the degree of evolution of the mafic mixing component. Because the Broken Top "mafic" mixing component had evolved up to ≈65% SiO₂ instead of ≈58% SiO₂ (i.e., +7 wt% higher), 50:50 mixtures with rhyolite could produce rhyodacitic (i.e., +7 wt% SiO₂ higher) instead of dacitic melts.

9. PETROGENESIS OF RHYOLITES

9.1 Three Sisters Rhyolites

Rhyolites associated with the Three Sisters are texturally and compositionally heterogeneous. Holocene rhyolites show relatively small degrees of geochemical variation (table 9.1), which is thought to have been produced through "small" amounts of crystal fractionation (Clark, 1983; Scott, 1987). Pleistocene low-silica ($\approx 72\%$ SiO₂) rhyolites are similar to the Holocene rhyolites, but range to more evolved high-silica (76% SiO₂) compositions (table 9.1). There are also subtle trace element distinctions between Holocene and Pleistocene rhyolites. The Holocene rhyolites tend to plot in more closely spaced, distinct groups, and have relatively higher incompatible element abundances (figure 9.1).

Although both rhyolite types have the same range in mineral chemistry and degrees of crystallinity, plagioclase phenocrysts from each type are texturally distinct. Both rhyolite types contain $\leq 20\%$ plagioclase, which has oscillatory zonation from $\approx An_{40}$ to $\approx An_{30}$. However, most plagioclase phenocrysts in all of the Holocene rhyolites contain 10-40% clear-to-light-brown coarse glass inclusions (figure 4.4C-D), but plagioclase from Pleistocene rhyolites rarely contains glass inclusions. This sieved texture in the Holocene rhyolites cannot be simply the product of magma mixing; assuming that the Holocene rhyolites were originally a Pleistocene rhyolite composition, the Holocene geochemical trends point **away** from both mafic and more silicic mixing components (figure 9.1). It is also not clear how magma mixing could have produced texturally distinct but compositionally identical phenocrysts. A rapid decrease in pressure might produce sieved plagio-clase (e.g., Nelson and Montana, 1989), but the Holocene plagioclase should be more sodic (Blundy and Shimizu, 1991) if it were produced through decompression of a hydrous melt. The plagioclase textures and geochemical variations indicate that the Holocene rhyolites were produced from a different magma chamber than the Pleistocene rhyolites.

Geochemical variations within both rhyolite suites can be explained through relatively small amounts of crystal fractionation. Although the range in geochemical variation and crystallinity is small, there is a rough correlation between variations in CaO, Al_2O_3 and FeO and degree of crystallinity for Three Sisters rhyolites (figure 9.2a). Simple mass balance models (table 9.2) show that variations between Pleistocene low-silica and high-silica rhyolite can be accounted for by the fractionation of 5-18 wt% of the observed minerals (plag + opx + mgt + ilmn + ap ± hb) in modal proportions. The smaller degrees of geochemical variation observed in the Holocene rhyolites also can be accounted for with 5% modal fractionation (cf., Clark, 1983; Scott, 1987). REE variations are consistent with these mass balance models. Small depletions in Eu and generally increasing REE abundances reflect plag + opx ± hb fractionation (figure 9.2b, table 9.2); small HREE and Zr depletions in high-silica rhyolite 3S139 require the addition of 0.4% zircon to

| | Middle | Hol | ocene | Plei | stocene | |
|--------------|---------------|----------|----------|----------|------------|-------|
| | Sister | Low | High | Low | High | |
| Sampl | e 3s134 | 35145 | 35060 | 35148 | 3\$139 | |
| SiO2 | 65.7 | 71.8 | 73.7 | 72.5 | 76.1 | |
| TiO2 | 0.73 | 0.35 | 0.29 | 0.33 | 0.14 | |
| A1203 | 16.2 | 14.9 | 14.1 | 14.5 | 13.4 | |
| FeO* | 5 03 | 2 20 | 1 75 | 2 06 | 1 16 | |
| MnO | 0.14 | 0 07 | 0.05 | 0.06 | 0.04 | |
| Mao | 0.14 | 0.53 | 0.03 | 0.00 | 0.11 | |
| MgO | 0.33 | 0.55 | 1 62 | 1.06 | 0.11 | |
| CaU Na 20 | 2.12 | 2.11 | 1.02 | 1.00 | 0.97 | |
| Nazu | 0.20 | 4.01 | 4.38 | 4.// | 4.40 | |
| K20 | 1.89 | 3.09 | 3.48 | 3.15 | 3.41 | |
| P205 | 0.25 | 0.09 | 0.07 | 0.07 | 0.04 | |
| Rb | 31 | 73 | 85 | 71 | 79 | |
| Ba | 607 | 755 | 809 | 742 | 876 | |
| Sr | 272 | 251 | 185 | 193 | 119 | |
| Cs | 0.85 | 1.93 | 2.20 | 2.02 | 1.94 | |
| Sc | 12.81 | 3,92 | 3.29 | 4,60 | 1.63 | |
| v | 12 | 29 | 20 | 31 | Q | |
| Ċo | 3 56 | 3 5 3 | 2 44 | 3 75 | 0 5 9 | |
| Ni | 19 | 16 | ~10 | 5.75 | ~10 | |
| C~ | 10 | 10 | ~10 | 9 | 10 | |
| | 200 | | 47 | 190 | C A | |
| 211 | 208 | 90 | 4./ | 180 | 04 | |
| La | 21.3 | 18.8 | 21.0 | 19.2 | 20.0 | |
| Ce | 53.8 | 43.1 | 38.4 | 38.6 | 44.6 | |
| Nd | 28.0 | 16.7 | 16.0 | 17.4 | 14.9 | |
| Sm | 6.75 | 2.97 | 3.07 | 3.06 | 2.63 | |
| Eu | 1.80 | 0.70 | 0.59 | 0.64 | 0.42 | |
| Tb | 1.27 | 0.54 | 0.41 | 0.41 | 0.39 | |
| Yb | 4.2 | 1.7 | 1.8 | 1.9 | 1.5 | |
| Lu | 0.62 | 0.27 | 0.31 | 0.29 | 0.24 | |
| La/Sn | 1.99 | 4.00 | 4.32 | 3.96 | 4.80 | |
| ΣHREE | 66.22 | 27.94 | 26.89 | 26.75 | 22.87 | |
| Eu/Eu | .* 0.77 | 0.68 | 0.60 | 0.65 | 0.49 | |
| 7 | 241 | 160 | 160 | 150 | 0.6 | |
| 41 | 241 | 100 | 100 | 100 | 2 2 | |
| | 7.0 | 4.0 | 4.0 | 4.3 | 3.2 | |
| Ga | 21 | 10 | 13 | 1/ | 15 | |
| I | 30 | 10 | 10 1 | 81 | 10 0 | |
| ND | 12.7 | 10.6 | 12.1 | 8.7 | 10.0 | |
| Ta | 0.62 | 0.68 | 0.79 | 0.67 | 0.74 | |
| Th | 2.9 | 7.2 | 8.2 | 7.0 | 7.3 | |
| U | 1.3 | 2.6 | 3.2 | 2.4 | 2.6 | |
| 35134: | Middle Sister | dacite, | Lane Me | sa. | | |
| 3S145: | Holocene low- | silica r | hyolite, | Carver | Lake dome | • |
| 35060: | Holocene high | -silica | rhyolite | , Rock M | esa. | |
| 35148: | Pleistocene 1 | ow-silic | a rhyoli | te, Cham | bers Lake | dome. |
| 35139: | Pleistocene h | igh-sili | ca rĥyol | ite, Obs | idian Cli | ffs. |

Table 9.1: Analyses of representative Three Sisters rhyolite types, showing range of geochemical variation observed in Holocene and Pleistocene rhyolites. Middle Sister dacite is shown for comparison. Note the extreme depletions in HREE and high La/Sm that characterize all Three Sisters rhyolites.



Figure 9.1: Compatible-incompatible element plots for the Three Sisters system. Holocene and Pleistocene rhyolites plot in distinct groups; note the consistently higher Sr in the Holocene rhyolites. Holocene rhyolites do not plot along mixing lines with observed compositions, and the geochemical differences are not the products of crystal fractionation. Fractionation of modal minerals can, however, account for intra-suite variations.



Figure 9.2: Comparisons of Holocene and Pleistocene Three Sisters rhyolites. A) Three Sisters rhyolites show a very rough correlation between CaO content and modal % phenocrysts in the rock, suggesting that fractionation or accumulation of the modal minerals may have controlled geochemical evolution. B) Chondrite-normalized REE patterns for the range of Holocene and Pleistocene rhyolite compositions. The most evolved rhyolites of each group are characterized by higher La/Sm and Eu/Eu*, and lower HREE. These features can be modeled with <1% apatite and <0.4% zircon fractionation, which is supported by mass balance calculations.</p>

| Miner | al | SiO2 | TiO2 | A1203 | FeO* | MnO | MgO | CaO | Na 20 | K20 | P205 |
|-------|------|------|------|-------|------|-----|------|------|-------|-----|------|
| Plag | An31 | 61.1 | | 24.7 | 0.3 | | 0.0 | 6.2 | 7.2 | 0.5 | |
| Орх | En58 | 52.4 | 0.1 | 0.4 | 24.8 | 1.1 | 20.2 | 0.9 | 0.1 | | |
| Mgt | | 0.1 | 9.6 | 1.8 | 82.7 | 0.5 | 1.4 | | | | |
| Ilmn | | 0.0 | 42.2 | | 51.5 | 0.6 | 2.5 | | | | |
| Apt | | | | | | | | 56.0 | | | 43.0 |
| Hb | | 42.1 | 3.8 | 10.4 | 14.7 | 0.3 | 12.9 | 10.7 | 2.3 | 0.3 | |

Pleistocene parental composition SiO2 TiO2 Al2O3 FeO* MnO MgO CaO Na2O K2O P2O5 3S148 72.5 0.33 14.5 2.06 0.06 0.61 1.86 4.77 3.15 0.07

Pleistocene 7% model: Fractionation of 6% Plag, 0.6% Opx, 0% Hb, 0.35% Mgt, 0.15% Ilmn, 0.03% Ap

| | SiO2 | TiO2 | A1203 | FeO* | MnO | MgO | CaO | Na20 | K2O | P205 |
|-----------|------|------|-------|------|------|------|------|------|------------|------|
| 7% model: | 74.7 | 0.23 | 14.0 | 1.58 | 0.05 | 0.21 | 1.30 | 4.58 | 3.20 | 0.05 |
| 35054 | 74.2 | 0.24 | 14.3 | 1.58 | 0.05 | 0.23 | 1.20 | 4.53 | 3.37 | 0.05 |

Pleistocene 18% model: Fractionation of 15% Plag, 2.2% Opx, 0.5% Hb, 0.35% Mgt, 0.35% Ilmn, 0.1% Ap

| | SiO2 | TiO2 | A1203 | FeO* | MnO | MgO | CaO | Na 20 | K20 | P205 |
|-----------|------|------|-------|------|------|------|------|-------|------|------|
| 18% Model | 72.6 | 0.32 | 14.7 | 2.08 | 0.07 | 0.61 | 1.86 | 4.73 | 2.86 | 0.07 |
| 35139 | 76.1 | 0.14 | 13.4 | 1.16 | 0.04 | 0.11 | 0.97 | 4.46 | 3.41 | 0.04 |

Holocene parental composition

| | SiO2 | TiO2 | A1203 | FeO* | MnO | MgO | CaO | Na 20 | K20 | P205 |
|-------|------|------|-------|------|------|------|------|-------|------|------|
| 3S145 | 71.8 | 0.35 | 14.9 | 2.20 | 0.07 | 0.53 | 2.11 | 4.61 | 3.09 | 0.09 |

Holocene 11% model: Fractionation of 9% Plag, 0.7% Opx, 0.5% Hb, 0.35% Mgt, 0.10% Ilmn, 0.06% Ap

| | SiO2 | TiO2 | A1203 | FeO* | MnO | MgO | CaO Na | 20 K20 | P205 |
|-----------|------|------|-------|------|------|------|----------|---------|------|
| 11% Model | 71.9 | 0.35 | 14.9 | 2.18 | 0.06 | 0.58 | 2.10 4. | 57 3.15 | 0.09 |
| 35060 | 73.7 | 0.29 | 14.1 | 1.75 | 0.05 | 0.41 | 1.62 4.3 | 38 3.48 | 0.07 |

Table 9.2: Crystal fractionation models for Three Sisters rhyolites. Models used mineral compositions from Three Sisters rhyolites 3S114 and 3S122. Removal of 7% to 18% of the observed modal minerals from the lowest silica rhyolite (3S148) reproduces the geochemical variations found in Three Sisters Pleistocene rhyolites. The range of composition in the Holocene rhyolites can be produced though the fractionation of 11% modal minerals from the least evolved Holocene rhyolite (3S145). These features strongly indicate that the members of each rhyolite type were erupted from a common magma chamber, which was compositionally zoned due to crystal fractionation. the fractionating assemblage. Trace element distributions also show that the Holocene rhyolites cannot be derived from the Pleistocene rhyolites through crystal fractionation (figure 9.1), and must have originated as a discrete batch of rhyolitic magma.

9.1.1 Crystal fractionation

Previous models have shown that differentiation to 65% SiO₂ at Middle Sister was primarily controlled by crystal fractionation, but the petrogenesis of andesites and dacites at South Sister was controlled by magma mixing. South Sister intermediate composition rocks could not have been parental to Three Sisters rhyolites, because the formation of these intermediate rocks required an independent source of rhyolitic magma. However, continued fractionation at Middle Sister may have resulted in the formation of Three Sisters rhyolite magmas and a 7% SiO₂ compositional gap (cf., Grove and Donnelly-Nolan, 1986).

Other High Cascade volcanos, which are also thought to represent fractionating magmatic systems, contain large (i.e., greater than the 7% SiO₂ gap in the Three Sisters) compositional gaps. The climactic eruption at Mt. Mazama contained a 9% SiO₂ compositional gap, which was produced through a combination of fractionation and accumulation of crystals in the underlying high-Sr andesitic magmas (Bacon, 1983; Druitt and Bacon, 1989). The Mt. Mazama rhyodacites can be successfully modeled by fractionation of the accumulated anhydrous minerals from andesitic compositions, with minor assimilation of granitic wall rocks (Bacon and Druitt, 1988). Rhyolites at Medicine Lake were apparently derived through extensive (50-60%) amphibole + plagioclase fractionation from andesite, with small amounts of wall-rock assimilation (Grove et al., 1982; Grove and Donnelly-Nolan, 1986). Due to the shallow slope of the amphibole liquidus, an 11% SiO₂ compositional gap developed at Medicine Lake with only small decreases in temperature of the system (Grove and Donnelly-Nolan, 1986). However, both Medicine Lake and Mt. Mazama are characterized by abundant intermediate composition rocks that are highly phyric (30-55% crystals) and contain significant amounts of amphibole (10-20%). Intermediate composition rocks in the Three Sisters generally contain $\leq 30\%$ phenocrysts and lack amphibole (Chapter 4), yet the Three Sisters system still contains a 7% SiO₂ compositional gap. It is possible that the Three Sisters rhyolites were produced through fractionation of Middle Sister dacites, but that the fractionating minerals remained on the walls of the magma chamber (cf. McBirney and Noyse, 1979) and were not erupted as cumulate intermediate rocks. This fractionation hypothesis can be tested through mass balance and trace element models.

Mass balance models indicate that Three Sisters rhyolites (table 9.1) cannot be derived from Middle Sister dacites through the fractionation of plag \pm opx \pm hb \pm cpx \pm mgt \pm ilmn \pm ap. Increases in K₂O necessitate >50% crystallization, as K₂O will remain essentially incompatible using the observed mineral compositions (table 9.3). The low abundance of Na₂O in the rhyolites

| Mineral | SiO2 | TiO2 | A1203 | FeO* | MnO | MgO | CaO | Na2O | K20 | P205 |
|----------|--------|------|-------|------|-----|------|------|------|-----|------|
| Plag An2 | 0 62.5 | | 22.3 | 0.2 | | 0.0 | 4.1 | 8.5 | 0.8 | |
| Opx En5 | 1 51.6 | 0.2 | 0.7 | 26.8 | 1.2 | 17.7 | 1.6 | 0.1 | | |
| Mgt | 0.1 | 14.5 | 2.2 | 78.1 | 0.7 | 1.4 | | | | |
| Ilmn | 0.0 | 44.4 | | 51.1 | 0.7 | 2.3 | | | | |
| Apt ≈ | | | | | | | 56.0 | | | 43.0 |
| Hb | 42.1 | 3.8 | 10.4 | 14.7 | 0.3 | 12.9 | 10.7 | 2.3 | 0.3 | |

Starting composition

| MS Dacite | SiO2 | TiO2 | A1203 | FeO* | MnO | MgO | CaO | Na2O | K20 | P205 |
|-----------|------|------|-------|------|------|-----|-----|------|-----|------|
| 35134 | 65.7 | 0.73 | 16.2 | 5.0 | 0.14 | 1.0 | 2.7 | 6.3 | 1.9 | 0.25 |

Best fit using anhydrous assemblage 83% An₂₀ Plag, 8% Opx, 8% Mgt, 1% Ap

| | SiO2 | TiO2 | A1203 | FeO* | MnO | MgO | CaO | Na2O | K20 | P205 |
|-----------|------|------|-------|------|------|-----|-----|------|-----|------|
| F = 0.40: | 73.9 | 0.35 | 14.1 | 2.1 | 0.13 | 0.5 | 1.6 | 5.6 | 2.9 | 0.09 |
| 35148 | 72.5 | 0.33 | 14.5 | 2.1 | 0.06 | 0.6 | 1.9 | 4.8 | 3.2 | 0.07 |

Model #2, Medicine Lake: After Grove and Donnelly-Nolan, 1986

| Mine | ral | SiO2 | TiO2 | A1203 | FeO* | MnO | MgO | CaO | Na20 | K20 | P205 |
|------|------|------|------|-------|------|-----|------|------|------|-----|------|
| Plag | An35 | 58.7 | | 25.6 | 0.5 | | 0.0 | 7.2 | 7.0 | 0.4 | |
| Opx | En51 | 51.6 | 0.2 | 0.7 | 26.8 | 1.2 | 17.7 | 1.6 | 0.1 | | |
| Mgt | | 0.1 | 14.5 | 2.2 | 78.1 | 0.7 | 1.4 | | | | |
| Apt | * | | | | | | | 56.0 | | | 43.0 |
| Hb | | 42.1 | 3.8 | 10.4 | 14.7 | 0.3 | 12.9 | 10.7 | 2.3 | 0.3 | |

Fractionation with above minerals, 66% An₃₅ Plag, 8% Opx, 17% Hb, 7% Mgt, 1% Ap

| | SiO2 | TiO2 | A1203 | FeO* | MnO | MgO | CaO | Na 20 | K20 | P205 |
|----------|------|------|-------|------|------|------|------|-------|-----|------|
| F = 0.50 | 75.9 | 0.07 | 14.4 | 1.2 | 0.10 | -0.9 | -0.4 | 7.1 | 2.9 | 0.04 |
| 35148 | 72.5 | 0.33 | 14.5 | 2.1 | 0.06 | 0.6 | 1.9 | 4.8 | 3.2 | 0.07 |

Model #3, Medicine Lake, same compositions as above: 69% An₃₅ Plag, 6% Opx, 15% Hb, 10% Mgt, 1% Ap

| | SiO2 | TiO2 | A1203 | FeO* | MnO | MgO | CaO | Na2O | K20 | P205 |
|----------|------|-------|-------|------|------|------|------|------|-----|------|
| F = 0.50 | 76.3 | -0.08 | 14.1 | 0.8 | 0.12 | -0.4 | -0.3 | 7.0 | 2.9 | 0.12 |
| 35148 | 72.5 | 0.33 | 14.5 | 2.1 | 0.06 | 0.6 | 1.9 | 4.8 | 3.2 | 0.07 |

Table 9.3: Simple mass balance calculations used to model petrogenesis of Three Sisters low-silica rhyolites through crystal fractionation of Middle Sister dacite. Calculation method explanation in table 7.1. Fractionation of an anhydrous assemblage (model 1) can only approximate 3S rhyolite compositions if unreasonably large amounts of anomalously sodic plagioclase are crystallized. Note that 60% crystallization with 83% An₂₀ plagioclase still results in anomalously high Na₂O in the modeled rhyolite; using An₃₀ plagioclase does not result in significant Na₂O depletions. Medicine Lake models use the fractionating assemblages of Grove and Donnelly-Nolan (1986), but fail to approximate 3S rhyolite compositions. Any model involving fractionation of >5% Hb results in CaO, MgO and Al₂O₃ depletions that are much greater than observed in 3S rhyolites. Hydrous and anhydrous mass balance models of Middle Sister dacites fail to reproduce Three Sisters rhyolite compositions, indicating that crystal fractionation is not a valid hypothesis for Three Sisters rhyolite petrogenesis. can not be modeled even with fractionating an assemblage containing >80% An_{20} plagioclase. Because plagioclase phenocrysts in 3S rhyolites are generally more calcic than An_{30} and 3S amphiboles contain $\leq 2\%$ Na₂O, it is not possible to deplete Na₂O to the observed levels by fractionation from Middle Sister dacites.

Amphibole fractionation is critical in the petrogenesis of rhyolitic magmas at Medicine Lake, and is thought to control the development of many intermediate-to-rhyolite compositional gaps (Grove and Donnelly-Nolan, 1986). Mass balance models involving >5% hornblende fractionation are anomalously depleted in MgO and CaO (Medicine Lake models, table 9.3), as are models containing >5% cpx. The inability to fit hornblende fractionation into the mass balance models supports the conclusion that amphibole was not a liquidus phase in the Three Sisters intermediate system, and fractionation of amphibole did not control the development of a compositional gap.

Apatite fractionation is critical to evaluating REE distributions, as apatite is the only mineral in these rocks that has high REE compatibility (Appendix E). Three Sisters rhyolites are characterized by low HREE ($\approx 10X$ chondrite), high La/Sm (4-5) and moderate Eu/Eu* (0.7-0.5) (figure 9.3a). These REE characteristics require a D_{La} = 1.1, D_{Ce} = 1.4, D_{Tb} = 2.4 and D_{Lu} = 2, with F \approx 0.4 (Rb & Th). Apatite is the only observed mineral that has high La and Ce distribution coefficients (≈ 10 -30) (Watson and Green, 1981; Appendix E); hornblende K_d's are <1 for La, and <2 for Ce. In situ fractionation of a LREE-silicate such as allanite or sphene (Gromet and Silver, 1983) would also result in significant Th depletions and LREE-HREE fractionation, which is not observed (table 9.1). In addition, mass balance calculations limit apatite to 1% of the fractionating assemblage. This amount of apatite fractionation is insufficient to deplete the LREE and HREE to observed levels in the rhyolites. Fractionation would thus have to involve unreasonably large amounts of hornblende (>50%, with apatite constrained to $\approx 1\%$) to produce the observed rhyolite REE patterns (figure 9.3a).

Petrogenetic models at Medicine Lake require <20% hornblende fractionation, which is supported by mass balance and petrographic constraints (Grove and Donnelly-Nolan, 1986). Andesite to rhyolite differentiation at Medicine Lake is characterized by increasing LREE and constant HREE (figure 9.3b), which is consistent with \approx 15% hb + \approx 1% ap fractionation. Dacite to rhyolite differentiation at the Three Sisters, however, involves large LREE and HREE depletions that can only be produced through >50% amphibole fractionation. There is no evidence to indicate that amphibole fractionation occurred in the Three Sisters system, which had lower water contents than either Mt. Mazama or Medicine Lake (Chapter 5) and lacks hornblende-bearing intermediate rocks. The observed rhyolite REE abundances cannot be produced by any reasonable model involving crystal fractionation from Middle Sister dacite. Unlike many other calc-alkaline arc volcanos (Grove and Donnelly-Nolan, 1986), the dacite-rhyolite compositional gap at the Three





Figure 9.3: Rare earth element plots for Three Sisters rhyolite petrogenesis. A) Comparison of evolved Middle Sister dacite REE pattern with least evolved Three Sisters rhyolite REE's. Derivation of the rhyolites through crystal fractionation from the dacite would require > 50% amphibole fractionation, as apatite is limited to 1% by mass balance. This model is not reasonable for the essentially anhydrous Three Sisters system. B) Variations in REE at Medicine Lake are consistent with fractionation of plag (≈70%) + thb (≈20%) + ap (≈1%) from andesite to produce rhyolite (Grove and Donnelly-Nolan, 1986). Note that the Medicine Lake model produces large increases in LREE, La/Sm, Eu/Eu*, and relatively constant HREE; Three Sisters dacite-rhyodacite differentiation is characterized by significant depletions in LREE and HREE, and constant Eu/Eu*, which is not consistent with the Medicine Lake model.

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Δ

B

Sisters cannot be the product of crystal fractionation. If the Three Sisters rhyolites cannot be produced through fractionation of less evolved magmas, then they must ultimately be derived from partial melts of crustal rocks.

9.1.2 Partial melting

Previous discussions have concluded that relatively dense, mafic crust is located from ≈ 30 to ≈ 45 km beneath this part of the High Cascades, with a more heterogeneous (but predominately mafic) crust between ≈ 12 and ≈ 30 km (e.g., Leaver et al., 1984; Riddihough et al., 1984; Chapter 2). Temperatures at middle (15-30 km) crustal depths are not certain, but temperatures of around 500°C at 10 km are justified through extrapolation of surficial heat flow values (Blackwell et al., 1982; 1990). Assuming that the "meta-mafic" country rock is at 500°C with a solidus at 1000°C, a specific heat of ≈ 0.25 cal/g °C, and a heat of fusion equivalent to the latent heat of crystallization of nearby mafic to intermediate magmas (at ≈ 1100 °C), then $\approx 50\%$ -25% of the volume of the intrusive rocks can be produced as 100% melt of the country rock. In reality, smaller volumes of intrusive rock are required to generate the same volume of melt, because melting will only involve 30%-50% of the country rock (cf. Bergantz, 1990). Although there is a large uncertainty on each of the values used in this calculation, it is reasonable to conclude that volumes of rhyolitic crustal melt of at least 1-10 km³ could be generated at mid-crustal depths under the High Cascades.

Although isotopic data are extremely limited for this part of the High Cascades, Three Sisters rhyolite 3S139 has the same relatively low 87 Sr/ 86 Sr (0.70363 ± 3) as Broken Top rhyolite 3S032 (0.70368 ± 2), indicating that isotopically evolved crust (i.e. \geq 0.705) does not underlie this area (table 3.3). Three Sisters rhyolites are characterized by flat HREE patterns, indicating that garnet and amphibole were not restitic minerals (Barker and Arth, 1976; Drummond and Defant, 1990). Garnet is a stable mineral in many basalt melting experiments at pressures >10 kbars (Ellis and Thompson, 1986; Rutter and Wyllie, 1988; Rapp, 1990), which strongly indicates that Three Sisters rhyolites were not produced at depths >35-40 km. Three Sisters rhyolites are also metaluminous, and only contain trace amount of normative corundum (cf. Chapter 6). Partial melts of peraluminous metasediments always produce peraluminous melts (Thompson, 1982; Conrad et al., 1988; Patiño-Douce and Johnston, 1990), and as such could not generate the Three Sisters rhyolites. Water saturated melts of basalt that leave amphibole in the residuum are also mildly to strongly peraluminous (Helz, 1976; Spulber and Rutherford, 1983; Ellis and Thompson, 1986; Beard and Lofgren, 1989, 1991; Winther, 1989), and do not resemble most silicic rock compositions.

Dehydration melting experiments (i.e. the only H_2O in the sample is contained in the hydrous minerals) of basalt have generally focused on the origin of island arc silicic rocks (e.g. tonalites and low- K_2O granitoids). These studies have used a variety of mafic starting materials: Unaltered MORB's and Hawaiian tholeiites (Holloway and Burnham, 1972; Helz, 1976; Spulber

and Rutherford, 1983; Wolf and Wyllie, 1989), tholeiitic metabasalts and metabasaltic andesite (Beard and Lofgren, 1989, 1991), and the synthetic CMASH system (Ellis and Thompson, 1986).

The rocks used by Beard and Lofgren (1991) were greenschist facies metavolcanics from the late Jurassic Smartville Complex of northern California (Beard and Day, 1987), which may be broadly correlative to the metavolcanic rocks that probably underlie this part of the High Cascades (Chapter 2). Samples were partially melted at pressures (1, 3, 6.9 kbars) and temperatures (850 °C to 1000 °C) that correspond to the most likely conditions for partial crustal melting underneath the High Cascades. Although the concentrations of TiO₂, Na₂O, K₂O, and P₂O₅ in the partial melts will be controlled by the concentrations of these elements in the parental rock, the metavolcanic rocks are reasonably similar to High Cascade mafic compositions.

The results of Beard and Lofgren (1991), using starting compositions that most closely resemble High Cascade compositions, are given in table 9.4. Partial melts produced by $\approx 15-25\%$ melting of the metabasalts at ≈ 900 °C, and $\approx 25\%$ melting of basaltic andesite at 950 °C, most closely resemble the composition of Three Sisters rhyolites. Overall enrichments in TiO₂ and P₂O₅ and depletions in alkalies systematically reflect the relative abundances of these elements in the metabasaltic rocks (Beard and Lofgren, 1991). The partial melts are also higher in CaO and FeO than the Three Sisters rhyolites, which is probably due to the relative enrichment of these elements in the starting compositions. The initial melt-forming reaction for the dehydration melts at 3 kbars is: 9 Qtz + 33 Amph + 5 Plag = 2 Mgt + 6 Opx + 20 Cpx + 19 Melt, and is terminal for both quartz and amphibole (Beard and Lofgren, 1991). Rocks that contain more CaO and FeO will thus produce melts with higher CaO and FeO. Above 900 °C, the melt-forming reaction is : 13 Plag + 1 Cpx + 1 Opx + 1 Ilmn + 1 Mgt = 18 Melt, and most of the melt is produced through the breakdown of plagioclase (Beard and Lofgren, 1991). Because the normative plagioclase composition is similar for both the metabasaltic and mafic High Cascade rocks ($\approx An_{50}$), differences in CaO content are probably controlled by amphibole and not by plagioclase.

Dehydration melting of metabasaltic rocks at pressures of 3 to 7 kbars and temperatures of ≈900 °C can produce $\approx20\%$ melt that has the general characteristics of Three Sisters rhyolites, once the differences in source composition are accounted for. The mineralogy of the metabasaltic source rock will change from an amphibole-rich greenstone (40-50% amphibole) to a pyroxene hornfels ($\approx50\%$ plag, $\approx40\%$ pyroxene, table 9.4) restite, which is observed in the Smartville Complex (Beard and Day, 1987). The experimental data of Beard and Lofgren (1991) show that apatite was also a restitic mineral up to 1000 °C; P₂O₅ contents in the melt increase with increasing temperature, and mass balance calculations indicate that the restite contains ≈0.5 wt.% apatite (table 9.4).

Trace elements from experimentally produced partial melts are not determined, due to the difficulty in analyzing small (≤ 100 mg), heterogeneous glass samples. These data are also not

| Sample: | 478 | | | | | | | 571 | | | | | 557 | | | | |
|---------|-------|-------|-------|------|--------|--------|----------|----------|---------|----------|------|------|------|------|------------|------|------|
| P (kbar |) | 3 | 3 | 3 | 6.9 | 6.9 | 6.9 | | 3 | 6.9 | 6.9 | 6.9 | | 3 | 3 | 3 | 6.9 |
| T (°C) | / | 900 | 950 | 1000 | 850 | 900 | 950 | | 900 | 850 | 900 | 950 | | 900 | 950 | 1000 | 1000 |
| Si02 | 52.5 | 76.6 | 73.1 | 62.9 | 75.6 | 72.3 | 71.0 | 51.4 | 73.2 | 74.7 | 73.0 | 70.4 | 57.0 | 76.2 | 72.6 | 69.1 | 69.1 |
| Ti02 | 1.74 | 0.43 | 0.67 | 1.56 | 0.16 | 0.59 | 0.64 | 1.55 | 0.49 | 0.12 | 0.41 | 0.64 | 0.60 | 0.40 | 0.74 | 0.90 | 0.93 |
| A1203 | 15.3 | 13.1 | 13.9 | 15.5 | 14.5 | 15.3 | 15.6 | 15.8 | 15.2 | 14.7 | 15.6 | 17.3 | 15.4 | 13.0 | 14.5 | 16.0 | 16.3 |
| Fe0* | 11.79 | 2.39 | 3.59 | 7.91 | 2.10 | 2.93 | 3.81 | 12.23 | 2.07 | 1.96 | 2.30 | 2.18 | 8.01 | 2.29 | 3.44 | 4.55 | 4.20 |
| MnQ | 0.26 | 0.08 | 0.10 | 0.16 | 0.06 | 0.06 | 0.06 | 0.26 | 0.08 | 0.08 | 0.07 | 0.09 | 0.17 | 0.08 | 0.05 | 0.08 | 0.04 |
| MaO | 4.74 | 0.49 | 0.83 | 2.08 | 0.24 | 0.87 | 0.47 | 4.42 | 0.47 | 0.38 | 0.72 | 0.30 | 5.52 | 0.60 | 0.69 | 0.51 | 0.50 |
| CaO | 10.90 | 2.72 | 3.40 | 5.61 | 2.76 | 3.76 | 3.63 | 8.95 | 2.11 | 2.42 | 3.20 | 2.67 | 9.20 | 2.67 | 3.13 | 3.71 | 3.80 |
| Na20 | 1.96 | 3.21 | 3.43 | 3.15 | 2.83 | 3.05 | 3.50 | 3.30 | 3.77 | 3.10 | 2.58 | 3.63 | 2.54 | 3.01 | 3.13 | 3.46 | 3.36 |
| K20 | 0.15 | 0.92 | 0.85 | 0.58 | 1.76 | 0.96 | 1.02 | 0.37 | 2.56 | 2.49 | 1.98 | 2.33 | 0.44 | 1.66 | 1.48 | 1.52 | 1.48 |
| P205 | 0.30 | 0.07 | 0.15 | 0.58 | na | 0.18 | 0.27 | 0.30 | 0.08 | 0.09 | 0.11 | 0.37 | 0.18 | 0.08 | 0.25 | 0.25 | 0.24 |
| H20 | ≈1 | 4.9 | 3.0 | 3.8 | 5.7 | 4.8 | 3.8 | ≈1 | 6.0 | 3.5 | 8.1 | 5.2 | ≈1 | 5.2 | 3.4 | 4.3 | 4.4 |
| XHELT | | 21 | 26 | 39 | 6 | 21 | 20 | | 16 | 7 | 12 | 17 | | 23 | 28 | 35 | 28 |
| | | | | | | | | | | | | | | | | | |
| Restite | 2 | | | | - | | | | | | | | | 70 | 70 | 70 | |
| Plg | | 40 | 36 | 29 | 36 | 37 | 38 | | 48 | 49 | 48 | 44 | | 39 | 38 | 50 | 33 |
| Срх | | 19 | 22 | 21 | | 23 | 24 | | 23 | | 24 | 20 | | 25 | 22 | 21 | 29 |
| Орх | | 11 | 7 | 6 | | 12 | 12 | | 0.5 | 70 | 8 | 2 | 10 | 8 | (| 10 | Y |
| Amph | 46 | | | | 46.2 | | | 50 | | 38 | Ir | | 40 | | | | |
| Qtz | | | | | 6 | | | | | , | - | 0 | | (F) | 1 | | |
| Mgt | | 4 | 4 | 4 | 1 | 3 | 2 | | 11 | 4 | 5 | 8 | | 2 | 0 | 4 | 2 |
| Ilmn | | 6 | 5 | 3 | 4 | 4 | 4 | | 1 | 3 | 4 | 2 | | | | | 6 |
| | 35060 | 35045 | 35148 | | Restit | e reca | lculated | to 100% | | | | | | | | | |
| sin2 | 51.8 | 56.0 | 72.5 | | Plag | 47 | | | 58 | | | | | | 52 | | |
| Ti02 | 1.02 | 1.12 | 0.33 | | Cox | 29 | | | 28 | | | | | | 30 | | |
| A1203 | 16.7 | 18.3 | 14.5 | | XqO | 15 | | | 0. | 4 | | | | | 10 | | |
| Fe0* | 8.39 | 7.41 | 2.06 | | Mat | 4 | | | 13 | | | | | | 8 | | |
| MnO | 0.15 | 0.13 | 0.06 | | Ilmn | 5 | | | 1 | | | | | | - | | |
| MgO | 8.79 | 3.66 | 0.61 | | Ap* | 0. | 56 | | 0. | 67 | | | | | 0. | 53 | |
| CaO | 9.14 | 7.70 | 1.86 | | | | | | | | | | | | | | |
| Na20 | 3.22 | 4.45 | 4.77 | | | | | | | | | | | | | | |
| K20 | 0.64 | 1.02 | 3.15 | | Ap* by | mass b | alance, | Ca0 = 55 | .4 wt.% | , P205 = | 42.1 | it.% | | | | | |
| P205 | 0.21 | 0.28 | 0.07 | | | | | | | | | | | | | | |

Table 9.4: Composition of dehydration melts of metabasaltic rocks, from Beard and Lofgren (1991). 3S148 is the least evolved Pleistocene Three Sisters rhyolite, with representative High Cascade mafic compositions shown for comparison. The general effect of increasing pressure is to lower the melt fraction and overall variations in melt composition, and to produce slightly more aluminous melts. Increasing temperature lowers SiO₂ and K₂O contents, with large increases in MgO, FeO, CaO and Al₂O₃ as more refractory phases (plag + px) melt. Dehydration melts of basalt 478 (6.9 kbars) and 571 (3 kbars) at 900°C, and basaltic andesite 557 (3 kbars) at 950°C, most closely resemble Three Sisters rhyolite 3S148. Elevated FeO and CaO contents of the melts relative to 3S148 probably reflects the enrichment of these element in the starting compositions; differences in TiO₂, alkalis and P₂O₅ between 3S148 and the melts are directly controlled by the source composition (Beard and Lofgren, 1991).

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reported for the starting compositions used in the dehydration melting experiments. Trace element abundances can, however, be estimated by using High Cascade compositions from this study, melting models constrained by Beard and Lofgren's data, and distribution coefficients representative of a silicic (i.e. low silica rhyolite) melt in equilibrium with an anhydrous restite. Equilibrium batch melting models of High Cascade basalt 3S069, using melt fractions and restite mineralogies from Beard and Lofgren (1991), support the origin of Three Sisters rhyolites through dehydration melting of metabasaltic rocks between 3 and 6.9 kbars (table 9.5). The models are limited by the uncertainties associated with most trace element distribution coefficients in low silica rhyolite melts (cf., Appendix E), and potential ranges in source rock composition. In general, successful trace element models require small amounts of apatite ($\leq 0.5\%$) in the restite to buffer REE abundances and produce flat HREE patterns, which is supported by mass balance calculations (table 9.5). Melts that strongly resemble Three Sisters rhyolites can be produced through 16-20% equilibrium batch melting of basalt 3S069, leaving a residual mineralogy of plag + px + oxides + apatite. Differences between observed and calculated trace element abundances are within reasonable uncertainties in trace element distribution coefficients and source composition (i.e., ± 10% variation in each parameter).

Holocene Three Sisters rhyolites generally contain more Al_2O_3 , FeO*, CaO and Sr, and lower MgO and Sc, than Pleistocene rhyolites (table 9.1). Basaltic dehydration melt reactions above 900°C are dominated by the dissolution of plagioclase (Beard and Lofgren, 1991). The geochemical differences between the Holocene and Pleistocene rhyolites may thus be due to slightly higher melting temperatures (\geq 900°C) during production of the Holocene rhyolites; enhanced dissolution of plagioclase would increase the melt in Al_2O_3 , CaO and Sr, and lower MgO and Sc through relatively lower degrees of augite dissolution (i.e., increased augite in the restite). Although the geochemical variations are too small to be rigorously modeled, there is no compelling evidence to indicate that the Holocene rhyolites were produced through partial melts of a crust that was compositionally distinct from the source of the Pleistocene rhyolites.

9.1.3 Summary of Three Sisters rhyolite petrogenesis

In summary, Three Sisters rhyolites cannot be produced through crystal fractionation from dacitic compositions. The Three Sisters silicic system lacks abundant amphibole, which is commonly required to produce compositional gaps and rhyolitic melts through crystal fractionation. Major and trace element models demonstrate that the Three Sisters rhyolites were most likely produced through dehydration melting of metabasaltic crust at 3-7 kbars ($\approx 10-25$ km) and around 900°C. Magma at ≈ 1100 °C intruded into an ≈ 500 °C crust at these depths can likely produce partial melt volumes roughly equivalent to the volume of the intrusions. Because the volume of rhyolite is much smaller than the volume of basaltic andesitic to andesitic rocks, this model is thermally valid.

| Model ; | #1, | Metabasalt | 571, | 3 | kbars, | 900°C |
|---------|-----|------------|------|---|--------|-------|
|---------|-----|------------|------|---|--------|-------|

| | Distribution Coefficients | | | | | | | | | | | | | |
|------------------|---------------------------|------------|------------|------------|----------|--------------|--------------|------------|------------|------------|--|--|--|--|
| | Restite | Rb | Sr | Sc | V | La | Sm | Eu | Tb | Lu | | | | |
| Plag | 0.58 | 0.3 | 4.4 | | | 0.3 | 0.1 | 2.5 | 0.2 | 0.1 | | | | |
| Срх | 0.28 | | 0.4 | 33 | 3 | 0.5 | 2.9 | 2.5 | 3.8 | 3.0 | | | | |
| Орх | 0.004 | | | 7 | 1 | 0.4 | 0.8 | 0.6 | 0.9 | 0.9 | | | | |
| Ap | 0.007 | | 1 | | | 15.0 | 46.0 | 26.0 | 38.0 | 14.0 | | | | |
| Mgt | 0.13 | | | 5 | 40 | | | | | | | | | |
| Ilmn | 0.01 | | | 1 | 10 | | | | | | | | | |
| Bulk D: | | 0.17 | 2.67 | 9.93 | 6.14 | 0.4 | 1.2 | 2.3 | 1.4 | 1.0 | | | | |
| Cs (3SC |)69) | 11 | 439 | 27.7 | 187 | 30.4 | 16.1 | 15.1 | 10.9 | 8.7 | | | | |
| Model # 35148 | 1, F=0.1 | 6 36 71 | 183 193 | 3.3 4.6 | 35 31 | 59.7 61.5 | 13.9 15.5 | 7.1 8.6 | 8.1 8.7 | 8.7 9.0 | | | | |

Model #2, Metabasalt 478, 6.9 kbars, 900°C

| Distribution Coefficients | | | | | | | | | | | | |
|---------------------------|----------|------|------|-------|------|------|------|------|------|------|--|--|
| | Restite | Rb | Sr | Sc | V | La | Sm | Eu | Tb | Lu | | |
| Plag | 0.47 | 0.3 | 4.4 | | | 0.3 | 0.1 | 2.5 | 0.2 | 0.1 | | |
| Срж | 0.29 | | 0.4 | 33 | 3 | 0.5 | 2.9 | 2.5 | 3.8 | 3.0 | | |
| Орх | 0.15 | | | 7 | 1 | 0.4 | 0.8 | 0.6 | 0.9 | 0.9 | | |
| Ap | 0.005 | | 1 | | | 15.0 | 46.0 | 26.0 | 38.0 | 14.0 | | |
| Mgt | 0.04 | | | 5 | 40 | | | | | | | |
| Ilmn | 0.05 | | | 1 | 10 | | | | | | | |
| Bulk D: | 1 | 0.14 | 2.19 | 10.87 | 3.12 | 0.4 | 1.2 | 2.1 | 1.5 | 1.1 | | |
| Св (350 |)69) | 11 | 439 | 27.7 | 187 | 30.2 | 16.1 | 15.1 | 10.9 | 8.7 | | |
| Model # | 2, F=0.2 | 0 34 | 226 | 3.1 | 70 | 55.7 | 13.5 | 8.0 | 7.8 | 7.9 | | |
| 35148 | | 71 | 193 | 4.6 | 31 | 61.5 | 15.5 | 8.6 | 8.7 | 9.0 | | |

Table 9.5: Estimated trace element abundances from partial melting of High Cascade basalt 3S069, using melt fractions and restite modal mineralogies from the dehydration melting experiments of Beard and Lofgren (1991). Model #1 uses a melt composition from metabasalt 571 that most closely matches the composition of Three Sisters rhyolite 3S148 (cf., table 9.4); the partial melt from metabasalt 478 (Model #2) also approximates 3S148. Both models use representative distribution coefficients for rhyodacitic magmas, and trace element abundances are calculated assuming equilibrium batch melting ($C_1 = C_s / [D (1 - F) + F]$). The correspondence between modeled and observed trace element abundances are within the uncertainty associated with the rhyodacitic distribution coefficients and reasonable variations in source composition. For example, models underestimate Rb abundances; lowering the plag Rb K_d to 0.1 and increasing Rb in the source to 15 ppm duplicates 3S rhyolite abundances. Residual apatite is required to buffer the REE's, and to produce flat HREE patterns. Models that lack residual apatite fail to account for observed P_2O_5 and REE trends. REE abundances in the models are also lower than observed in 3S rhyolites, but a 10% increase in the source concentration duplicates 3S rhyolite abundances. Partial melting of High Cascade basalts leaving a restite of essentially plag + pyroxene can generate Three Sisters rhyolite compositions.

Holocene rhyolites were probably generated at slightly higher temperatures than the Pleistocene rhyolites, resulting in a relatively larger plagioclase component in the Holocene melts. Variations within the Holocene and Pleistocene rhyolite suites can be accounted for by 6-15% crystal fractionation of modal minerals, but variations between these suites cannot be modeled through either crystal fractionation or mixing with any known composition. The Holocene rhyolites represent a compositionally distinct batch of magma, which thus had the potential to interact with the South Sister system differently than the Pleistocene rhyolites.

9.2 Broken Top System Rhyolites

Rhyolites in the Broken Top system are restricted in occurrence to domes and pyroclastic deposits associated with the Tumalo volcanic center (Chapter 3). Although the ages of most of these rhyolites are poorly constrained, the main pulse of rhyolitic volcanism occurred around 380 ± 20 ka. However, previous discussions have demonstrated the necessity of rhyolite magma mixing in the petrogenesis of Broken Top rhyodacites and dacites (Chapter 8). Rhyolitic magma must thus have been present in the Broken Top system to at least ≈ 200 ka, which is the age of the base of Broken Top volcano (Chapter 3).

Broken Top rhyolites are aphyric to sparsely phyric and contain $\leq 1\%$ phenocrysts of plagioclase (An₂₅₋₂₀) > hypersthene (En₄₀) > amphibole \geq magnetite + ilmenite > apatite > zircon. In contrast to the more phyric Three Sisters rhyolites, Broken Top rhyolites contain more sodic plagioclase (3S = An₄₀₋₃₀) and less magnesian hypersthene (3S \approx En₅₅) (figure 4.6). Broken Top rhyolites are also compositionally distinct from Three Sisters rhyolites (table 9.6), and contain lower amounts of TiO₂, MgO, CaO, Sr, and Co. Broken Top rhyolites also contain significantly more Na₂O, Sc, Zn, REE's (esp. HREE) and HFSE's than Three Sisters rhyolites (figure 9.4). These compositional distinctions have been used in previous petrogenetic models (Chapters 7-8) to show that rhyolites from each magmatic system did not interact with less evolved magmas from the other system.

9.2.1 Crystal fractionation

Models of Broken Top intermediate system petrogenesis have demonstrated the necessity of rhyolite mixing to evolve most dacites and rhyodacites; high Na₂O dacite is the most evolved Broken Top composition that can be attributed to predominantly crystal fractionation. If the Broken Top rhyolites were produced through crystal fractionation, then high Na₂O dacite should have been the parental composition. Amphibole fractionation over a small temperature interval has been demonstrated to produce compositional gaps between dacite and rhyolite at Medicine Lake volcano (Grove and Donnelly-Nolan, 1986). A possible model for Broken Top magmatism could

| SiO2 TiO2 A12O3 FeO* MnO MgO CaO Na2O K2O P2O5 | 38098 71.1 0.40 14.9 2.74 0.08 0.30 1.40 6.13 2.77 0.07 | 38037 69.6 0.54 15.4 3.31 0.07 0.36 1.64 5.38 2.93 0.11 | 35XBP 74.2 0.16 14.2 1.82 0.07 0.08 0.81 4.46 3.47 0.03 | 38032 74.5 0.15 13.7 1.70 0.06 0.01 0.71 5.26 3.48 0.02 | Avg 74.6 0.15 13.9 1.63 0.06 0.05 0.69 5.18 3.49 0.02 | $1\sigma \\ \pm 0.2 \\ \pm 0.01 \\ \pm 0.05 \\ \pm 0.00 \\ \pm 0.05 \\ \pm 0.02 \\ \pm 0.03 \\ \pm 0.03 \\ \pm 0.00 $ | 38054 74.2 0.24 14.3 1.58 0.05 0.23 1.20 4.53 3.37 0.05 |
|---|--|--|---|--|--|--|---|
| Rb Ba Sr Cs | 58 721 135 2.04 | 59 736 181 0.90 | 84 794 65 3.10 | 86 796 52 1.50 | 84 796 50 1.37 | ± 2 ± 11 ± 3 ± 0.27 | 74 840 144 1.30 |
| SC V Co Ni Cr Zn | 9.25 1 1.59 10 2 135 | 8.73 15 2.50 10 <5 54 | 4.29 0.50 11 105 | 4.25 0 0.54 10 <3 84 | 4.30 3 0.66 10 1 79 | ± 0.18 ± 4 ± 0.15 ± 1 ± 1 ± 1 ± 11 | 2.51 3 1.47 <1 <2 49 |
| La Ce Nd Sm Eu Tb Yb Lu La/Sm ZHREE Eu/Eu* | 24.4 61.3 30.8 6.46 1.28 1.28 4.2 0.66 2.38 67.67 0.56 | 22.2 53.7 24.0 5.99 1.28 0.84 3.5 0.50 2.34 50.02 0.66 | 28.7 59.7 26.5 5.95 0.83 0.94 4.2 0.58 3.05 57.77 0.42 | 25.3 52.6 20.4 4.77 0.64 0.78 3.7 0.53 3.35 50.62 0.40 | 23.2 52.5 20.4 4.42 0.61 0.73 3.3 0.50 3.32 46.74 0.41 | $\begin{array}{c} \pm & 2.8 \\ \pm & 8.8 \\ \pm & 2.3 \\ \pm & 0.44 \\ \pm & 0.03 \\ \pm & 0.2 \\ \pm & 0.2 \\ \pm & 0.03 \\ \pm & 0.19 \\ \pm & 3.39 \\ \pm & 0.03 \end{array}$ | $ 19.8 \\ 36.2 \\ 14.0 \\ 2.62 \\ 0.49 \\ 0.35 \\ 1.6 \\ 0.24 \\ 4.77 \\ 22.50 \\ 0.59 \\ \end{array} $ |
| Zr Hf Ga Y Nb Ta Th U | 274 8.3 20 37 15.5 0.90 5.5 2.2 | 308 8.4 19 30 16.8 1.00 5.4 2.2 | 231 6.7 20 36 16.9 1.20 8.1 2.7 | 222 6.5 20 31 17.8 1.00 7.9 2.7 | 217 6.6 20 28 17.4 1.02 8.1 2.5 | $\begin{array}{c} \pm & 6 \\ \pm & 0.5 \\ \pm & 2 \\ \pm & 3 \\ \pm & 0.7 \\ \pm & 0.03 \\ \pm & 0.4 \\ \pm & 0.2 \end{array}$ | 180 4.6 15 16 11.5 0.77 6.4 2.6 |

3S098: Squaw Ck. dome, high Na₂O rhyodacite.
3S037: E. Tam MacArthur Rim, main series rhyodacite.
3SXBP: Average Bend Pumice; bulk pumice sample.
3S032: Three Ck. Butte dome, from summit.
AVG: Average of 8 TVC rhyolites, excluding 3SXBP.
3S054: Pleistocene Three Sisters rhyolite, E. flank Devils Hill.

Table 9.6: Representative analyses of Broken Top system rhyodacites and rhyolites, with Three Sisters rhyolite 3S054 for comparison. 3S098 and 3S037 show range in composition of Broken Top system rhyodacites, which represents range of compositions parental to the rhyolites. Note that 3S032 is representative of average TVC rhyolite compositions, but that many elemental abundances in 3SXBP are outside of average TVC rhyolite compositions. This is most apparent for Na₂O and Cs, which are highly succeptable to deuteric alteration or weathering. 3SXBP is also relatively enriched in REE and Ta, which may indicate that the Bend Pumice was somwhat more evolved than other TVC rhyolites.

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Figure 9.4: Chondrite normalized REE patterns for Broken Top rhyolite petrogenesis. A) Plot of high Na₂O and main trend rhyodacites, with Broken Top and Three Sisters rhyolites shown for comparison (table 9.6). If these rhyodacites were parental to Broken Top rhyolites, then fractionation involved significant REE depletions. These depletions cannot be modeled from high Na₂O dacites and rhyodacites, as models require unrealistically high bulk distribution coefficients for the REE's. B) Modeled REE abundances resulting from fractionation of main trend rhyodacite 3S037, using a mass-balance constrained assemblage. All fractionation models that reproduce observed major element trends fail to produce observed REE depletions. Crystal fractionation is not a valid hypothesis for the origin of Broken Top rhyolites.

B

thus involve production of rhyolite through crystal fractionation of dacite, with subsequent magma mixing and fractionation to produce the main dacite-rhyodacite trend.

Mass balance calculations show that the Broken Top rhyolites cannot be produced through crystal fractionation of high Na₂O dacite (table 9.7). Anhydrous fractionation involving 90% An_{20} plagioclase cannot lower Na₂O contents to observed rhyolite levels without severely depleting the melt in CaO and Al₂O₃. The addition of amphibole to the fractionating assemblage results in further CaO and Al₂O₃ depletions, and raises FeO, MnO and MgO contents in the modeled rhyolite (table 9.7). Fractionation of high Na₂O dacite, which was derived predominantly through crystal fractionation, does not produce Broken Top rhyolite compositions.

Mass balance models using rhyodacite from the main Broken Top series can produce Broken Top rhyolites (table 9.7). However, it has been demonstrated that the petrogenesis of these rhyodacites involves $\approx 30\%$ mixing with Broken Top rhyolite. If rhyolite was a prerequisite for the evolution of the main series rhyodacites, then the rhyolite must have been independently produced and thus these rhyodacites could not have been parental to the Broken Top rhyolites.

Trace element distributions also clearly demonstrate that fractionation of the main series rhyodacites could not have resulted in Broken Top rhyolite compositions. Observed enrichments in incompatible elements required a minimum of 30% crystallization, but mass balance constraints limit crystallization to 22%. The crystallizing mineral assemblage is dominated by plagioclase, which is capable of producing the observed depletions in Sr and Eu (rhyodacite distribution coefficients from table 9.5). However, REE distributions are inconsistent with crystal fractionation of the mass balance constrained mineral assemblage (figure 9.4b). Rhyodacite to rhyolite fractionation is characterized by depletions in all the REE's, but bulk distribution coefficients are $\approx 50\%$ too low to have depleted the REE's in the rhyolite. Crystal fractionation of 82% plag, 7% opx, and 1% ap results in elevated REE contents in the modeled rhyolite (figure 9.4b). In conclusion, Broken Top rhyolites cannot be produced through crystal fractionation of either type of Broken Top rhyodacite. These rhyolites must thus also be the products of partial crustal melts.

9.2.2 Partial melting

Three Sisters rhyolites were produced through $\approx 20\%$ melting of metabasaltic crustal rocks, leaving a plagioclase + pyroxene + apatite restite. The restite mineralogy resulted in REE patterns that have relatively high La/Sm (≈ 4), modest Eu/Eu* (≈ 0.6) and low HREE (cf. 3S054 in figure 9.4a). Relative to Three Sisters rhyolites, Broken Top rhyolites have similar major element concentrations, but have elevated REE contents, lower La/Sm (≈ 3) and larger Eu/Eu* (≈ 0.4) (table 9.6). These REE distributions indicate the residual mineralogy probably contained more plagioclase and less apatite than observed for Three Sisters rhyolites. The restites from dehydration melting of metabasaltic rocks (Beard and Lofgren, 1991) always contained $\leq 60\%$ plag $\leq 15\%$ mgt +

| | | SiO2 | TiO2 | A1203 | FeO* | MnO | MgO | CaO | Na 20 | K2O | P205 |
|------|------|------|------|-------|------|-----|------|------|-------|------------|------|
| Plag | An35 | 58.7 | | 25.6 | 0.5 | | 0.0 | 7.2 | 7.0 | 0.4 | |
| Plag | An20 | 62.5 | | 22.3 | 0.2 | | 0.0 | 4.1 | 8.5 | 0.8 | |
| Орх | En55 | 52.4 | 0.1 | 0.4 | 24.8 | 1.1 | 20.2 | 0.9 | 0.1 | | |
| Mgt | | 0.1 | 9.6 | 1.8 | 82.7 | 0.5 | 1.4 | | | | |
| Ilmn | | 0.0 | 42.2 | | 51.5 | 0.6 | 2.5 | | | | |
| Apt | | | | | | | | 56.0 | | | 43.0 |
| Hb | | 42.1 | 3.8 | 10.4 | 14.7 | 0.3 | 12.9 | 10.7 | 2.3 | 0.3 | |

 Starting composition: high Na₂O rhyodacite, models 1 & 2

 SiO2
 TiO2
 Al2O3
 FeO*
 MnO
 MgO
 CaO
 Na2O
 K2O
 P2O5

 3S098
 71.1
 0.40
 14.9
 2.74
 0.08
 0.30
 1.40
 6.13
 2.77
 0.07

Model #1: 28% anhydrous fractionation of 90% An₂₀ Plag, 4% Opx, 3.5% Mgt, 2% Ilmn, 0.5% Ap

| | SiO2 | TiO2 | A1203 | FeO* | MnO | MgO | CaO | Na 20 | K2O | P205 |
|-----------|------|------|-------|------|------|------|------|-------|------------|------|
| Model #1: | 75.3 | 0.14 | 13.2 | 1.96 | 0.08 | 0.10 | 0.56 | 5.63 | 3.44 | 0.02 |
| AVG RHYO | 74.6 | 0.15 | 13.9 | 1.63 | 0.06 | 0.05 | 0.69 | 5.18 | 3.49 | 0.02 |

Model #2: 28% hydrous fractionation of 90% An₂₀ Plag, 4% Hb, 4% Mgt, 1.5% Ilmn, 0.5% Ap

| | SiO2 | TiO2 | A1203 | FeO* | MnO | MgO | CaO | Na20 | K20 | P205 |
|-----------|------|------|-------|------|------|------|------|------|------|------|
| Model #2: | 75.4 | 0.15 | 13.0 | 2.04 | 0.09 | 0.19 | 0.43 | 5.61 | 3.44 | 0.02 |
| AVG RHYO | 74.6 | 0.15 | 13.9 | 1.63 | 0.06 | 0.05 | 0.69 | 5.18 | 3.49 | 0.02 |

Model #3: 22% fractionation of main series rhyodacite 3S037, 81.5% An₃₀ Plag, 7% Opx, 7% Mgt, 3.5% Ilmn, 1% Ap

| | | | St | tartin | g comp | ositio | a | | | |
|-----------|------|------|-------|--------|--------|--------|------|------|------------|------|
| | SiO2 | TiO2 | A1203 | FeO* | MnO | MgO | CaO | Na20 | K2O | P205 |
| 35037 | 69.9 | 0.54 | 15.4 | 3.31 | 0.07 | 0.36 | 1.64 | 5.38 | 2.93 | 0.11 |
| Model #3: | 74.2 | 0.14 | 14.2 | 1.73 | 0.05 | 0.05 | 0.66 | 5.19 | 3.55 | 0.03 |
| AVG RHYO | 74.6 | 0.15 | 13.9 | 1.63 | 0.06 | 0.05 | 0.69 | 5.18 | 3.49 | 0.02 |

Table 9.7: Mass balance models involving crystal fractionation of Broken Top system rhyodacites to produce Broken Top rhyolite. Although most rhyodacite petrogenesis requires $\approx 30\%$ mixing of rhyolite magma, the Broken Top rhyodacites that were most directly related to crystal fractionation are part of a high Na₂O trend. These rhyodacites (e.g., 3S098) are the most likely parental compositions for rhyolite differentiation. Simple mass balance models (#1 & #2) demonstrate that fractionation of 90% An₂₀ plagioclase will not produce the required depletions in Na₂O. The fractionation of hornblende in place of orthopyroxene (model #2) also results in CaO and Al₂O₃ depletions that are too large, and FeO and MgO depletions that are too small, relative to average rhyolite compositions. Broken Top rhyolites cannot be produced by using a main series rhyodacite (e.g., 3S037) as the parental composition. Orthopyroxene can also be replaced by up to 2% hornblende and still result in rhyolite compositions at 22% fractionation. Note the large amounts of plagioclase (An₃₀) and Fe-Ti oxides required by the model, which will strongly affect trace element distributions.

ilmn, and $cpx \ge opx$. Equilibrium batch melting models (table 9.8) using these mineralogical constraints fail to reproduce Broken Top rhyolite trace element abundances using typical basaltic starting compositions. These models always resulted in melt concentrations of Sr and V were too high, and Rb and REE that were too low, when elemental abundances representative of High Cascade mafic rocks were used as the starting composition (table 9.8). The source rocks of Broken Top rhyolite partial melts must be depleted in Sr, Sc and V, and contain elevated REE's with higher La/Sm, relative to High Cascade mafic rocks.

There are few detailed partial melting experiments over a range of pressures (3-10 kbars) and water contents (highly undersaturated) that would be appropriate for the High Cascades that use intermediate composition igneous rocks as starting materials. A detailed study by Conrad et al. (1988) on Taupo volcanic zone silicic magmatism utilized dacitic and greywacke starting compositions, which were melted at water-saturated and water-undersaturated conditions at 10 kbars. Partial melts of greywacke were always strongly peraluminous, and do not resemble High Cascade silicic rock. Partial melts of dacite under water-saturated to moderately undersaturated (X_{H2O} = 0.5) conditions also resulted in peraluminous rhyolitic melts. Metaluminous rhyolitic melts were only produced at water-undersaturated conditions (X_{H2O} = 0.25), which corresponds to ≈4 wt.% H₂O (Conrad et al., 1988).

This water content is probably higher than expected for the High Cascades magma system; dehydration melts typically occur at ≈ 1 wt.% H₂O, if the source rock contains $\approx 50\%$ amphibole (cf., Beard and Lofgren, 1991). The overall effect of elevated water content will be to increase the solubility of anorthitic plagioclase in the melt, resulting in higher Al₂O₃ and CaO contents (Conrad et al., 1988). Partial melts of Taupo dacite are higher in Al₂O₃ and CaO, and lower in Na₂O, than Broken Top rhyolites (table 9.9). Although some of these differences were likely controlled by differences in the source composition, partial melting of the Taupo dacite under relatively elevated water contents would result in more anorthitic plagioclase in the melt, and thus higher Al₂O₃ and CaO, and lower Na₂O. In order to account for these differences, it is likely that the generation of Broken Top rhyolites occurred at lower water contents (< 4 wt% H₂O) than used by Conrad et al. (1988). It is thus highly unlikely that Broken Top melting occurred with residual amphibole, because amphibole was only stable in the residua at >4 wt.% H₂O (X_{H2O} = >0.25) in the Conrad et al. experiments. Lower water contents might also stabilize apatite in the residua by limiting P₂O₅ diffusion into the melt (Harrison and Watson, 1984), which is supported by low P₂O₅ contents and REE characteristics of the Broken Top rhyolites (table 9.9).

Major element concentrations of Taupo dacite partial melts at 850°C, \approx 45% melting, are broadly similar to Broken Top rhyolites, once the effects of presumedly higher experimental water contents are accounted for (table 9.9). Trace element models, however, do not directly support the origin of Broken Top rhyolites though partial melting of dacitic rocks with the constraints of Equilibrium Batch Melting: $C_{l} = C_{s} / [D (1-F) + F]$

| | Distribution Coefficients | | | | | | | | | | | | |
|--------------------|---------------------------|-------------|-----------|------------|---------|--------------|--------------|------|--------------|--------------|--|--|--|
| | Restite | Rb | Sr | Sc | V | La | Sm | Eu | Tb | Lu | | | |
| Plag | 0.6 | 0.3 | 4.4 | | | 0.3 | 0.1 | 2.5 | 0.2 | 0.1 | | | |
| Cpx | 0.14 | | 0.4 | 33 | 3 | 0.5 | 2.9 | 2.5 | 3.8 | 3.0 | | | |
| Opx | 0.12 | | | 7 | 1 | 0.4 | 0.8 | 0.6 | 0.9 | 0.9 | | | |
| Ap | 0 | | 1 | | | 15.0 | 46.0 | 26.0 | 38.0 | 14.0 | | | |
| Mgt | 0.09 | | | 5 | 40 | | | | | | | | |
| Ilmn | 0.05 | | | 1 | 10 | | | | | | | | |
| Bulk D: | | 0.18 | 2.70 | 5.96 | 4.64 | 0.3 | 0.6 | 1.9 | 0.7 | 0.6 | | | |
| Св (350 | 69) | 11 | 439 | 27.7 | 187 | 30.2 | 16.1 | 15.1 | 10.9 | 8.7 | | | |
| Model # Cl (3S0 | (1, F = 0) | .2 32 86 | 186 52 | 5.6 4.3 | 48 3 | 68.9 81.1 | 24.4 24.2 | 8.7 | 13.9 16.6 | 13.0 16.4 | | | |

| | Distribution Coefficients | | | | | | | | | | | | |
|--------------------|---------------------------|-------------|----------|------|--------|--------------|--------------|------------|--------------|--------------|--|--|--|
| | Restite | Rb | Sr | Sc | V | La | Sm | Eu | Tb | Lu | | | |
| Plag | 0.6 | 0.3 | 4.4 | | | 0.3 | 0.1 | 2.5 | 0.2 | 0.1 | | | |
| Срх | 0.14 | | 0.4 | 33 | 3 | 0.5 | 2.9 | 2.5 | 3.8 | 3.0 | | | |
| Орх | 0.12 | | | 7 | 1 | 0.4 | 0.8 | 0.6 | 0.9 | 0.9 | | | |
| Ap | 0 | | 1 | | | 15.0 | 46.0 | 26.0 | 38.0 | 14.0 | | | |
| Mgt | 0.09 | | | 5 | 40 | | | | | | | | |
| Ilmn | 0.05 | | | 1 | 10 | | | | | | | | |
| Bulk D: | | 0.18 | 2.70 | 5.96 | 4.64 | 0.3 | 0.6 | 1.9 | 0.7 | 0.6 | | | |
| Cs (est | imated) | 30 | 125 | 22 | 15 | 35.0 | 16.1 | 15.1 | 13.0 | 11.0 | | | |
| Model # C1 (350 | ≇2, F = 0 032) | .2 87 86 | 53 52 | 4.4 | 4 3 | 79.8 81.1 | 24.4 24.2 | 8.7 8.6 | 16.6 16.6 | 16.4 16.4 | | | |

Table 9.8: Equilibrium batch melting models for Broken Top rhyolite petrogenesis, using mineralogical constraints from Beard and Lofgren (1991). Model #1: Restites from dehydration melts of metabasaltic rocks contain ≤60% plag and ≤15% mgt + ilmn (cf., Table 9.4; Beard and Lofgren, 1991). These mineral proportions are insufficent to lower Sr and V in a basaltic partial melt to rhyolite abundances, even at 10% melting. Note that apatite in the restite would result in additional HREE depletions, but would likely increase La/Sm to rhyolitic values. Model #2: Estimated source composition constrained by 60% plag and 15% mgt + ilmn in the restite, 20% melting. Note the low abundances of Sr and V necessary in the source to produce the observed rhyolitic concentrations; variations in F from 0.1 to 0.3 results in small variations in Sr and V abundances, but large variations in less compatible elements. Models #1 and #2 clearly demonstrate that Broken Top rhyolites were not produced through partial melts of metabasaltic rocks. Relative to basalts, potential source rocks must be depleted in Sr, Sc, and V, and contain elevated REE's with higher La/Sm.

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| | SiO2 | TiO2 | A1203 | FeO* | MnO | MgO | CaO | Na 20 | K2O | P205 |
|----------|------|------|-------|------|-----|-----|-----|-------|------------|------|
| TVZ exp. | 65.2 | 0.60 | 15.7 | 5.0 | 0.1 | 2.5 | 5.3 | 3.6 | 1.7 | 0.10 |
| 35107 | 63.0 | 1.10 | 16.2 | 5.8 | 0.1 | 1.8 | 4.2 | 5.5 | 1.9 | 0.33 |

Run 617A, X_{H20}=0.25, T=850°C, Residual Mineralogy: 44.5% Glass, 6.6% Qtz, 31.7% An₄₄ Plag, 8.8% Cpx, 8.3% Opx, 0.2% Ilmn

| | SiO2 | TiO2 | A1203 | FeO* | MnO | MgO | CaO | Na 20 | K2O | P205 |
|------------|------|------|-------|------|------|------|-----|-------|------------|------|
| Glass 617A | 72.0 | 0.29 | 15.3 | 2.1 | 0.05 | 0.49 | 2.6 | 3.6 | 3.5 | 0.10 |
| Avg BT | 74.6 | 0.15 | 13.9 | 1.6 | 0.06 | 0.05 | 0.7 | 5.2 | 3.5 | 0.02 |

| | | Model | #1, | distril | oution | coeff: | icients | 5 | | |
|---------|----------|-------|------|---------|--------|--------|---------|------|------|------|
| | Restite | Rb | Sr | Sc | V | La | Sm | Eu | Tb | Lu |
| Plag | 0.65 | 0.3 | 4.4 | | | 0.3 | 0.1 | 2.5 | 0.2 | 0.1 |
| Срх | 0.18 | | 0.4 | 33 | 3 | 0.5 | 2.9 | 2.5 | 3.8 | 3.0 |
| Opx | 0.17 | | | 7 | 1 | 0.4 | 0.8 | 0.6 | 0.9 | 0.9 |
| Ap | 0 | | 1 | | | 15.0 | 46.0 | 26.0 | 38.0 | 14.0 |
| Mgt | 0 | | | 5 | 40 | | | | | |
| Ilmn | 0.004 | | | 1 | 10 | | | | | |
| Bulk D | • | 0.20 | 2.93 | 7.13 | 0.75 | 0.4 | 0.7 | 2.2 | 0.9 | 0.8 |
| | | | | | | | | | | |
| Cs (3S) | 107) | 37 | 375 | 16.3 | 98 | 70.0 | 33.0 | 21.4 | 21.3 | 15.8 |
| Model : | #1 F=0.4 | 5 67 | 181 | 3.7 | 114 | 109.2 | 38.7 | 12.9 | 22.1 | 18.3 |
| C1 (3S | 032) | 86 | 52 | 4.3 | 3 | 81.1 | 24.2 | 8.6 | 16.6 | 16.4 |

| | | Model | <pre>lodel #2, distribution coefficients</pre> | | | | | | | |
|--|----------|-------|--|------|------|------|------|------|---------------|------|
| R | estite | Rb | Sr | Sc | V | La | Sm | Eu | \mathbf{Tb} | Lu |
| Plag | 0.64 | 0.3 | 4.4 | | | 0.3 | 0.1 | 2.5 | 0.2 | 0.1 |
| Срх | 0.16 | | 0.4 | 33 | 3 | 0.5 | 2.9 | 2.5 | 3.8 | 3.0 |
| Opx | 0.15 | | | 7 | 1 | 0.4 | 0.8 | 0.6 | 0.9 | 0.9 |
| Ap | 0.02 | | 1 | | | 15.0 | 46.0 | 26.0 | 38.0 | 14.0 |
| Mgt | 0.01 | | | 5 | 40 | | | | | |
| Iĺmn | 0.02 | | | 1 | 10 | | | | | |
| Bulk D: | | 0.19 | 2.90 | 6.40 | 1.23 | 0.6 | 1.6 | 2.6 | 1.6 | 1.0 |
| | | | | | | | | | | |
| Cs (3S10 | 7) | 37 | 375 | 16.3 | 98 | 70.0 | 33.0 | 21.4 | 21.3 | 15.8 |
| W | | | 102 | 4 1 | 07 | 00.0 | 25 0 | 11 2 | 16 0 | 16 0 |
| Model #2 | , F=0.45 | 6/ | 183 | 4.1 | 87 | 88.0 | 25.0 | 11.3 | 10.0 | 10.2 |
| CI (3S03 | 2) | 86 | 52 | 4.3 | 5 | 81.1 | 24.2 | 8.6 | 10.0 | 10.4 |
| Model #3 came regidual mineralogy as in #2 | | | | | | | | | | |
| Ca estim | atod. | A7 | 105 | 17 | 5 | 64.5 | 32.0 | 16.0 | 22 0 | 16.0 |
| CB CBCIM | aceu. | -27 | 100 | 11 | 5 | 0410 | 52.0 | 10.0 | 22.0 | *0.0 |
| Model #3 | F=0.45 | 85 | 51 | 4.3 | 4 | 81.1 | 24.2 | 8.4 | 16.5 | 16.4 |
| C1 (3S03 | 2) | 86 | 52 | 4.3 | 3 | 81.1 | 24.2 | 8.6 | 16.6 | 16.4 |

Table 9.9: Partial melt models for Broken Top rhyolite petrogenesis, using experimental data on melts of Taupo volcanic zone dacite from Conrad et al., 1988. Average TVZ dacite used in melting experiments is broadly similar to dacite 3S107, but is more magnesian and has a higher Ca/Na than High Cascades dacites. Melting experiment 617A (850°C, 45% melt) was water-undersaturated (≈4 wt% H₂O) and resulted in an anhydrous restite and a rhyolitic melt. Differences between experimental melt and Broken Top rhyolite compositions are probably a function of differences in source composition and high water content of the partial melts stabilizing more calcic plagioclase. Trace element models show that the experimental restite is unlikely to have produced Broken Top rhyolite compositions. The REE's in model #1 are too high, and probably indicate residual apatite (model #2). Model #3 uses estimated trace element abundances that result in a Broken Top rhyolite composition, using an experimentally-constrained melt fraction and residual mineralogy.

Conrad et al., 1988. Sr abundances in the modeled melt are extremely high, even when rhyolitic distribution coefficients are used in the models (table 9.9). Major element distributions indicate that if partial melting occurred at lower water contents, the residuum should contain a more calcic plagioclase, which will have a lower distribution coefficient for Sr and thus higher Sr contents in the melt. The source of Broken Top rhyolites was limited to ~100 ppm Sr (model #3), if ~60% of the residuum contains plagioclase. Apatite must also have been a residual mineral, to produce the observed La/Sm and REE patterns in the Broken Top rhyolites (model #2, table 9.9). In addition, ilmenite \pm magnetite must be more abundant (i.e., >0.2%) in the residuum to account for the low TiO₂, FeO* and V contents in the Broken Top rhyolites. Even when the effects of lower water content and increased abundances of minor mineral phases are accounted for, it is highly unlikely that partial melts of dacitic rocks could have produced the Broken Top rhyolites. In general, the estimated source composition (model #3, Table 9.9) needs to be somewhat more evolved than typical High Cascades dacites.

The climactic eruption of Mt. Mazama ejected medium-grained tonalitic blocks as large as 4 m, which contain 1-5 mm thick veins of rhyolitic glass (Bacon and Druitt, 1988). The tonalitic blocks are thought to represent fragments of the deep walls of the Mazama magma chamber, which had remained at high temperatures (≤ 950 °C) for sufficient time to produce rhyolitic partial melts (Druitt and Bacon, 1989). The compositions of these rhyolitic partial melts are very similar to the composition of Broken Top rhyolites (table 9.10 and figure 9.5). Silicic rocks at Mt. Mazama are thought to have been produced mainly through crystal fractionation of High Cascades andesite (Bacon and Druitt, 1988), and are thus petrogenetically distinct from Broken dacites and rhyodacites. These tonalites may be analogous to older silicic plutonic rocks that potentially underlie this part of the Cascade Range, which may be correlative with plutonic rocks of the Klamath Mountains or related to older episodes of Cascades arc magmatism (Chapter 2-3).

The tonalite blocks had an original mineralogy dominated by plag + opx + cpx + mgt + ilmn, accessory ap + zr, and minor amounts of qtz + hb + bio + kfelds. Partial melting resulted in dissolution of hb + kfelds + bio (\pm plag + qtz), with a restite of relict plag + qtz and new kfelds + opx + cpx + mgt + ilmn (Druitt and Bacon, 1989). Recrystallized, euhedral Fe-Ti oxides in tonalite blocks with high melt fractions (\approx 10-35%) yield equilibration temperatures of 950-980°C, which suggests that melting occurred at very low water contents (Druitt and Bacon, 1989). Rhyolitic melts in the tonalite blocks likely represent dehydration melts that were produced at relatively high temperatures (\geq 950°C) and low pressures (\approx 2-5 kbars).

Although the Mazama rhyolites are chemically similar to the Broken Top rhyolites (table 9.10), there are significant differences between the compositions of these rhyolites that must be explained. The Mazama rhyolites are more enriched in K_2O , Rb, and Cs than possible through $\approx 50\%$ partial melting of a tonalitic source, which Bacon and Druitt attribute selective enrichment

| | Broken | Top | Mazama | : 1023 | Mazama | : 799 |
|----------------------|-------------|-------------|--------|--------|--------|-------|
| Sample | Avg | 1σ | CGR | Glass | CGR | Glass |
| SiO2 | 74.6 ± | 0.2 | 67.4 | 75.8 | 67.3 | 75.4 |
| TiO2 | 0.15 ± | 0.01 | 0.57 | 0.49 | 0.53 | 0.58 |
| A1203 | 13.9 ± | 0.1 | 15.9 | 11.9 | 16.2 | 12.2 |
| FeO* | $1.63 \pm$ | 0.05 | 3.28 | 2.02 | 3.24 | 2.14 |
| MnO | 0.06 ± | 0.00 | 0.05 | 0.04 | 0.06 | 0.03 |
| MgO | $0.05 \pm$ | 0.05 | 1.52 | 0.24 | 1.42 | 0.23 |
| CaO | 0.69 ± | 0.02 | 3.40 | 0.66 | 3.51 | 0.67 |
| Na2O | 5.18 ± | 0.18 | 5.13 | 4.03 | 5.15 | 3.91 |
| K20 | $3.49 \pm$ | 0.03 | 2.21 | 4.68 | 2.26 | 4.65 |
| P205 | 0.02 ± | 0.00 | 0.13 | 0.04 | 0.13 | 0.07 |
| Rb | 84 ± | 2 | 41 | 117 | 47 | 113 |
| Ba | 796 ± | 11 | 671 | 567 | 674 | 547 |
| Sr | 50 ± | 3 | 390 | 54 | 379 | 52 |
| Cs | $1.37 \pm$ | 0.27 | 2.50 | 7.70 | 3.00 | 7.90 |
| Sc | 4.30 ± | 0.18 | 9.32 | 7.48 | 8.95 | 9.03 |
| v | 3.25 ± | 3.56 | - | - | 64 | - |
| Co | 0.66 ± | 0.15 | 7.80 | 4.20 | 7.80 | 4.10 |
| Ni | $11.13 \pm$ | 0.93 | - | - | 13 | - |
| Cr | 0.50 ± | 0.66 | 13 | 2 | 13 | - |
| Zn | 79 ± | 11 | 38 | 40 | 57 | 50 |
| La | 23.2 ± | 2.8 | 17.0 | 22.0 | 19.0 | 28.0 |
| Ce | 52.53 ± | 8.78 | 25.0 | 46.0 | 38.0 | 63.0 |
| Nd | 20.36 ± | 2.34 | 19.0 | 22.0 | 22.0 | 29.0 |
| Sm | 4.42 ± | 0.44 | 4.20 | 5.60 | 4.90 | 7.40 |
| Eu | $0.61 \pm$ | 0.03 | 0.90 | 0.50 | 0.90 | 0.49 |
| Tb | 0.73 ± | 0.08 | 0.51 | 0.83 | 0.63 | 1.03 |
| Yb | 3.3 ± | 0.2 | 2.0 | 3.1 | 2.1 | 4.1 |
| Lu | 0.50 ± | 0.03 | 0.33 | 0.45 | 0.32 | 0.61 |
| La/Sm | $3.32 \pm$ | 0.19 | 2.56 | 2.48 | 2.45 | 2.39 |
| ΣHREE | 46.72 ± | 3.39 | 30.59 | 46.35 | 33.31 | 60.32 |
| Eu/Eu* | 0.41 ± | 0.03 | 0.68 | 0.27 | 0.58 | 0.20 |
| Zr | 217 ± | 6 | 195 | 202 | 197 | 367 |
| Hf | 6.6 ± | 0.5 | 5.1 | 5.8 | 5.0 | 10.6 |
| Ga | 20 ± | 2 | - | - | 19 | |
| Y | 28 ± | 3 | 25 | 30 | 25 | 38 |
| Nb | $17.4 \pm$ | 0.7 | 5.0 | | 5.2 | - |
| Ta | $1.02 \pm$ | 0.03 | 0.43 | 0.70 | 0.46 | 0.95 |
| Th | 8.10 ± | 0.42 | 4.4 | 8.2 | 4.7 | 11.3 |
| U | 2.48 ± | 0.18 | 1.8 | 3.5 | 2.2 | 5.0 |
| | | | | | | |
| Estimated melt fract | tion (Th | $K_{-} = 0$ | 01) | 0.54 | | 0.42 |
| | | | / | | | |

Table 9.10: Comparison of average Broken Top rhyolite with glass compositions from partially melted tonalitic xenoliths (CGR) from the climactic eruption of Mt. Mazama (Bacon and Druitt, 1988). Rhyolitic glass from tonalite 1023 most closely resembles Broken Top rhyolite compositions. Relative depletions in Broken Top rhyolites of TiO₂, FeO and MgO can be modeled by adding 1.3% ilmn and 2.8% cpx to the Mazama restite. Large enrichments in K₂O, Rb and Cs are not consistent with incompatible element enrichments during partial melting, and probably indicate enrichment during hydrothermal alteration. Melt fractions calculated by assuming that Th was incompatible and immobile during hydrothermal alteration. Note how relatively small variations in melt fraction and starting composition between 799 and 1023 result in large REE variations in the melt.





Figure 9.5: Chondrite normalized REE plots comparing Mt. Mazama and Broken Top rhyolites.
A) Comparison of Mt. Mazama tonalitic xenoliths (799 and 1023) and rhyolitic partial melts from these xenoliths (799G and 1023G)(Bacon and Druitt, 1988). Melts occurred at water-undersaturated conditions at ≈950°C and 2-5 kbars, with a plag + px + fe-oxide + ap residuum. B) Comparison of Mazama rhyolite 1023G with Broken Top and Three Sisters rhyolites. Mazama rhyolite 1023G closely resembles Broken Top rhyolites; small variations in La/Sm and Eu/Eu* can be accounted for by small variations in the amount of residual apatite and in the anorthite content of residual plagioclase. Three Sisters rhyolites are petrogenetically distinct, and were produced through partial melts of basaltic rocks.

Δ

during alteration. Oxygen isotopic exchange between plagioclase and quartz in the tonalitic blocks indicates that hydrothermal alteration of these rocks preceded high temperature (>900°C) isotopic re-equilibration (Bacon et al., 1989). If melting first occurred along altered joints and selvages, it would be possible to selectively enrich the melt in relatively mobile K_2O , Rb and Cs. Assuming that Rb and K_2O behave incompatibly, partial melts (F=0.54) of tonalite 1023 should contain 76 ppm Rb and 4.01 wt.% K_2O , which corresponds to Broken Top rhyolite compositions (table 9.10). Additional depletions in TiO₂, FeO* and MgO in the Mazama rhyolite could be produced by the addition of 1.3% ilmn and 2.8% cpx in the restite.

Mazama rhyolites (esp. 1023G) contain REE and HFSE abundances that are very close to the composition of Broken Top rhyolites (figure 9.5a). Although 0.3% residual apatite is required to balance P_2O_5 , the slightly lower La/Sm in the Mazama rhyolite probably reflects a relatively lower amount of apatite in the tonalitic residuum. The slightly larger Eu/Eu* and lower Al_2O_3 + Na_2O in the Mazama rhyolite may also reflect a larger amount of sodic plagioclase in the restite.

9.2.3 Summary of Broken Top rhyolite petrogenesis

Broken Top rhyolites are spatially, temporally, mineralogically and compositionally distinct from Three Sisters rhyolites. Crystal fractionation of high Na₂O dacites is not a valid model for Broken Top rhyolite petrogenesis, and other fractionation models require the pre-existence of rhyolite to create the parental compositions. Trace element abundances also invalidate many petrogenetic models, in part because amphibole was not a significant fractionating phase. The Broken Top rhyolites are petrogenetically distinct from high silica melts at Medicine Lake and Mt. Mazama, which were produced essentially through crystal fractionation.

Partial melts of metabasaltic rocks do not result in Broken Top rhyolite trace element compositions. Although experimental data are scarce for partial melts of intermediate composition rocks at low water contents, dehydration melts of dacitic rocks do not apparently produce Broken Top rhyolite compositions. Enrichments in incompatible trace elements, and depletions of compatible trace elements, in the Broken Top rhyolites requires the melting of a source more evolved than typical High Cascade dacites.

Rhyolitic partial melts in tonalitic ejecta from the climactic eruption of Mt. Mazama are very similar to Broken Top rhyolite compositions. The Mazama rhyolites were apparently produced through $\approx 50\%$ dehydration melts of the tonalite at ≈ 950 °C and 2-5 kbars pressure. Geochemical differences between the Mazama and Broken Top rhyolites can be accounted for by alteration in the Mazama source rocks, and small variations in the residual mineralogy. The Mazama tonalites may be analogous to older silicic plutons that likely underlie this part of the central High Cascades. The Broken Top rhyolites could have been produced independently from the High Cascades magma system through partial melts of older silicic plutons that likely occur in the upper-to-middle(?) crust.

10. SUMMARY OF CENTRAL HIGH CASCADE PETROGENESIS

10.1 Petrogenesis of Volcanoes in the Three Sisters Area Since ≈600 ka

Continental arc magmatism in the Three Sisters region of the central High Cascades reflects a complex interrelationship between crystal fractionation, magma recharge, and the mixing of rhyolitic crustal melts. Although these relationships are certainly not unique to the High Cascades (cf., Hildreth and Moorbath, 1988), significant temporal and spatial variations in the expression of these petrogenetic processes have developed in only a 600 km² area of the Oregon Cascades arc since \approx 500 ka.

Previous studies of central Cascades mafic petrogenesis (Hughes, 1983, 1990; Priest et al., 1983; Hughes and Taylor, 1986; Conrey and Sherrod, 1988; Conrey, 1991) have generally recognized that rocks older than ≈ 5 Ma are often petrogenetically distinct from rocks younger than ≈ 2 Ma, which also may be distinct from rocks younger than ≈ 0.7 Ma. This study further supports the conclusions of previous researchers that there are no discernable spatial or temporal controls on mafic petrogenesis for rocks younger than ≈ 0.7 Ma within the Three Sisters region of the Oregon High Cascades.

Major and trace element variations for Three Sisters area mafic rocks clearly indicate that clinopyroxene and titanomagnetite are early fractionating phases, which is the general characteristic of relatively high pressure (7-10 kbars), water-undersaturated (≈ 1 wt.%) mafic systems. Although experimental data on phase relationships and trace element partitioning are available for anhydrous, 1 atm. mafic systems, there are only limited data for high pressure (7-10 kbars), water-undersaturated systems that most closely correspond to the High Cascades mafic system. Even with these limitations, there are certain conclusions on mafic petrogenesis reached in this study.

Crystal fractionation coupled with magma recharge was an important petrogenetic process for mafic rocks in the study area. In essence, any long-lived $(10^5-10^6 \text{ years})$ mafic magma system must have had a recharge rate equal to the crystallization + eruption ± assimilation rate, in order to maintain thermal and mass balance. It is reasonable to conclude that the High Cascades mafic system has been recharged with relatively primitive, mantle-derived melts, because:

1) Parental mafic compositions have been erupted throughout the last ≈ 700 k.y.

- The subduction-related processes that result in mafic arc magmatism have continued to operate during the last ≈700 k.y.
- 3) High Cascades mafic petrogenesis cannot be explained by a model involving closed system mafic fractionation or assimilation.

The primary effect of mafic recharge is to shift the evolving system back towards parental compositions, resulting in compatible element buffering with incompatible element enrichments. This feature is commonly observed in the central Cascades mafic system, and in many other magmatic

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arcs (cf., Conrey, 1990). It is also likely that recharge involved different types of primary basaltic melts, which may have been produced by different degrees of peridotite melting, mantle metasomatism, or compositional heterogeneities in the mantle source (e.g., Hughes, 1990).

Recharge rates that result in a steady-state magma system (i.e. recharge = crystallization + eruption rates) apparently limits differentiation in that system to roughly 58% SiO₂ (cf., Nielsen, 1988); magma systems that evolve andesite and dacite through crystal fractionation are probably closed to mafic recharge. Rocks with >58% SiO₂ are rarely observed in the High Cascades except at large, stratovolcanic centers; the most abundant rocks in the High Cascades, including the Plio-Pleistocene mafic platform, are basaltic andesite (e.g., Hughes and Taylor, 1986). This observation also supports the hypothesis that the High Cascades mafic magma system has remained open to large amounts of recharge through time. This mafic system most likely resided at deep crustal levels (35-45 km), and ponded at the crust-mantle boundary due to density contrasts (cf., Marsh, 1985; Hildreth and Moorbath, 1988).

A large, recharging, mafic magma system located in the deep crust has sufficient heat to generate partial melts of adjacent mafic crustal rocks. If 10-30% melting occurs at water-undersaturated conditions, then these melts are broadly rhyolitic. Many of the geochemical features of the High Cascades mafic system, which cannot be explained by recharge + fractionation, can be accounted for by periodically mixing 10-20 vol.% High Cascades rhyolite into the mafic magma system. Mixing also should have resulted in crystal fractionation, because the rhyolitic magma was likely cooler ($\approx900^{\circ}$ C) than the mafic magma (>1100°C).

In summary, a large mafic magma system probably extends along the length of the Oregon Cascades arc at the base of the crust. Parental basaltic magmas of this system were generated through subduction-induced melting of upper mantle peridotite, which was possibly metasomatized and undoubtedly heterogenous. Periodic recharge coupled with crystal fractionation of plag + ol + pyroxene + Fe-Ti oxides can explain many of the geochemical features of the High Cascades mafic system since \approx 700 ka. Other geochemical features are consistent with periodically mixing 10-20% rhyolitic crustal melts into the mafic magma system. Although it is possible that portions of this mafic system could remain isolated from the main system in both space and time, there is no evidence to suggest that any part of the central High Cascades mafic system has been petrogenetically distinct since \approx 700 ka. The petrogenetically distinct silicic magma systems found in this part of the High Cascades thus evolved from the same mafic magma system.

In contrast to the mafic system, spatial and temporal petrogenetic controls are observed for intermediate composition rocks in this part of the High Cascades. These controls are best expressed in the petrogenesis of the Three Sisters. North Sister, which is the oldest of the Three Sisters ($\leq \approx 350$ ka), is an essentially mafic stratovolcano; the most evolved compositions at North Sister are several small dikes of low silica andesite (E.M. Taylor, pers. comm., 1987). At Middle

Sister (younger than NS, ≥ 100 ka.), intermediate rocks were produced almost exclusively through the fractionation of plag + px + Fe-Ti oxides \pm ol \pm ap, resulting in elevated Na₂O and REE's to 66% SiO₂. In addition, the evolving system was essentially closed to mafic recharge, or else fractionation would not have proceeded beyond ≈58% SiO₂. It is not clear why fractionation was limited to 66% SiO₂ at Middle Sister. Other High Cascade volcanoes, such as Mt. Mazama and Medicine Lake, evolved >66% SiO₂ magmas through crystal fractionation and minor amounts of silicic wall-rock assimilation from similar mafic systems. A lack of evolution above 66% SiO₂ may indicate that the Middle Sister magma system was smaller and more closed to recharge than these other systems. If the Middle Sister system was smaller, then magma volumes could have been insufficient to maintain a fractionating chamber beyond 65% SiO₂ before conductive heat loss resulted in $\approx 50\%$ crystallization (i.e. a non-contiguous magma (cf., Bergantz, 1990)). Small amounts of rhyolite mixing (i.e., Mazama) or amphibole fractionation (i.e. Medicine Lake), neither of which occurred during Middle Sister petrogenesis, may also be a prerequisite to differentiation above 66% SiO₂. It is important to note that rhyolitic magmas were present < 10 km to the southeast in the contemporaneous (~380-200 ka) Broken Top system, but these rhyolites did not interact with either the North or Middle Sister magma system. The progression from NS to MS probably reflects the temporal evolution of a relatively closed magma system from essentially mafic (NS) to more andesitic and dacitic compositions (MS) through the fractionation of anhydrous minerals (figure 10.1).

At ≈100 ka, volcanism had shifted ≈5 km south of Middle Sister to South Sister, and the first rhyolites were erupted in the Three Sisters system. In contrast to the fractionation origin of the Middle Sister intermediate series, the petrogenesis of South Sister is characterized by the mixing of $\geq \approx 40\%$ rhyolite into an essentially mafic (58%-60%?) system to produce intermediate composition rocks to 66% SiO₂; mixing 30% rhyolite with fractionation-produced dacite would have resulted in compositions more silicic than 66% SiO₂, which are not observed at South Sister. If dacitic magmas produced through crystal fractionation (i.e. MS) were present in the Three Sisters system at ≈ 100 ka, they did not mix with rhyolitic magma to produce rocks with >66% SiO₂, nor were they erupted at South Sister. The petrogenetic changes that occurred at ≈100 ka between Middle and South Sister were not caused exclusively by the addition of rhyolitic partial melts into a Middle Sister-type fractionating magma system; the main magma system was re-established ≈ 5 km to the south, became more mafic, and generated rhyolitic magmas through dehydration melting of metabasaltic crustal rocks. These features are consistent with the emplacement of a new, essentially mafic system into the thermally-mature middle crust beneath South Sister at ≈ 100 ka (figure 10.1). It is speculated that the South Sister mafic system was probably quite similar to the older North Sister mafic system, but that rhyolitic partial melts were produced at South Sister because the mafic system was emplaced into crust that was relatively hotter due to Middle Sister magmatism. South



Figure 10.1: Schematic diagram of petrogenetic development in the Three Sisters region of the Oregon High Cascades. Arrows indicate direction of mass and heat transfer, solid lines are eruption paths, and boxes represent fractionating magma bodies. Vertical and horizontal distances are schematic and do not imply actual positions of these magma systems.

Sister can be thought of as a primarily bimodal system, with intermediate composition magmatism controlled by magma mixing and crystal fractionation. With increasing time, more evolved intermediate compositions will likely become available to mix with the rhyolite at South Sister, resulting in dacitic-to-rhyodacitic magmas and a reduction in the observed 66%-72% compositional gap (i.e., ≈Broken Top system).

Crystal fractionation to intermediate compositions ($\approx 65\%$ SiO₂) combined with $\approx 30\%$ rhyolite mixing resulted in dacitic-to-rhyodacitic magmatism in the Broken Top system from ≈ 500 ka to at least ≈ 200 ka. In contrast to Middle Sister, only a few Broken Top rocks that evolved to 65% SiO₂ through anhydrous crystal fractionation were erupted; differentiation above 63% SiO₂ generally required the mixing of $\approx 30\%$ rhyolite into the magma system. In contrast to South Sister, dacitic magmas were available for mixing with 30-35\% rhyolite, resulting in the evolution of rhyodacite and only a small ($\approx 2\%$ SiO₂) compositional gap in the Broken Top system (figure 10.1)

Rhyolites from the Three Sisters and Broken Top systems were produced from partial melts of different crustal compositions. Observed major and trace element abundances in both of these rhyolites are not consistent with evolution from intermediate compositions through crystal fractionation. Three Sisters rhyolites most closely resemble dehydration melts of metabasaltic rocks, with $\approx 20\%$ melt produced at ≈ 3.7 kbars and ≈ 900 °C. Although intra-suite variations (72-76% SiO₂) can be accounted for through 6-15% fractionation of modal minerals, inter-suite variations between Holocene and Pleistocene rhyolites cannot be related through crystal fractionation or magma mixing. Holocene rhyolites were generated from the same type of metabasaltic crustal rocks as Pleistocene rhyolites, but melting probably occurred at slightly higher temperatures and resulted in a larger plagioclase component in the Holocene melts. The compositional differences between the Holocene and Pleistocene rhyolites may thus reflect an increase in crustal temperature with time due to continued magmatism beneath South Sister.

In contrast to the Three Sisters rhyolites, Broken Top rhyolites were likely produced through partial melts of tonalitic crustal rocks, and do not resemble dehydration melts of metabasaltic rocks, or melts of associated restites. Although there are only limited experimental data on the waterundersaturated melting of silicic crustal rocks, partial melts from tonalitic blocks in the climactic ejecta of Mt. Mazama closely resemble Broken Top rhyolites. The Mt. Mazama rhyolites were apparently produced through 40-50% dehydration melts of tonalite at \approx 950°C and \approx 2-5 kbars. The tonalites at Mt. Mazama were also produced through crystal fractionation, and could be analogous to older silicic intrusions that may underlie the Broken Top system.

Most, and possibly all, of the Broken Top rhyolite was associated with the Tumalo volcanic center and erupted at 380 ± 20 ka. The continuous eruption of rhyodacites and dacites from ≈ 600 ka to at least 200 ka in the Broken Top system, however, indicates that rhyolitic magmas must have been present in this system from at least ≈ 600 ka to ≈ 200 ka. Rhyolitic volcanism in the Three

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Sisters system, however, is constrained to the last ≈ 100 k.y. There may thus have been a ≈ 100 k.y. hiatus between Broken Top and Three Sisters rhyolitic volcanism, although more age data are needed to evaluate this relationship. Stratigraphic and K-Ar age data in this study clearly demonstrate that petrogenetically distinct rhyolites can be produced, and can control the development of distinct intermediate composition magma systems, in contemporaneous and adjacent parts of a 600 km² area of the Oregon High Cascades.

10.2 The Occurrence of Rhyolite in the Oregon High Cascades

Rhyolitic volcanism of any significant magnitude or duration in the Oregon High Cascades is restricted to the Three Sisters region. No rocks with >68% SiO₂ have been reported from Mt. Hood (figure 10.2) (Wise, 1969; Priest and Vogt, 1982). Rhyolite at Mt. Jefferson is restricted to several small ($<1 \text{ km}^3$) domes; partial crustal melts at Mt. Jefferson are primarily rhyodacitic (Conrey, 1991). Mt. Mazama produced large amounts of rhyodacite through crystal fractionation and minor crustal assimilation (Bacon and Druitt, 1988), but rhyolite is restricted at Mt. Mazama to small melt zones in tonalitic xenoliths.

Appreciable volumes of rhyolite have been erupted from Newberry volcano since ≈500 ka (MacLeod and Sherrod, 1988). However, Newberry is petrogenetically distinct from other major volcanoes in the Oregon High Cascades, and is not strictly a High Cascade volcano (Waters, 1962; Higgins, 1973; Goles and Lambert, 1990):

- Newberry is predominantly a bimodal shield volcano, and lacks the abundant intermediate composition rocks that characterize High Cascade stratovolcanoes (cf., Beyer, 1973; Higgins, 1973; MacLeod et al., 1982; MacLeod and Sherrod, 1988).
- 2) Newberry rhyolites are petrologically similar to a series of rhyolitic domes and flows that extend ≈300 km to the southeast (MacLeod et al., 1975; MacLeod and Sherrod, 1988; Goles and Lambert, 1990; Draper, 1991). These rhyolites become progressively younger as they approach Newberry, and are thought to be associated with the Brothers Fault Zone (Walker, 1974; MacLeod et al., 1975; Draper, 1991). Volcanism associated with the Brothers Fault Zone since ≈10 Ma is probably related to Basin & Range extension and not subduction (Carlson and Hart, 1987; Draper, 1991).
- 3) Newberry is located at the convergence of 3 major fault zones (figure 10.3), in an area of thinner crust than the rest of the Oregon High Cascades (37 km vs. 44 km; Fitterman, 1988; Leaver et al., 1984). Mafic eruption rates at Newberry are appreciably higher than in the adjacent part of the High Cascades (Sherrod and Smith, 1990).

Newberry is located in an area that has apparently undergone greater amounts of Pleistocene crustal extension than any part of the Oregon High Cascades. Even if Newberry has a petrogenetic component related to subduction, Newberry petrogenesis must have been strongly influenced by the



Figure 10.2: Distribution of silicic Quaternary vents in the Oregon High Cascades, and chondritenormalized REE patterns for the most evolved compositions at each volcano. See text for data sources. Rhyolite in the High Cascades is restricted to the Three Sisters area, where two distinct types of rhyolite are erupted. Rhyolite at Newberry is not directly related to High Cascade volcanism, and is probably influenced by Basin and Range extension.



Figure 10.3: Faults and vent alignments of central Oregon, modified from MacLeod and Sherrod, 1988. Approximate boundaries of the regional gravity gradient from Thiruvathukal et al., 1970. FZ = Fault Zone. Note the change in structural trend from north to northwest that occurs east of the Three Sisters.

relatively higher heat flow associated with Basin and Range-style crustal thinning and increased mafic volcanism (cf., MacLeod and Sherrod, 1988). Although the relationships between crustal extension and petrogenesis are poorly constrained, intuitively an increase in crustal heat flow should favor the production of rhyolitic crustal melts (cf., Marsh, 1985; Bergantz, 1989).

10.3 Tectonic Controls on Continental Arc Petrogenesis

Arc-length variations in the petrogenesis of continental calc-alkaline volcanic rocks have been reported for the Cascades (Hughes et al., 1980), Aleutians (Kay et al., 1982), Central America (Carr et al., 1982) and the Andes (Thorpe et al., 1982; Hildreth and Moorbath, 1988). These studies have shown that variations in subduction rate, angles of plate convergence, age of volcanism, crustal thickness and composition can strongly influence the petrogenesis of continental arc rocks (i.e., Arculus and Johnson, 1978).

The Oregon High Cascades south of Mt. Hood are thought to represent a single tectonic segment of the High Cascades magmatic arc (Guffanti and Weaver, 1988). Although crustal thickness remains essentially constant at 44 km along the length of this segment (Leaver et al., 1984), it is likely that a small change in crustal composition or age occurs between Mt. Jefferson and the Three Sisters area (Chapter 2). Two rhyodacites ($\approx 70\%$ SiO₂) at Mt. Jefferson, which were produced through partial melts of mafic crustal rocks, have 87 Sr/ 86 Sr ≤ 0.7034 (Conrey, 1991). These rhyodacites are petrogenetically similar to rhyolites in the Three Sisters, which were produced through dehydration melts of metabasalt and have 87 Sr/ 86 Sr = 0.7036 (table 3.3). These data support the argument that metabasaltic crust in the northern part of this arc segment is potentially younger (early Cenozoic to late Mesozoic) than metabasaltic crust in the southern part of the segment (early Mesozoic to late Paleozoic) (e.g., Riddihough et al., 1986). However, the isotopic (and presumably geochemical) differences between these crustal types are relatively small, compared to crustal variations reported in other studies (i.e. ${}^{87}\text{Sr}/{}^{86}\text{Sr} = 0.7036-0.7057$, Hildreth and Moorbath, 1988). In addition, rhyodacites and granitoid xenoliths at Mt. Mazama also have ⁸⁷Sr/⁸⁶Sr = 0.7037 (Bacon et al., 1989), yet Mt. Mazama did not erupt rhyolite. It is unlikely that the petrogenetic differences between the Three Sisters area and other Oregon High Cascade volcanoes originated from these small presumed changes in crustal age or composition.

The Three Sisters region apparently is located on crust of similar composition and thickness as the rest of the Oregon Cascades arc segment. There is no evidence to indicate that subduction rates or convergence angles were different in the Three Sisters area than for the rest of this arc segment. Mafic extrusion rates have also remained relatively constant within this arc segment since ≈ 2 Ma (Sherrod and Smith, 1990). In spite of these similarities, the Three Sisters area is petrogenetically distinct from the other stratovolcanic centers in this segment of the Cascades arc. Previous models involving large-scale tectonic or crustal controls on the petrogenesis of continental arc volcanism cannot explain the petrologic features of the Three Sisters area. In addition, differences between the Broken Top and Three Sisters systems have developed contemporaneously from the same mafic magma system, in the same 600 km² area of the Cascades arc. Clearly, these features cannot be the result of regional tectonic or crustal heterogeneities.

10.4 A Paradigm of Crustal Stress and Igneous Petrogenesis

The Oregon High Cascades are located in an area of regional east-west extension (Zoback and Zoback, 1980), which is in part related to the dynamics of oblique subduction of the Juan de Fuca plate (Wells et al., 1984; Rodgers, 1985; Pezzopane and Weldon, 1990). North-south alignments of Quaternary vents in the Three Sisters area parallel the strike of older (≤ 5 Ma) normal faults, especially north and west of the Three Sisters (figure 10.3; cf., Bacon, 1985, figure 2). These orientations reflect a Quaternary east-west extensional stress regime (cf., Nakamura, 1977). East and southeast of the Three Sisters, vent alignments strike to the northwest and parallel the trend of the Tumalo Fault Zone (figure 10.3), which was active in the Quaternary (Peterson et al., 1976; Taylor, 1981).

It has long been recognized that the Brothers Fault Zone (BFZ) may intersect the Oregon High Cascades in the Three Sisters area, and that the BFZ is related to Basin and Range extension (Lawrence, 1976; Bacon, 1983, 1985; Guffanti and Weaver, 1988). The timing of BFZ impingement into the Cascades, however, is not clear. BFZ impingement might have influenced development of the Oregon Cascades graben at ≈5.4 Ma (Guffanti and Weaver, 1988), but there is evidence that the Cascades graben may not be a Basin and Range structure (Taylor, 1981; Smith, 1986; Smith et al., 1987). The Oregon Cascades graben developed well north of the Basin and Range province, and there is little evidence of major crustal extension in the intervening area (Smith et al., 1987). The location of the graben near the Basin and Range province may be coincidental; discontinuous intra-arc grabens can develop under a variety of tectonic or volcanic conditions (see summary by Smith, 1986), and their formation does not require back-arc extension. It is also possible that the formation of the Cascades graben may be related to differential crustal loading between the Miocene Cascades volcanic axis and the Deschutes Basin (Smith et al., 1989). In addition, if rhyolitic volcanism in southeastern Oregon was related to northwesterly progression of the BFZ (Walker, 1974; MacLeod et al., 1975), then ≈5 Ma volcanism was located ≈100 km east of Newberry volcano (figure 10.4). These data suggest that the BFZ did not impinge upon the Oregon Cascades arc until later than 5 Ma, and that the Oregon Cascades graben may not be directly related to Basin and Range extension (Smith, 1986; Smith et al., 1987).

Newberry volcano is thought to represent the termination of northwest-younging rhyolitic volcanism associated with the BFZ (Walker, 1974; MacLeod et al., 1975; MacLeod and Sherrod, 1988), based in part on the abundance of ≤ 1 Ma to Holocene rhyolitic eruptions located at



Figure 10.4: Age of rhyolitic domes and flows ≤10 Ma in eastern Oregon plotted against distance from McKay Butte (Newberry volcano), using age data from MacLeod et al., 1975, Fiebelkorn et al., 1983, and this study. The "Newberry" trend (triangles) may have separated from the main trend at ≈3 Ma, while the main trend continued on to Benham Butte and the Three Sisters area.

Newberry. MacLeod et al. (1975) concluded that the rate of northwesterly age progression changed at ≈ 5 Ma from ≈ 3 cm/yr to ≈ 1 cm/yr (figure 10.4), and that the lower rate terminated at Newberry volcano. However, the timing of this rate change may be as young as $\approx 2-3$ Ma, given the degree of variation (r=0.91) in the age-position data (figure 10.4). In addition, Fiebelkorn et al. (1983) report a K-Ar age of 1.8 ± 0.8 Ma on sanidine from Benham Butte (figure 10.3), which has a reversed paleomagnetic direction (i.e. >0.72 Ma) and a composition more similar to Newberry rhyolite than adjacent High Cascades rhyolite (Appendix D). Benham Butte is also located ≈ 30 km northwest of Newberry, and thus cannot simply be related to the younger "Newberry" trend. However, the linear extrapolation of a best-fit (r=0.91) line through the older (≈ 3 cm/yr) trend intersects the position of Benham Butte at its age (figure 10.4). It is speculated that the Newberry trend is superimposed upon, or a branch of, the older(?) ≈ 3 cm/yr trend, and that rhyolitic volcanism at Newberry is only in part related to the BFZ; either Cascade magmatism or extension associated with the Walker Rim-Tumalo Fault Zones could have influenced the development of Newberry volcano.

If the ≈ 3 cm/yr trend continued northwest of Newberry and on past Benham Butte, it would have intersected the Three Sisters area (Walker, 1974) at ≈ 500 ka (figure 10.4). This corresponds to the age of early Broken Top system silicic volcanism. Although this relationship may be coincidental, it probably indicates that rhyolitic volcanism in this area may be a direct result of more recent impingement of "BFZ-tectonics + volcanism" onto the Oregon High Cascades. This does not imply that the BFZ actually intersects the High Cascades or merges with the Tumalo Fault Zone, as the last motion on the BFZ probably occurred prior to ≈ 3 Ma (MacLeod et al., 1982; Hawkins et al., 1988). Instead, this relationship suggests that the regional tectonic forces that controlled faulting and volcanism along the BFZ could have begun to operate in this part of the High Cascades at ≈ 500 ka.

It is not clear whether the BFZ represents a zone of dextral shear related to Basin and Range extension (Lawrence, 1976) or transtensional motion linked to oblique subduction of the Juan de Fuca Plate (Gutmanis, 1989; Pezzopane and Weldon, 1990). In either case, the Three Sisters region is apparently located in a zone of greater crustal extension than occurs within 75 km to the north or south (figure 10.3). The western boundary of the central High Cascades is marked by north to north-northwest striking faults and vent alignments, which were produced through both older extensional events (i.e., \approx 5.4 Ma graben formation) and a regional east-west extensional stress field in the Quaternary (e.g., Zoback and Zoback, 1980; Smith et al., 1987). However, east and south of Broken Top volcano, faults and vent alignments strike to the northwest (figure 10.3). The northeast-to-northwest change in strike of the Walker Rim-Tumalo Fault Zones at Newberry volcano, coupled with essentially northern strike along and west of the High Cascades axis, must have resulted in an area of relatively greater crustal extension east of the Three Sisters in order to accommodate this shift from east-west to northeast-southwest extension.

The effects of crustal extension on the petrogenesis of calc-alkaline volcanic rocks are poorly constrained. Empirical (Nakamura, 1977; Hildreth, 1981) and theoretical (Shaw, 1980) models of the effects of crustal stress on magmatism generally conclude that extension favors the vertical ascent of mafic magma in the crust. Bimodal (i.e. lacking intermediate composition rocks) volcanism is often characteristic of extensional terranes (e.g., Smith, 1979; Hildreth, 1981). The production of rhyolitic crustal melts should be enhanced by increased mafic intrusion accompanying extension, although magma residence times in the crust are insufficient to develop intermediate compositions through fractionation or mixing. However, these general models were developed for volcanic systems in highly extended terranes, such as the Basin and Range. The High Cascades represent a slightly extensional terrane, and the proposed variations in crustal extension are small relative to the adjacent Basin and Range province.

10.5 A Tectonomagmatic Model of Three Sisters Area Petrogenesis

The above general concepts can be used to construct a hypothesis on the effects of smallscale changes in the magnitude of crustal extension. If a zone of northeast-southwest extension intersects a region of overall east-west extension, then the area of intersection is relatively more extended that surrounding area, regardless of the absolute amount of extension. Mafic intrusion should be somewhat favored in the more extended area, resulting in a higher potential for recharge of crustal magma chambers and additional thermal input into the crust. Elevated recharge rates and wall rock temperatures should favor the production of fractionating magma chambers, which would have larger volumes and less conductive heat loss and could thus differentiate to more evolved (i.e. intermediate) compositions. In addition, these factors would enhance the production of rhyolitic crustal melts, which could then interact with the fractionating magma chambers.

This hypothesis explains many of the observed temporal and spatial controls on petrogenesis in the Broken Top and Three Sisters systems. The Broken Top system is characterized from ≈ 500 ka to ≤ 200 ka by significant amounts of andesite to rhyodacite, which were produced through a combination of crystal fractionation and rhyolite mixing. However, the only large eruption (≥ 10 km³) of rhyolite in this system occurred at ≈ 380 ka, and was associated with the Tumalo volcanic center. Some geological event must have triggered this production of anomalously large volumes of rhyolitic crustal melt. The most efficient process would have been an increase in crustal heat-flow produced by mafic intrusion into the thermally mature middle(?) crust. It is thus speculated that the impingement of an extensional stress regime related to Brothers Fault Zone tectonics onto the central High Cascades magmatic arc at $\approx 500-400$ ka triggered the production of anomalously large volumes of Broken Top rhyolite. After the eruption of this rhyolite at ≈ 380 ka, the locus of Broken Top system volcanism shifted west of the Tumalo volcanic center to the Tam MacArthur Rim area. Eruption of rocks as silicic as rhyodacite continued in the Broken Top system to at least 200 ka, indicating that sufficient volumes of rhyolitic magma remained in the system to mix with a differentiating intermediate system.

Although North Sister is the oldest of the Three Sisters stratovolcanoes (\leq 340 ka), remnants of older mafic composite volcanoes (Little Brother, Sphinx Butte, The Wife) have been exposed beneath the Three Sisters. It is thus likely that mafic magmatism in the Three Sisters system occurred contemporaneously with silicic magmatism in the Broken Top system prior to \approx 340 ka. Three Sisters mafic magmatism became focused at North Sister around 340 ka, after the large rhyolite eruption in the Broken Top system. Differentiation at North Sister was limited to low silica andesite, probably due to a significant amount of recharge into the mafic magma system. Between \approx 340 and \approx 100 ka, the locus of Three Sisters volcanism shifted from North Sister \approx 2 km south to Middle Sister. It is possible that this shift was tectonically controlled, but there are no data that directly support this hypothesis. Andesite and dacite were produced at Middle Sister through crystal fractionation from mafic compositions, indicating that the magma system was essentially closed to mafic recharge. Migration of the magma system to shallower crustal depths (\approx 20-30 km?) might have effectively closed the system to mafic recharge. Although Middle Sister volcanism was probably contemporaneous with Broken Top volcanism, there is no evidence that rhyolite from the Broken Top system interacted with the North or Middle Sister systems.

At ≈ 100 ka, the locus of Three Sisters volcanism shifted ≈ 5 km south to South Sister. Although data are limited, stratigraphic relations suggest the shift from Middle to South Sister volcanism occurred over a short interval of time. Limited stratigraphic relationships also indicate that Broken Top system volcanism had essentially ceased by ≈ 100 ka. South Sister petrogenesis is characterized by the first eruptions of rhyolite in the Three Sisters system, and the mixing of rhyolite into a fractionating magma system that was more mafic than the Middle Sister system. The fractionating magma system associated with South Sister probably represents a new batch of mafic magma, which had likely segregated from the regional Cascades mafic system into the thermally mature middle(?) crust at ≈ 100 ka. The shift in High Cascade rhyolitic magmatism from the Broken Top system to South Sister is also consistent with northwesterly age progression of tectonically-influenced rhyolitic magmatism in this area. Holocene rhyolites at South Sister are petrogenetically distinct from Pleistocene South Sister rhyolites, and probably reflect a small increase in the temperature of crustal melting with continued volcanic activity at South Sister.

11. CONCLUSIONS

Petrogenetic models in this thesis have demonstrated that fractionation of plagioclase, pyroxene, Fe-Ti oxides \pm apatite (PPFA) can effectively produce intermediate composition magmas from typical High Cascade mafic compositions. In order to fractionate to intermediate compositions, the relative amount of recharge into the system must be lower than in an essentially mafic magma system; magma systems having recharge rates capable of sustaining a steady state system (i.e., recharge = fractionation + eruption) are generally restricted to $\leq 58\%$ SiO₂. Low recharge rates at Middle Sister thus enabled PPFA fractionation to proceed to intermediate compositions, but the degree of magma evolution at Middle Sister was limited to $\approx 66\%$ SiO₂ due to the relatively small size of the differentiating magma system.

PPFA fractionation can also explain some of the intermediate composition rocks ($\leq 65\%$ SiO₂) in the Broken Top system. However, most of the intermediate composition rocks in the Broken Top system were produced through mixing 20%-35% rhyolite into the fractionating magma system. The rhyolite in this system cannot be produced from less evolved compositions through crystal fractionation, and was likely produced through partial melts of tonalitic crustal rocks. Magma mixing coupled with PPFA fractionation in the Broken Top system also resulted in a small compositional gap from 72-74% SiO₂.

In contrast to the Broken Top system, rhyolite associated with the Three Sisters system was produced through partial melts of mafic amphibolite. Rhyolitic magma chambers also developed in the Three Sisters system, and evolved from 72% to 76% SiO₂ through PPFA + minor amphibole fractionation. Intermediate composition rocks at South Sister were produced by mixing $\geq 40\%$ rhyolitic magma into a PPFA-fractionating mafic system. In spite of the marked petrogenetic differences between Middle and South Sister, the degree of differentiation at South Sister also was limited to $\approx 66\%$ SiO₂. Magma mixing coupled with PPFA fractionation in the South Sister system also resulted in a compositional gap from 66-72% SiO₂.

Intermediate composition rocks at South Sister and in the Broken Top system reflect the degree of evolution in the less evolved magma that is mixed with rhyolite. At South Sister, the less evolved magma was essentially mafic (SiO₂ \approx 58-60%). Mixing (+ PPFA fractionation) between roughly equal amounts of essentially mafic and rhyolitic magma can only generate magmas with \leq 66% SiO₂. In contrast, the less evolved magma in the Broken Top system was essentially intermediate (SiO₂ \approx 60-65%). Mixing (+ PPFA fractionation) between intermediate and rhyolitic magma can result in compositions with \leq 72% SiO₂. Although intermediate composition magmas were present during the construction of Middle Sister, the lack of rhyolitic magma precluded differentiation beyond 66% SiO₂ through magma mixing (i.e., Broken Top system).

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The petrogenesis of the Three Sisters region of the Oregon High Cascades can be summarized in the following sequence:

- Volcanism from ≈600(?) to ≈350 ka occurred at the Tumalo volcanic center and Todd Lake volcano. Rhyolite was produced in the Broken Top system through the partial melting of tonalitic crustal rocks, with sufficient volumes of melt produced to erupt ≥10 km³ of rhyolite at ≈380 ka. Intermediate composition magmas were produced in the Broken Top system through the mixing of 20-35% rhyolite into a PPFA-fractionating magma system. Rhyodacites in the early Broken Top system also closely resemble rhyodacites in the Deschutes Formation, which suggests that similar petrogenetic processes occurred in the Pliocene Oregon Cascades (cf., Conrey, 1991).
- 2) Volcanism from ≈350 to ≈100 ka shifted to the west, and was localized at Tam MacArthur Rim and Broken Top, and at North-Middle(?) Sister. Petrogenesis at Tam MacArthur Rim and Broken Top volcano involved the same petrogenetic processes as earlier Broken Top system magmatism. However, petrogenesis in the contemporaneous Three Sisters system involved the development of an intermediate composition system (Middle Sister) from the Cascades mafic system (North Sister) through PPFA fractionation. Although rhyolitic magmas were present in the Broken Top system, these rhyolites did not interact with the Three Sisters system.
- 3) Volcanism younger than ≈100 ka was probably localized at South Sister. Rhyolites were first erupted in the Three Sisters system at this time, and were produced through partial melts of mafic amphibolites. In contrast to Middle Sister, South Sister petrogenesis was controlled by mixing ≥40% rhyolite into a magma system that was more mafic than at Middle Sister.

This study has clearly shown that profound changes in the petrogenesis of silicic rocks can occur within a relatively small area of a continental arc over a period of $\approx 10^5$ years. Previous models of continental arc petrogenesis have demonstrated the ability of marked or regional changes in tectonic setting, crustal thickness, and crustal composition to control regional variations in the petrogenesis of mafic and silicic volcanic rocks. However, these parameters did not change in the 20x30 km² area of the Three Sisters region, and thus could not have controlled petrogenesis in this area within the last ≈ 600 k.y.

The subtle influences of crustal extension on the petrogenesis of igneous rocks are poorly understood. Bimodal volcanism is often thought to be characteristic of moderately extended terranes such as the Basin and Range, and highly extended rifts such as the Rio Grande Rift erupt prodigious amounts of basaltic lava. But what about areas that are less extended than the Basin and Range? Most continental arcs, like the Oregon Cascades, are located in areas of modest crustal extension (i.e. σ_3 is perpendicular to the arc axis), and many contain discontinuous intra-arc grabens. Could small variations in the orientation or magnitude of crustal extension have influenced the petrogenesis of these arcs? Empirically, these variations in extension could effect the magma supply rate into the crust, which would control the crustal heat flux. Relatively small increases in extension could thus result in relatively small increases in the volume of mafic magma intruding into the crust. However, when these increases are superimposed on a previously steadystate magma system (e.g., North-Middle Sister), the resulting perturbations may result in significant petrogenetic shifts in the system. Increases in crustal heat may trigger significant amounts of crustal melting, and result in the establishment of rhyolitic magma chambers. The increase in heat would also favor the establishment of longer lived (i.e. more differentiated) magma chambers (e.g., Broken Top system), because the thermal contrast between the magma and wall rock would be lowered. Recharge by mafic magmas during extension could also shift a fractionating intermediate magma system back towards more mafic compositions (e.g., South Sister). All of these features are characteristic of petrogenesis in the Three Sisters region, as is the location of this region in an area of slightly greater crustal extension than the rest of the Oregon Cascades. The amount of petrogenetic diversity that occurred within this area since ≈ 600 ka requires the addition of small scale variations in crustal extension to the list of geological processes that can control continental arc petrogenesis.

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APPENDICES

APPENDIX A: ANALYTICAL METHODS

Instrumental Neutron Activation Analysis

Facilities at the Oregon State University Radiation Center were used to obtain trace element abundances through instrumental neutron activation analysis (INAA). Experimental design follows the procedures outlined in Laul (1979) and Hughes (1983). Two analytical runs were made for this study: 71 samples (3S001-3S080, 3S140-3S144) in the spring of 1987, and 65 samples (3S086-3S152) in the spring of 1988. All samples were prepared from the fresh interiors of field samples, crushed in an alumina-ceramic jaw-crusher to ≈0.5 cm, split, and ≈20 g powdered in a alumina-ceramic shatter box. In-house standards CRB-3 (equivalent to U.S.G.S. BCR-1) and CSG-1 (Champion Stock Granodiorite), and N.B.S. standard 1633-a (Fly ash) were used as standards. About 0.8 g of rock powder was sealed in 2/5d polyvials, which were then sealed into a 2d polyvial. The 2d polyvials were double-stacked in the rotating rack of the O.S.U. TRIGA reactor, and irradiated for 6 hours at 1 MW. Gamma-ray counts were obtained using high (≈20%) efficiency Ge(Li) detectors coupled with 4096-channel analyzers. Data storage and reduction was accomplished with DOSbased desk top computers. First counts of 4k to 10k seconds occurred 6 to 14 days after irradiation, and obtained Na, Ba, La, Sm, Yb, Lu, and U abundances. Long counts of 15k to 20k seconds occurred 30 to 45 days after irradiation, and measured Fe, Sc, Co, Ni, Rb, Sr, Cs, Ce, Nd, Eu, Tb, Zr, Hf, Ta and Th abundances.

Duplicate sample analyses were obtained in each analytical run, and are given in table A-1. The high precision of the INAA technique is shown by the small degree of variation shown in analyzing sample splits under different irradiation and counting conditions. Accuracy was tested by running a CRB-3 standard as an unknown; elemental abundances were generally within $\pm 5\%$ of accepted standard values.

X-Ray Fluorescence

X-ray fluorescence (XRF) analyses were accomplished during the summer of 1988 at Washington State University. Previously crushed samples (≈ 0.5 cm) were split, and ≈ 15 g were ground in a tungsten carbide (WC) shatter box for 2 minutes. 3.5000 ± 3 g of rock powder and 7.0000 ± 3 g of lithium tetraborate flux were combined and mechanically homogenized for 10 minutes. The powder was then fused in graphite crucibles at 1000°C for 35 minutes, cooled, reground and refused. The glass beads were then ground flat with 240 and 600 SiC grit, cleaned in an ultrasonic cleaner and washed in lab-grade alcohol. Glass beads were analyzed in a fully automated Rigaku 3370 X-ray spectrometer, with all analyses accomplished in one week-long run.

| | | | | | | | INAA | INAA | XRF |
|----------|-------------|------------------|-------|--------|-------|--------|-------|-------|--------|
| | XRF - | F INAA | INAA | INAA | XRF | Std | 1987 | 1988 | 1988 |
| | 35046 | 3S046B | 3S136 | 3S136B | 35136 | CRB-3* | CRB3 | CRB3 | CRB3 |
| SiO2 | 69.5 | 69.5 | | | | 54.35 | | | 55.1 |
| TiO2 | 0.55 | 0.55 | | | | 2.22 | | | 2.25 |
| A12O3 | 15.2 | 15.3 | | | | 13.63 | | | 13.8 |
| FeO* | 3.30 | 3.27 | 5.41 | 5.28 | 5.47 | 12.07 | 12.01 | 12.23 | 12.34 |
| MnO | 0.09 | 0.09 | | | | 0.18 | | | 0.22 |
| MgO | 0.54 | 0.51 | | | | 3.45 | | | 3.64 |
| CaO | 1.85 | 1.82 | | | | 6.95 | | | 7.16 |
| Na2O | 5.98 | 5.97 | 5.80 | 5.90 | 5.80 | 3.27 | 3.26 | 3.29 | 3.40 |
| K20 | 2.88 | 2.88 | | | | 1.69 | | | 1.74 |
| P2O5 | 0.11 | 0.11 | | | | 0.36 | | | 0.36 |
| TOTAL | 101.27 | 101.39 | | | | 98.17 | | | 100.00 |
| Rb | 60 | 60 | 20 | 30 | 28 | 47.1 | 35 | n/a | 48 |
| Ba | 740 | 725 | 610 | 612 | 583 | 678 | 626 | 705 | 634 |
| Sr | 182 | 182 | 360 | 390 | 392 | 330 | 310 | 320 | 326 |
| Cs | 1.90 | 1.80 | 0.95 | 0.82 | | 0.97 | 0.99 | 1.00 | |
| Sc | 9.45 | 9.31 | 15.75 | 14.86 | 15 | 32.8 | 33.04 | 32.1 | 28 |
| V | 1 | 7 | | | | 404 | | | 403 |
| Co | 2.89 | 2.89 | 7.12 | 6.59 | | 36.3 | 36.1 | 36.6 | |
| Ni | 8 | 8 | <40 | <60 | 4 | 13 | <50 | 30 | 3 |
| Cr | <4 | <4 | 3 | 4 | 0 | 16 | 14 | 13 | 29 |
| Zn | 89 | 86 | 191 | 275 | 95 | 129 | 130 | 136 | 148 |
| La | 25.9 | 25.8 | 22.0 | 21.6 | | 25 | 25.7 | 24.8 | |
| Ce | 55.1 | 54.6 | 53.2 | 48.1 | | 53.7 | 55.2 | 54.9 | |
| Nd | 29.5 | 28.6 | 26.0 | 26.2 | | 28.7 | 32.3 | 28.4 | |
| Sm | 6.30 | 6.37 | 6.27 | 6.35 | | 6.58 | 6.64 | 6.7 | |
| Eu | 1.36 | 1.31 | 1.82 | 1.75 | | 1.96 | 1.97 | 1.91 | |
| TD | 1.03 | 1.01 | 1.13 | 1.15 | | 1.05 | 1.01 | 1.1 | |
| YD | 4.0 | 4.0 | 3.2 | 3.3 | | 3.39 | 3.37 | 3.37 | |
| Lu | 0.64 | 0.61 | 0.49 | 0.49 | | 0.51 | 0.49 | 0.52 | |
| Zr | 307 | 307 | 130 | 180 | 203 | 131 | 150 | 190 | 173 |
| HI | 8.5 | 8.4 | 5.8 | 5.3 | | 4.9 | 5.1 | 5.2 | 0.1 |
| Ga | 20 | 18 | | | | | | | 21 |
| 1 hTh | 177 | 17 4 | | | | | | | 34 |
| C T a | 1/./ | 1/.4 | 0 50 | 0 50 | | 0 70 | 0 0 | 0 01 | 12.0 |
| Th | U.72 E 2 | U.94 E A | 0.39 | 0.58 | | 6.04 | 6 10 | 0.01 | |
| TT | 2.0 | 2.4 | 2.0 | 2.0 | | 1 71 | 1 7 | 1 6 | |
| 0 | 2.0 | ~ • ` | 1.4 | ±• ± | | 1./1 | ±•/ | 1.0 | |

*CRB-3 is equivalent to U.S.G.S. Standard BCR-1; values for BCR-1 from Flanagan (1976).

Table A-1: Replicate analyses for samples 3S046 (1987) and 3S136 (1988) using INAA. Sample splits were irradiated at different levels in the reactor, counted on different Ge(Li) detectors and reduced with different calibration standards. Note the high degree of precision indicated by these analyses. Accuracy for INAA and XRF shown by blind analysis of standard CRB-3 (i.e., BCR-1). Note the close correspondence between analyzed and reported values for CRB-3; actual abundance of FeO* in CRB-3 may be 12.25 wt. % (R. Walker, pers. comm., 1988).

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A 2:1 flux-to-rock ratio was used to increase the precision of trace element analyses. However, this ratio results in larger matrix corrections during XRF analysis, which may introduce a bias towards higher SiO_2 (+0.5 wt.%) and lower FeO*, CaO, and TiO₂ (-0.5 wt.% total) in W.S.U. analyses (Conrey, 1991). Precision and accuracy of the XRF analyses was evaluated through replicate analysis of sample 3S046, and analysis of standard CRB-3 as an unknown (table A-1). These analyses confirm the relative degree of precision and accuracy reported by Conrey (1991) for W.S.U. XRF analyses.

Electron Microprobe

Mineral compositions were determined with the fully automated Cameca SX-50 microprobe at O.S.U. during 1989-1990. Twenty one samples were selected for detailed microprobe analysis, which represented the major geochemical trends observed in the study area for basalt to rhyolite. In addition, mineral separates from 5 Quaternary pyroclastic units in the Bend area were analyzed with the microprobe.

Operating conditions of the microprobe were 15 kV accelerating voltage with a current of 20 nA. A beam width of 2μ m was used for all minerals except plagioclase, which used a 5μ m beam width. Calibrations used mineral standards, and were checked during each analytical run by successful analysis of standard reference minerals. Mineral compositions and structural formulae were calculated for all minerals except amphiboles and Fe-Ti oxides with standard Cameca software. Amphibole compositions were calculated using the program AMPREC.FOR by Spear and Kimball (1984), which calculates an average value of Fe³⁺ based on the structural limits of Fe³⁺ in the appropriate composition amphibole. Fe-Ti oxides were recalculated using the method of Carmichael (1967).

APPENDIX B: Visual estimates of phenocryst abundances and additional petrographic information for studied rocks. 0.1 = Trace amounts, other numbers = estimated percents. Phno: Total phenocryst abundance. An: Plagioclase An determined through the Michel-Levy method. Zon: Plagioclase zoning, W = weak (<10 An), M = moderate (10-20 An), S = strong (> 20 An), N = normal, R = reverse. Siv: % of plagioclase phenocrysts with coarse glass inclusions. Frt: % plagioclase phenocrysts with evidence of resorption or fine glass inclusions on rim. GM: Groundmass features: C = coarse (> 0.2 mm), M = medium (0.1-0.2 mm), F = fine (< 0.1 mm), X = holocrystalline, V, hyalocrystalline, G = glassy, D = devitrified. Other notes: Rsbd: resorbed; Gloms: Glomerocrysts

| SAMPLE | Plag | Срх | Орх | 01 | Hb | Opq | Ap Z | r | Phno | An | Zon | Siv | Frt | GM | Additional notes |
|--------|------|-----|------|-----|-----|-----|-------|-----|------|----|-----|-----|-----|----|---------------------------------|
| 3SXBP | 0.5 | - | 0.3 | | 0.1 | 0.1 | 0.1 0 |).1 | 1.0 | 22 | | 0 | | G | Rhyolite pumice |
| 35001 | 0.1 | | 0.1 | | | 0.1 | | | 0.1 | | | 0 | | G | Rhyodacite pumice |
| 35003 | 1.0 | 0.5 | 0.5 | 0.5 | | 0.5 | | | 3.0 | 46 | WN | 10 | | G | Andesite pumice |
| 35005 | 10.0 | 0.5 | 2.0 | | | 1.0 | 0.1 | | 13.6 | 45 | WN | 50 | | G | Rhyodacite pumcie |
| 35008 | 2.0 | 1.0 | | | | 1.0 | 0.1 | | 4.1 | 45 | WN | 10 | | MX | Gloms, coarse plag sieved |
| 35009 | 20.0 | 2.0 | 1.0 | | 0.1 | 1.0 | 0.1 | | 25.1 | 42 | MN | 5 | 1 | MD | Devitrified, cpx & opx oxidized |
| 3S010 | 7.0 | 2.0 | | | | 2.0 | | | 11.0 | 48 | MN | 50 | 5 | MV | Aphyr. rhyo(?) xenoliths |
| 3S010X | 5.0 | | | | | 1.0 | | | 6.0 | 30 | WN | 5 | | CD | Devit, FeMg's -> Opaques |
| 3S011 | 7.0 | 2.0 | | 1.0 | | 5.0 | | | 15.0 | 80 | SN | 50 | 10 | FX | Sieved=An35, Felty GM |
| 3S012 | 0.5 | | | | | | 0.1 | | 0.6 | 20 | WN | 0 | | FV | Devitrified |
| 3S015 | 20.0 | 2.0 | | 1.0 | | 2.0 | | | 25.0 | 46 | WN | 75 | | СХ | Sieved=An80 cpx rimmed |
| 3S017 | 3.0 | 0.5 | | 1.0 | | 0.5 | | | 5.0 | 68 | WN | 25 | 10 | FX | Pilotax; tr. ol Rimmed |
| 3S019 | 20.0 | 3.0 | 2.0 | | 0.1 | | 0.1 | | 25.2 | 45 | WN | 25 | | MX | Sieved=An30 |
| 3S022 | 15.0 | 0.5 | | 3.0 | | 0.5 | | | 19.0 | 68 | SN | 25 | 25 | FX | An68-30; Felty |
| 3S023 | 15.0 | 1.0 | 5.0 | | | 2.0 | 0.5 | | 23.5 | 22 | WR | 5 | 100 | CX | Wk Fritted Rims, Oxidized Opx |
| 3S027 | 20.0 | 0.1 | 0.1 | | | 0.1 | | | 20.3 | 22 | | 5 | | DG | Eutaxitic, Devitrified |
| 35028 | 0.5 | 0.1 | 0.1 | 0.1 | | 0.1 | | | 1.0 | 45 | WN | 10 | | G | Andesite pumice |
| 35029 | 3.0 | 0.5 | 1.0 | | | 0.5 | 0.1 0 | .1 | 5.2 | 22 | MN | 0 | | MX | Cpx rimmed |
| 35030 | 10.0 | 0.1 | 10.0 | | | 7.0 | | | 27.1 | 55 | SN | 0 | 100 | CX | Subophitic, Plg Rims |
| 3S031 | 5.0 | 3.0 | 0.1 | 2.0 | 0.1 | 2.0 | | | 12.2 | 48 | WN | 50 | 25 | CX | Opacitic Hb, Ol rimmed (opx?) |
| 35032 | 0.5 | | | | | 0.1 | | | 0.6 | 20 | WN | 0 | | FD | Devitrified |
| 3S033 | 3.0 | 0.1 | | 1.0 | | 0.5 | | | 4.6 | 52 | WN | 25 | 25 | FΧ | Ol rimmed (opx?) |
| 35034 | 1.0 | | | | | 0.1 | | | 1.1 | 12 | WN | 0 | | FD | Devitrified |
| 35035 | 1.0 | | | | | 0.1 | | | 1.1 | 20 | WN | 0 | | FD | Devitrified |
| 35036 | 1.0 | 0.5 | 0.1 | | | 0.5 | 0.1 | | 2.4 | 33 | MN | 10 | 50 | FV | Sanid, Qtx, Bio. An40-33 |
| 3S037 | 5.0 | 0.1 | 2.0 | | | 1.0 | 0.1 | | 8.2 | 22 | MN | 5 | | FX | An38-22 |
| 3S038 | 2.0 | 1.0 | | 0.1 | | 0.5 | | | 3.6 | 75 | WN | 50 | 10 | MX | Cpx Rimmed (opx?) |
| 3S040A | 7.0 | 0.1 | 1.0 | | | 1.0 | 0.1 | | 9.2 | 38 | WN | 25 | 25 | FX | An44-34 Siv, Embayed opx |
| 35042 | 1.0 | | 0.1 | | | 0.5 | 0.1 | | 1.7 | 15 | WN | 0 | | FV | Freshest TVC Rhyolite Dome |
| 35043 | 0.1 | | | | | | | | 0.1 | | | 0 | | G | Fresh obsidian |
| 35044 | 10.0 | | 0.5 | | | | | | 10.5 | 40 | WN | 0 | | G | Dacitic pumice |
| 35045 | 10.0 | 0.1 | | 1.0 | | 1.0 | | | 12.1 | 46 | SN | 10 | 40 | MX | An73-46, Skeletal ol |
| 35046 | 1.0 | | 0.1 | | | 0.5 | 0.1 | | 2.7 | 36 | WN | 50 | | FV | Coarsely sieved plag |

| SAMPLE | Plag | Срх | Орх | 01 | Hb | Opq | Ap | Zr | Phno | An | Zon | Siv | Frt | GM | Additional notes |
|--------|------|-----|-----|------|-----|------|-----|-----|------|----|-----|-----|-----|----|----------------------------------|
| 35048 | 0.5 | | 0.1 | | | | | | 0.5 | | | 50 | | G | Rhyodacite pumice, ≈3SO46 |
| 35049 | 10.0 | 1.0 | | 5.0 | | 5.0 | | | 21.0 | 49 | SN | 10 | | MX | Visible banding; An78-48 |
| 35050 | 10.0 | 1.0 | 0.5 | | | 1.0 | 0.1 | | 12.6 | 39 | MN | 50 | | FV | An44-39 |
| 35051 | 7.0 | 2.0 | 0.5 | | | 1.0 | 0.1 | | 10.6 | 35 | WN | 5 | | FV | Banded; Oscillatory Plag |
| 35052 | 7.0 | 3.0 | 1.0 | | | 0.1 | 0.5 | | 11.6 | 32 | SN | 50 | | MX | An55-32; Glomerocrysts |
| 35053 | 15.0 | 0.1 | 3.0 | | | 1.0 | 0.1 | | 19.2 | 36 | WN | 70 | | MV | Biotite Shreds in GM |
| 35054 | 5.0 | | 1.0 | | | 0.5 | 0.4 | | 6.9 | 22 | MN | 0 | | MV | An26-22; Glomerocrysts |
| 35057 | 15.0 | 0.1 | 2.0 | | | 1.0 | 0.1 | | 18.2 | 36 | WN | 90 | | MV | Hyalopilitic dome |
| 35058 | 10.0 | | 1.0 | | 0.1 | 0.5 | 0.5 | | 12.1 | 35 | WR | 10 | 10 | MV | An30-35; Oxy opx, Rsbd Plg |
| 35059 | 7.0 | 3.0 | 1.0 | | | 0.1 | 0.5 | | 11.6 | 32 | SN | 25 | 20 | MX | AN55-32; Gloms, Ol rims (opx?) |
| 35060 | 15.0 | | 3.0 | | 0.1 | 1.0 | 0.1 | | 19.2 | 20 | WN | 75 | 10 | FV | Tr. Biotite; Highly sieved plag |
| 35061 | 15.0 | 2.0 | 1.0 | | | 2.0 | 0.1 | | 20.1 | 24 | WN | 25 | | FV | Cpx Rimmed: Sieved=An41 |
| 35062 | 10.0 | | 1.0 | | | 0.5 | 0.1 | | 11.6 | 23 | WN | 25 | | MX | Biotite; An28-23; Gloms |
| 35063 | 25.0 | 5.0 | | | | 1.0 | | | 31.0 | 28 | WN | 5 | | FV | GM%'s: An34-28, Cpx rims (opx?) |
| 35064 | 25.0 | 2.0 | 2.0 | | | 5.0 | 2.0 | | 36.0 | 35 | WN | 50 | | CV | An35-30; Gloms; Zonation |
| 35065 | 25.0 | 2.0 | 2.0 | | | 5.0 | 2.0 | | 36.0 | 35 | WN | 50 | | CV | An35-30; Gloms; Zonation |
| 35066 | 10.0 | | | 10.0 | | 10.0 | | | 30.0 | 48 | MN | 10 | 90 | CX | Fritted Plg, resorbed Ol |
| 35067 | 25.0 | 7.0 | 2.0 | | | 5.0 | 0.1 | | 39.1 | 42 | WR | 50 | 5 | FV | An36-42; Glomerocrysts |
| 35069 | 0.1 | | | 10.0 | | 3.0 | | | 13.1 | 75 | WN | 0 | | CX | Embayed & Opacitic Ol |
| 35073 | 7.0 | 1.0 | 1.0 | | | 0.5 | | | 9.5 | 45 | WN | 25 | | FX | Phenos An35 |
| 35074 | 7.0 | 2.0 | | 1.0 | | 2.0 | | | 12.0 | 55 | MN | 50 | 25 | MV | An65-55; Sieved=An55-35; O1 Rims |
| 35075 | 3.0 | 0.1 | 0.5 | | | | | 0.1 | 3.7 | 20 | WN | 10 | | FX | Glomerocrysts |
| 35078 | 0.1 | | | | | | | | 0.1 | | | 0 | | FV | Aphyric, Hyalopilitic |
| 35079 | 0.5 | 0.1 | 0.1 | | | 0.1 | | | 0.5 | | | 0 | | G | Rhyodacite pumice |
| 35085 | 2.0 | 0.5 | 0.1 | | | 0.5 | 0.1 | | 3.2 | 45 | WN | 25 | | FX | Gloms, Lgst Plags sieved |
| 35090 | 20.0 | 1.0 | 1.0 | 3.0 | | 1.0 | | | 26.0 | 45 | WN | 75 | | CX | Sieved=An35-25 |
| 35091 | 2.0 | | | | | 0.1 | | | 2.1 | 35 | WN | 50 | | FV | Almost Aphyric |
| 35092 | 5.0 | 1.0 | | 0.5 | | 1.0 | | | 7.5 | 45 | WN | 50 | | FV | Sieved=An20; O1 w/ opx Rims |
| 35093 | 7.0 | 1.0 | | 3.0 | | 1.0 | | | 12.0 | 34 | MN | 90 | 10 | CX | An45-34; Rsbd Plag |
| 35094 | 1.0 | 0.5 | 0.1 | | | 0.1 | | | 1.0 | 45 | WN | 0 | | G | Rhyodacite pumice |
| 35095 | 5.0 | 2.0 | 0.1 | | | 1.0 | 0.1 | | 8.2 | 37 | WN | 0 | | FV | Moderately Vesicular |
| 35096 | 7.0 | 2.0 | 0.5 | | | 1.0 | 0.1 | | 10.6 | 33 | WN | 0 | | FV | Glomerocrysts |
| 35097 | 10.0 | 3.0 | 0.1 | 0.5 | | 1.0 | 0.1 | | 14.7 | 35 | MN | 25 | | FV | Sieved=An62; Oxidized Opx |
| 35098 | 7.0 | 0.1 | 1.0 | | | 0.5 | 0.1 | | 8.7 | 30 | WN | 25 | | FV | Highly pleochroic opx |
| 35099 | 7.0 | 0.5 | | 1.0 | | 0.1 | | | 8.6 | 64 | WN | 0 | | FX | Glomerocrysts |
| 35100 | 10.0 | 2.0 | 1.0 | | | 1.0 | 0.1 | | 14.1 | 24 | WN | 10 | | FV | An30-24; Glomerocrysts |
| 35101 | 5.0 | 1.0 | | 0.5 | | 0.5 | | | 7.0 | 42 | SN | 50 | 10 | FX | Sieved=An80 |
| 35102 | 1.0 | 0.5 | 0.1 | | | 0.1 | | | 1.7 | 80 | WR | 0 | | FX | Cognate Xenoliths |
| 35104 | 10.0 | 0.1 | 1.0 | | | 0.5 | 0.1 | 0.1 | 11.8 | 35 | WN | 10 | | FV | Devitrified; Zr incs Plg |
| 35105 | 20.0 | 1.0 | | 3.0 | | 2.0 | | | 26.0 | 40 | MN | 25 | | FX | An52-40, Skeletal Ol |
| 35106 | 20.0 | 2.0 | | 3.0 | | 1.0 | | | 26.0 | 52 | WN | 10 | | FX | An58-52; Glomerocrysts |
| SAMPLE | Plag | Срх | Орх | 01 | Hb | Opq | Ap | Zr | Phno | An | Zon | Siv | Frt | GM | Additional notes |
|--------|------|-----|-----|-----|-----|-----|-----|-----|------|----|-----|-----|-----|----|--------------------------------|
| 3S107 | 7.0 | | 1.0 | | | 0.5 | 0.1 | | 8.6 | 40 | WN | 5 | | FV | An45-40 |
| 3S108 | 7.0 | 2.0 | 1.0 | | | 1.0 | 0.5 | | 11.5 | 32 | MN | 5 | | FX | An44-32 |
| 3S109 | 2.0 | 0.5 | | 0.1 | | 0.1 | | | 2.7 | 60 | MN | 50 | | FX | AutoBxx; Ol Opacitic |
| 3S110 | 7.0 | 2.0 | 0.5 | | | 1.0 | 0.1 | | 10.6 | 35 | WN | 10 | | FV | Hyalopilitic |
| 3S111 | 10.0 | 3.0 | 1.0 | | | 2.0 | 0.1 | | 16.1 | 38 | WN | 5 | | MV | Glomerocrysts |
| 3S112 | 15.0 | 3.0 | 1.0 | | | 1.0 | | | 20.0 | 41 | MN | 10 | | CX | An58-41; Px Rimmed; Gloms |
| 3S113 | 15.0 | | 3.0 | | 0.1 | 1.0 | 0.1 | | 19.2 | 36 | WN | 0 | | FV | Oxidized Opx: Oscil Plg |
| 3S114 | 12.0 | | 2.0 | | | 0.5 | 0.1 | | 14.6 | 33 | MN | 90 | | FV | An33-20 |
| 3S115 | 7.0 | 2.0 | 0.5 | | | 1.0 | 0.1 | | 10.6 | 36 | WN | 10 | | CX | Cpx rimmed |
| 3S116 | 0.5 | | | | | 0.1 | | | 0.6 | 20 | WN | 0 | | FV | Devitrified |
| 3S117 | 0.5 | | | | | 0.1 | | | 0.6 | 20 | WN | 0 | | FV | Devitrified |
| 3S118 | 20.0 | | 2.0 | | | 0.5 | | | 22.5 | 27 | WN | 5 | | G | Slightly hydrated glass |
| 3S119 | 15.0 | 2.0 | 1.0 | | | 1.0 | 0.1 | | 19.1 | 45 | MR | 50 | | MV | An36-45: Mafic Clot: Gloms |
| 3\$120 | 20.0 | 3.0 | 1.0 | | | 3.0 | 0.1 | | 27.1 | 38 | SN | 10 | | FX | An70-38: Gloms: Opgs rimmed |
| 3S122 | 20.0 | | 3.0 | | | 0.5 | 0.1 | 0.1 | 23.7 | 26 | MN | 90 | | FV | Trace of Biotite in GM |
| 3S123 | 15.0 | 3.0 | 2.0 | 0.1 | | 0.5 | 0.1 | | 20.7 | 37 | WN | 10 | | FX | Ol Rims; Brecciated |
| 35124 | 15.0 | | 2.0 | | 0.5 | 2.0 | 0.1 | | 19.6 | 30 | MN | 10 | | FV | Glomerocrysts, Abund Hb |
| 3S125 | 10.0 | 2.0 | 1.0 | | | 1.0 | 0.1 | | 14.1 | 36 | WN | 5 | | FV | Glomerocrysts |
| 3S126 | 7.0 | | 1.0 | | | 0.5 | 0.1 | | 8.6 | 42 | WN | 10 | | FV | Weathered |
| 3S127 | 7.0 | | 2.0 | | 0.1 | 0.5 | 0.1 | | 9.7 | 33 | MN | 90 | | FV | An41-33 |
| 3S128 | 1.0 | 0.5 | 0.1 | | | 0.5 | 0.1 | | 2.2 | 45 | WN | 0 | | MX | Felty |
| 3S130 | 0.5 | 0.1 | 0.1 | | | 0.1 | 0.5 | | 0.3 | 40 | WN | 50 | 25 | FX | Tr Bio in plg, Fritted |
| 3S131 | 10.0 | 0.5 | | 5.0 | | 3.0 | 0.1 | | 18.6 | 45 | WN | 100 | | CX | Dikty; Ol Zoned, Wk Sievs |
| 3S132 | 1.0 | 0.5 | 0.1 | | | 0.5 | | | 2.1 | 38 | WN | 0 | | MV | Glassy; Mafic Xenoliths |
| 3S133 | 5.0 | | | 1.0 | | 1.0 | | | 7.0 | 60 | MN | 10 | | CV | An70-60; Seriate; Ol Rims |
| 3S134 | 10.0 | 1.0 | 3.0 | | | 2.0 | 0.2 | | 16.2 | 43 | WN | 5 | | MV | Cpx rimmed (opx?) |
| 3S135 | 5.0 | 1.0 | 1.0 | | | 1.0 | 0.1 | | 8.1 | 42 | WN | 0 | | FV | Cpx rimmed (opx?) |
| 3S136 | 7.0 | 1.0 | 3.0 | | | 1.0 | 0.5 | | 12.5 | 33 | WN | 0 | | FV | Brn AP; Gloms; cpx rims (opx?) |
| 3S137 | 7.0 | 0.5 | 1.0 | | | 2.0 | 0.2 | | 10.7 | 37 | MN | 5 | | FV | cpx rims, Brown apatite |
| 3S138 | 10.0 | 0.5 | 1.0 | | | 1.0 | 0.1 | | 12.6 | 44 | WN | 0 | | FV | cpx rims, Glomerocrysts |
| 3S139 | 0.1 | | | | 0.1 | | | | 0.1 | | WN | 0 | | G | Fresh Obsidian |
| 3S140 | 0.5 | | | | | 0.1 | | | 0.6 | 15 | WN | 10 | 10 | FV | Consertal Plg+Gls Xenos |
| 3S145 | 20.0 | 0.1 | 3.0 | | 0.1 | 1.0 | 0.1 | | 24.0 | 29 | WN | 90 | | FV | Complex oscillatory zoning |
| 3S146 | 7.0 | 2.0 | 0.5 | | | 1.0 | | | 10.5 | 33 | MN | 5 | | FX | Cpx rims, Low An plg |
| 3S147 | 7.0 | 1.0 | 1.0 | | | 2.0 | 0.1 | | 11.1 | 38 | WN | 5 | | CX | Cpx rimmed |
| 35148 | 5.0 | 0.5 | 1.0 | 0.1 | 0.5 | 0.1 | 0.1 | | 8.3 | 21 | MN | 50 | 25 | FV | Ol Xenos, V. complx Gloms |
| 35149 | 30.0 | 0.5 | 0.1 | 2.0 | | 1.0 | | | 33.6 | 48 | SN | 75 | 10 | CV | Cpx Rimmed, An70-48 |
| 3\$150 | 10.0 | 2.0 | 0.5 | 0.1 | | 0.5 | 0.1 | | 13.2 | 38 | SN | 50 | | MV | Hb Relicts, Mafic clots |
| 3\$151 | 10.0 | 0.5 | 1.0 | | | 1.0 | | | 12.5 | 42 | WN | 10 | | FV | Slight hydration |
| 3S152 | 3.0 | 0.5 | | 0.1 | | 0.5 | | | 4.1 | 34 | MN | 25 | 5 | CX | Skeletal Ol in GM |

Appendix C-1: Feldspar compositions determined through electron microprobe analysis. Mol percents anorthite (An), albite (An) and orthoclase (Or) calculated using standard CAM-ECA software. GM: groundmass crystal; glom: glomerocryst; acic: acicular crystal; siv: coarse glass inclusions (sieved); intmed: significant reversal in core to rim zoning.

| Sample | Si02 | A1203 | FeO | MaD | CaO | Na20 | 120 | Reff | TOTAL | Am. | Ab | Or |
|--------------------|---------|-------|------|------|--------|------|------|------|----------------|----------------|-------|-------|
| 3SXBP core | 62 51 | 23 52 | 0 19 | 0 00 | 1. 1.6 | 8 41 | 0 55 | 0 00 | 00 73 | 21 05 | 7/ 85 | 3 21 |
| SCADI COTC | 42 17 | 23.2/ | 0.20 | 0.00 | / 75 | 0.41 | 0.55 | 0.07 | 99.13 | 21.75 | 74.03 | 7 / 2 |
| JOADE TTHE | (2.1) | 23.24 | 0.20 | 0.01 | 4.33 | 0.70 | 0.60 | 0.09 | YY.JI | 20.90 | (2.00 | 3.42 |
| JOAD -:- | 02.14 | 23.39 | 0.25 | 0.02 | 4.40 | 0.42 | 0.58 | 0.09 | 99.30 | 21.95 | 74.09 | 3.30 |
| 35XBP F1m | 01.55 | 22.30 | 0.14 | 0.00 | 4.06 | 8.77 | 0.62 | 0.14 | 97.57 | 19.65 | 76.78 | 5.58 |
| SSUUSa core | 57.69 | 25.83 | 0.46 | 0.02 | 7.79 | 6.66 | 0.27 | 0.02 | 98.74 | 38.63 | 59.78 | 1.59 |
| 3S003a rim | 56.41 | 25.03 | 0.44 | 0.02 | 7.82 | 7.41 | 0.30 | 0.09 | 97.53 | 36.26 | 62.10 | 1.65 |
| 3S003b core | 56.85 | 25.83 | 0.49 | 0.01 | 7.91 | 6.63 | 0.28 | 0.05 | 98.05 | 39.10 | 59.25 | 1.65 |
| 3S003b rim | 59.10 | 26.95 | 0.42 | 0.02 | 7.84 | 6.51 | 0.27 | 0.04 | 101.14 | 39.32 | 59.10 | 1.59 |
| 3S005a core | 57.54 | 26.55 | 0.35 | 0.00 | 8.49 | 6.26 | 0.37 | 0.04 | 99.59 | 41.90 | 55.91 | 2.19 |
| 3S005a rim | 56.73 | 25.57 | 0.33 | 0.00 | 7.78 | 6.63 | 0.41 | 0.05 | 97.50 | 38.38 | 59.24 | 2.38 |
| 3S005b core | 56.57 | 26.71 | 0.38 | 0.01 | 8.62 | 6.24 | 0.37 | 0.03 | 98.93 | 42.38 | 55.47 | 2.16 |
| 3S005b rim | 55.30 | 25.99 | 0.35 | 0.01 | 8.75 | 6.78 | 0.37 | 0.02 | 97.56 | 40.77 | 57.18 | 2.05 |
| 3S009a core | 58.59 | 25.90 | 0.33 | 0.01 | 7.81 | 6.87 | 0.29 | 0.04 | 99.84 | 37.94 | 60.37 | 1.69 |
| 3S009a rim | 59.53 | 24.19 | 0.23 | 0.02 | 5.83 | 7.49 | 0.63 | 0.04 | 97.96 | 28 94 | 67 35 | 3 71 |
| 3S009b core | 57.99 | 26.01 | 0.26 | 0 01 | 7 67 | 6 59 | 0 31 | 0 03 | 98 87 | 38 41 | 50 75 | 1 84 |
| 35009b rim | 80.68 | 23 83 | 0 25 | 0 01 | 5 30 | 7 03 | 0 74 | 0 11 | 08 85 | 25 81 | 60 00 | 1.20 |
| 350179 0009 | /6 03 | 33 11 | 0.67 | 0.0/ | 16 // | 2 03 | 0.05 | 0.01 | 70.03 | 23.01 | 19 20 | 9.27 |
| 35017a core | 55 /0 | 24 52 | 0.07 | 0.04 | 0.44 | 2.0J | 0.03 | 0.01 | 99.20 | /0 11 | 10.20 | 2.01 |
| 35017b conc | 51 74 | 20.32 | 0.90 | 0.07 | 10 76 | 2.00 | 0.34 | 0.02 | 90.50 | 40.11 | 49.00 | 2.01 |
| | /4 7/ | 29.00 | 0.04 | 0.00 | 12.70 | 3.90 | 0.10 | 0.00 | 90.09 | 04.01 | 33.39 | 0.60 |
| 30017b F1m | 40.74 | 32.00 | 0.03 | 0.07 | 10.50 | 2.07 | 0.07 | 0.00 | 99.03 | 81.24 | 18.30 | 0.40 |
| SSUITC CORE | 49.05 | 51.49 | 0.70 | 0.06 | 15.02 | 2.80 | 0.06 | 0.00 | 99.17 | 74.52 | 25.13 | 0.35 |
| SSU1/c rim | 53.22 | 28.47 | 0.88 | 0.08 | 11.63 | 4.54 | 0.16 | 0.00 | 98.98 | 58.06 | 41.00 | 0.94 |
| 3S017d core | 46.50 | 33.32 | 0.59 | 0.04 | 16.70 | 1.88 | 0.05 | 0.01 | 99.08 | 82.86 | 16.83 | 0.32 |
| 3S017d rim | 45.82 | 33.71 | 0.56 | 0.05 | 17.23 | 1.56 | 0.03 | 0.02 | 98.98 | 85.73 | 14.07 | 0.20 |
| 3S017e core siv | 49.91 | 30.73 | 0.66 | 0.05 | 13.98 | 3.15 | 0.07 | 0.00 | 98.56 | 70.71 | 28.87 | 0.42 |
| 3S017e edge siv | 50.24 | 30.16 | 0.65 | 0.06 | 13.39 | 3.70 | 0.10 | 0.00 | 98.31 | 66.25 | 33.16 | 0.60 |
| 3S017e rim | 52.78 | 28.45 | 0.87 | 0.09 | 11.48 | 4.52 | 0.20 | 0.04 | 98.44 | 57.68 | 41.11 | 1.21 |
| 3S017f core siv | 46.61 | 33.56 | 0.52 | 0.07 | 16.85 | 1.68 | 0.05 | 0.00 | 99.33 | 84.52 | 15.20 | 0.28 |
| 3S017f edge siv | 46.18 | 33.18 | 0.65 | 0.06 | 16.70 | 1.97 | 0.04 | 0.00 | 98.78 | 82.22 | 17.57 | 0.22 |
| 3S017f rim | 54.54 | 23.12 | 4.14 | 2.53 | 9,12 | 4.39 | 0.43 | 0.05 | 98.31 | 51.92 | 45.20 | 2.88 |
| 3S017 30µm | 52.60 | 29.57 | 0.80 | 0.07 | 11.94 | 4.21 | 0.15 | 0.00 | 99.32 | 60.51 | 38.62 | 0.88 |
| 3SO17 50µm acic GM | 53.12 | 27.77 | 1 11 | 0 10 | 11 00 | 4 82 | 0 21 | 0 03 | 08 24 | 55 31 | 43 44 | 1 24 |
| 3S017 50µm acic G | 53 56 | 27 82 | 1 13 | 0 12 | 11 22 | 4.88 | 0 18 | 0.00 | 08 00 | 55 37 | 43 57 | 1 06 |
| 350289 0000 | 57 54 | 25 33 | 0.46 | 0.04 | 7 28 | 6.07 | 0.70 | 0.00 | 07 07 | 35 05 | 62 28 | 1 77 |
| 35020a core | 57 52 | 27.5 | 0.40 | 0.04 | 9 95 | 4 17 | 0.30 | 0.04 | 101 00 | /2 51 | 54 21 | 1 20 |
| 35020a 1 m | 57 / 1 | 27.43 | 0.45 | 0.07 | 7 93 | 6.47 | 0.23 | 0.00 | 09 /7 | 42.31 | 50.21 | 1.47 |
| 300200 0000 | 57.41 | 22.02 | 0.42 | 0.04 | 7.02 | 0.03 | 0.28 | 0.05 | 98.47 | 38.80 | 39.33 | 1.0/ |
| 300200 F1m | 27.39 | 22.02 | 0.34 | 0.06 | 1.31 | 7.49 | 0.30 | 0.05 | 98.00 | 34.04 | 03.0/ | 1.09 |
| SSUSUA core | 42.78 | 34.30 | 0.54 | 0.01 | 17.51 | 1.50 | 0.03 | 0.06 | 99.79 | 86.42 | 15.58 | 0.20 |
| SSUSUA FIM | 54.51 | 28.01 | 1.16 | 0.02 | 10.65 | 5.07 | 0.30 | 0.01 | 99.52 | 52.76 | 45.50 | 1.74 |
| 3SU30b core | 45.22 | 34.14 | 0.51 | 0.04 | 17.49 | 1.49 | 0.03 | 0.05 | 98.96 | 86.53 | 13.32 | 0.15 |
| 3S030b rim | 47.87 | 32.39 | 0.60 | 0.00 | 15.78 | 2.53 | 0.11 | 0.03 | 99.31 | 77.01 | 22.35 | 0.64 |
| 3S030c siv | 50.15 | 30.74 | 0.84 | 0.05 | 14.36 | 3.39 | 0.17 | 0.02 | 99.7 1 | 69.40 | 29.63 | 0.98 |
| 3S030c edge siv | 54.35 | 27.93 | 1.06 | 0.04 | 10.87 | 5.01 | 0.26 | 0.04 | 99.56 | 53.70 | 44.77 | 1.53 |
| 3S030c rim | 59.95 | 24.35 | 0.74 | 0.00 | 6.30 | 7.16 | 0.70 | 0.06 | 99.26 | 31.36 | 64.49 | 4.15 |
| 3S030 30µm GM | 55.31 | 27.62 | 0.84 | 0.00 | 10.10 | 5.17 | 0.38 | 0.03 | 99.45 | 50.76 | 46.96 | 2.28 |
| 3S030 50µm GM | 51.88 | 29.75 | 0.99 | 0.02 | 12.92 | 4.13 | 0.22 | 0.02 | 99.92 | 62.57 | 36.19 | 1.24 |
| 3S030 60µm acic GM | 1 54.30 | 27.89 | 0.95 | 0.01 | 10.85 | 4.79 | 0.29 | 0.00 | 99.07 | 54.65 | 43.62 | 1.73 |
| 3S036a core | 64.71 | 20.95 | 0.48 | 0.02 | 2.30 | 7.68 | 3.27 | 0.17 | 99.58 | 11.47 | 69.17 | 19.37 |
| 3S036a rim | 56.67 | 26.87 | 0.41 | 0.04 | 8.78 | 5.82 | 0.46 | 0.07 | 99.12 | 44.22 | 53.04 | 2.75 |
| 3S036b core | 57.97 | 26.34 | 0.37 | 0.05 | 7.96 | 6 45 | 0.56 | 0.06 | 99.75 | 39 24 | 57 50 | 3.26 |
| 3S036b rim | 58.40 | 25.77 | 0.34 | 0.03 | 7 38 | 6 42 | 0.58 | 0 07 | 08 08 | 37 50 | 59 02 | 3 48 |
| 350366 0058 | 47 25 | 23 82 | 0.52 | 0.05 | 0 30 | 5 21 | 0.20 | 0 01 | 86 45 | 48 76 | 40 43 | 1 80 |
| 35036c rim | 56 42 | 27 21 | 0.65 | 0.00 | 0 33 | 5 17 | 0.29 | 0.01 | 00.45 | 47.61 | 47 78 | 4 61 |
| 35036 50um acic G | 1 57 85 | 25 86 | 0.05 | 0.04 | 8.02 | 4 21 | 0.53 | 0.05 | 00 30 | 47.01 | 56 50 | 3 20 |
| | 1 50 44 | 2/ 72 | 0.70 | 0.01 | 6.02 | 6.00 | 0.00 | 0.04 | 77.JU 00 5/ | 40.30 32 ZE | 62 97 | 6 79 |
| | 47.00 | 10 40 | 0.11 | 0.00 | 0.01 | 0.99 | 1 70 | 0.07 | 99.34 | 21.00 | 02.0/ | 4./0 |
| 33030 30µm GM | 01.00 | 17.00 | 0.51 | 0.02 | 2.09 | 0.12 | 1.79 | 0.03 | 90.90 | 21.04 | 02.24 | 12.02 |
| 30000 /JM 8010 | 02.74 | 22.01 | 0.00 | 0.02 | 4.58 | 1.20 | 2.25 | 0.09 | 99.92 | 21.86 | 04.95 | 13.21 |
| SSU428 CORE | 02.80 | 22.58 | 0.15 | 0.00 | 5.87 | 8.67 | 0.67 | 0.09 | 98.80 | 19.00 | 11.09 | 5.90 |
| 55042a rim | 65.07 | 22.85 | 0.26 | 0.00 | 4.09 | 8.49 | 0.61 | 0.10 | 99.48 | 20.28 | 76.11 | 3.61 |
| 3S042b core | 61.77 | 23.59 | 0.20 | 0.00 | 5.00 | 8.06 | 0.57 | 0.04 | 99.23 | 24.67 | 71.97 | 3.37 |
| 3SO42b rim | 62.80 | 22.36 | 0.20 | 0.00 | 3.80 | 8.71 | 0.65 | 0.11 | 98.63 | 18.68 | 77.53 | 3.79 |
| 3SO42 30µm acic | 64.88 | 21.24 | 0.43 | 0.00 | 2.28 | 8.61 | 1.33 | 0.31 | 99.07 | 11.71 | 80.15 | 8.14 |

| Sample | Si02 | AL203 | Fe0 | HgO | CaO | Na20 | K20 | BaO | TOTAL | An | Ab | Or |
|----------------------------------|----------------|-------|------|------|-------|------|------|------|----------------|----------------|-------|------|
| 3S042 50µm acic GM | 66.09 | 20.85 | 0.21 | 0.00 | 1.94 | 9.20 | 1.30 | 0.22 | 99.82 | 9.65 | 82.68 | 7.68 |
| 3S042 60µm | 62.71 | 21.37 | 0.30 | 0.00 | 2.58 | 8.57 | 1.15 | 0.24 | 96.92 | 13.25 | 79.69 | 7.06 |
| 35042 GM 8010 | 65.06 | 21.12 | 0.37 | 0.00 | 2.24 | 8.92 | 1.14 | 0.24 | 99.10 | 11.37 | 81.77 | 6.86 |
| 3SO45a glom rim | 53 84 | 28 50 | 0.40 | 0.00 | 11 10 | 1.70 | 0.05 | 0.00 | 99.30 | 55 31 | 14.20 | 1 /0 |
| 3S045b core | 50.56 | 31.30 | 0.52 | 0.11 | 13.98 | 3.27 | 0.09 | 0.00 | 99.83 | 69.91 | 29.56 | 0.53 |
| 3S045b rim | 56.25 | 27.07 | 1.00 | 0.05 | 9.08 | 5.67 | 0.40 | 0.00 | 99.53 | 45.80 | 51.78 | 2.42 |
| 3SO45c core | 20.69 | 16.52 | 0.00 | 0.07 | 17.17 | 1.84 | 0.04 | 0.00 | 56.33 | 83.58 | 16.22 | 0.20 |
| 3S045c edge | 47.43 | 33.32 | 0.55 | 0.09 | 16.48 | 1.94 | 0.04 | 0.00 | 99.85 | 82.20 | 17.55 | 0.25 |
| 35045C rim 35045 300m acie CM | 57.00 | 27.28 | 1.09 | 0.04 | 9.59 | 5.66 | 0.34 | 0.00 | 99.66 | 47.37 | 50.62 | 2.01 |
| 3SO45 30µm acic GM | 55 36 | 27 52 | 1 25 | 0.04 | 0 88 | 4.91 | 0.25 | 0.02 | 99.01 | 24.10 40 04 | 44.40 | 2 14 |
| 3S045 40µm acic GM | 56.91 | 25.99 | 1.12 | 0.06 | 8.50 | 6.10 | 0.44 | 0.01 | 99.13 | 42.36 | 55.01 | 2.62 |
| 3S049 core | 49.02 | 31.73 | 0.49 | 0.05 | 15.07 | 2.80 | 0.12 | 0.00 | 99.29 | 74.29 | 24.99 | 0.72 |
| 3S049 intmed | 53.81 | 29.08 | 0.55 | 0.08 | 12.73 | 3.68 | 0.17 | 0.00 | 100.10 | 65.01 | 33.97 | 1.02 |
| 3S049 rim | 49.21 | 31.93 | 0.68 | 0.03 | 15.02 | 2.72 | 0.14 | 0.01 | 99.73 | 74.74 | 24.46 | 0.80 |
| 35058a introd | 56.31 | 24.81 | 0.34 | 0.00 | 6.88 | 7.01 | 0.50 | 0.06 | 98.03 | 34.11 | 62.93 | 2.96 |
| 3S058a rim | 58 78 | 22.30 | 0.20 | 0.01 | 6 01 | 7 18 | 0.31 | 0.05 | 92.17 | 27.00 | 63 70 | 2.73 |
| 3S058b core siv | 57.64 | 26.29 | 0.22 | 0.00 | 7.87 | 6.52 | 0.38 | 0.09 | 99.00 | 39.11 | 58.64 | 2.25 |
| 3S058b rim | 60.60 | 24.43 | 0.26 | 0.00 | 5.84 | 7.54 | 0.54 | 0.09 | 99.29 | 29.01 | 67.80 | 3.19 |
| 3S058 125M acic GM | 58.83 | 25.98 | 0.20 | 0.00 | 7.19 | 6.84 | 0.37 | 0.04 | 99.45 | 35.91 | 61.87 | 2.22 |
| 3SU61a core | 54.23 | 28.56 | 0.30 | 0.00 | 10.51 | 5.14 | 0.27 | 0.00 | 99.02 | 52.20 | 46.20 | 1.60 |
| 35061b core siv | 04.0/ 58 17 | 23.00 | 0.38 | 0.00 | 5.51 | 1.20 | 0.80 | 0.06 | 102.07 | 27.55 | 67.54 | 4.92 |
| 3S061b rim | 55.31 | 27.34 | 0.62 | 0.04 | 9 73 | 5 55 | 0.40 | 0.01 | 99.20 | 48 13 | 40 73 | 2.09 |
| 3S061 50µm GM | 56.64 | 26.67 | 0.53 | 0.02 | 9.03 | 5.91 | 0.41 | 0.04 | 99.25 | 44.67 | 52.91 | 2.43 |
| 3S069a core | 49.40 | 31.66 | 0.73 | 0.09 | 15.01 | 2.78 | 0.11 | 0.00 | 99.78 | 74.40 | 24.96 | 0.64 |
| 3S069a rim | 53.14 | 14.55 | 4.89 | 8.01 | 14.34 | 2.53 | 0.30 | 0.05 | 97.82 | 74.38 | 23.76 | 1.87 |
| 3S069b core | 49.98 | 30.88 | 0.51 | 0.12 | 14.37 | 3.35 | 0.10 | 0.02 | 99.32 | 69.93 | 29.50 | 0.57 |
| 350690 F1m 35069 30um CM acic | 48.00 | 30 18 | 1 10 | 0.09 | 15.32 | 2.95 | 0.12 | 0.00 | 99.83 | 15.15 | 25.55 | 0.69 |
| 3S069 50µm GM | 50.87 | 30.53 | 0.94 | 0.02 | 13.47 | 3.66 | 0.20 | 0.00 | 99.29 | 66 41 | 32 31 | 1 28 |
| 3S074a core siv | 56.25 | 26.58 | 0.46 | 0.01 | 9.12 | 5.92 | 0.26 | 0.00 | 98.59 | 45.30 | 53.20 | 1.51 |
| 3S074a rim | 53.11 | 28.58 | 1.06 | 0.07 | 11.58 | 4.31 | 0.44 | 0.01 | 99.16 | 58.17 | 39.21 | 2.63 |
| 3S074b glom core | 50.29 | 30.81 | 0.52 | 0.09 | 14.02 | 3.18 | 0.07 | 0.04 | 99.03 | 70.56 | 28.99 | 0.44 |
| 3S074b rim | 52.53 | 28.74 | 0.84 | 0.07 | 12.09 | 4.02 | 0.35 | 0.01 | 98.64 | 61.15 | 36.76 | 2.09 |
| 3S074c core SIV | 40.29 | 37.50 | 0.82 | 0.00 | 15.40 | 2.02 | 0.17 | 0.01 | 99.05 | 79 05 | 25.51 | 0.97 |
| 3S074d core rsbd | 58.36 | 25.48 | 0.56 | 0.02 | 7.90 | 6.33 | 0.48 | 0.02 | 99.20 | 39.64 | 57.52 | 2.84 |
| 3S074d edge rsbd | 58.01 | 25.56 | 0.55 | 0.01 | 8.00 | 6.05 | 0.90 | 0.04 | 99.13 | 39.96 | 54.67 | 5.37 |
| 3S074d rim | 51.01 | 28.54 | 0.74 | 0.06 | 11.98 | 3.95 | 0.33 | 0.00 | 96.59 | 61.41 | 36.59 | 2.00 |
| 3S074 30µm GM | 53.67 | 28.11 | 1.03 | 0.06 | 11.18 | 4.69 | 0.41 | 0.00 | 99.15 | 55.49 | 42.11 | 2.40 |
| 3SU74 3Uµm 3SU74 50µm ania CM | 55.59 | 27.65 | 1.44 | 0.05 | 11.15 | 4.35 | 0.45 | 0.04 | 98.71 | 56.99 | 40.26 | 2.75 |
| 3S079a core | 57 87 | 27.40 | 0.04 | 0.05 | 8 45 | 5.90 | 0.33 | 0.00 | 99.31 | 62.02 | 56 12 | 1 25 |
| 3S079a rim | 58.62 | 26.57 | 0.40 | 0.03 | 8.41 | 6.51 | 0.22 | 0.00 | 100.75 | 41.13 | 57.59 | 1.28 |
| 3S079b core | 59.03 | 26.08 | 0.37 | 0.01 | 7.90 | 6.69 | 0.25 | 0.05 | 100.37 | 38.93 | 59.63 | 1.45 |
| 3S079b rim | 59.81 | 24.99 | 0.41 | 0.03 | 7.14 | 7.40 | 0.27 | 0.07 | 100.11 | 34.25 | 64.20 | 1.55 |
| 3S079c core | 59.54 | 25.60 | 0.28 | 0.04 | 7.38 | 6.83 | 0.26 | 0.00 | 99.93 | 36.78 | 61.66 | 1.56 |
| | 52 11 | 24.02 | 0.45 | 0.00 | 12 66 | 1.22 | 0.24 | 0.04 | 98.04 | 30.27 | 37 32 | 1.38 |
| 3S090a rim | 49.22 | 31.73 | 0.66 | 0.06 | 14.58 | 2.93 | 0.09 | 0.05 | 99.31 | 72.97 | 26.52 | 0.51 |
| 3S090b core | 47.58 | 32.79 | 0.55 | 0.05 | 15.89 | 2.24 | 0.08 | 0.03 | 99.20 | 79.31 | 20.23 | 0.46 |
| 3S090b edge | 48.68 | 31.79 | 0.71 | 0.08 | 15.04 | 2.78 | 0.08 | 0.00 | 99.15 | 74.58 | 24.95 | 0.47 |
| 3S090b rim | 55.73 | 26.89 | 0.71 | 0.06 | 9.42 | 5.78 | 0.32 | 0.01 | 98.92 | 46.49 | 51.65 | 1.85 |
| SSUYUC CORE SIV | 40.88 | 33.01 | 0.69 | 0.04 | 16.23 | 2.12 | 0.07 | 0.02 | 99.05 | 80.58 | 19.03 | 0.40 |
| 3S090c rim | 48 82 | 31 55 | 0.61 | 0.04 | 16 77 | 2.86 | 0.17 | 0.00 | 90.07 08 87 | 73 63 | 25 83 | 0.54 |
| 3S090d core | 50.21 | 31.45 | 0.56 | 0.06 | 14.38 | 3.12 | 0.07 | 0.02 | 99.86 | 71.51 | 28,10 | 0.39 |
| 3S090d intmed | 49.70 | 32.33 | 0.66 | 0.05 | 15.24 | 2.78 | 0.06 | 0.00 | 100.81 | 74.93 | 24.72 | 0.36 |
| 3S090d rim | 55.49 | 28.23 | 0.63 | 0.03 | 10.29 | 5.38 | 0.23 | 0.04 | 100.32 | 50.71 | 47.96 | 1.33 |
| 3SU9De core siv | 50.07 | 26.93 | 4.13 | 1.30 | 13.21 | 3.03 | 0.54 | 0.06 | 99.26 | 68.33 | 28.33 | 3.34 |
| 35090f core siv | 50.95 /8 ZO | 51.00 | 0.66 | 0.06 | 15.82 | 5.60 | 0.06 | 0.06 | 100.19 | 67.72 | 51.95 | 0.33 |
| 3S090f edge siv | 49.29 | 32.05 | 0.60 | 0.05 | 15.08 | 2.76 | 0.51 | 0.05 | 90 01 | 74 84 | 24 75 | 0.30 |
| 3S090f rim | 51.14 | 31.02 | 0.52 | 0.02 | 13.63 | 3.64 | 0.09 | 0.00 | 100.06 | 67.10 | 32.38 | 0.52 |
| 3S090 30µm acic GM | 52.66 | 29.13 | 0.84 | 0.04 | 11.88 | 4.62 | 0.23 | 0.02 | 99.41 | 57.91 | 40.76 | 1.33 |
| 3S090 30µm acic GM | 55.40 | 27.60 | 1.12 | 0.04 | 10.10 | 5.45 | 0.27 | 0.00 | 99.99 | 49.82 | 48.59 | 1.60 |
| 55090 50µm acic GM | 55.53 | 29.31 | 0.66 | 0.05 | 11.62 | 4.65 | 0.17 | 0.05 | 100.03 | 57.42 | 41.56 | 1.02 |

| Sample | Si02 | AL203 | Fe0 | MgO | CaO | Na20 | K20 | BaO | TOTAL | An | Ab | Or |
|---------------------|----------------|-------|------|------|-------|------|------|------|--------|-------|-------|-------|
| 35090 30µm acic GM | 54.86 | 28.23 | 0.86 | 0.07 | 10.48 | 5.25 | 0.24 | 0.00 | 99.99 | 51.68 | 46.91 | 1.41 |
| 3S090 50µm acic GM | 52.98 | 29.29 | 0.69 | 0.08 | 11.66 | 4.42 | 0.18 | 0.00 | 99.29 | 58.67 | 40.27 | 1.06 |
| 3S090 60µm GM | 52.55 | 28.72 | 0.74 | 0.04 | 11.71 | 4.53 | 0.17 | 0.00 | 98.47 | 58.22 | 40.78 | 1.00 |
| 3S098a core | 56.40 | 26.89 | 0.39 | 0.02 | 8.64 | 6.33 | 0.22 | 0.01 | 98.89 | 42.44 | 56.25 | 1.31 |
| SSU98a intmed | 56.81 | 27.26 | 0.30 | 0.02 | 8.97 | 6.07 | 0.23 | 0.01 | 99.67 | 44.36 | 54.31 | 1.33 |
| 35098b cono | 61.50 | 24.09 | 0.36 | 0.00 | 5.23 | 8.03 | 0.45 | 0.06 | 99.73 | 25.76 | 71.58 | 2.65 |
| 350900 core | 61.19 | 24.30 | 0.24 | 0.03 | 5.00 | 0.01 | 0.43 | 0.05 | 99.98 | 27.46 | 70.05 | 2.49 |
| 35098c core | 61 05 | 24.32 | 0.20 | 0.01 | 5 43 | 7 80 | 0.40 | 0.05 | 99.93 | 20.21 | 71.00 | 2.19 |
| 3S098c rim | 62.16 | 23.46 | 0.32 | 0.07 | 4.76 | 8.16 | 0.51 | 0.06 | 90.77 | 23 64 | 73 33 | 3 03 |
| 3S098 100µm | 60.23 | 24.40 | 0.38 | 0.01 | 5.93 | 7.84 | 0.36 | 0.06 | 99.22 | 28.85 | 69.07 | 2.08 |
| 38098 50µm acic GM | 63.33 | 23.02 | 0.48 | 0.00 | 3.92 | 8.49 | 0.68 | 0.18 | 100.09 | 19.52 | 76.47 | 4.02 |
| 3S112a core siv | 48.79 | 32.82 | 0.54 | 0.06 | 15.64 | 2.46 | 0.08 | 0.00 | 100.38 | 77.53 | 22.01 | 0.46 |
| 3S112a edge siv | 50.49 | 31.14 | 0.47 | 0.09 | 13.81 | 3.37 | 0.14 | 0.00 | 99.51 | 68.79 | 30.38 | 0.83 |
| 3S112a rim | 57.06 | 26.31 | 0.83 | 0.02 | 8.57 | 6.20 | 0.49 | 0.00 | 99.47 | 42.06 | 55.07 | 2.88 |
| SS112D CORE glom | 54.60 | 27.82 | 0.69 | 0.06 | 10.37 | 5.22 | 0.22 | 0.00 | 98.98 | 51.65 | 47.05 | 1.30 |
| 351120 0000 | 54.17 | 28.43 | 0.67 | 0.00 | 11.02 | 4.74 | 0.21 | 0.00 | 99.30 | 55.49 | 45.24 | 1.2/ |
| 3\$112c rim | 55 31 | 20.90 | 0.00 | 0.07 | 10.02 | 4.00 | 0.22 | 0.00 | 00.5/ | 50.01 | 42.01 | 2 17 |
| 3S112 30µm acic GM | 56.50 | 26.78 | 1.04 | 0.02 | 9.07 | 5.86 | 0.41 | 0.06 | 99.72 | 44 98 | 52 61 | 2.17 |
| 3S112 30µm acic GM | 57.82 | 26.33 | 0.76 | 0.00 | 8.04 | 6.20 | 0.46 | 0.07 | 99.67 | 40.60 | 56.67 | 2.74 |
| 3S114a core | 56.95 | 26.81 | 0.49 | 0.03 | 9.06 | 5.92 | 0.29 | 0.06 | 99.61 | 45.05 | 53.24 | 1.71 |
| 3S114a rim | 56.74 | 26.36 | 0.22 | 0.03 | 8.45 | 6.44 | 0.31 | 0.04 | 98.57 | 41.26 | 56.95 | 1.79 |
| 3S114b core siv | 61.27 | 24.22 | 0.26 | 0.01 | 6.07 | 7.02 | 0.66 | 0.07 | 99.58 | 31.01 | 64.95 | 4.04 |
| 3S114b rim | 57.90 | 26.25 | 0.30 | 0.03 | 8.06 | 6.69 | 0.39 | 0.04 | 99.66 | 39.07 | 58.69 | 2.24 |
| 3S114 50µm acic | 57.17 | 26.89 | 0.35 | 0.01 | 8.77 | 6.06 | 0.32 | 0.01 | 99.59 | 43.59 | 54.54 | 1.88 |
| 3S115a odre | 57 77 | 20.01 | 0.52 | 0.05 | 10.65 | 5.13 | 0.13 | 0.00 | 99.93 | 53.00 | 46.21 | 0.79 |
| 39115a rim | 55 18 | 29.39 | 0.50 | 0.03 | 10.52 | 4.49 | 0.12 | 0.00 | 100.20 | 57.07 | 40.20 | 1.02 |
| 3S115b core | 54.24 | 28.67 | 0.63 | 0.04 | 10.32 | 5 11 | 0.17 | 0.00 | 90.75 | 53.46 | 40.92 | 1 14 |
| 3S115b rim | 57.96 | 26.40 | 0.57 | 0.02 | 8.11 | 6.51 | 0.33 | 0.01 | 99.91 | 39.99 | 58.08 | 1.94 |
| 3S115c core siv | 54.64 | 27.84 | 0.49 | 0.04 | 10.16 | 5.30 | 0.21 | 0.01 | 98.68 | 50.79 | 47.94 | 1.27 |
| 3\$115c rim | 56.23 | 27.42 | 0.74 | 0.03 | 9.29 | 5.83 | 0.25 | 0.00 | 99.77 | 46.13 | 52.41 | 1.47 |
| 3S115 30µm GM | 55.15 | 28.45 | 0.69 | 0.03 | 10.40 | 5.11 | 0.22 | 0.01 | 100.06 | 52.22 | 46.45 | 1.33 |
| 3S115 40µm acic | 57.12 | 26.30 | 0.90 | 0.19 | 8.57 | 6.06 | 0.33 | 0.04 | 99.52 | 42.99 | 55.02 | 1.99 |
| 3S115 acic 20µm | 59.69 | 25.09 | 0.88 | 0.03 | 6.48 | 7.15 | 0.53 | 0.06 | 99.91 | 32.33 | 64.54 | 3.13 |
| 35122a core | 47 77 | 24.41 | 0.17 | 0.00 | 0.00 | 7.52 | 0.56 | 0.08 | 98.80 | 29.80 | 66.92 | 3.28 |
| 3S122h core siv | 59 48 | 25.10 | 0.31 | 0.00 | 7 05 | 6 80 | 0.00 | 0.00 | 99.30 | 20.04 | 61 88 | 4.00 |
| 3S122b rim | 60.36 | 25.35 | 0.29 | 0.00 | 7.03 | 6.96 | 0.53 | 0.11 | 100 62 | 34 68 | 62 19 | 3 13 |
| 3S122c core siv | 62.84 | 23.84 | 0.14 | 0.00 | 5.21 | 7.65 | 0.69 | 0.06 | 100.43 | 26.22 | 69.63 | 4.15 |
| 3S122c rim | 60.49 | 25.53 | 0.20 | 0.02 | 6.68 | 6.93 | 0.48 | 0.03 | 100.36 | 33.77 | 63.33 | 2.90 |
| 3S122d core siv | 59.53 | 25.43 | 0.21 | 0.00 | 6.92 | 6.86 | 0.51 | 0.08 | 99.54 | 34.71 | 62.24 | 3.05 |
| 3S122d rim | 61.07 | 24.65 | 0.29 | 0.00 | 6.24 | 7.20 | 0.53 | 0.10 | 100.08 | 31.35 | 65.48 | 3.17 |
| 3S122e core | 54.12 | 28.71 | 0.35 | 0.00 | 11.01 | 4.85 | 0.23 | 0.05 | 99.32 | 54.86 | 43.75 | 1.39 |
| 3S122e r1m | 60.35 | 25.50 | 0.26 | 0.01 | 6.79 | 6.69 | 0.47 | 0.06 | 100.13 | 34.90 | 62.23 | 2.87 |
| 35123a core | 55 17 | 29.09 | 0.50 | 0.05 | 11.45 | 4.72 | 0.18 | 0.02 | 99.50 | 56.66 | 42.29 | 1.05 |
| 3\$123h core | 55 08 | 20.02 | 0.05 | 0.05 | 10.23 | 5 36 | 0.33 | 0.00 | 99.79 | 50.43 | 47.01 | 2 16 |
| 3S123b rim | 54.34 | 28.33 | 0.56 | 0.06 | 10.66 | 5.09 | 0.37 | 0.02 | 99.35 | 52.79 | 45.61 | 1.60 |
| 3S123c core | 53.32 | 28.86 | 0.45 | 0.07 | 11.32 | 4.94 | 0.19 | 0.01 | 99.16 | 55.27 | 43.61 | 1.13 |
| 3S123c rim | 56.17 | 27.05 | 0.58 | 0.04 | 9.26 | 5.82 | 0.41 | 0.01 | 99.33 | 45.67 | 51.95 | 2.39 |
| 3S123d core glom | 53.83 | 29.05 | 0.50 | 0.05 | 11.21 | 4.93 | 0.16 | 0.00 | 99.73 | 55.14 | 43.92 | 0.95 |
| 3S123d rim glom | 62.82 | 21.84 | 0.73 | 0.02 | 4.33 | 7.01 | 1.62 | 0.05 | 98.43 | 22.85 | 66.94 | 10.20 |
| 3S123d rim glom | 56.56 | 26.99 | 0.74 | 0.00 | 9.00 | 5,90 | 0.49 | 0.05 | 99.74 | 44.41 | 52.73 | 2.86 |
| 3S123 30µm acic GM | 57.67 | 25.46 | 1.14 | 0.37 | 7.76 | 6.58 | 0.58 | 0.04 | 99.59 | 38.11 | 58.47 | 3.42 |
| 35123 JUJUM acto GM | 50 /2 | 24.94 | 0.70 | 0.01 | 0./9 | 7.17 | 0.08 | 0.08 | 99.02 | 33.01 | 03.04 | 3.95 |
| 3\$131a edge siv | 50.42 | 24 77 | 0.65 | 0.05 | 7 10 | 7 05 | 0.10 | 0.03 | 99.00 | 70.94 | 62 31 | 3 03 |
| 3S131a rim | 50.53 | 30.65 | 0.58 | 0.04 | 13.83 | 3.44 | 0.08 | 0.05 | 99.20 | 68.64 | 30.92 | 0.45 |
| 3S131b core siv | 50.21 | 31.05 | 0.62 | 0.05 | 14.19 | 3.45 | 0.12 | 0.02 | 99.70 | 68.99 | 30.32 | 0.69 |
| 3\$131b rim | 53.79 | 27.46 | 0.95 | 0.03 | 10.67 | 5.14 | 0.23 | 0.03 | 98.29 | 52.70 | 45.97 | 1.34 |
| 3\$131c core siv | 55.22 | 27.08 | 0.61 | 0.10 | 9.81 | 5.55 | 0.16 | 0.03 | 98.57 | 48.91 | 50.12 | 0.97 |
| 3S131c edge siv | 52.02 | 29.46 | 0.80 | 0.07 | 12.61 | 4.10 | 0.15 | 0.04 | 99.24 | 62.41 | 36.72 | 0.87 |
| 55131c rim | 61.13 | 23.28 | 0.48 | 0.02 | 4.97 | 7.77 | 0.63 | 0.06 | 98.33 | 25.13 | 71.10 | 3.77 |
| 35131 SUMM GM | 54.76 | 27.64 | 0.83 | 0.07 | 10.18 | 5.21 | 0.27 | 0.04 | 99.00 | 51.07 | 47.31 | 1.62 |
| 39131 30µm 6M | J4.03 61 79 | 27.10 | 0.90 | 0.07 | 9.92 | 2.39 | 0.28 | 0.04 | Y0.00 | 40.72 | 49.0/ | 5.5/ |
| 35131 75M acic GM | 54.65 | 27.06 | 1.00 | 0.11 | 9.40 | 5 53 | 0.94 | 0.11 | 98.60 | 48 55 | 49 78 | 1.66 |
| 3S134a core | 60.01 | 25.38 | 0.29 | 0.00 | 6.68 | 7.22 | 0.23 | 0.01 | 99.81 | 33.39 | 65.26 | 1.35 |

| Sample | Si02 | AL203 | FeO | MgO | CaO | Na20 | K20 | BaO | TOTAL | An | Ab | Or |
|--------------------|-------|-------|------|------|------|------|------|------|--------|-------|-------|------|
| 3S134a rim | 58.77 | 25.97 | 0.30 | 0.01 | 7.37 | 6.84 | 0.24 | 0.02 | 99.52 | 36.77 | 61.78 | 1.45 |
| 3S134b core glom | 59.03 | 25.80 | 0.38 | 0.00 | 7.41 | 6.86 | 0.35 | 0.04 | 99.87 | 36.64 | 61.32 | 2.04 |
| 3S134b rim | 57.68 | 25.24 | 0.40 | 0.00 | 7.34 | 6.38 | 0.27 | 0.00 | 97.32 | 38.21 | 60.11 | 1.68 |
| 3S134 30µm acic GM | 60.51 | 23.97 | 0.64 | 0.00 | 5.53 | 7.69 | 0.59 | 0.07 | 98.99 | 27.45 | 69.06 | 3.49 |
| 3S134 50µm GM | 59.65 | 24.71 | 0.51 | 0.01 | 6.03 | 7.45 | 0.42 | 0.09 | 98.87 | 30.10 | 67.38 | 2.52 |
| 3S137a core | 55.65 | 27.10 | 0.47 | 0.03 | 9.35 | 6.04 | 0.17 | 0.01 | 98.82 | 45.65 | 53.38 | 0.97 |
| 3S137a rim | 55.34 | 27.02 | 0.35 | 0.03 | 8.96 | 5.75 | 0.20 | 0.02 | 97.66 | 45.68 | 53.09 | 1.23 |
| 3S137b core | 56.49 | 26.75 | 0.40 | 0.02 | 8.81 | 6.25 | 0.18 | 0.02 | 98.92 | 43.32 | 55.61 | 1.07 |
| 3S137b rim | 57.36 | 26.33 | 0.48 | 0.01 | 8.37 | 6.65 | 0.31 | 0.05 | 99.55 | 40.28 | 57.95 | 1.77 |
| 3S137c core | 56.34 | 26.99 | 0.35 | 0.01 | 9.20 | 6.08 | 0.20 | 0.03 | 99.20 | 45.02 | 53.81 | 1.18 |
| 3S137c rim | 56.61 | 25.95 | 0.48 | 0.04 | 8.15 | 6.48 | 0.22 | 0.04 | 97.98 | 40.47 | 58.25 | 1.28 |
| 3S137 50µm acic GM | 58.46 | 25.54 | 0.49 | 0.01 | 7.43 | 6.91 | 0.38 | 0.04 | 99.26 | 36.45 | 61.31 | 2.24 |
| 3S137 50µm acic GM | 58.89 | 25.59 | 0.53 | 0.02 | 7.06 | 7.08 | 0.41 | 0.05 | 99.61 | 34.68 | 62.93 | 2.39 |
| 3S137 100µm acic | 58.64 | 25.73 | 0.52 | 0.01 | 7.20 | 6.99 | 0.37 | 0.09 | 99.54 | 35.50 | 62.36 | 2.15 |
| Lava Is. 1 core | 62.81 | 23.28 | 0.19 | 0.00 | 4.72 | 8.22 | 0.56 | 0.05 | 99.83 | 23.30 | 73.43 | 3.27 |
| Lava Is. 1 rim | 62.19 | 22.19 | 0.21 | 0.00 | 4.11 | 8.90 | 0.62 | 0.14 | 98.35 | 19.60 | 76.85 | 3.55 |
| Lava Is. 2 core | 63.03 | 23.34 | 0.21 | 0.00 | 4.67 | 8.33 | 0.56 | 0.09 | 100.23 | 22.90 | 73.85 | 3.25 |
| Lava Is. 2 rim | 64.86 | 23.98 | 0.18 | 0.00 | 4.53 | 8.07 | 0.52 | 0.05 | 102.19 | 22.95 | 73.93 | 3.12 |

Appendix C-2: Compositions of pyroxenes determined by electron microprobe analysis. FeO and Fe₂O₃, and mol percents wollastonite (Wo), enstatite (En) and ferrosilite (Fs) calculated with standard CAMECA software. Groundmass crystals denoted by size with sample number. Glom: glomerocryst; edge: analysis next to overgrowth rims.

| | Si02 | Ti02 | AL203 | FeO | Fe203 | MnO | MgO | CaO | Na20 | Cr203 | TOTAL | Wo | En | Fs |
|----------------|---------------|------|-------|-------|-------|------|-------|-------|------|-------|--------|-------|-------|-------|
| 3SXBPa core | 51.31 | 0.42 | 1.18 | 10.25 | 1.61 | 0.49 | 13.85 | 19.43 | 0.30 | 0.00 | 98.84 | 41.26 | 40.92 | 17.82 |
| 3SXBPa rim | 50.90 | 0.38 | 1.10 | 8.86 | 2.70 | 0.51 | 14.23 | 19.53 | 0.31 | 0.00 | 98,52 | 41.88 | 42.43 | 15.69 |
| 3SXBPb core | 49.72 | 0.09 | 0.27 | 31.92 | 2.80 | 1.97 | 13.27 | 1.35 | 0.05 | 0.00 | 101.44 | 2.91 | 39.89 | 57.20 |
| 3SXBPb rim | 49.07 | 0.10 | 0.22 | 31.78 | 2.24 | 1.98 | 12.97 | 1.29 | 0.05 | 0.00 | 00.70 | 2.83 | 30 47 | 57 70 |
| 3SXBPc core | 49.97 | 0.14 | 0.29 | 32.56 | 0.67 | 1.90 | 13 27 | 1 24 | 0 03 | 0 03 | 100 09 | 2 66 | 30 61 | 57 73 |
| 3SXBPc rim | 40 10 | 0 11 | 0 22 | 31 69 | 2 34 | 1 81 | 13 26 | 1 23 | 0.05 | 0 01 | 00 01 | 2 68 | 40 23 | 57 00 |
| 3SYRD | 51 17 | 0 00 | 3 04 | 8 1/ | 1 00 | 0.60 | 14 64 | 10 56 | 0.05 | 0.01 | 100 61 | /1 8/ | 40.23 | 1/ 60 |
| TSYRD | 51 18 | 0.0% | 2 64 | 7 30 | 2 / 6 | 0.50 | 15 24 | 10 /0 | 0.47 | 0.00 | 100.01 | /1 4/ | 43.30 | 17 01 |
| 350030 0000 | 51 13 | 0.64 | 1 8/ | 7 63 | 7 /0 | 0.50 | 1/ 7/ | 10.71 | 0.45 | 0.00 | 100.20 | 41.04 | 43.30 | 17 01 |
| | 51 99 | 0.04 | 1.04 | 6.00 | 7 55 | 0.02 | 14.74 | 19.71 | 0.42 | 0.01 | 100.13 | 42.23 | 43.93 | 13.01 |
| 35003a 1111 | 51.00 | 0.42 | 1.07 | 0.90 | 3.33 | 0.39 | 13.70 | 19.02 | 0.34 | 0.00 | 100.22 | 41.31 | 40.24 | 12.45 |
| JSUUJD COre | 51.00 | 0.47 | 1.20 | 9.10 | 1.09 | 0.75 | 14.55 | 19.12 | 0.39 | 0.05 | 99.21 | 40.60 | 42.98 | 10.42 |
| 35003D F1M | 51.62 | 0.49 | 1.32 | 8.54 | 1.49 | 0.48 | 14.89 | 19.48 | 0.36 | 0.00 | 98.67 | 41.24 | 43.85 | 14.91 |
| SSUUSC COre | 52.55 | 0.19 | 1.10 | 18.64 | 1.07 | 0.95 | 23.32 | 1.42 | 0.00 | 0.02 | 99.24 | 2.90 | 65.98 | 31.12 |
| SSUUSC rim | 52.11 | 0.26 | 0.55 | 14.59 | 4.67 | 0.93 | 25.22 | 1.49 | 0.03 | 0.00 | 99.86 | 3.06 | 72.05 | 24.89 |
| 3S003d core | 52.37 | 0.25 | 0.64 | 19.28 | 1.28 | 1.17 | 22.68 | 1.48 | 0.01 | 0.00 | 99.18 | 3.02 | 64.38 | 32.60 |
| 3\$003d rim | 50.66 | 0.25 | 0.67 | 12.68 | 7.05 | 1.01 | 25.28 | 1.46 | 0.04 | 0.03 | 99.12 | 3.09 | 74.31 | 22.60 |
| 3S005a core | 51.58 | 0.25 | 0.59 | 23.76 | 2.11 | 0.88 | 19.67 | 1.56 | 0.05 | 0.01 | 100.44 | 3.25 | 56.81 | 39.94 |
| 3S005a rim | 52.56 | 0.22 | 0.61 | 25.96 | 0.00 | 0.94 | 18.74 | 1.51 | 0.03 | 0.00 | 100.56 | 3.10 | 53.66 | 43.24 |
| 3S005b core | 51.25 | 0.21 | 0.40 | 24.78 | 2.20 | 0.99 | 18.89 | 1.47 | 0.03 | 0.00 | 100.21 | 3.06 | 54.91 | 42.03 |
| 3S005b rim | 52.40 | 0.19 | 0.49 | 27.14 | 0.00 | 0.93 | 18.28 | 1.58 | 0.01 | 0.00 | 101.02 | 3.23 | 51.98 | 44.79 |
| 3S005c core | 50.83 | 0.36 | 1.04 | 12.55 | 1.95 | 0.61 | 12.83 | 18.43 | 0.32 | 0.00 | 98.92 | 39.58 | 38.35 | 22.07 |
| 3S005c rim | 51.47 | 0.35 | 0.96 | 13.66 | 0.00 | 0.48 | 12.41 | 18.78 | 0.24 | 0.00 | 98.35 | 39.88 | 36.67 | 23.45 |
| 3S005d core | 50.68 | 0.39 | 1.18 | 12.15 | 1.85 | 0.60 | 12.79 | 18.85 | 0.27 | 0.04 | 98.81 | 40.45 | 38.18 | 21.37 |
| 3S005d rim | 50.30 | 0.34 | 1.10 | 11.06 | 2.63 | 0.59 | 12.77 | 19.25 | 0.30 | 0.00 | 98.33 | 41.75 | 38.52 | 19.74 |
| 3S009a core | 52.36 | 0.28 | 0.97 | 10.42 | 1.21 | 0.58 | 13.97 | 10 03 | 0 31 | 0.00 | 100 03 | 41 56 | 40 52 | 17 92 |
| 35009a rim | 52 16 | 0.27 | 1 10 | 10.26 | 1 40 | 0 55 | 13 87 | 20 00 | 0.37 | 0.05 | 80 00 | 41.90 | 40.52 | 17 60 |
| 35009b core | 51 71 | 0 52 | 1 69 | 10 33 | 1 27 | 0.55 | 13.67 | 10 80 | 0.36 | 0.01 | 00 00 | 41.70 | 40.41 | 17.80 |
| 35000b cim | 52 10 | 0.21 | 0.81 | 12 02 | 0 11 | 0.41 | 13.07 | 10.75 | 0.30 | 0.01 | 00 09 | 42.0J | 79 00 | 20.47 |
| 350096 cono | 52 5/ | 0.21 | 1 00 | 20 35 | 2 /5 | 0.01 | 13.10 | 1 / 7 | 0.20 | 0.00 | 101 7/ | 7 00 | JO.09 | 20.03 |
| 30000 core | 53 67 | 0.14 | 0.70 | 20.33 | 0.37 | 0.70 | 10 90 | 1.42 | 0.03 | 0.00 | 101.34 | 2.00 | 03.34 | 33.30 |
| 350090 1111 | 50.50 | 0.10 | 7 77 | 24.14 | 1 00 | 0.90 | 19.09 | 1.42 | 0.03 | 0.02 | 100.03 | 2.09 | 20.27 | 40.84 |
| 35017a core | 50.59 | 0.95 | 3.21 | 7.02 | 1.99 | 0.39 | 14.45 | 20.00 | 0.38 | 0.05 | 99.87 | 42.99 | 43.22 | 13.78 |
| 350178 F1m | 49.00 | 1.13 | 3.72 | 7.19 | 2.93 | 0.39 | 14.30 | 19.49 | 0.34 | 0.05 | 99.78 | 42.51 | 43.56 | 13.93 |
| JSUI/D Core | 49.70 | 1.14 | 4.32 | 1.02 | 5.04 | 0.34 | 17.18 | 15.92 | 0.28 | 0.19 | 99.92 | 34.47 | 51.75 | 13.79 |
| 350176 FIM | 48.80 | 1.21 | 4.86 | 8.79 | 2.68 | 0.33 | 15.19 | 17.14 | 0.29 | 0.05 | 99.40 | 37.76 | 46.54 | 15.70 |
| 3SU1/c core | 50.87 | 0.77 | 5.15 | 7.26 | 1.54 | 0.20 | 14.96 | 20.27 | 0.31 | 0.10 | 99.40 | 43.22 | 44.37 | 12.41 |
| 3SO17c rim | 50.65 | 0.84 | 3.38 | 6.62 | 2.96 | 0.21 | 15.20 | 20.13 | 0.35 | 0.05 | 100.40 | 43.18 | 45.37 | 11.45 |
| 3S028 core | 51.79 | 0.40 | 1.17 | 8.38 | 1.93 | 0.63 | 14.59 | 19.83 | 0.40 | 0.02 | 99.12 | 42.05 | 43.04 | 14.92 |
| 3S028a core | 51.95 | 0.42 | 1.18 | 8.60 | 1.72 | 0.55 | 14.76 | 19.84 | 0.35 | 0.00 | 99.37 | 41.75 | 43.20 | 15.05 |
| 3S028a rim | 49.87 | 0.47 | 1.34 | 5.40 | 6.04 | 0.59 | 15.34 | 19.33 | 0.42 | 0.01 | 98.82 | 42.62 | 47.06 | 10.32 |
| 3S028b core | 50 .83 | 0.64 | 1.93 | 9.26 | 1.84 | 0.64 | 14.23 | 19.00 | 0.38 | 0.01 | 98.75 | 40.84 | 42.55 | 16.62 |
| 3S028b rim | 52.50 | 0.40 | 1.26 | 9.78 | 0.30 | 0.52 | 14.55 | 19.68 | 0.36 | 0.02 | 99.36 | 41.03 | 42.21 | 16.77 |
| 3S028c core | 52.53 | 0.26 | 0.71 | 17.91 | 1.77 | 0.97 | 23.75 | 1.37 | 0.02 | 0.00 | 99.29 | 2.79 | 67.22 | 30.00 |
| 3S028c rim | 51.97 | 0.27 | 0.69 | 14.64 | 5.01 | 0.90 | 25.12 | 1.45 | 0.04 | 0.02 | 100.12 | 2.99 | 72.00 | 25.01 |
| 3S030 20µm | 52.22 | 0.78 | 0.92 | 12.44 | 0.18 | 0.40 | 14.15 | 18.31 | 0.35 | 0.05 | 99.79 | 38.13 | 41.00 | 20.87 |
| 3S030 30µm | 51.72 | 0.69 | 1.93 | 9.11 | 0.84 | 0.30 | 15.92 | 18.18 | 0.30 | 0.09 | 99.06 | 38.13 | 46.46 | 15.41 |
| 3\$030 GM 30µm | 52.05 | 0.61 | 2.19 | 8.77 | 1.13 | 0.24 | 16.37 | 18.15 | 0.29 | 0.14 | 99.94 | 37.85 | 47.48 | 14.68 |
| 3S030 Ophitic | 53.87 | 0.34 | 1.20 | 17.03 | 0.30 | 0.42 | 24.90 | 2.12 | 0.04 | 0.03 | 100.24 | 4.21 | 68.75 | 27.05 |
| 3S030a core | 53.00 | 0.39 | 1.01 | 19.92 | 0.00 | 0.48 | 22.59 | 2.10 | 0.04 | 0.04 | 99.55 | 4.25 | 63.56 | 32.20 |
| 3S030a rim | 51.46 | 0.21 | 0.13 | 28.28 | 0.00 | 0.84 | 16 83 | 1 56 | 0 04 | 0.06 | 00 40 | 3 26 | 40 08 | 47 66 |
| 35036a core | 53 21 | 0 40 | 1 23 | 17.61 | 1 30 | 0.67 | 24 34 | 1 62 | 0 05 | 0 03 | 100 46 | 3 26 | 68 05 | 28 60 |
| 35036a rim | 53 07 | 0 35 | 1 08 | 16 95 | 1 67 | 0.62 | 26 73 | 1 50 | 0.05 | 0.03 | 100.40 | 3 21 | 60.00 | 27 50 |
| 350364 30µm | 51 82 | 0.58 | 1 56 | 8 37 | 1 68 | 0.33 | 15 /8 | 10 28 | 0.02 | 0.02 | 00.00 | 40 48 | 15 74 | 1/ 28 |
| 35036 30µm | 53 38 | 0.35 | 1 16 | 17 24 | 0 44 | 0.55 | 76 77 | 1 79 | 0.52 | 0.00 | 00 71 | 3 57 | 68 7/ | 27 70 |
| 35036 50µm | 52 49 | 0.40 | 1 71 | 13 25 | 2 9/ | 0.71 | 24.12 | 8 20 | 0.03 | 0.00 | 101 47 | 17 14 | 60.07 | 21.10 |
| 35030 J0µm | 5/ 17 | 0.00 | 0.92 | 15.70 | 0.50 | 0.44 | 21.42 | 2 07 | 0.21 | 0.00 | 101.03 | 11.10 | 70 99 | 21.00 |
| 3045 20mi | 50 5/ | 0.43 | 7 47 | 7 77 | 0.39 | 0.51 | 23.03 | 2.03 | 0.03 | 0.00 | 00.79 | 4.01 | 10.00 | 17 (9 |
| 3045a core | 51 77 | 0.00 | 3.03 | 10.5/ | 2.14 | 0.24 | 10.00 | 10.40 | 0.33 | 0.03 | YY.40 | 37.31 | 41.01 | 12.00 |
| 330438 F1m | 51.23 | 0.02 | 3.06 | 10.04 | 0.00 | 0.30 | 14.47 | 17.45 | 0.47 | 0.03 | 90.30 | 37.90 | 43./1 | 10.39 |
| JSU43D CORE | 22.00 | 0.49 | 1.48 | 10.27 | 0.00 | 0.46 | 14.55 | 19.47 | 0.55 | 0.00 | 99.90 | 40.50 | 42.06 | 17.44 |
| 350450 F1M | 51.69 | 0.80 | 2.54 | 10.65 | 0.00 | 0.46 | 14.18 | 19.05 | 0.33 | 0.04 | 99.51 | 40.17 | 41.58 | 18.25 |
| 55049 50µm | 52.35 | 0.60 | 1.81 | 7.98 | 0.16 | 0.22 | 16.77 | 18.81 | 0.21 | 0.23 | 99.14 | 38.75 | 48.07 | 13.18 |
| 3S049 30µm | 48.48 | 1.52 | 4.10 | 1.38 | 6.56 | 0.46 | 15.64 | 20.56 | 0.71 | 0.05 | 99.46 | 46.99 | 49.72 | 3.29 |
| 3SO49a core | 50.65 | 0.82 | 3.50 | 6.07 | 1.45 | 0.20 | 15.77 | 19.74 | 0.36 | 0.49 | 99.05 | 42.39 | 47.11 | 10.50 |

| | sin2 | TiO2 | AI 203 | FeO | Fe203 | Meth | Ha0 | CeO | No20 | Ce203 | TOTAL | Uo | En Ec |
|---------------|-------|------|--------|-------|-------|------|-------|-------|------|-------|--------|--------------|-------------|
| 35040a rim | 50 40 | 0 01 | 3 75 | 6 02 | 1 0/ | 0 27 | 15 27 | 20 52 | 0 77 | 0 75 | 09 97 | // 05 | /5 /9 10 /7 |
| 35049b cono | 50.37 | 0.44 | / 75 | / 53 | 7.04 | 0.23 | 14 11 | 20.02 | 0.33 | 0.33 | 90.07 | 44.03 | 43.40 10.47 |
| 330490 COTE | 50.27 | 0.00 | 4.33 | 4.52 | 2.02 | 0.20 | 10.11 | 20.00 | 0.30 | 0.79 | 99.20 | 43.3/ | 48.63 8.00 |
| 35049D r1m | 51.94 | 0.48 | 2.96 | 6.00 | 0.71 | 0.21 | 17.44 | 18.34 | 0.38 | 0.25 | 98.71 | 38.66 | 51.12 10.22 |
| 3S058 50µm | 51.65 | 0.13 | 0.46 | 25.17 | 0.49 | 1.03 | 19.18 | 1.08 | 0.02 | 0.00 | 99.21 | 2.23 | 55.34 42.43 |
| 3S058a core | 52.68 | 0.11 | 0.33 | 24.86 | 0.00 | 1.24 | 19.04 | 0.86 | 0.01 | 0.00 | 99.13 | 1.80 | 55.50 42.70 |
| 3S058a rim | 52.43 | 0.12 | 0.39 | 24.43 | 0.00 | 1.17 | 20.01 | 0.92 | 0.01 | 0.02 | 99.50 | 1.88 | 57.11 41.01 |
| 3S061 glom | 51.52 | 0.69 | 2.22 | 7.23 | 1.89 | 0.36 | 15.25 | 21.19 | 0.35 | 0.00 | 100.70 | 42.66 | 44.82 12.52 |
| 3S061 glom | 53.75 | 0.21 | 0.63 | 18.46 | 0.00 | 0.65 | 24.10 | 1.52 | 0.03 | 0.04 | 00.30 | 3.03 | 67.11 29 86 |
| 3S061 atom | 53.88 | 0.24 | 0.81 | 16.41 | 1.03 | 0 71 | 25 66 | 1 30 | 0 00 | 0 00 | 101 03 | 2 76 | 70 74 26 50 |
| 3S061a core | 54 37 | 0 27 | 0.63 | 16 52 | 0.83 | 0.52 | 25 76 | 1 63 | 0.05 | 0.00 | 100 58 | 3 21 | 70 58 26 21 |
| 35061a rim | 53 78 | 0 27 | 0.03 | 16 73 | 0.72 | 0.62 | 25 /0 | 1 3/ | 0.05 | 0.00 | 00 80 | 2.67 | 70 36 26 07 |
| 35061b core | 51 5/ | 0 60 | 2 / 8 | 0 14 | 0.72 | 0.02 | 15 00 | 10 11 | 0.03 | 0.00 | 09 90 | /0.3/ | // 33 45 5/ |
| 35061b core | 51 47 | 0.09 | 0.77 | 77 44 | 0.22 | 1.04 | 17.05 | 7.07 | 0.31 | 0.00 | 90.09 | 40.24 | 44.22 13.34 |
| 30010 Fill | 51.07 | 0.40 | 0.37 | 23.00 | 0.30 | 1.00 | 17.95 | 3.93 | 0.00 | 0.00 | 99.48 | 8.15 | 51.80 40.05 |
| 350610 by r1m | 52.30 | 0.52 | 1.07 | 7.34 | 1.02 | 0.37 | 15.47 | 20.42 | 0.34 | 0.00 | 100.05 | 42.58 | 44.87 12.55 |
| 3SU61c core | 52.85 | 0.38 | 1.41 | 17.65 | 1.01 | 0.57 | 24.03 | 1.80 | 0.04 | 0.01 | 99.75 | 3.64 | 67.61 28.76 |
| 3S061c rim | 53.63 | 0.29 | 0.93 | 16.45 | 1.33 | 0.51 | 25.52 | 1.40 | 0.03 | 0.00 | 100.09 | 2.79 | 70.80 26.41 |
| 3S074 50μm | 53.20 | 0.21 | 0.34 | 20.29 | 0.92 | 0.73 | 22.51 | 1.86 | 0.06 | 0.00 | 100.11 | 3.75 | 63.16 33.10 |
| 3S074a core | 51.44 | 0.81 | 2.66 | 7.42 | 1.39 | 0.25 | 15.37 | 20.00 | 0.34 | 0.00 | 99.67 | 42.22 | 45.14 12.64 |
| 3S074a rim | 48.99 | 1.20 | 5.02 | 6.65 | 2.77 | 0.27 | 14.79 | 19.45 | 0.32 | 0.05 | 99.49 | 42.82 | 45.29 11.90 |
| 3S074b core | 50.91 | 0.94 | 3.05 | 7.65 | 1.73 | 0.39 | 15.16 | 19.41 | 0.39 | 0.02 | 99.65 | 41.50 | 45.08 13.42 |
| 3S074b rim | 50.42 | 0.96 | 4.14 | 6.15 | 1.94 | 0.22 | 15.29 | 20.50 | 0.28 | 0.38 | 100.26 | 43.86 | 45.51 10.64 |
| 3S079a core | 53.33 | 0.46 | 1.65 | 17.08 | 1.22 | 0.72 | 24.45 | 2.06 | 0.03 | 0.00 | 100.99 | 4.12 | 68.06 27.82 |
| 3S079a rim | 54.47 | 0.39 | 1.27 | 15.23 | 2 46 | 0 59 | 26 63 | 1 53 | 0.05 | 0 00 | 102 63 | 3 01 | 72 73 24 26 |
| 35079b core | 52 26 | 0.63 | 2 02 | 0 62 | 1 68 | 0 64 | 1/ 58 | 10 51 | 0.00 | 0.00 | 101 37 | 40.81 | /2 /3 16 76 |
| 35079b cim | 51 58 | 0.05 | 1 35 | 6 80 | 3 57 | 0.67 | 15 12 | 20 14 | 0.39 | 0.04 | 100.30 | 40.01 | 42.43 10.70 |
| 350796 6000 | 57 /6 | 0.47 | 1.33 | 0.07 | 0.04 | 0.02 | 1/ 97 | 20.10 | 0.39 | 0.05 | 100.20 | 42.04 | 44.09 12.47 |
| 350790 core | 53.40 | 0.45 | 1.17 | 9.00 | 7.70 | 0.55 | 14.0/ | 20.43 | 0.31 | 0.07 | 101.07 | 41.01 | 42.10 10.24 |
| 35079C F11 | 52.43 | 0.42 | 1.15 | 7.04 | 3.30 | 0.00 | 12.44 | 19.98 | 0.38 | 0.00 | 101.28 | 41.75 | 44.89 13.37 |
| JSU/YO COPE | 23.42 | 0.29 | 0.65 | 20.25 | 1.51 | 1.10 | 22.19 | 1.54 | 0.05 | 0.00 | 101.43 | 3.08 | 63.52 33.40 |
| 3SU/9d r1m | 53.48 | 0.31 | 0.66 | 20.66 | 0.17 | 1.12 | 22.69 | 1.44 | 0.03 | 0.00 | 100.54 | 2.88 | 63.11 34.01 |
| 3S090 20µm | 50.27 | 1.18 | 3.10 | 10.23 | 1.59 | 0.38 | 15.04 | 17.38 | 0.32 | 0.01 | 99.49 | 37.31 | 44.92 17.77 |
| 3S090a core | 53.32 | 0.46 | 2.16 | 14.88 | 1.70 | 0.41 | 26.14 | 1.69 | 0.03 | 0.02 | 100.80 | 3.39 | 72.75 23.87 |
| 3S090a rim | 53.88 | 0.34 | 0.81 | 17.04 | 1.30 | 0.55 | 24.97 | 1.84 | 0.06 | 0.00 | 100.78 | 3.66 | 69.04 27.30 |
| 3S090b glom | 50.60 | 1.04 | 3.54 | 7.61 | 2.04 | 0.31 | 15.08 | 19.49 | 0.36 | 0.03 | 100.11 | 41.77 | 44.97 13.27 |
| 3S090b rim | 50.14 | 1.09 | 3.37 | 7.94 | 1.56 | 0.29 | 14.55 | 19.57 | 0.36 | 0.03 | 98.90 | 42.33 | 43.78 13.89 |
| 3S090c core | 52.47 | 0.50 | 1.77 | 8.08 | 0.70 | 0.26 | 14.96 | 20.79 | 0.33 | 0.01 | 99.86 | 43.20 | 43.26 13.54 |
| 3S090c rim | 49.21 | 1.06 | 4.02 | 7.43 | 3.70 | 0.26 | 14.35 | 19.45 | 0.35 | 0.08 | 99.90 | 42.83 | 43.96 13.21 |
| 3\$098 | 52.01 | 0.20 | 0.64 | 14.51 | 0.78 | 0.87 | 11.69 | 19.40 | 0.28 | 0.01 | 100.38 | 40.70 | 34.11 25.90 |
| 35098 glom | 51.20 | 0.23 | 0.46 | 27.13 | 0.13 | 1.30 | 17.27 | 1.60 | 0.04 | 0.00 | 99.34 | 3.34 | 50.23 46.43 |
| 3S098a core | 51.16 | 0 23 | 0 40 | 20 04 | 0 20 | 1 28 | 16 12 | 1 63 | 0 05 | 0 00 | 100 11 | 3 42 | 46 98 49 60 |
| 35008a rim | 51 37 | 0.21 | 0.40 | 28 /0 | 0.20 | 1 24 | 16 60 | 1.50 | 0.05 | 0.06 | 00.97 | 3.42 | 40.70 47.00 |
| 35008b core | 51 38 | 0.21 | 0.33 | 27 11 | 0.60 | 1 25 | 17 /9 | 1.70 | 0.00 | 0.00 | 100 0/ | J.17 Z 11 | 40.20 40.JJ |
| 75008b aim | 51.00 | 0.10 | 0.00 | 21.11 | 0.09 | 1.20 | 17.40 | 1.47 | 0.03 | 0.05 | 100.04 | 7.00 | 17 07 10 00 |
| 350900 F11 | 51.02 | 0.20 | 0.40 | 47 57 | 0.30 | 1.40 | 10.11 | 1.00 | 0.05 | 0.00 | 99.00 | 3.00 | 47.03 49.09 |
| 350900 Core | 51.45 | 0.35 | 0.00 | 13.37 | 1.00 | 0.70 | 11.90 | 19.20 | 0.32 | 0.02 | YY. 30 | 40.92 | 33.41 23.0/ |
| 35098C rim | 51.95 | 0.20 | 0.62 | 14.41 | 0.35 | 0.71 | 11.79 | 19.33 | 0.31 | 0.00 | 99.75 | 40.66 | 54.49 24.84 |
| 35112 | 54.35 | 0.29 | 0.89 | 17.03 | 0.96 | 0.47 | 25.40 | 1.80 | 0.04 | 0.00 | 101.23 | 3.54 | 69.56 26.89 |
| 3S112a core | 52.07 | 0.72 | 1.97 | 9.91 | 1.12 | 0.32 | 15.69 | 17.98 | 0.36 | 0.04 | 100.18 | 37.61 | 45.68 16.72 |
| 3S112a rim | 51.52 | 0.87 | 1.32 | 13.25 | 0.44 | 0.36 | 14.36 | 17.12 | 0.27 | 0.06 | 99.55 | 35.88 | 41.86 22.26 |
| 3S112b core | 51.30 | 0.78 | 2.23 | 9.02 | 1.62 | 0.34 | 14.98 | 18.91 | 0.38 | 0.00 | 99.56 | 40.18 | 44.29 15.53 |
| 3S112b edge | 51.59 | 0.66 | 2.14 | 9.38 | 0.85 | 0.33 | 15.75 | 18.08 | 0.29 | 0.14 | 99.19 | 38.02 | 46.05 15.93 |
| 3S112b rim | 52.88 | 0.75 | 3.12 | 13.08 | 0.00 | 0.39 | 12.73 | 16.77 | 0.78 | 0.03 | 100.53 | 37.27 | 39.36 23.37 |
| 3S112c core | 53.50 | 0.37 | 1.08 | 17.45 | 1.15 | 0.66 | 24.70 | 1.63 | 0.02 | 0.00 | 100.55 | 3.25 | 68.55 28.20 |
| 3S112c rim | 52.14 | 0.54 | 0.53 | 23.38 | 0.44 | 0.63 | 20.25 | 1.92 | 0.06 | 0.00 | 99.89 | 3.93 | 57.69 38.39 |
| 3S112d core | 53.88 | 0.30 | 0.86 | 17.04 | 1.48 | 0.51 | 25.04 | 1.89 | 0.02 | 0.02 | 101.03 | 3.75 | 69.08 27.17 |
| 3S112d rim | 52.22 | 0.62 | 0.85 | 20.15 | 0.00 | 0.59 | 18.11 | 7.14 | 0.12 | 0.03 | 99.82 | 14.71 | 51.92 33.37 |
| 3S112e core | 52.57 | 0.54 | 1.39 | 9.14 | 0.43 | 0 38 | 16.15 | 18 49 | 0 20 | 0 02 | 00 38 | 38 21 | 46.42 15 37 |
| 3S112e rim | 51.69 | 0 72 | 2 40 | 7 00 | 0.90 | 0 25 | 15 65 | 19 46 | 0 30 | 0 12 | 00 47 | 40 83 | 45 68 13 49 |
| 3S112f core | 54 03 | 0 30 | 1 15 | 17 22 | 0 00 | 0 /3 | 25 15 | 1 86 | 0 01 | 0 00 | 100 25 | 3 67 | 40 11 27 22 |
| 30112f core | 57 78 | 0.37 | 0.06 | 17 03 | 1 20 | 0.43 | 2/ 49 | 1 99 | 0.05 | 0.00 | 100.23 | 3.07 | 49 79 37 /4 |
| 3S11/ alono | 52 26 | 0.30 | 0.90 | 20.07 | 1 74 | 0.33 | 24.00 | 7.01 | 0.05 | 0.04 | 00.0/ | 1 12 | 60 07 7/ 04 |
| 7c11/ | 52.20 | 0.15 | 0.07 | 20.93 | 0.00 | 0.00 | 21.33 | 2.01 | 0.05 | 0.00 | 99.94 | 4.12 | 50.92 34.90 |
| 7011/0 0000 | 52.00 | 0.14 | 0.50 | 23.31 | 0.00 | 0.98 | 20.90 | 1.05 | 0.00 | 0.00 | YY.04 | 2.10 | 57.20 30.04 |
| J51148 COPE | 52.34 | 0.10 | 0.36 | 24.21 | 0.59 | 1.15 | 20.18 | 0.89 | 0.05 | 0.00 | 77.85 | 1.85 | 57.58 40.59 |
| 551148 rim | 52.35 | 0.17 | 0.70 | 22.08 | 0.00 | 0.76 | 21.28 | 0.99 | 0.01 | 0.00 | 98.34 | 2.05 | 61.15 36.83 |
| 55115 50µm | 52.76 | 0.48 | 1.06 | 20.86 | 0.00 | 0.72 | 22.19 | 1.78 | 0.01 | 0.01 | 99.86 | 3.60 | 62.37 34.04 |
| 3S115a core | 53.49 | 0.35 | 1.23 | 18.56 | 1.43 | 0.74 | 24.07 | 1.47 | 0.04 | 0.00 | 101.38 | 2.94 | 66.93 30.13 |
| 3S115a rim | 53.94 | 0.32 | 0.98 | 18.36 | 0.34 | 0.52 | 24.68 | 1.37 | 0.04 | 0.04 | 100.58 | 2.72 | 68.07 29.22 |
| 3S115b core | 52.11 | 0.70 | 1.81 | 9.76 | 0.39 | 0.44 | 14.82 | 19.38 | 0.32 | 0.00 | 99.73 | 40.40 | 42.99 16.61 |
| 3S115b rim | 52.44 | 0.64 | 1.65 | 9.71 | 0.74 | 0.33 | 15.09 | 19.29 | 0.35 | 0.00 | 100.22 | 40.08 | 43.63 16.29 |
| 3S115c core | 52.51 | 0.66 | 1.81 | 9.54 | 0.88 | 0.45 | 15.14 | 19.38 | 0.34 | 0.03 | 100.73 | 40.17 | 43.66 16.18 |
| 3S115c rim | 52.24 | 0.36 | 0.40 | 23.26 | 0.00 | 0.99 | 18.26 | 3.99 | 0.10 | 0.04 | 99.62 | 8.25 | 52.57 39.18 |
| 3S122 50µm | 51.10 | 0.16 | 1.06 | 25.49 | 0.89 | 1.09 | 18.68 | 1.01 | 0.02 | 0.00 | 99.50 | 2.11 | 54.41 43.48 |

| | Si02 | TiO2 | Al203 Fe | 0 Fe203 | HnD | MgO | CaO | Na20 | Cr203 | TOTAL | Mo | En | Fs |
|---------------------|-------|------|-----------|---------|------|-------|-------|------|-------|--------|-------|---------|-------|
| 3S122a core | 51.07 | 0.10 | 0.37 26.0 | 0.00 | 1.37 | 18.01 | 1.04 | 0.03 | 0.02 | 98.01 | 2.19 | 52.78 4 | 5.03 |
| 3S122a rim | 52.64 | 0.09 | 0.93 21.6 | 6 0.26 | 0.82 | 21.99 | 1.03 | 0.01 | 0.02 | 99.45 | 2.10 | 62.21 3 | 5.70 |
| 3S123 glom | 53.72 | 0.33 | 1.00 18.0 | 8 0.70 | 0.64 | 24.46 | 1.56 | 0.05 | 0.00 | 100.53 | 3.12 | 67.78 2 | 9.11 |
| 3S123 core | 51.31 | 0.67 | 2.23 7.4 | 5 2.04 | 0.34 | 14.95 | 20.40 | 0.30 | 0.00 | 99.67 | 43.15 | 43.99 | 12.86 |
| 3S123 50µm | 51.96 | 0.57 | 1.76 8.3 | 9 1.54 | 0.38 | 15.11 | 19.80 | 0.34 | 0.00 | 99.84 | 41.54 | 44.10 | 14.36 |
| 3S123 50µm | 53.76 | 0.30 | 0.63 18.5 | 9 0.68 | 0.57 | 24.11 | 1.75 | 0.04 | 0.00 | 100.43 | 3.49 | 66.75 2 | 29.77 |
| 3S123 glom | 52.02 | 0.70 | 1.92 8.7 | 9 0.99 | 0.34 | 14.84 | 20.25 | 0.28 | 0.03 | 100.15 | 42.17 | 42.99 | 14.84 |
| 3S123a core | 53.50 | 0.33 | 0.94 17.3 | 9 1.83 | 0.54 | 24.74 | 1.72 | 0.01 | 0.00 | 101.00 | 3.43 | 68.64 2 | 27.93 |
| 3S123a rim | 53.58 | 0.36 | 1.08 18.4 | 1 1.07 | 0.52 | 24.31 | 1.61 | 0.01 | 0.02 | 100.99 | 3.21 | 67.36 | 29.44 |
| 3S123b core | 52.07 | 0.56 | 1.76 8.4 | 3 1.32 | 0.32 | 15.02 | 20.06 | 0.33 | 0.01 | 99.88 | 41.97 | 43.73 | 14.30 |
| 3S123b rim | 51.89 | 0.58 | 1.67 8.1 | 7 1.73 | 0.40 | 15.02 | 20.23 | 0.28 | 0.02 | 99.99 | 42.30 | 43.70 | 14.00 |
| 3S131 20µm | 51.83 | 0.53 | 0.78 20.9 | 6 0.00 | 0.58 | 19.09 | 5.03 | 0.07 | 0.06 | 98.93 | 10.39 | 54.86 3 | 34.75 |
| 3S131 50µm | 52.77 | 0.48 | 0.76 19.9 | 4 0.05 | 0.53 | 21.17 | 3.92 | 0.07 | 0.00 | 99.68 | 7.94 | 59.68 | 32.38 |
| 3S131 50µm | 52.32 | 0.35 | 0.44 22.5 | 2 0.76 | 0.62 | 19.18 | 4.19 | 0.04 | 0.01 | 100.43 | 8.56 | 54.52 3 | 36.92 |
| 3S131a core | 50.03 | 1.04 | 3.96 6.7 | 3 2.51 | 0.20 | 14.57 | 20.54 | 0.33 | 0.02 | 99.93 | 44.44 | 43.85 | 11.72 |
| 3S131a rim | 51.80 | 0.70 | 1.49 11.0 | 3 1.12 | 0.36 | 15.61 | 17.38 | 0.24 | 0.02 | 99.75 | 36.22 | 45.25 | 18.54 |
| 3S131a rim | 51.37 | 0.97 | 1.45 13.1 | 4 0.00 | 0.39 | 14.11 | 17.31 | 0.31 | 0.00 | 99.05 | 36.44 | 41.33 | 22.23 |
| 3S131b core | 50.60 | 0.97 | 3.70 7.4 | 9 1.25 | 0.27 | 14.69 | 20.37 | 0.29 | 0.09 | 99.70 | 43.47 | 43.62 | 12.92 |
| 3S131b rim | 51.49 | 0.89 | 1.25 14.4 | 1 0.13 | 0.47 | 13.36 | 17.45 | 0.29 | 0.00 | 99.74 | 36.62 | 39.00 | 24.38 |
| 3S134 core | 51.65 | 0.35 | 1.19 13.3 | 6 0.93 | 0.74 | 12.80 | 18.57 | 0.30 | 0.02 | 99.90 | 39.19 | 37.57 | 23.24 |
| 3S134 50µm | 51.70 | 0.24 | 0.55 26.5 | 1 0.01 | 1.10 | 18.15 | 1.58 | 0.01 | 0.02 | 99.86 | 3.27 | 52.17 | 44.56 |
| 3S134a core | 51.42 | 0.19 | 0.45 26.0 | 4 0.33 | 1.16 | 18.09 | 1.63 | 0.02 | 0.00 | 99.33 | 3.38 | 52.40 | 44.22 |
| 3S134a rim | 52.16 | 0.33 | 0.88 24.0 | 5 0.23 | 0.97 | 19.80 | 1.65 | 0.05 | 0.00 | 100.12 | 3.39 | 56.51 | 40.10 |
| 3S134b core | 51.15 | 0.31 | 0.75 28.3 | 0.37 | 1.16 | 16.64 | 1.66 | 0.04 | 0.00 | 100.37 | 3.47 | 48.41 | 48.12 |
| 3 S134 b rim | 51.55 | 0.21 | 0.67 26.7 | 2 0.06 | 1.22 | 17.73 | 1.61 | 0.05 | 0.00 | 99.81 | 3.34 | 51.29 | 45.37 |
| 3S134c glom | 51.98 | 0.32 | 1.03 14.0 | 3 0.00 | 0.66 | 12.50 | 18.62 | 0.29 | 0.00 | 99.41 | 39.22 | 36.62 | 24.16 |
| 3 S 134c rim | 51.15 | 0.41 | 1.37 13.2 | 8 0.95 | 0.78 | 12.38 | 18.60 | 0.34 | 0.01 | 99.26 | 39.74 | 36.79 | 23.47 |
| 3S134d core | 51.89 | 0.37 | 0.90 25.3 | 4 0.88 | 0.93 | 19.07 | 1.48 | 0.05 | 0.01 | 100.92 | 3.06 | 54.67 | 42.28 |
| 3S134d rim | 51.40 | 0.22 | 0.62 26.2 | 3 0.27 | 1.13 | 17.92 | 1.61 | 0.06 | 0.00 | 99.46 | 3.35 | 52.04 | 44.61 |
| 3S137 core | 52.30 | 0.34 | 1.47 20.3 | 8 1.22 | 0.80 | 22.11 | 1.65 | 0.03 | 0.00 | 100.30 | 3.36 | 62.84 | 33.80 |
| 3S137a core | 51.32 | 0.52 | 1.53 8.6 | 0 2.79 | 0.50 | 14.37 | 19.96 | 0.34 | 0.01 | 99.93 | 42.41 | 42.48 | 15.10 |
| 3S137a edge | 51.48 | 0.52 | 1.80 9.8 | 7 1.67 | 0.50 | 13.70 | 19.78 | 0.41 | 0.00 | 99.73 | 42.14 | 40.60 | 17.26 |
| 3S137a rim | 52.17 | 0.27 | 0.89 21.7 | 0 1.09 | 0.81 | 21.24 | 1.65 | 0.03 | 0.01 | 99.85 | 3.37 | 60.59 | 36.04 |
| LI-A core | 50.55 | 0.13 | 0.22 34.0 | 3 0.33 | 1.91 | 12.79 | 1.37 | 0.01 | 0.00 | 101.33 | 2.89 | 37.68 | 59.43 |
| LI-A edge | 49.53 | 0.12 | 0.30 31.2 | 6 3.28 | 2.14 | 13.45 | 1.41 | 0.03 | 0.07 | 101.58 | 3.06 | 40.48 | 56.47 |
| LI-a rim | 49.75 | 0.04 | 0.22 23.8 | 8 1.37 | 3.52 | 7.22 | 14.52 | 0.14 | 0.00 | 100.63 | 31.58 | 21.83 | 46.59 |
| LI-B edge | 49.76 | 0.09 | 0.39 32.5 | 5 1.68 | 1.88 | 12.99 | 1.45 | 0.02 | 0.01 | 100.84 | 3.13 | 38.94 | 57.93 |
| LI-b rim | 49.56 | 0.11 | 0.67 20.2 | 3 2.31 | 2.50 | 6.87 | 17.86 | 0.32 | 0.04 | 100.47 | 39.53 | 21.15 | 39.32 |
| LI-C core | 52.57 | 0.30 | 0.88 11.2 | 3 1.48 | 0.78 | 14.28 | 18.67 | 0.38 | 0.00 | 100.56 | 38.97 | 41.46 | 19.57 |
| LI-c rim | 52.12 | 0.33 | 0.96 10.3 | 8 1.61 | 0.66 | 14.02 | 19.29 | 0.40 | 0.00 | 99.77 | 40.69 | 41.14 | 18.18 |

Appendix C-3: Composition of amphiboles determined through electron microprobe analysis. Capitol letter suffix on sample number indicates analysis was from the core of the crystal, lower case suffix indicates rim was analyzed. Size of the crystal is approximate. Iron partitioning and cation abundances determined through program AMPREC.FOR (Spear and Kimball, 1984), using the "average Fe" algorithm. Water calculated on the basis of 2(OH, F, Cl) and 15 cations; F and Cl abundances by microprobe analysis. BP samples from the top (t), base (b) or lower reworked zone (l) of the Bend Pumice (Hill, 1985), whole rock analysis 3SXBP in appendix D-1.

| | 3\$009 | 3\$053A | 3S053a | 3S053B | 3S053b | 3S053C | 3S053c | 3S053D | 3S053E | 3S060A | 3\$060a | 3\$060B | 3S060C | 3\$060c | 3\$060D | 3S060d | 3S060E |
|---------|--------|---------|--------|--------|--------|--------|--------|--------|--------|--------|---------|---------|--------|---------|---------|--------|--------|
| Wt. % | 100µm | 250µm | rim | 100µm | rim | 200µm | rim | 100µm | 50µm | 250µm | rim | 200µm | 250µm | rim | >250µm | rim | 200µm |
| Si02 | 43.18 | 44.18 | 43.45 | 47.80 | 47.29 | 44.67 | 44.22 | 46.01 | 45.52 | 42.68 | 41.41 | 43.23 | 42.49 | 42.79 | 47.99 | 48.08 | 47.23 |
| Ti02 | 2.76 | 1.47 | 1.81 | 1.30 | 1.98 | 2.08 | 2.17 | 2.18 | 1.89 | 3.17 | 3.78 | 3.34 | 3.38 | 2.79 | 1.13 | 1.20 | 1.63 |
| AL203 | 9.19 | 9.94 | 9.28 | 5.57 | 8.03 | 9.40 | 9.85 | 7.76 | 7.92 | 11.77 | 11.51 | 10.37 | 10.62 | 10.64 | 5.99 | 5.96 | 6.98 |
| FeO | 12.32 | 12.74 | 10.19 | 14.74 | 15.77 | 12.93 | 12.06 | 11.52 | 11.52 | 11.61 | 8.73 | 8.95 | 9.12 | 8.07 | 12.22 | 12.10 | 11.23 |
| Fe203 | 4.32 | 3.32 | 6.10 | 2.45 | 0.00 | 1.25 | 1.66 | 4.50 | 5.23 | 0.97 | 5.70 | 5.13 | 5.22 | 6.23 | 4.77 | 4.48 | 5.35 |
| FeO* | 16.21 | 15.73 | 15.68 | 16.95 | 15.77 | 14.05 | 13.55 | 15.57 | 16.22 | 12.48 | 13.86 | 13.57 | 13.82 | 13.68 | 16.52 | 16.13 | 16.05 |
| MnO | 0.32 | 0.33 | 0.20 | 0.48 | 0.28 | 0.19 | 0.18 | 0.25 | 0.37 | 0.16 | 0.19 | 0.19 | 0.23 | 0.28 | 0.60 | 0.55 | 0.32 |
| MgO | 11.98 | 13.10 | 13.71 | 13.48 | 13.12 | 13.73 | 14.11 | 13.21 | 13.12 | 13.81 | 13.28 | 13.89 | 13.74 | 13.96 | 13.99 | 14.09 | 13.34 |
| Ca0 | 10.32 | 10.77 | 10.94 | 10.87 | 10.71 | 10.73 | 10.66 | 10.91 | 10.73 | 11.24 | 11.20 | 11.13 | 11.17 | 11.35 | 10.33 | 10.39 | 10.90 |
| Na20 | 2.30 | 1.57 | 1.81 | 0.98 | 1.56 | 1.60 | 1.67 | 1.49 | 1.45 | 1.89 | 2.01 | 1.94 | 1.90 | 2.02 | 0.95 | 1.04 | 1.10 |
| K20 | 0.44 | 0.34 | 0.43 | 0.74 | 0.41 | 0.29 | 0.30 | 0.39 | 0.36 | 0.43 | 0.36 | 0.37 | 0.41 | 0.40 | 0.29 | 0.30 | 0.44 |
| H20 | 1.68 | 1.74 | 1.88 | 1.96 | 1.92 | 1.91 | 1.90 | 1.85 | 1-71 | 1.92 | 1.93 | 1.95 | 1.91 | 1.97 | 2.00 | 1.96 | 1.96 |
| F | 0.59 | 0.54 | 0.21 | 0.04 | 0.24 | 0.21 | 0.25 | 0.34 | 0.60 | 0.24 | 0.17 | 0.17 | 0.22 | 0.10 | 0.00 | 0.12 | 0.12 |
| CL | 0.07 | 0.09 | 0.08 | 0.19 | 0.09 | 0.05 | 0.02 | 0.06 | 0.07 | 0.02 | 0.01 | 0.03 | 0.01 | 0.03 | 0.12 | 0.07 | 0.07 |
| Total | 99.47 | 100.14 | 100.10 | 100.59 | 101.41 | 99.02 | 99.03 | 100.47 | 100.49 | 99.90 | 100.29 | 100.68 | 100.42 | 100.63 | 100.37 | 100.33 | 100.67 |
| | | | | | | | | | | | | | | | | | |
| | | | | | | | | | | | | | | | | | |
| Cations | 5 | | | | | | | | | 4 97 | 6 00 | (20 | (22 | 1 21 | 7 04 | 7 00 | (00 |
| Si | 6.46 | 6.53 | 6.42 | 7.04 | 6.89 | 6.62 | 6.54 | 6.74 | 6.69 | 6.27 | 6.08 | 6.29 | 0.22 | 0.24 | 7.01 | 7.02 | 0.00 |
| Ti | 0.31 | 0.16 | 0.20 | 0.14 | 0.22 | 0.23 | 0.24 | 0.24 | 0.21 | 0.35 | 0.42 | 0.57 | 0.37 | 0.51 | 0.12 | 0.15 | 0.18 |
| Aliv | 1.54 | 1.47 | 1.58 | 0.96 | 1.12 | 1.38 | 1.46 | 1.26 | 1.31 | 1.75 | 1.92 | 1./1 | 1.78 | 1./6 | 0.99 | 0.98 | 1.12 |
| Alvī | 0.09 | 0.26 | 0.03 | 0.01 | 0.26 | 0.27 | 0.26 | 0.08 | 0.07 | 0.31 | 0.07 | 0.07 | 0.05 | 0.07 | 0.04 | 0.05 | 0.08 |
| Fe2+ | 1.54 | 1.57 | 1.27 | 1.81 | 1.92 | 1.61 | 1.50 | 1.41 | 1.41 | 1.43 | 1.07 | 1.09 | 1.12 | 0.98 | 1.49 | 1.47 | 1.3/ |
| Fe3+ | 0.49 | 0.37 | 0.67 | 0.28 | 0.00 | 0.14 | 0.18 | 0.50 | 0.59 | 0.11 | 0.64 | 0.56 | 0.57 | 0.69 | 0.55 | 0.50 | 0.59 |
| Mn | 0.04 | 0.04 | 0.03 | 0.06 | 0.04 | 0.02 | 0.02 | 0.03 | 0.05 | 0.02 | 0.02 | 0.02 | 0.03 | 0.04 | 0.07 | 0.07 | 0.04 |
| Mg | 2.67 | 2.89 | 3.02 | 2.96 | 2.85 | 3.03 | 3.11 | 2.88 | 2.88 | 3.02 | 2.90 | 5.01 | 5.00 | 5.05 | 5.05 | 5.07 | 2.90 |
| Са | 1.66 | 1.71 | 1.73 | 1.72 | 1.67 | 1.70 | 1.69 | 1.71 | 1.69 | 1.77 | 1.76 | 1.74 | 1.75 | 1.// | 1.02 | 1.05 | 1.70 |
| Na(M4) | 0.21 | 0.00 | 0.06 | 0.02 | 0.05 | 0.00 | 0.00 | 0.15 | 0.12 | 0.00 | 0.15 | 0.14 | 0.10 | 0.12 | 0.08 | 0.09 | 0.15 |
| Na(A) | 0.67 | 0.45 | 0.46 | 0.26 | 0.39 | 0.46 | 0.48 | 0.28 | 0.29 | 0.54 | 0.45 | 0.41 | 0.44 | 0.45 | 0.19 | 0.21 | 0.10 |
| ĸ | 0.08 | 0.06 | 0.08 | 0.14 | 0.08 | 0.06 | 0.06 | 0.07 | 0.07 | 0.08 | 0.07 | 0.07 | 0.08 | 0.07 | 0.05 | 0.06 | 0.05 |
| OH | 1.70 | 1.73 | 1.88 | 1.93 | 1.87 | 1.89 | 1.88 | 1.83 | 1.70 | 1.88 | 1.92 | 1.91 | 1.90 | 1.95 | 1.97 | 1.95 | 1.95 |
| CL | 0.02 | 0.02 | 0.02 | 0.05 | 0.02 | 0.01 | 0.01 | 0.02 | 0.02 | 0.01 | 0.00 | 0.01 | 0.00 | 0.01 | 0.03 | 0.02 | 0.02 |
| F | 0.28 | 0.25 | 0.10 | 0.02 | 0.11 | 0.10 | 0.12 | 0.16 | 0.28 | 0.11 | 0.08 | 0.08 | 0.10 | 0.05 | 0.00 | 0.06 | 0.06 |

| | | | | | | | | | | | | | | | | | 1 |
|----------|-------|---------|---|--------|--------|--------|-------|--------|--------|--------|-------|--------|--------|--------|--------|--------|--------|
| | 38090 | 3\$090a | 3 \$090a | BPtA | BPta | BPtB | BPtb | BPtC | BPtc | BPtD | BPtd | 8PbA | BPba | BPbB | BPbb | BPbC | BPbc |
| Wt. % | glom | glom | 2nd pt | 250µm | rim | >250µm | rim | 200µm | rim | 150µm | rim | 250µm | rim | 200µm | rîm | 250µm | rim |
| Si02 | 41.45 | 42.21 | 42.72 | 43.17 | 42.54 | 42.10 | 41.35 | 42.15 | 43.20 | 41.42 | 41.35 | 41.70 | 42.18 | 41.71 | 42.47 | 41.64 | 40.55 |
| Ti02 | 3.72 | 3.88 | 3.75 | 2.60 | 3.15 | 3.52 | 3.34 | 3.29 | 3.13 | 3.20 | 3.29 | 3.86 | 3.45 | 3.02 | 2.63 | 3.03 | 3.53 |
| AL203 | 10.86 | 11.00 | 10.78 | 8.58 | 9.03 | 9.39 | 9.34 | 10.00 | 10.05 | 9.75 | 9.65 | 10.53 | 10.07 | 9.70 | 8.99 | 9.36 | 9.74 |
| FeO | 8.08 | 8.26 | 12.05 | 19.81 | 21.31 | 19.21 | 19.03 | 15.92 | 19.14 | 18.25 | 17.51 | 14.22 | 14.34 | 20.01 | 19.62 | 19.44 | 22.40 |
| Fe203 | 4.03 | 3.93 | 0.00 | 3.00 | 3.54 | 2.37 | 2.61 | 3.62 | 0.00 | 3.30 | 3.71 | 3.95 | 4.24 | 4.24 | 4.15 | 4.12 | 3.08 |
| Fe0* | 11.70 | 11.80 | 12.05 | 22.51 | 24.50 | 21.35 | 21.39 | 19.18 | 19.14 | 21.22 | 20.85 | 17.78 | 18.15 | 23.83 | 23.36 | 23.14 | 25.16 |
| MnO | 0.16 | 0.23 | 0.13 | 0.71 | 0.70 | 0.59 | 0.72 | 0.52 | 0.52 | 0.64 | 0.71 | 0.37 | 0.51 | 0.80 | 0.72 | 0.70 | 0.75 |
| Maû | 14 47 | 14 47 | 14 22 | 9.25 | 7.40 | 9.28 | 8.91 | 10.81 | 10.06 | 9.16 | 8.89 | 10.83 | 10.84 | 7.78 | 8.54 | 7.54 | 6.13 |
| Call | 10 96 | 10 90 | 11 05 | 9 67 | 9 44 | 9 81 | 9 71 | 10.17 | 10.06 | 9.93 | 9.79 | 10.30 | 10.11 | 9.34 | 9.39 | 9.69 | 9.69 |
| Na20 | 2.27 | 2.50 | 2.50 | 2.02 | 2.13 | 2.17 | 2.10 | 2.22 | 2.07 | 2.31 | 2.13 | 2.26 | 2.12 | 2.19 | 1.96 | 2.16 | 2.21 |
| K20 | 0 32 | 0 35 | 0.35 | 0 40 | 0 37 | 0.32 | 0.33 | 0.29 | 0.28 | 0.32 | 0.34 | 0.32 | 0.28 | 0.31 | 0.31 | 0.35 | 0.38 |
| H20 | 1 93 | 2.02 | 1.96 | 1.92 | 1.89 | 1.89 | 1.80 | 1.88 | 1.77 | 1.94 | 1.93 | 1.87 | 1.79 | 1.83 | 1.88 | 1.82 | 1.88 |
| F | 0 14 | 0 00 | 0 13 | 0.08 | 0 10 | 0.14 | 0.26 | 0.21 | 0.46 | 0.00 | 0.00 | 0.24 | 0.39 | 0.22 | 0.13 | 0.22 | 0.05 |
| ĊI. | 0 03 | 0.03 | 0.04 | 0.08 | 0.08 | 0.06 | 0.06 | 0.04 | 0.05 | 0.05 | 0.05 | 0.01 | 0.03 | 0.07 | 0.07 | 0.03 | 0.07 |
| Total | 98 40 | 00 78 | 00 60 | 101 29 | 101 70 | 100.85 | 99.56 | 101.10 | 100.79 | 100.26 | 99.32 | 100.47 | 100.32 | 101.21 | 100.85 | 100.08 | 100.45 |
| locat | | | ,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,, | | | | | | | | | | | | | | |
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| Cations | | | | | | | | | | | | | | | | | |
| ci | 6 16 | 6 18 | 6 20 | 6 52 | 6 15 | 6 37 | 6 35 | 6 20 | 6 1.8 | 6 31 | 6 33 | 6 23 | 6 31 | 6 34 | 6 45 | 6 30 | 6 20 |
| эт т; | 0.10 | 0.10 | 0.42 | 0.30 | 0.36 | 0.40 | 0.30 | 0.37 | 0 35 | 0.37 | 0.38 | 0.43 | 0.30 | 0.34 | 0 30 | 0 35 | 0.41 |
| Alix | 1.8/ | 1 82 | 1 70 | 1 48 | 1 55 | 1 63 | 1 65 | 1 71 | 1 52 | 1 70 | 1 67 | 1 78 | 1 69 | 1 66 | 1 55 | 1 61 | 1 72 |
| Alui | 0.04 | 0.08 | 0 17 | 0.04 | 0.07 | 0.05 | 0.05 | 0.05 | 0.26 | 0.05 | 0 08 | 0 08 | 0.08 | 0.08 | 0.06 | 0 00 | 0.06 |
| Eo2+ | 1 00 | 1 01 | 1 .0 | 2 50 | 2 72 | 2 44 | 2 44 | 1 00 | 2 40 | 2 34 | 2 24 | 1 77 | 1 79 | 2 55 | 2 49 | 2 48 | 2 89 |
| For the | 0.46 | 0.44 | 0.00 | 0 35 | 0 30 | 0 27 | 0 31 | 0 41 | 0 00 | 0 37 | 0 43 | 0 45 | 0 48 | 0 49 | 0 48 | 0 49 | 0 37 |
| Mo | 0.40 | 0.44 | 0.00 | 0.00 | 0.00 | 0.08 | 0.00 | 0.07 | 0.00 | 0.08 | 0.00 | 0.05 | 0.40 | 0.10 | 0.40 | 0 00 | 0 10 |
| Ma | 7 20 | 7 14 | 7 12 | 2 09 | 1 47 | 2 00 | 2 0/ | 2 10 | 2 25 | 2 08 | 2 03 | 2 41 | 2 42 | 1 76 | 1 03 | 1 73 | 1 42 |
| mg | 3.20 | 1 71 | 1 75 | 1 54 | 1.07 | 1 50 | 1 40 | 1 67 | 1 62 | 1 62 | 1 61 | 1 65 | 1 62 | 1.52 | 1.53 | 1 50 | 1 61 |
| | 0.11 | 0.14 | 0.0/ | 0.09 | 0.17 | 0.00 | 0.00 | 0.00 | 0.05 | 0 10 | 0.15 | 0 17 | 0.16 | 0.15 | 0 12 | 0 10 | 0 14 |
| Na(M4) | 0.11 | 0.10 | 0.04 | 0.00 | 0.17 | 0.09 | 0.09 | 0.09 | 0.05 | 0.10 | 0.15 | 0.17 | 0.10 | 0.15 | 0.12 | 0.19 | 0.14 |
| Na(A) | 0.55 | 0.55 | 0.07 | 0.51 | 0.40 | 0.55 | 0.04 | 0.55 | 0.00 | 0.50 | 0.49 | 0.49 | 0.40 | 0.49 | 0.40 | 0.40 | 0.02 |
| K | 0.06 | 0.07 | 0.07 | 0.07 | 0.07 | 0.00 | 0.07 | 0.00 | 0.00 | 1.00 | 1.00 | 1 00 | 1 01 | 4 00 | 1.00 | 1.00 | 1.04 |
| OH | 1.93 | 1.99 | 1.93 | 1.94 | 1.95 | 1.92 | 1.86 | 1.89 | 1.// | 1.99 | 1.99 | 1.08 | 1.01 | 1.00 | 1.92 | 1.09 | 1.90 |
| Cl | 0.01 | 0.01 | 0.01 | 0.02 | 0.02 | 0.02 | 0.02 | 0.01 | 0.01 | 0.01 | 0.01 | 0.01 | 0.01 | 0.02 | 0.02 | 0.01 | 0.02 |
| F | 0 07 | 0.00 | 0.06 | 0.04 | 0.05 | 0.07 | 0.15 | 0.10 | 0.22 | 0.00 | 0.00 | 0.11 | 0.18 | 0.11 | 0.06 | 0.11 | 0.05 |

| | BPbD | BPbd | BPbE | BPLA | BPla | BPIB | BPlb | BPIC | BPlc | LIA | LIa | LIB | LIP | LIC | LIC |
|---------|--------|--------------|--------------|--------|--------|--------|-------|--------|-------|--------|--------|--------|--------|--------|-------|
| Wt. % | >250µп | n rim | 100µm | >250µл | n rim | >250µm | rim | 250µm | rim | 200µm | rim | 250µm | rim | 200µm | rim |
| Si02 | 41.54 | 42.25 | 41.71 | 41.90 | 42.31 | 42.23 | 41.50 | 42.42 | 40.04 | 42.03 | 42.50 | 42.30 | 42.86 | 41.90 | 41.35 |
| T102 | 3.85 | 3.50 | 2.96 | 2.97 | 2.92 | 3.03 | 2.80 | 2.98 | 3.32 | 2.73 | 2.76 | 3.23 | 2.75 | 3.33 | 3.47 |
| AL203 | 10.29 | 10.22 | 9.44 | 9.17 | 8.78 | 8.60 | 8.70 | 8.79 | 9.88 | 8.61 | 9.23 | 9.28 | 9.81 | 9.86 | 10.07 |
| FeO | 14.95 | 13.38 | 20.70 | 19.12 | 19.65 | 22.86 | 21.89 | 20.66 | 20.10 | 21.10 | 18.92 | 17.33 | 19.82 | 17.00 | 16.55 |
| Fe203 | 3.16 | 4.44 | 3.44 | 3.46 | 2.70 | 2.21 | 3.01 | 2.55 | 4.25 | 2.61 | 5.26 | 2.63 | 0.00 | 3.87 | 4.60 |
| FeO* | 17.80 | 17.38 | 23.79 | 22.24 | 22.08 | 24.85 | 24.59 | 22.95 | 23.92 | 23.45 | 23.65 | 19.70 | 19.82 | 20.48 | 20.69 |
| MnO | 0.51 | 0.40 | 0.75 | 0.75 | 0.65 | 0.74 | 0.93 | 0.73 | 0.78 | 0.81 | 0.69 | 0.67 | 0.58 | 0.57 | 0.46 |
| MgO | 10.85 | 11.05 | 7.94 | 8.32 | 8.70 | 7.69 | 7.08 | 8.74 | 6.31 | 8.21 | 7.75 | 10.28 | 9.98 | 9.36 | 8.85 |
| CaO | 10.35 | 10.40 | 9.53 | 9.94 | 9.76 | 9.18 | 9.43 | 9.42 | 9.63 | 9.80 | 9.16 | 10.06 | 9.85 | 9.99 | 9.99 |
| Na2O | 2.19 | 2.22 | 2.22 | 2.06 | 2.21 | 2.08 | 1.97 | 2.12 | 2.25 | 2.04 | 2.18 | 2.19 | 2.12 | 2.27 | 2.13 |
| K20 | 0.32 | 0.36 | 0.33 | 0.35 | 0.28 | 0.30 | 0.35 | 0.30 | 0.35 | 0.38 | 0.32 | 0.30 | 0.32 | 0.29 | 0.32 |
| H20 | 1.84 | 1.73 | 1.82 | 1.85 | 1.69 | 1.87 | 1.83 | 1.86 | 1.86 | 1.85 | 1.88 | 1.83 | 1.85 | 1.82 | 1.76 |
| F | 0.27 | 0.52 | 0.24 | 0.17 | 0.49 | 0.14 | 0.13 | 0.18 | 0.03 | 0.14 | 0.10 | 0.29 | 0.26 | 0.28 | 0.38 |
| CL | 0.04 | 0.05 | 0.05 | 0.06 | 0.09 | 0.05 | 0.08 | 0.05 | 0.08 | 0.08 | 0.08 | 0.04 | 0.04 | 0.05 | 0.04 |
| Total | 100.17 | 100.52 | 101.12 | 100.10 | 100.21 | 100.97 | 99.70 | 100.80 | 98.90 | 100.37 | 100.83 | 100.42 | 100.23 | 100.59 | 99.96 |
| | | | | | | | | | | | | | | | |
| | | | | | | | | | | | | | | | |
| Cations | 5 | | | | | | | | | | | | | | |
| Si | 6.24 | 6.29 | 6.36 | 6.41 | 6.47 | 6.48 | 6.45 | 6.46 | 6.26 | 6.46 | 6.45 | 6.39 | 6.49 | 6.34 | 6.29 |
| Ti | 0.44 | 0.39 | 0.34 | 0.34 | 0.34 | 0.35 | 0.33 | 0.34 | 0.39 | 0.32 | 0.32 | 0.37 | 0.31 | 0.38 | 0.40 |
| Aliv | 1.76 | 1.71 | 1.64 | 1.59 | 1.53 | 1.52 | 1.55 | 1.54 | 1.73 | 1.54 | 1.55 | 1.61 | 1.51 | 1.68 | 1.71 |
| Alvi | 0.06 | 0.09 | 0.06 | 0.07 | 0.06 | 0.03 | 0.05 | 0.04 | 0.09 | 0.02 | 0.10 | 0.04 | 0.24 | 0.08 | 0.09 |
| Fe2+ | 1.88 | 1.66 | 2.64 | 2.44 | 2.50 | 2.95 | 2.84 | 2.64 | 2.62 | 2.73 | 2.40 | 2.20 | 2.51 | 2.13 | 2.11 |
| Fe3+ | 0.36 | 0.51 | 0.40 | 0.41 | 0.32 | 0.24 | 0.36 | 0.28 | 0.51 | 0.29 | 0.60 | 0.29 | 0.00 | 0.45 | 0.52 |
| Mn | 0.07 | 0.05 | 0.10 | 0.10 | 0.08 | 0.10 | 0.12 | 0.09 | 0.10 | 0.11 | 0.09 | 0.09 | 0.07 | 0.08 | 0.06 |
| Mg | 2.43 | 2.45 | 1.81 | 1.90 | 1.98 | 1.76 | 1.64 | 1.98 | 1.47 | 1.88 | 1.75 | 2.31 | 2.25 | 2.11 | 2.01 |
| Са | 1.67 | 1.55 | 1.56 | 1.63 | 1.60 | 1.51 | 1.57 | 1.54 | 1.62 | 1.62 | 1.49 | 1.63 | 1.60 | 1.62 | 1.63 |
| Na(M4) | 0.11 | 0.19 | 0.11 | 0.12 | 0.12 | 0.06 | 0.09 | 0.08 | 0.20 | 0.04 | 0.25 | 0.08 | 0.02 | 0.17 | 0.19 |
| Na(A) | 0.53 | 0.45 | 0.55 | 0.49 | 0.54 | 0.55 | 0.50 | 0.55 | 0.48 | 0.56 | 0.40 | 0.57 | 0.61 | 0.50 | 0.44 |
| K | 0.06 | 0.07 | 0.06 | 0.07 | 0.06 | 0.06 | 0.07 | 0.06 | 0.07 | 0.08 | 0.06 | 0.06 | 0.06 | 0.06 | 0.06 |
| 011 | | | | | | | | | _ / | | | | | | |
| UH | 1.86 | 1.74 | 1.87 | 1.90 | 1.74 | 1.92 | 1.92 | 1.90 | 1.96 | 1.91 | 1.93 | 1.85 | 1.87 | 1.85 | 1.81 |
| Cl | 1.86 | 1.74 0.01 | 1.87 0.01 | 1.90 | 1.74 | 1.92 | 1.92 | 1.90 | 1.96 | 1.91 | 1.93 | 1.85 | 1.87 | 1.85 | 0.01 |

Appendix C-4: Composition of spinels determined through microprobe analysis. FeO, Fe₂O₃, and mol fraction ülvospinel (XÜsp) calculated using the method of Carmichael (1967). Xlvd: sample had visible exsolution lamallae.

| | | Si 02 | Ti02 | A1203 | FeO | Fe203 | MnO | MgO | Cr203 | TOTAL | XÜsp |
|----------------|-----------------|--------------|-------|-------|-------|-------|------|-------|-------|--------|------|
| 3SX8P | | 0.10 | 14.53 | 1.21 | 42.48 | 38.56 | 0.83 | 0.57 | 0.00 | 98.28 | 42.4 |
| 3SX8P | | 0.05 | 14.08 | 1.21 | 43.01 | 41.10 | 0.82 | 0.46 | 0.00 | 100.73 | 39.9 |
| 3SX8P | 150µm | 0.14 | 14.66 | 1.28 | 43.43 | 39.63 | 0.83 | 0.50 | 0.13 | 100.59 | 42.0 |
| 3SX8P | 200µm | 0.11 | 14.42 | 1.35 | 43.24 | 40.31 | 0.93 | 0.50 | 0.11 | 100.97 | 41.0 |
| 3SX8P | 150µm | 0.10 | 14.66 | 1.29 | 43.50 | 39.66 | 0.75 | 0.50 | 0.07 | 100.52 | 41.8 |
| 3SX8P | 40 <i>u</i> m | 0.09 | 14.48 | 1.25 | 43 30 | 40 17 | 0.83 | 0 44 | 0.06 | 100 70 | 41 2 |
| 3SX8P | 150µm | 0.12 | 14.52 | 1.19 | 43.50 | 40 30 | 0.68 | 0 53 | 0 02 | 100 94 | 41 4 |
| 3SX8P | 50µm | 0.10 | 16.70 | 2.41 | 43 55 | 34 04 | 0.87 | 1 72 | 0.02 | 100.36 | 46.8 |
| 35003 | 5 opini | 0 08 | 13 67 | 2 80 | 40.65 | 60 80 | 0.67 | 2 10 | 0.04 | 100.30 | 38 0 |
| 35003 | | 0.05 | 14 23 | 2 75 | 40.05 | 40.07 | 0.40 | 2 / 2 | 0.04 | 101.05 | 30.0 |
| 30005 | xtvd | 0.05 | 12 76 | 1 00 | 40.71 | 40.21 | 0.65 | 1 08 | 0.06 | 08 64 | 34.2 |
| 30005 | xtvd | 0.10 | 11 63 | 2 56 | 30 32 | 41.00 | 0.50 | 1 37 | 0.00 | 00 21 | 30.0 |
| 30017 | 30.00 | 0.00 | 10 60 | 1 71 | /5 90 | 43.37 | 0.19 | 1.37 | 0.07 | 99.21 | 57.0 |
| 30077 | Jopan | 0.13 | 1/ 03 | 2 26 | 43.09 | 20.00 | 0.49 | 1.0 | 0.21 | 100 71 | 70.7 |
| 3020 | | 0.11 | 14.00 | 2.20 | 42.00 | 70 40 | 0.39 | 1.47 | 0.01 | 00.71 | JY.2 |
| 3020 | 100.00 | 0.05 | 14.22 | 2.10 | 41.04 | 30.00 | 0.00 | 1.00 | 0.00 | 90.33 | 40.5 |
| 35020 | 150µm | 0.10 | 17 44 | 2.00 | 40.99 | 39.22 | 0.49 | 1.00 | 0.01 | 99.03 | 40.8 |
| 35020 | 150µm | 0.00 | 13.00 | 2.24 | 40.23 | 41.27 | 0.60 | 2.18 | 0.03 | 100.27 | 38.2 |
| 39020 | 50m | 0.09 | 14.33 | 2.29 | 42.24 | 38.77 | 0.51 | 1.19 | 0.08 | 99.49 | 40.8 |
| 35028 | 100 <i>µ</i> .m | 0.06 | 14.51 | 2.20 | 41.08 | 38.44 | 0.63 | 1.85 | 0.05 | 98.88 | 41.2 |
| 35030 | | 0.05 | 18.39 | 1.40 | 44.15 | 31.01 | 0.44 | 2.18 | 0.96 | 98.56 | 52.1 |
| 35036 | | 0.11 | 16.54 | 1.85 | 45.48 | 35.53 | 0.53 | 0.56 | 0.00 | 100.58 | 46.9 |
| 3 \$042 | | 0.04 | 13.56 | 0.97 | 41.70 | 39.45 | 0.61 | 0.39 | 0.00 | 96.89 | 39.8 |
| 3S042 | 20m | 2.15 | 0.24 | 0.52 | 28.70 | 62.19 | 0.05 | 0.03 | 0.04 | 93.96 | 10.7 |
| 3S045 | 20 <i>µ</i> .m | 0.10 | 15.53 | 0.57 | 44.04 | 36.84 | 0.34 | 0.30 | 0.02 | 97.73 | 45.7 |
| 3S045 | 120µm | 0.07 | 2.03 | 20.60 | 21.50 | 20.02 | 0.32 | 9.38 | 24.04 | 97.95 | 5.2 |
| 3\$049 | 20 <i>µ</i> .m | 0.07 | 10.00 | 0.77 | 35.35 | 48.22 | 0.28 | 2.57 | 0.02 | 97.28 | 29.2 |
| 3S058 | 50µ.m | 0.09 | 8.49 | 1.82 | 37.58 | 51.29 | 0.54 | 0.86 | 0.07 | 100.74 | 24.3 |
| 3S061 | 120µm xivd | 0.07 | 13.44 | 2.57 | 41.04 | 42.58 | 0.50 | 2.02 | 0.02 | 102.24 | 36.9 |
| 3\$069 | 20µ.m | 0.17 | 22.37 | 0.14 | 48.82 | 21.56 | 0.11 | 1.00 | 0.60 | 94.77 | 66.9 |
| 3\$069 | 30µm | 0.07 | 1.78 | 3.51 | 27.85 | 50.79 | 0.43 | 3.19 | 11.52 | 99.14 | 5.2 |
| 3\$069 | 30µm | 0.04 | 1.44 | 2.26 | 27.06 | 61.95 | 0.51 | 3.27 | 2.84 | 99.37 | 4.2 |
| 3 \$074 | 30µm | 0.01 | 18.14 | 0.23 | 43.56 | 30.19 | 0.28 | 1.44 | 0.08 | 93.94 | 54.3 |
| 3\$079 | 100µm | 0.08 | 14.55 | 2.19 | 40.95 | 39.36 | 0.69 | 2.14 | 0.09 | 100.05 | 40.8 |
| 3S079 | 100µm | 0.09 | 13.97 | 3.11 | 39.87 | 40.63 | 0.70 | 2.82 | 0.03 | 101.21 | 38.5 |
| 3\$090 | 60µm | 1.19 | 15.35 | 2.78 | 40.34 | 32.08 | 0.44 | 1.81 | 0.03 | 94.02 | 52.9 |
| 3\$098 | 100µm | 0.07 | 17.49 | 1.35 | 45.59 | 35.85 | 0.86 | 1.17 | 0.00 | 102.38 | 48.4 |
| 3\$098 | 100µm | 0.08 | 17.32 | 1.30 | 45.38 | 35.73 | 0.72 | 1.15 | 0.01 | 101.68 | 48.3 |
| 3\$098 | 75µm | 0.08 | 17.66 | 1.38 | 45.82 | 35.20 | 0.85 | 1.05 | 0.01 | 102.05 | 49.0 |
| 3S112 | 30 <i>µ</i> .m | 0.07 | 20.46 | 1.53 | 47.45 | 24.66 | 0.38 | 1.20 | 2.15 | 97.89 | 58.7 |
| 35114 | 50µm | 0.07 | 9.63 | 1.84 | 37.93 | 49 74 | 0.52 | 1 43 | 0 02 | 101 18 | 27 1 |
| 35115 | 30 <i>µ</i> m | 0 11 | 23 68 | 0.65 | 51 17 | 20 20 | 0.37 | 0 34 | 0.00 | 06 52 | 60 8 |
| 30173 | 50µm | 0 05 | 1/ /0 | 2 17 | 12 11 | 30 61 | 0.7 | 1 / 2 | 0.00 | 100 63 | 40.6 |
| 30123 | 50µm | 0.05 | 14.47 | 2 00 | 12 03 | 30 30 | 0.47 | 1 20 | 0.00 | 100.03 | 40.0 |
| 30123 | 100um | 0.07 | 7/ 33 | 1 06 | 50 49 | 10 3/ | 0.30 | 1 20 | 0.00 | 07 21 | 70 7 |
| 30131 | 100µm | 0.07 | 24.33 | 0.07 | 57 17 | 15.07 | 0.43 | 1.20 | 0.12 | 91.21 | 70.2 |
| 30131 | 100µm | 0.07 | 19 55 | 1 99 | 72.12 | 70./0 | 0.47 | 1.10 | 0.15 | 90.33 | 12.2 |
| 30134 | 30um | 0.07 | 10.00 | 1.15 | 43.04 | 30.49 | 0.75 | 1.12 | 0.00 | YO.JU | 53.0 |
| 30134 | 100um | 0.15 | 10.20 | 1.10 | 43.32 | 31.32 | 0.82 | 0.88 | 0.00 | 91.09 | JJ.4 |
| 3513/ | i uuuuu | 0.09 | 19.89 | 1.04 | 40.98 | 30.57 | 0.76 | 1.41 | 0.01 | 100.75 | 22.0 |
| 55157 | 20µm | 0.10 | 20.05 | 0.86 | 47.45 | 50.61 | 0.75 | 1.25 | 0.00 | 101.07 | 56.2 |
| Lava | IS XIVO | 0.10 | 12.10 | 1.09 | 40.48 | 44.76 | 1.85 | 0.19 | 0.06 | 100.62 | 54.7 |
| Lava | is Xivd | 0.09 | 10.28 | 1.29 | 38.65 | 48.14 | 2.06 | 0.19 | 0.05 | 100.74 | 29.5 |

Appendix C-5: Composition of ilmenites determined through microprobe analysis. FeO, Fe₂O₃, and mol fraction of hematite component (XR₂O₃) calculated using the method of Carmichael (1967). N/a: not analyzed; xlvd: Sample had visible exsolution lamallae.

| | | Si 02 | TiOZ | A1203 | FeO | Fe203 | MnO | MgO | Cr203 | Nb205 | TOTAL | XR203 |
|----------------|------------------|--------------|-------|-------|-------|-------|------|------|-------|-------|--------|-------|
| 3S003 | 200 <i>µL</i> m | 0.00 | 44.25 | 0.00 | 36.64 | 17.70 | 0.50 | 1.50 | 0.06 | 0.07 | 100.72 | 16.73 |
| 3S003 | 100µum | 0.01 | 43.57 | 0.00 | 36.05 | 18.30 | 0.54 | 1.47 | 0.08 | 0.02 | 100.04 | 17.43 |
| 3 \$005 | >200µ.m | 0.03 | 48.90 | 0.00 | 39.96 | 6.31 | 0.63 | 1.92 | 0.11 | 0.08 | 97.94 | 6.16 |
| 3S005 | 300 <i>µ.</i> m | 0.00 | 48.90 | 0.00 | 39.98 | 7.83 | 0.74 | 1.86 | 0.00 | n/a | 99.30 | 7.42 |
| 3S 005 | 10µ.m | 0.06 | 49.62 | 0.00 | 40.98 | 6.55 | 0.69 | 1.69 | 0.00 | n/a | 99.59 | 6.19 |
| 3S 005 | | 0.01 | 49.07 | 0.00 | 39.97 | 8.99 | 0.80 | 1.88 | 0.02 | 0.10 | 100.84 | 8.40 |
| 3\$009 | xtvd | 0.05 | 26.07 | 0.02 | 18.91 | 45.99 | 1.45 | 1.77 | 0.00 | n/a | 94.26 | 46.84 |
| 3 \$028 | 50µ.m | 0.00 | 44.15 | 0.00 | 35.58 | 18.47 | 0.51 | 2.05 | 0.09 | 0.10 | 100.94 | 17.37 |
| 3 \$028 | 30µ.m | 0.01 | 44.41 | 0.00 | 35.16 | 17.64 | 0.68 | 2.31 | 0.06 | 0.02 | 100.27 | 16.63 |
| 3\$028 | 30 <i>µ</i> .m | 0.03 | 44.04 | 0.00 | 35.31 | 17.85 | 0.77 | 1.99 | 0.07 | 0.04 | 100.10 | 16.90 |
| 3S058 | 50µ.m | 0.00 | 42.68 | 0.00 | 34.73 | 19.24 | 0.76 | 1.60 | 0.05 | n/a | 99.06 | 18.44 |
| 3s079 | 15µm | 0.03 | 44.85 | 0.00 | 33.50 | 17.54 | 0.80 | 3.38 | 0.02 | 0.06 | 100.18 | 16.36 |
| 3S079 | 30µ.m | 0.00 | 44.64 | 0.00 | 33.00 | 17.77 | 0.93 | 3.48 | 0.01 | 0.00 | 99.83 | 16.62 |
| 3\$090 | 10 <i>µ</i> .m | 0.17 | 36.65 | 0.00 | 32.37 | 21.99 | 0.15 | 0.25 | 0.02 | 0.02 | 91.62 | 22.99 |
| 3\$098 | 100 <i>µ</i> .m | 0.01 | 48.60 | 0.00 | 39.90 | 8.87 | 1.08 | 1.53 | 0.02 | 0.18 | 100.19 | 8.39 |
| 3S114 | 50µ.m | 0.03 | 42.16 | 0.04 | 33.00 | 20.60 | 0.56 | 2.49 | 0.00 | n/a | 98.88 | 19.68 |
| 3S115 | 40 <i>µ</i> .m | 0.03 | 49.09 | 0.00 | 39.47 | 9.73 | 0.65 | 2.26 | 0.00 | 0.11 | 101.33 | 9.01 |
| 3S115 | 60µ.m | 0.04 | 49.03 | 0.00 | 39.15 | 10.11 | 0.64 | 2.41 | 0.02 | 0.00 | 101.40 | 9.36 |
| 3S123 | 125 <i>µ</i> .m | 0.05 | 45.79 | 0.07 | 35.99 | 14.99 | 0.50 | 2.66 | 0.04 | 0.00 | 100.09 | 14.18 |
| 3S123 | 30µ.m | 0.06 | 45.58 | 0.00 | 36.70 | 14.29 | 0.62 | 2.06 | 0.01 | 0.00 | 99.31 | 13.54 |
| 3S134 | 150µum | 0.00 | 48.51 | 0.00 | 39.45 | 7.95 | 0.92 | 1.82 | 0.00 | 0.07 | 98.73 | 7.58 |
| 3\$137 | ≈200 <i>µ</i> um | 0.05 | 48.02 | 0.00 | 38.23 | 11.08 | 0.93 | 2.25 | 0.00 | 0.10 | 100.66 | 10.34 |
| 3S137 | ≈200 <i>µ</i> .m | 0.03 | 47.90 | 0.00 | 38.35 | 11.46 | 0.97 | 2.10 | 0.00 | 0.03 | 100.84 | 10.68 |
| 3SXBP | 50µ.m | 0.02 | 40.39 | 0.00 | 33.58 | 23.16 | 0.96 | 1.03 | 0.04 | n/a | 99.19 | 22.31 |
| 3SXBP | | 0.03 | 48.17 | 0.02 | 35.40 | 10.09 | 0.76 | 4.02 | 0.00 | 0.00 | 98.49 | 9.50 |
| 3SX8P | | 0.02 | 46.49 | 0.08 | 33.51 | 11.34 | 0.91 | 4.18 | 0.03 | n/a | 96.56 | 11.00 |
| 3SX8P | 200µum, | 0.03 | 47.68 | 0.09 | 35.09 | 11.58 | 0.84 | 3.95 | 0.01 | n/a | 99.27 | 10.95 |
| 3SXBP | 100µ.m | 0.03 | 49.69 | 0.00 | 40.99 | 6.40 | 1.42 | 1.29 | 0.08 | 0.24 | 100.13 | 6.12 |
| 3SXB P | 50µ.m | 0.02 | 49.03 | 0.00 | 38.12 | 9.83 | 1.03 | 2.80 | 0.14 | 0.08 | 101.05 | 9.23 |
| Lava I | s. xivd | 0.00 | 46.51 | 0.00 | 37.49 | 12.64 | 3.47 | 0.48 | 0.08 | 0.17 | 100.83 | 12.03 |

Appendix D-1: Compositions and locations of Quaternary rocks from the Three Sisters area. Total is the original XRF total; major elements reported as normalized to 100% anhydrous. Blank space indicates element was not analysed, 0 indicates not detected by XRF, < indicates upper limit of element in this sample by INAA. La/Sm, Eu/Eu* and total HREE use chondrite normalized values (Anders and Ebihara, 1982). Label is the symbol used in geochemical plots.

| Sample | 3S001 | 3\$003 | 3\$005 | 3S007 | 35008 | 3 S009 | 3S010 | 3S010X | 3s011 | 3s012 |
|--------|-------|--------|--------|--------------|--------|---------------|--------|--------|--------|--------|
| SiO2 | 71.9 | 61.5 | 68.0 | 71.9 | 63.0 | 68.2 | 59.0 | 69.7 | 55.3 | 74.5 |
| T102 | 0.21 | 1.28 | 0.72 | 0.25 | 1.12 | 0.61 | 1.08 | 0.44 | 1.28 | 0.15 |
| A1203 | 16.4 | 17.6 | 15.6 | 15.3 | 16.7 | 16.1 | 17.4 | 15.1 | 17.6 | 14.1 |
| Fe0* | 2.33 | 6.53 | 4.24 | 2.38 | 5.62 | 3.44 | 7.10 | 3.83 | 7.78 | 1.68 |
| MnO | 0.10 | 0.22 | 0.09 | 0.08 | 0.13 | 0.09 | 0.13 | 0.11 | 0.14 | 0.06 |
| MgO | 0.12 | 1.50 | 0.45 | 0.47 | 1.73 | 0.87 | 3.02 | 0.29 | 4.39 | 0.17 |
| CaO | 0.82 | 5.85 | 2.12 | 1.41 | 4.32 | 2.35 | 6.52 | 1.59 | 8.05 | 0.71 |
| Na2O | 3.75 | 5.16 | 4.64 | 4.02 | 6.06 | 5.56 | 4.11 | 5.97 | 4.69 | 5.06 |
| K20 | 5.49 | 1.65 | 3.25 | 3.14 | 1.82 | 2.58 | 1.29 | 2.76 | 0.98 | 3.46 |
| P205 | 0.04 | 0.47 | 0.19 | 0.05 | 0.29 | 0.16 | 0.23 | 0.09 | 0.30 | 0.02 |
| TOTAL | 98.06 | 99.92 | 99.86 | 96.72 | 101.25 | 101.16 | 100.62 | 100.70 | 100.96 | 101.78 |
| Rb | 77 | 26 | 77 | 67 | 46 | 52 | 24 | 61 | 15 | 82 |
| Ba | 986 | 519 | 755 | 776 | 474 | 732 | 431 | 774 | 341 | 797 |
| Sr | 68 | 396 | 209 | 130 | 407 | 237 | 503 | 157 | 575 | 54 |
| Cs | 2.70 | 1.20 | 1.40 | 1.70 | 1.20 | 0.95 | 0.50 | 0.37 | 0.50 | 0.80 |
| Sc | 6.32 | 18.33 | 9.52 | 4.88 | 14.81 | 8.72 | 20.18 | 12.30 | 26.34 | 4.06 |
| V | 0 | 105 | 25 | 0 | 71 | 23 | 369 | 51 | 194 | 0 |
| Co | 4.93 | 11.60 | 5.87 | 1.54 | 8.11 | 3.72 | 18.00 | 1.27 | 23.80 | 0.54 |
| Ni | 18 | 7 | 12 | 13 | 4 | 8 | 13 | 10 | 16 | 12 |
| Cr | 4 | 4 | 3 | 2 | 0 | 0 | 14 | 0 | 63 | 0 |
| Zn | 69 | 108 | 66 | 69 | 70 | 65 | 62 | 105 | 59 | 88 |
| La | 27.4 | 20.8 | 25.2 | 21.8 | 18.9 | 25.9 | 16.7 | 25.5 | 15.6 | 23.0 |
| Ce | 62.7 | 49.8 | 57.2 | 44.9 | 38.0 | 57.7 | 33.1 | 58.9 | 34.1 | 48.1 |
| Nd | 26.1 | 23.9 | 27.1 | 20.7 | 21.3 | 26.0 | 18.5 | 31.0 | 17.8 | 20.8 |
| Sm | 5.69 | 6.13 | 6.27 | 4.71 | 5.28 | 6.02 | 4.34 | 7.15 | 4.54 | 4.26 |
| Eu | 0.75 | 1.71 | 1.49 | 0.78 | 1.39 | 1.40 | 1.27 | 1.66 | 1.46 | 0.59 |
| Tb | 0.85 | 0.88 | 0.97 | 0.71 | 0.81 | 0.88 | 0.65 | 1.23 | 0.68 | 0.70 |
| Yb | 4.0 | 2.9 | 3.7 | 3.1 | 2.8 | 3.7 | 2.1 | 4.8 | 2.3 | 3.1 |
| Lu | 0.56 | 0.43 | 0.51 | 0.48 | 0.45 | 0.53 | 0.34 | 0.72 | 0.36 | 0.47 |
| La/Sm | 3.04 | 2.14 | 2.54 | 2.92 | 2.26 | 2.72 | 2.43 | 2.25 | 2.17 | 3.41 |
| ΣHREE | 54.47 | 45.99 | 54.09 | 44.59 | 44.50 | 52.75 | 34.36 | 71.32 | 36.57 | 44.02 |
| Eu/Eu* | 0.40 | 0.86 | 0.72 | 0.50 | 0.80 | 0.71 | 0.89 | 0.68 | 0.98 | 0.41 |
| Zr | 264 | 203 | 293 | 253 | 173 | 277 | 158 | 285 | 152 | 218 |
| Hf | 7.1 | 5.2 | 7.9 | 6.1 | 4.6 | 7.5 | 3.8 | 8.5 | 3.7 | 6.1 |
| Ga | 23 | 20 | 24 | 19 | 19 | 20 | 20 | 20 | 19 | 22 |
| Y | 37 | 30 | 34 | 32 | 31 | 33 | 25 | 39 | 26 | 29 |
| Nb | 20.8 | 12.9 | 18.6 | 14.8 | 11.4 | 16.1 | 10.6 | 18.2 | 9.8 | 16.5 |
| Ta | 1.16 | 0.72 | 1.01 | 0.75 | 0.64 | 0.98 | 0.49 | 0.97 | 0.59 | 1.01 |
| Th | 7.6 | 3.3 | 7.5 | 5.5 | 4.1 | 5.8 | 2.2 | 5.9 | 1.9 | 7.7 |
| U | 2.5 | 1.2 | 2.6 | 1.8 | 1.4 | 1.9 | 0.8 | 1.8 | <2 | 2.2 |
| Series | TVC | TVC | TVC | TVC | BT | TVC | TVC | TVC | TVC | TVC |
| Label | + | + | + | + | + | + | + | + | + | + |

Lat 44.10908 44.11018 44.10745 44.06482 44.04606 44.07912 44.09950 44.09950 44.10699 44.10996 -Long 121.4869 121.4891 121.4631 121.2865 121.5324 121.5522 121.5401 121.5401 121.5326 121.5319

3S001 Pumice lapilli on top of the Bull Sp. Inlier, = Hydrated Bend Pumice 3S003 Pumcie from Shevlin Park Tuff at W. margin of Bull Sp. Inlier 3S005 Pumice from Desert Spring Tuff on E. flank of Bull Sp. Inlier 3S007 Pumice lapilli @ base Pilot Bu; less evolved than Bottle Ck. 3S024 3S008 Main older An flow, Tumalo Ck. valley 3S009 Summit of Bearwallow Bu. dome 3S010 Triangle Hill, nonvesiculated part of bomb: Flow=3S073 3S010X Hb-bearing xenolith from Triangle Hill ejecta 3S011 Flow 0.75 mi ENE of Triangle Hill, under 5900'Cinder cone 3S012 Dome @ 5400', 3/4 mi ENE Triangle Hill

| Sample | 3s014 | 3s01 5 | 3s017 | 3s018 | 3s019 | 3s022 | 3s023 | 3s024 | 3s027 | 38028 |
|--|--|--|--|--|--|--|--|--|--|---|
| Si02 | 67.6 | 57.4 | 58.4 | 71.6 | 67.7 | 56.0 | 67.8 | 71.4 | 65.9 | 58.5 |
| Ti02 | 0.30 | 1.01 | 1.15 | 0.23 | 0.66 | 1.17 | 0.63 | 0.26 | 0.92 | 1.51 |
| Al203 | 20.5 | 18.3 | 17.4 | 16.3 | 16.1 | 17.5 | 16.1 | 16.2 | 16.2 | 17.4 |
| FeO* | 3.03 | 6.69 | 7.25 | 2.36 | 3.64 | 7.16 | 3.59 | 2.51 | 4.67 | 7.88 |
| MnO | 0.10 | 0.12 | 0.13 | 0.07 | 0.10 | 0.13 | 0.10 | 0.08 | 0.11 | 0.16 |
| MgO | 0.20 | 3.30 | 3.13 | 0.22 | 0.88 | 5.08 | 0.90 | 0.22 | 1.05 | 2.26 |
| Ca0 | 1.03 | 7.70 | 6.59 | 1.02 | 2.50 | 7.64 | 2.48 | 1.16 | 2.82 | 5.36 |
| Na20 | 3.66 | 5.44 | 4.33 | 4.44 | 5.37 | 3.88 | 5.49 | 4.30 | 5.53 | 4.78 |
| K20 | 2.63 | 1.17 | 1.34 | 3.05 | 2.72 | 1.14 | 2.52 | 3.12 | 2.44 | 1.02 |
| P205 | 0.06 | 0.22 | 0.31 | 0.04 | 0.19 | 0.26 | 0.17 | 0.04 | 0.21 | 0.37 |
| TOTAL | 95.58 | 100.06 | 101.57 | 98.23 | 99.63 | 100.92 | 101.25 | 98.13 | 100.67 | 100.09 |
| Rb | 62 | 23 | 27 | 73 | 56 | 26 | 56 | 74 | 52 | 16 |
| Ba | 1253 | 379 | 470 | 859 | 739 | 312 | 698 | 830 | 687 | 503 |
| Sr | 163 | 536 | 525 | 125 | 253 | 429 | 249 | 113 | 271 | 498 |
| Cs | 1.60 | 0.60 | 1.20 | 2.00 | 2.10 | 1.00 | 0.70 | 2,10 | 0.50 | 0.80 |
| Sc | 6.03 | 21.60 | 21.24 | 4.86 | 8.77 | 22.30 | 8.22 | 4.99 | 13.21 | 21.82 |
| V | 0 | 186 | 182 | 17 | 20 | 183 | 32 | 1 | 61 | 159 |
| Co | 1.08 | 19.33 | 21.00 | 0.85 | 3.60 | 25.90 | 3.79 | 1.05 | 6.88 | 15.38 |
| Ni | 13 | 15 | 11 | 9 | 5 | 64 | 7 | 11 | 9 | 3 |
| Cr | 0 | 33 | 16 | 0 | 0 | 111 | <4 | 2 | 6 | 10 |
| Zn | 92 | 62 | 85 | 87 | 81 | 60 | 75 | 84 | 90 | 66 |
| La | 25.8 | 14.9 | 19.7 | 27.9 | 27.5 | 14.6 | 24.4 | 24.3 | 22.6 | 19.3 |
| Ce | 57.8 | 32.1 | 43.4 | 48.2 | 58.6 | 32.5 | 48.7 | 48.2 | 49.9 | 39.9 |
| Nd | 24.8 | 17.1 | 21.0 | 25.7 | 27.9 | 17.0 | 24.9 | 23.0 | 25.6 | 24.1 |
| Sm | 5.51 | 4.06 | 5.38 | 5.50 | 6.45 | 4.32 | 5.42 | 5.09 | 6.26 | 6.24 |
| Eu | 0.95 | 1.26 | 1.47 | 0.83 | 1.43 | 1.30 | 1.26 | 0.84 | 1.61 | 1.92 |
| Tb | 0.84 | 0.61 | 0.76 | 0.86 | 0.93 | 0.67 | 0.80 | 0.78 | 1.03 | 0.96 |
| Yb | 3.7 | 2.0 | 2.4 | 3.8 | 3.7 | 2.4 | 3.2 | 3.4 | 3.6 | 3.2 |
| Lu | 0.55 | 0.32 | 0.36 | 0.56 | 0.56 | 0.37 | 0.50 | 0.54 | 0.57 | 0.49 |
| La/Sm | 2.96 | 2.32 | 2.31 | 3.20 | 2.69 | 2.13 | 2.84 | 3.01 | 2.28 | 1.95 |
| ZHREE | 52.38 | 32.41 | 38.74 | 53.73 | 54.74 | 37.14 | 47.79 | 49.69 | 56.80 | 51.02 |
| Eu/Eu* | 0.52 | 0.95 | 0.84 | 0.46 | 0.68 | 0.91 | 0.71 | 0.50 | 0.77 | 0.93 |
| Zr Hf Ga Y Nb Ta Th U | 342 8.3 23 36 19.7 1.05 7.6 2.3 | 147 3.4 17 22 9.1 0.46 2.0 <2 | 165 4.2 25 11.0 0.54 2.7 1.1 | 281 7.0 21 15.6 0.89 6.4 2.3 | 287 7.5 19 35 17.9 1.04 5.5 2.0 | 146 3.6 17 24 11.9 0.58 2.5 0.9 | 275 6.5 17 32 17.3 0.84 5.1 1.8 | 270 6.7 23 34 15.7 0.86 6.2 2.2 | 251 6.8 19 37 16.4 0.83 5.0 1.7 | 176 4.6 24 32 9.5 0.61 3.0 0.9 |
| Series Label | TVC + | TVC + | TVC + | TVC | TVC | TVC | TVC | TVC | BT + | TVC |

Lat 44.10882 44.10375 44.09579 44.09827 44.07912 44.05953 44.05549 44.06806 44.03800 43.97077 -Long 121.5522 121.5545 121.5558 121.5510 121.5522 121.5038 121.5459 121.6021 121.5225 121.5857

3S014 Pumice lapilli 1mi NW from Triangle Hill: Hydrated 3S015 Cinder cone @ 6260', 1 mi W of Triangle Hill 3S017 Cinder cone @ 6415', 1 mi SW of Triangle Hill 3S018 Pumice lapilli 0.5 mi W Triangle Hill, ≈3S024 but hydrated 3S019 Summit of dome 0.5 mi N of Bearwallow Butte 3S022 Cinder cone @ 5635', 2 mi ESE Bearwallow Bu. 3S023 Summit of dome 1.5 mi ESE of Bearwallow Butte 3S024 Poorly consolidated pumice lapilli @ 6600' on Bottle Ck.; hydrated 3S027 Whole rock analysis of welded ignimbrite in N. wall Tumalo Ck. @ 5200' 3S028 Mafic pumice from Century Drive Tuff, 2 mi NW of Wanoga Bu.

| Sample Si02 Ti02 Al203 Fe0* MnO MgO Ca0 Na20 K20 P205 T0TAL | 35029 69.2 0.49 15.7 3.39 0.10 0.44 1.90 6.25 2.31 0.11 100.33 | 35030 54.3 1.01 18.9 7.14 0.13 4.72 9.00 4.03 0.66 0.14 100.26 | 35031 55.3 1.62 17.2 9.15 0.16 3.33 7.28 4.76 0.94 0.28 100.01 | 38032 74.5 0.15 13.7 1.70 0.06 0.00 0.71 5.26 3.48 0.02 99 .92 | 35033 56.0 1.29 17.3 8.26 0.15 3.80 7.49 3.61 1.02 0.31 99.53 | 38034 74.7 0.16 13.7 1.57 0.06 0.02 0.72 5.25 3.46 0.02 100.17 | 35035 74.7 0.15 13.8 1.57 0.06 0.00 0.00 0.70 5.12 3.52 0.02 99.7 6 | 35036 64.9 1.02 16.0 4.94 0.11 1.33 3.43 5.54 2.23 0.33 100.53 | 35037 69.6 0.54 15.4 3.31 0.07 0.36 1.64 5.38 2.93 0.11 100.22 | 35038 55.5 1.86 16.6 9.06 0.16 3.60 7.24 4.40 1.07 0.42 99.89 |
|--|--|--|--|---|--|--|---|--|--|---|
| Rb | 43 | 12 | 16 | 86 | 16 | 82 | 83 | 41 | 59 | 18 |
| Ba | 598 | 203 | 304 | 796 | 3 83 | 818 | 802 | 649 | 736 | 321 |
| Sr | 215 | 565 | 471 | 52 | 611 | 50 | 49 | 363 | 181 | 474 |
| Cs | 0.30 | 0.60 | 0.70 | 1.50 | 0.60 | 1.60 | 1.10 | 0.50 | 0.90 | 0.90 |
| Sc | 9.18 | 25.67 | 25.45 | 4.25 | 22.71 | 4.28 | 4.03 | 13.60 | 8.73 | 28.16 |
| V | 9 | 198 | 264 | 0 | 203 | 2 | 1 | 58 | 15 | 259 |
| Co | 1.76 | 25.60 | 24.12 | 0.54 | 19.97 | 0.53 | 0.51 | 7.14 | 2.50 | 24.84 |
| Ni | 5 | 24 | 4 | 10 | 5 | 10 | 12 | 9 | 10 | 7 |
| Cr | <5 | 100 | 5 | <3 | 23 | <3 | <3 | <5 | <5 | 21 |
| Zn | 82 | 48 | 74 | 84 | 67 | 64 | 73 | 80 | 54 | 74 |
| La | 20.7 | 7.3 | 12.6 | 25.3 | 14.2 | 24.2 | 23.0 | 24.1 | 22.2 | 16.5 |
| Ce | 43.8 | 17.0 | 28.2 | 52.6 | 29.6 | 53.0 | 51.2 | 50.4 | 53.7 | 37.1 |
| Nd | 22.3 | 10.2 | 16.0 | 20.4 | 17.0 | 21.8 | 21.5 | 27.9 | 24.0 | 20.8 |
| Sm | 5.73 | 2.97 | 4.42 | 4.77 | 4.15 | 4.71 | 4.83 | 6.59 | 5.99 | 5.58 |
| Eu | 1.40 | 1.11 | 1.55 | 0.64 | 1.30 | 0.63 | 0.60 | 1.65 | 1.28 | 1.71 |
| Tb | 0.92 | 0.46 | 0.69 | 0.78 | 0.56 | 0.78 | 0.72 | 0.95 | 0.84 | 0.85 |
| Yb | 3.5 | 1.5 | 2.4 | 3.7 | 2.0 | 3.5 | 3.3 | 3.4 | 3.5 | 2.9 |
| Lu | 0.53 | 0.25 | 0.34 | 0.53 | 0.28 | 0.54 | 0.51 | 0.51 | 0.50 | 0.43 |
| La/Sm | 2.28 | 1.55 | 1.80 | 3.35 | 2.16 | 3.24 | 3.01 | 2.31 | 2.34 | 1.87 |
| ZHREE | 52.46 | 24.67 | 36.64 | 50.62 | 30.11 | 50.12 | 46.92 | 52.10 | 50.02 | 45.21 |
| Eu/Eu* | 0.73 | 1.13 | 1.06 | 0.40 | 0.98 | 0.40 | 0.38 | 0.77 | 0.66 | 0.93 |
| Zr | 230 | 93 | 130 | 222 | 142 | 219 | 219 | 250 | 308 | 158 |
| Hf | 6.0 | 1.9 | 3.2 | 6.5 | 2.9 | 6.5 | 6.1 | 6.3 | 8.4 | 3.9 |
| Ga | 22 | 18 | 20 | 20 | 20 | 18 | 19 | 18 | 19 | 22 |
| Y | 34 | 16 | 26 | 31 | 24 | 32 | 30 | 34 | 30 | 30 |
| Nb | 13.1 | 5.5 | 8.5 | 17.8 | 8.3 | 18.0 | 17.8 | 15.7 | 16.8 | 12.1 |
| Ta | 0.64 | 0.29 | 0.45 | 1.00 | 0.39 | 1.04 | 0.98 | 0.80 | 1.00 | 0.66 |
| Th | 4.0 | 1.1 | 1.4 | 7.9 | 1.5 | 7.8 | 7.7 | 4.0 | 5.4 | 1.7 |
| U | 1.4 | <1 | 0.7 | 2.7 | 0.8 | 2.5 | 2.3 | 1.6 | 2.2 | <2 |
| Series Label | TVC + | NSMS | NSMS | TVC | TVC | TVC | TVC | TMR | TMR | TMR |

Lat 44.14281 44.17489 44.18962 44.15244 44.17484 44.17766 44.15535 44.09165 44.09669 44.09550 -Long 121.6828 121.6691 121.6654 121.5760 121.5944 121.6021 121.6090 121.6216 121.6176 121.6310

3S029 Dome at Squaw Creek Falls, 5600'

3S030 BaLa(33) under Shevlin Park Tuff, NF Squaw Ck @ 4800'

3S031 BaLa(32), N. Fork Squaw Ck, 4760'

3S032 3 Creek Butte, summit

3S033 Flow at southern base of 3 Creek Butte

3S034 Melvin Butte, summit

3S035 Summit of dome 1.25 mi S. of Melvin Butte

3S036 E. Tam MacArthur Rim, NW23, 68001

3S037 N. side of saddle on E Tam Rim, 0.25 mi NE of 3S036

3S038 Flow exposed near SW bank of 3 Creeks Lake, 6700'

| Sample Si02 Ti02 Al203 FeO* MnO MgO Ca0 Na20 K20 P205 TOTAL | 38040 69.4 0.56 15.3 3.39 0.09 0.54 1.97 5.60 2.85 0.12 100.71 | 35041 68.1 0.59 16.9 3.61 0.10 0.49 1.87 4.92 2.79 0.12 97.99 | 3s042 74.5 0.15 13.9 1.64 0.06 0.03 0.67 5.18 3.49 0.02 100.27 | 38043 69.5 0.55 15.3 3.33 0.09 0.48 1.85 5.61 2.93 0.11 100.52 | 35044 66.6 0.90 16.6 5.10 0.15 1.29 3.17 3.04 2.46 0.27 97.10 | 38045 56.0 1.12 18.3 7.40 0.13 3.66 7.70 4.45 1.02 0.28 100.47 | 38046 69.5 15.2 3.30 0.09 0.54 1.85 5.98 2.88 0.11 101.27 | 38048 69.4 0.55 15.4 3.37 0.09 0.46 1.81 5.79 2.89 0.11 100.49 | 38049 53.0 1.37 17.6 8.06 0.14 6.11 8.41 3.73 1.17 0.41 100.70 | 38050 68.9 0.62 15.5 3.32 0.11 0.65 2.04 5.93 2.49 0.15 100.64 |
|--|--|---|--|--|--|--|---|--|--|--|
| Rb | 60 | 58 | 83 | 61 | 42 | 15 | 60 | 59 | 21 | 51 |
| Ba | 715 | 701 | 788 | 707 | 573 | 364 | 740 | 753 | 475 | 711 |
| Sr | 194 | 174 | 49 | 184 | 295 | 719 | 182 | 182 | 675 | 220 |
| Cs | 2.50 | 1.60 | 1.50 | 1.70 | 0.90 | 0.60 | 1.90 | 1.90 | 0.40 | 1.50 |
| Sc | 9.65 | 9.28 | 4.56 | 9.13 | 12.42 | 22.34 | 9.45 | 9.30 | 23.01 | 10.36 |
| V | 19 | 0 | 8 | 7 | 36 | 183 | 1 | 19 | 187 | 4 |
| Co | 3.35 | 2.89 | 0.92 | 2.94 | 3.63 | 21.80 | 2.89 | 2.88 | 30.50 | 2.32 |
| Ni | 10 | 8 | 12 | 10 | 8 | 12 | 8 | 10 | 86 | 9 |
| Cr | <5 | <4 | 1 | <5 | 3 | 40 | <4 | <5 | 138 | <5 |
| Zn | 83 | 85 | 68 | 86 | 98 | 79 | 89 | 87 | 74 | 90 |
| La Ce Nd Sm Eu Tb Yb Lu La/Sm ZHREE Eu/Eu* | 25.9 55.7 27.0 6.55 1.33 1.00 4.2 0.60 2.50 59.85 0.62 | 28.4 55.0 32.1 7.64 1.37 1.23 4.2 0.65 2.35 66.34 0.54 | 19.4 44.0 17.6 3.66 0.60 0.59 3.0 0.46 3.35 41.27 0.49 | 24.8 52.8 6.18 1.27 0.99 3.8 0.61 2.53 58.09 0.62 | 22.8 46.2 27.6 6.73 1.77 1.08 3.8 0.58 2.14 59.08 0.79 | 15.4 34.2 19.3 4.30 1.35 0.63 1.8 0.29 2.26 30.95 0.97 | 25.9 55.1 29.5 6.30 1.36 1.03 4.0 0.64 2.60 60.73 0.64 | 26.2 54.4 26.0 6.42 1.32 1.03 4.2 0.63 2.58 61.51 0.62 | 23.8 53.3 27.1 6.24 1.81 0.84 2.2 0.37 2.41 39.80 0.91 | 25.0 52.8 26.0 6.34 1.45 0.99 3.9 0.57 2.49 57.19 0.69 |
| Zr | 300 | 335 | 217 | 306 | 267 | 137 | 307 | 311 | 177 | 261 |
| Hf | 8.5 | 8.4 | 6.6 | 8.2 | 6.2 | 3.1 | 8.5 | 8.5 | 4.1 | 6.9 |
| Ga | 17 | 22 | 18 | 20 | 19 | 17 | 20 | 20 | 19 | 17 |
| Y | 37 | 45 | 24 | 37 | 44 | 20 | 37 | 39 | 26 | 36 |
| Nb | 16.3 | 18.6 | 16.3 | 16.9 | 18.7 | 7.0 | 17.7 | 17.0 | 16.2 | 15.8 |
| Ta | 1.01 | 0.96 | 1.00 | 0.88 | 0.85 | 0.50 | 0.92 | 0.97 | 0.80 | 0.91 |
| Th | 5.4 | 5.5 | 8.2 | 5.3 | 3.6 | 2.0 | 5.6 | 5.4 | 3.1 | 4.7 |
| U | 2.2 | 2.2 | 2.6 | 2.0 | 1.2 | 0.6 | 2.0 | 2.3 | 0.9 | 1.7 |
| Series Label | TMR + | TMR + | TVC | BT + | BT + | TMR + | TMR + | TMR | BT + | BT + |

Lat 44.08856 44.09084 44.14904 44.07083 44.07349 44.09102 44.09242 44.09223 44.05410 44.03938 -Long 121.5932 121.6124 121.6241 121.6524 121.6535 121.6347 121.6364 121.6397 121.6615 121.6736

3S040 Glassy flow, 1.5 mi SE of 3 Creeks Lake, 6640' 3S041 Pumice lapilli from vent breccia on E. Tam Rim, 7000' 3S042 Dome 2 mi SW of Melvin Butte 3S043 Glassy base of flow in N. Fork Tumalo Ck, 7040' 3S044 Pumice lapilli from tuff cone in N. Fork Tumalo Ck, 7160' 3S045 Flow under 3S046 a 7320', Tam Rim 3S046 Upper Tam Rim, black glassy flow a 7400' 3S048 Pumice lapilli from vent breccia on upper Tam Rim, 7440' 3S049 Cinder cone 1 mi SE of Ball Butte, 7229' 3S050 Flow 1 mi ENE of Todd Lake, 6700', =BT-390

| Sample Si02 Ti02 Al203 FeO* MnO MgO CaO Na20 K20 P205 TOTAL | 38051 70.6 0.43 15.3 2.92 0.08 0.33 1.40 5.77 2.68 0.08 100.61 | 35052 65.3 0.97 16.1 4.70 0.10 1.40 3.73 5.14 2.26 0.28 100.79 | 35053 72.7 0.33 14.6 2.02 0.05 0.50 1.90 4.60 3.20 0.08 100.65 | 38054 74.2 0.24 14.3 1.58 0.05 0.23 1.20 4.53 3.37 0.05 100.95 | 38057 72.7 0.32 14.6 2.00 0.05 0.51 1.88 4.52 3.21 0.08 100.63 | 3s058 73.7 0.27 14.3 1.73 0.05 0.29 1.39 4.83 3.25 0.06 100.35 | 3s059 65.3 0.96 16.1 4.75 0.10 1.35 3.64 4.88 2.30 0.27 100.23 | 35060 73.7 0.29 14.1 1.75 0.05 0.41 1.62 4.38 3.48 0.07 100.64 | 35061 64.6 0.96 16.5 4.82 0.11 1.52 3.98 4.92 2.07 0.26 101.00 | 38062 73.8 0.26 14.3 1.71 0.05 0.26 1.37 4.64 3.24 0.06 100.53 |
|--|--|--|--|--|--|--|--|--|--|--|
| Rb | 54 | 46 | 76 | 74 | 76 | 73 | 47 | 85 | 40 | 68 |
| Ba | 706 | 608 | 783 | 840 | 775 | 831 | 633 | 809 | 584 | 812 |
| Sr | 148 | 344 | 224 | 144 | 221 | 165 | 338 | 185 | 386 | 162 |
| Cs | 0.60 | 1.60 | 1.90 | 1.30 | 1.80 | 1.60 | 1.30 | 2.20 | 1.50 | 1.20 |
| Sc | 8.94 | 12.36 | 3.80 | 2,51 | 3.69 | 2.73 | 11.91 | 3.29 | 12.44 | 2.60 |
| V | 3 | 70 | 16 | 3 | 17 | 3 | 84 | 20 | 85 | 17 |
| Co | 1.76 | 8.92 | 2.92 | 1.47 | 2.80 | 1.80 | 8.45 | 2.44 | 8.54 | 1.69 |
| Ni | 7 | 32 | <1 | <1 | <1 | | 7 | <10 | 5 | |
| Cr | <5 | 1 | 3 | <2 | 2 | <3 | 3 | 2 | <7 | <3 |
| Zn | 70 | 68 | 53 | 49 | 51 | 51 | 61 | 47 | 75 | 48 |
| La | 19.0 | 20.4 | 20.0 | 19.8 | 19.4 | 20.7 | 20.4 | 21.0 | 19.6 | 19.9 |
| Ce | 44.4 | 43.8 | 37.2 | 36.2 | 36.6 | 37.3 | 42.9 | 38.4 | 41.6 | 34.9 |
| Nd | 20.0 | 21.0 | 16.0 | 14.0 | 15.2 | 15.0 | 22.0 | 16.0 | 21.0 | 13.7 |
| Sm | 4.94 | 5.05 | 3.00 | 2.62 | 2.98 | 2.75 | 5.09 | 3.07 | 4.64 | 2.67 |
| Eu | 1.19 | 1.44 | 0.65 | 0.49 | 0.63 | 0.54 | 1.38 | 0.59 | 1.39 | 0.52 |
| тb | 0.79 | 0.74 | 0.41 | 0.35 | 0.40 | 0.37 | 0.74 | 0.41 | 0.66 | 0.34 |
| Yb | 3.7 | 2.6 | 1.9 | 1.6 | 1.9 | 1.8 | 2.7 | 1.8 | 2.5 | 1.6 |
| Lu | 0.52 | 0.40 | 0.26 | 0.24 | 0.26 | 0.24 | 0.39 | 0.31 | 0.38 | 0.25 |
| La/Sm | 2.43 | 2.55 | 4.21 | 4.77 | 4.11 | 4.75 | 2.53 | 4.32 | 2.67 | 4.71 |
| ΣHREE | 50.62 | 40.51 | 25.82 | 22.50 | 25.61 | 23.87 | 40.68 | 26.89 | 37.71 | 22,59 |
| Eu/Eu* | 0.72 | 0.87 | 0.67 | 0.59 | 0.66 | 0.61 | 0.83 | 0.60 | 0.92 | 0.62 |
| Zr | 265 | 207 | 171 | 180 | 172 | 192 | 208 | 168 | 195 | 190 |
| Hf | 7.0 | 5.4 | 4.5 | 4.6 | 4.5 | 4.9 | 5.4 | 4.5 | 5.0 | 4.8 |
| Ga | 21 | 16 | 14 | 15 | 15 | 16 | 18 | 13 | 19 | 15 |
| Y | 32 | 27 | 19 | 16 | 19 | 17 | 26 | 18 | 24 | 16 |
| Nb | 16.7 | 11.1 | 11.5 | 11.5 | 11.8 | 10.9 | 11.7 | 12.1 | 12.9 | 11.5 |
| Та | 0.81 | 0.76 | 0.76 | 0.77 | 0.78 | 0.74 | 0.75 | 0.79 | 0.77 | 0.73 |
| Th | 4.8 | 4.4 | 7.3 | 6.4 | 7.2 | 6.4 | 4.5 | 8.2 | 4.1 | 6.2 |
| U | 2.0 | 1.9 | 2.9 | 2.6 | 2.7 | 2.5 | 1.6 | 3.2 | 1.4 | 2.4 |
| Series | BT | SS |
| Label | + | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 | 0 |

Lat 44.03220 44.03188 44.03361 44.03898 44.05077 44.05012 44.05280 44.05680 44.05075 44.04664 -Long 121.6724 121.7404 121.7555 121.7495 121.7572 121.7587 121.7702 121.8034 121.8001 121.7935

3S051 Flow 0.75 mi E. of Todd Lake, under 3S050, ≈BT-391 3S052 End of flow from SS, N. Sparks Lk. at Century Or. 3S053 Southernmost Holocene Oevils Hill Oome 3S054 SE flank of Oevils Hill, 5800' 3S057 Northernmost Holocene Devils Hill Oome 3S058 Upper Devils Hill, next to Holocene Dome @ 6600' 3S059 Flow along Upper Moraine Lk. Trail, 6500' 35060 Rock Mesa Holocene dome, SE margin 35061 Columnar flow margin, 0.75 mi SE of LeConte cone

3S062 Kaleetan Butte, SW flank, 6200'

| Sample | 35064 | 38065 | 3 S066 | 38067 | 35069 | 38073 | 3 s074 | 35075 | 35078 | 35079 |
|-----------------|--------------|--------------|---------------|--------------|--------------|--------------|---------------|--------------|--------------|--------------|
| Si02 | 62.7 | 63.1 | 54.5 | 64.2 | 51.8 | 58.6 | 55.6 | 74.9 | 73.7 | 67.0 |
| Ti02 | 1.03 | 1.01 | 1.31 | 1.03 | 1.02 | 1.08 | 1.42 | 0.14 | 0.21 | 0.81 |
| Al203 | 16.6 | 16.6 | 16.7 | 16.1 | 16.7 | 17.5 | 17.6 | 13.7 | 14.7 | 17.1 |
| FeO* | 5.63 | 5.45 | 7.89 | 5.44 | 8.39 | 7.04 | 8.46 | 1.56 | 1.62 | 4.12 |
| MnO | 0.11 | 0.11 | 0.14 | 0.11 | 0.15 | 0.13 | 0.14 | 0.05 | 0.06 | 0.13 |
| MgO | 1.97 | 1.94 | 6.43 | 1.73 | 8.79 | 3.07 | 3.67 | 0.08 | 0.24 | 0.83 |
| CaO | 4.73 | 4.65 | 7.58 | 3.88 | 9.14 | 6.57 | 7.88 | 0.66 | 1.49 | 2.39 |
| Na20 | 4.88 | 4.87 | 3.92 | 5.41 | 3.22 | 4.04 | 3.57 | 4.78 | 4.27 | 4.33 |
| K20 | 1.89 | 1.97 | 1.18 | 2.22 | 0.64 | 1.28 | 1.02 | 3.53 | 3.14 | 2.16 |
| P205 | 0.23 | 0.23 | 0.33 | 0.30 | 0.21 | 0.23 | 0.35 | 0.02 | 0.06 | 0.16 |
| TOTAL | 100.34 | 100.15 | 99.94 | 101.68 | 100.31 | 100.99 | 100.27 | 101.28 | 99.81 | 98.01 |
| Rb | 37 | 39 | 21 | 49 | 11 | 24 | 17 | 87 | 66 | 44 |
| Ba | 530 | 539 | 343 | 633 | 173 | 441 | 350 | 801 | 826 | 616 |
| Sr | 408 | 403 | 437 | 304 | 439 | 494 | 578 | 42 | 192 | 278 |
| Cs | 1.30 | 1.30 | 0.70 | 1.60 | 0.60 | 0.60 | 0.60 | 1.60 | 1.50 | 1.60 |
| Sc | 14.67 | 14.48 | 21.55 | 14.40 | 27.74 | 20.73 | 27.75 | 4.34 | 1.98 | 11.05 |
| V | 136 | 120 | 161 | 106 | 187 | 178 | 219 | 9 | 8 | 36 |
| Co | 12.24 | 12.09 | 32.22 | 10.06 | 39.74 | 18.55 | 23.30 | 0.69 | 1.32 | 2.12 |
| Ni | 25 | 40 | 119 | 6 | 148 | 9 | 15 | 12 | <1 | 7 |
| Cr | 6 | 6 | 217 | 3 | 451 | 17 | 45 | <4 | <3 | 3 |
| Zn | 60 | 67 | 65 | 78 | 54 | 68 | 60 | 71 | 60 | 87 |
| La | 17.9 | 18.1 | 15.7 | 24.9 | 9.5 | 16.0 | 17.6 | 19.4 | 19.2 | 19.8 |
| Ce | 37.7 | 38.8 | 34.9 | 55.0 | 21.2 | 33.3 | 36.2 | 43.4 | 36.1 | 39.9 |
| Nd | 17.0 | 18.5 | 17.9 | 28.4 | 11.3 | 19.0 | 21.0 | 17.5 | 13.0 | 22.1 |
| Sm | 4.28 | 4.38 | 4.49 | 6.41 | 3.18 | 4.41 | 5.12 | 3.84 | 2.65 | 6.06 |
| Eu | 1.28 | 1.30 | 1.41 | 1.58 | 1.12 | 1.28 | 1.67 | 0.54 | 0.56 | 1.37 |
| Tb | 0.62 | 0.66 | 0.68 | 0.99 | 0.51 | 0.68 | 0.77 | 0.64 | 0.35 | 0.87 |
| Yb | 2.0 | 2.3 | 2.2 | 3.7 | 1.6 | 2.2 | 2.6 | 3.1 | 1.6 | 3.4 |
| Lu | 0.33 | 0.37 | 0.35 | 0.54 | 0.28 | 0.35 | 0.38 | 0.46 | 0.24 | 0.50 |
| La/Sm | 2.64 | 2.61 | 2.21 | 2.45 | 1.89 | 2.29 | 2.17 | 3.19 | 4.57 | 2.06 |
| ZHREE | 32.93 | 36.45 | 35.78 | 55.40 | 27.14 | 35.78 | 40.53 | 42.62 | 22.50 | 50.23 |
| Eu/Eu* | 0.92 | 0.90 | 0.96 | 0.75 | 1.06 | 0.88 | 0.99 | 0.42 | 0.66 | 0.70 |
| Zr | 176 | 179 | 156 | 248 | 105 | 160 | 161 | 201 | 158 | 231 |
| Hf | 4.5 | 4.6 | 3.7 | 6.7 | 2.3 | 3.8 | 3.8 | 6.1 | 4.0 | 5.5 |
| Ga | 17 | 20 | 15 | 18 | 15 | 17 | 21 | 22 | 17 | 21 |
| Y | 23 | 24 | 24 | 33 | 18 | 24 | 28 | 26 | 16 | 36 |
| Nb | 11.7 | 12.7 | 13.6 | 15.6 | 8.3 | 9.6 | 11.1 | 17.4 | 10.6 | 14.3 |
| Ta | 0.70 | 0.75 | 0.88 | 0.96 | 0.48 | 0.50 | 0.59 | 1.04 | 0.73 | 0.66 |
| Th | 3.9 | 4.0 | 2.2 | 5.0 | 1.2 | 2.4 | 2.0 | 8.0 | 5.8 | 3.8 |
| U | 1.3 | 1.5 | 0.8 | 1.6 | 0.5 | 0.9 | 0.6 | 2.3 | 2.3 | 1.4 |
| Series Label | SS o | SS o | SS o | BT + | BT + | TVC + | TVC + | TVC + | NSMS | TVC + |

Lat 44.02272 44.01727 44.02367 44.02404 44.02697 44.09801 44.11785 44.11552 44.21842 44.09376 -Long 121.7752 121.7729 121.7649 121.7215 121.7249 121.5211 121.5429 121.5510 121.8312 121.4315

3S064 Kokostick Butte, S flank, 5680*

3S065 Small dome 1 mi SSW of Kokostick, 5820'

3S066 LeConte Flow, 2.5 mi SE of vent, along Century Dr. 3S067 SW Flank TLV, 5600', along Century Drive 3S069 Cayuse flow, 2.5 mi SSW of vent, at Century Dr.

3S073 Flow 1 mi E. of Triangle Hill, 4900'

3S074 Cinder cone 1 mi NNW Triangle Hill, 5000' 3S075 Glaciated dome 1 mi NW Triangle Hill, 5400'

3S078 Summit of glaciated dome 1 mi SE of Condon Butte

3S079 Pumice lapilli from fall deposit under Shevlin Park Tuff in Columbia Canyon

| Sample Si02 Ti02 Al203 FeO* MnO MgO Ca0 Na20 K20 P205 T0TAL | 3S085 63.4 1.13 16.1 5.70 0.12 1.70 3.90 5.44 2.11 0.38 100.97 | 35086 75.6 0.13 14.2 1.07 0.04 0.09 0.93 4.22 3.40 0.04 98.21 | 38090 55.2 1.14 19.2 7.44 0.14 4.02 7.74 3.85 0.88 0.35 101.45 | 38091 69.5 0.54 15.3 3.33 0.09 0.52 1.83 5.96 2.88 0.11 101.19 | 35092 57.0 1.39 16.9 8.34 0.15 3.27 6.84 4.67 1.12 0.31 100.79 | 35093 57.5 1.25 17.4 7.60 0.14 3.37 6.73 4.63 1.17 0.27 101.19 | 35094 67.2 0.79 16.0 3.79 0.12 0.90 2.46 6.15 2.27 0.20 100.52 | 38095 62.5 1.34 15.8 6.47 0.14 1.68 4.10 5.26 2.09 0.48 100.70 | 35096 63.8 1.18 16.2 5.66 0.13 1.47 5.56 1.87 0.35 101.31 | 35097 60.0 1.33 16.8 6.67 0.14 2.29 5.40 5.40 5.22 1.49 0.39 100.84 |
|--|--|---|--|--|--|--|---|--|--|--|
| Rb | 41 | 74 | 11 | 61 | 19 | 23 | 44 | 42 | 37 | 29 |
| Ba | 634 | 861 | 402 | 719 | 377 | 401 | 655 | 581 | 517 | 458 |
| Sr | 393 | 115 | 617 | 179 | 485 | 478 | 292 | 334 | 326 | 446 |
| Cs | 1.15 | 1.84 | 0.48 | 2.11 | 0.59 | 0.34 | 1.44 | 1.54 | 0.85 | 0.61 |
| Sc V Co Ni Cr Zn | 16.60 93 10.52 5 7 | 1.98 0 0.92 11 4 58 | 19.44 163 22.24 42 60 | 9.44 0 2.90 10 4 145 | 27.65 247 22.54 4 22 | 24.62 212 21.66 9 41 | 11.81 29 2.18 8 2 163 | 17.44 93 10.70 2 <8 | 15.99 51 7.66 6 <6 200 | 20.15 134 12.65 5 13 |
| La | 24.7 | 20.0 | 19.1 | 25.8 | 15.0 | 15.6 | 23.9 | 24.1 | 18.8 | 18.7 |
| Ce | 58.7 | 43.5 | 41.2 | 63.6 | 35.9 | 35.2 | 60.0 | 60.0 | 44.3 | 39.1 |
| Nd | 31.6 | 16.3 | 23.6 | 31.3 | 19.7 | 20.6 | 29.4 | 32.3 | 25.7 | 24.8 |
| Sm | 6.96 | 2.77 | 5.28 | 6.48 | 4.99 | 4.82 | 6.50 | 7.58 | 6.36 | 5.98 |
| Eu | 1.82 | 0.44 | 1.51 | 1.33 | 1.61 | 1.46 | 1.63 | 1.96 | 1.71 | 1.72 |
| Tb | 1.18 | 0.41 | 0.75 | 1.22 | 0.84 | 0.84 | 1.03 | 1.22 | 1.04 | 1.01 |
| Yb | 3.7 | 1.5 | 2.5 | 4.4 | 2.5 | 2.7 | 3.5 | 3.9 | 3.6 | 3.1 |
| Lu | 0.58 | 0.26 | 0.40 | 0.63 | 0.42 | 0.38 | 0.53 | 0.64 | 0.55 | 0.51 |
| La/Sm | 2.24 | 4.56 | 2.28 | 2.51 | 1.90 | 2.04 | 2.32 | 2.01 | 1.87 | 1.97 |
| ZHREE | 60.68 | 23.92 | 40.25 | 66.41 | 42.78 | 42.49 | 54.99 | 64.34 | 56.30 | 52.04 |
| Eu/Eu* | 0.77 | 0.49 | 0.88 | 0.59 | 0.96 | 0.89 | 0.75 | 0.77 | 0.80 | 0.85 |
| Zr | 239 | 97 | 168 | 306 | 154 | 157 | 242 | 244 | 201 | 176 |
| Hf | 6.7 | 3.3 | 3.8 | 9.1 | 3.9 | 3.9 | 6.8 | 6.6 | 5.5 | 4.5 |
| Ga | 19 | 17 | 16 | 20 | 21 | 20 | 20 | 21 | 21 | 20 |
| Y | 35 | 16 | 25 | 36 | 26 | 25 | 36 | 38 | 33 | 31 |
| Nb | 13.3 | 10.9 | 10.9 | 16.4 | 10.2 | 11.6 | 15.7 | 16.6 | 14.3 | 11.8 |
| Ta | 0.79 | 0.69 | 0.59 | 0.93 | 0.53 | 0.51 | 0.78 | 0.94 | 0.71 | 0.59 |
| Th | 4.2 | 7.3 | 1.9 | 5.9 | 1.7 | 1.8 | 4.4 | 4.1 | 3.3 | 2.8 |
| U | 1.8 | 4.5 | <3 | 1.9 | <3 | <3 | 1.7 | 1.5 | 1.4 | <3 |
| Series Label | TVC | SS | TVC | TVC | BT | BT + | BT + | BT | BT | BT + |

Lat 44.12728 44.13327 43.93606 44.12042 44.10968 44.09774 44.09323 44.09194 44.09279 44.09439 -Long 121.5849 121.5873 121.5941 121.5706 121.6563 121.6661 121.6704 121.6743 121.6761 121.6813

3S085 Flow 1.5 mi SSW of 3 Creek Butte, 5400' 3S086 Reworked pumice lapilli 1.25 mi SSW of 3 Ck. Bu., 5400' 3S090 Edison Butte, SE flank, 5240' 3S091 Glaciated dome remnant, 2 mi NW of Triangle Hill 3S092 Ridge a7020',NE16, Snow Ck. cirque 3S093 Flow in Snow Creek Cirque, 7040' 3S094 Lapilli fall tuff, Snow Creek cirque, 7350' 3S095 Dike, Snow Creek cirque, 7560', AnId6 3S096 Flow, Snow Creek cirque, 7480', AnLa5 3S097 Dike, 1.25 mi NE BT, 7280', AnId4

| Sample SiO2 TiO2 Al2O3 FeO* MnO CaO Na2O K2O P2O5 TOTAL | 38098 71.1 0.40 14.9 2.74 0.08 0.30 1.40 6.13 2.77 0.07 100.71 | 35099 56.5 1.54 16.7 8.50 0.16 3.49 7.12 4.49 1.02 0.31 100.38 | 3s100 64.9 0.97 16.3 4.85 0.14 1.22 3.26 6.22 1.99 0.30 100.96 | 3s101 57.8 1.21 17.3 7.40 0.14 3.29 6.65 4.81 1.17 0.29 100.90 | 35102 60.4 1.36 16.1 7.22 0.13 2.40 5.10 4.99 1.78 0.40 100.67 | 3S103H 56.0 1.65 16.8 8.99 0.16 3.34 6.90 4.84 1.02 0.37 100.12 | 3s103s 66.9 0.80 16.7 3.76 0.13 0.89 2.39 5.92 2.23 0.20 99.88 | 35104 68.6 0.62 15.8 3.38 0.10 0.61 2.13 6.09 2.43 0.14 100.35 | 35105 54.8 1.22 18.9 7.46 0.14 3.77 8.53 4.16 0.81 0.20 100.67 | 3s106 55.9 0.99 18.3 7.09 0.13 4.58 7.82 3.98 0.93 0.22 100.73 |
|---|--|--|--|--|--|---|--|--|--|--|
| Rb Ba Sr Cs | 58 721 135 2.04 | 19 300 440 0.66 | 39 548 300 1.29 | 20 412 564 0.84 | 35 523 462 0.65 | 18 347 534 0.54 | 42 637 283 1.34 | 49 705 236 1.58 | 16 243 515 0.27 | 17 357 542 0.19 |
| Sc V Co Ni Cr Zn | 9.25 1 1.59 10 2 135 | 27.96 243 21.63 7 27 | 14.07 24 3.84 6 4 207 | 23.92 189 20.17 7 37 104 | 20.41 163 16.37 12 12 170 | 27.01 234 22.45 2 7 118 | 12.18 11 2.23 7 3 158 | 10.48 11 2.44 5 4 87 | 25.52 203 22.90 21 73 | 21.77 164 25.44 50 92 |
| La Ce Nd Sm Eu Tb Yb Lu La/Sm La/Sm Eu/Eu* | 24.4 61.3 30.8 6.46 1.28 1.28 4.2 0.66 2.38 67.67 0.56 | 13.4 31.7 20.0 4.93 1.57 0.79 2.7 0.47 1.72 44.22 0.96 | 24.6 62.6 34.0 7.84 2.04 1.45 4.8 0.77 1.98 77.55 0.75 | 19.3 39.9 22.3 5.29 1.52 0.81 2.8 0.39 2.30 42.64 0.87 | 22.9 52.4 30.6 6.76 1.75 0.97 3.2 0.52 2.14 51.98 0.80 | 14.4 32.2 19.4 5.01 1.63 0.80 2.6 0.45 1.81 43.33 0.98 | 24.7 62.1 33.3 6.89 1.68 1.07 3.9 0.58 2.26 59.29 0.74 | 24.5 63.2 30.5 6.69 1.51 1.14 3.9 0.56 2.31 60.16 0.67 | 9.7 22.8 13.3 3.68 1.31 0.64 2.3 0.34 1.66 35.10 1.05 | 13.6 30.4 17.8 3.83 1.22 0.60 2.1 0.31 2.24 32.36 0.96 |
| Zr Hf Ga Y Nb Ta Th U | 274 8.3 20 37 15.5 0.90 5.5 2.2 | 150 3.7 20 27 10.6 0.46 1.5 <2 | 272 7.7 19 42 20.0 1.03 3.8 1.3 | 154 4.0 16 28 9.4 0.45 2.1 0.9 | 214 5.6 22 32 13.0 0.66 3.4 1.2 | 139 3.4 20 27 10.4 0.40 1.9 <3 | 252 7.1 21 37 16.9 0.85 4.8 1.5 | 260 7.5 19 36 15.7 0.91 5.0 1.9 | 113 2.7 20 22 7.7 0.35 1.3 <3 | 136 3.2 19 8.4 0.37 1.4 1.3 |
| Series Label | BT + | BT + | BT + | BT + | TVC + | TVC? | TVC? | BT + | BT + | BT + |

Lat 44.09975 44.09906 44.11087 44.09380 44.05756 43.92009 43.92009 44.04378 44.04255 44.04789 -Long 121.6817 121.6775 121.6897 121.6342 121.5622 121.5881 121.5881 121.6962 121.7003 121.7019

3S098 Flow along Squaw Creek, 7020', RdLa5 3S099 Flow along Squaw Creek, 7140' 3S100 Flow along Squaw Creek, 6600', DaLa2 3S101 Basal Tam Rim flow, W. shore of 3 Creek Lk, 6820' 3S102 Flow ridge @ 6212', 1.5 mi SSW of Bearwallow Bu. 3S103M Mafic pumice from ignimbrite 1.3 mi SSE of Edison Bu, 4900' 3S103S Silicic pumice from ignimbrite 1.3 mi SSE of Edison Bu, 4900' 3S104 Flow @ 6120', 1 mi NNW Todd Lk., ≈3S050 3S105 Flow @ 6160', 1 mi NNW Todd Lk. 3S106 Flow @ Crater Ck. Falls, 6320'

| Sample | 35107 | 3s108 | 35109 | 35110 | 3s111 | 35112 | 3s113 | 35114 | 3s115 | 3s116 |
|--|--|--|--|--|--|---|--|--|--|--|
| Si02 | 63.0 | 63.9 | 56.5 | 69.6 | 64.1 | 60.0 | 72.0 | 72.8 | 62.9 | 74.6 |
| Ti02 | 1.10 | 1.20 | 1.37 | 0.62 | 1.04 | 1.24 | 0.34 | 0.34 | 1.24 | 0.15 |
| Al203 | 16.2 | 15.7 | 17.1 | 15.3 | 16.2 | 16.7 | 14.8 | 14.5 | 16.3 | 13.9 |
| FeO* | 5.84 | 5.74 | 8.37 | 3.00 | 5.19 | 6.85 | 2.22 | 2.04 | 5.98 | 1.66 |
| MnO | 0.13 | 0.13 | 0.15 | 0.10 | 0.12 | 0.13 | 0.06 | 0.06 | 0.14 | 0.06 |
| MgO | 1.84 | 1.49 | 3.52 | 0.64 | 1.65 | 2.76 | 0.43 | 0.49 | 1.68 | 0.01 |
| CaO | 4.23 | 3.73 | 7.17 | 1.83 | 4.02 | 5.80 | 1.76 | 1.85 | 4.10 | 0.67 |
| Na20 | 5.45 | 5.42 | 4.61 | 6.24 | 5.26 | 4.76 | 5.09 | 4.70 | 5.59 | 5.40 |
| K20 | 1.87 | 2.19 | 0.95 | 2.60 | 2.05 | 1.58 | 3.15 | 3.22 | 1.71 | 3.46 |
| P205 | 0.33 | 0.40 | 0.19 | 0.13 | 0.27 | 0.27 | 0.09 | 0.08 | 0.39 | 0.02 |
| TOTAL | 101.03 | 100.77 | 100.46 | 101.23 | 100.43 | 100.78 | 101.17 | 100.79 | 100.96 | 100.86 |
| Rb | 37 | 45 | 17 | 52 | 40 | 28 | 67 | 77 | 34 | 84 |
| Ba | 558 | 581 | 275 | 713 | 572 | 458 | 851 | 788 | 495 | 7 83 |
| Sr | 375 | 311 | 497 | 196 | 383 | 418 | 205 | 216 | 355 | 51 |
| Cs | 1,10 | 0.94 | 0.50 | 1.91 | 1.05 | 0.34 | 1.75 | 2.15 | 0.71 | 1.54 |
| Sc V Co Ni Cr Zn | 16.27 98 9.63 4 8 135 | 16.10 71 8.71 3 4 181 | 28.50 269 24.90 3 24 | 11.23 20 1.95 10 3 146 | 13.54 95 8.89 15 3 129 | 20.29 166 19.13 13 24 | 3.69 15 2.35 24 3 <20 | 3.85 11 2.92 <20 3 82 | 17.70 80 9.24 4 3 180 | 4.36 6 0.73 11 1 89 |
| La | 21.8 | 23.8 | 10.0 | 26.6 | 20.0 | 16.3 | 22.5 | 20.8 | 20.5 | 23.0 |
| Ce | 51.5 | 62.2 | 23.9 | 68.7 | 44.5 | 36.4 | 49.8 | 44.9 | 48.7 | 73.4 |
| Nd | 28.6 | 33.9 | 13.8 | 34.3 | 20.5 | 20.2 | 19.5 | 14.7 | 28.8 | 18.4 |
| Sm | 6.50 | 7.47 | 3.75 | 7.00 | 4.88 | 4.59 | 3.48 | 3.19 | 6.55 | 4.39 |
| Eu | 1.58 | 1.86 | 1.36 | 1.55 | 1.40 | 1.45 | 0.75 | 0.66 | 1.86 | 0.66 |
| Tb | 1.00 | 1.36 | 0.61 | 1.25 | 0.84 | 0.73 | 0.56 | 0.49 | 1.18 | 0.84 |
| Yb | 3.6 | 4.4 | 2.3 | 4.4 | 2.7 | 2.7 | 2.1 | 1.9 | 3.5 | 3.2 |
| Lu | 0.51 | 0.61 | 0.31 | 0.63 | 0.37 | 0.39 | 0.34 | 0.28 | 0.56 | 0.51 |
| La/Sm | 2.12 | 2.01 | 1.68 | 2.40 | 2.59 | 2.24 | 4.08 | 4.12 | 1.98 | 3.31 |
| ZHREE | 54.21 | 68.77 | 33.53 | 67.05 | 42.18 | 40.46 | 32.44 | 28.14 | 59.11 | 48.90 |
| Eu/Eu* | 0.74 | 0.72 | 1.08 | 0.65 | 0.85 | 0.95 | 0.65 | 0.63 | 0.83 | 0.43 |
| Zr Hf Ga Y Nb Ta Th U | 222 5.8 21 34 14.4 0.71 3.3 1.5 | 247 7.3 24 40 17.3 1.02 4.3 1.7 | 116 2.8 22 23 6.6 0.32 1.6 <2 | 271 8.3 19 38 16.6 1.00 5.6 2.0 | 198 5.3 20 25 12.1 0.73 4.1 1.5 | 172 4.4 17 25 9.1 0.60 3.0 0.9 | 223 6.3 13 20 13.7 0.82 6.6 2.3 | 172 5.0 18 12.0 0.82 8.0 2.9 | 191 5.9 18 37 13.5 0.68 3.2 1.2 | 217 7.2 21 17.0 1.06 8.5 2.5 |
| Series | BT | BT | BT | BT | SS | ss | ss | SS | BT | TVC + |
| Label | + | + | + | + | o | o | o | o | + | |

Lat 44.04482 44.09182 44.09411 44.09658 44.11459 44.10692 44.09029 44.08407 44.06765 43.93969 -Long 121.7048 121.7167 121.7221 121.7246 121.7333 121.7293 121.7399 121.7361 121.7227 121.5987

3S107 Dike along Crater Ck., 6000' DaId10 3S108 Flow on NW arm BT, 7560' DaLa1 3S109 Dike NW arm BT, 7440', 0.5 mi NE Green Lk. AnId2 3S110 Dome NW arm BT 7200', 0.5 mi NE Green Lk. RdIp4 3S111 Flow on W. Fork Park Ck., 7040' 3S112 Flow 1 mi N. of Green Lk., 6990', on trail. AnLa1 3S113 Dome W. shore Green Lk., 6600'. RdLa3 3S114 Newberry Dome, E. flank a Green Lk., 6600' 3S115 Flow on SW BT arm, 6800', 1 mi SE Green Lk. DaLa8 3S116 Edison Butte, W. flank near summit, 5440'

| Sample | 35117 | 35118 | 35119 | 3s120 | 3s122 | 3s123 | 3s124 | 3s125 | 35126 | 35127 |
|--|--|--|--|--|--|--|--|--|--|--|
| Si02 | 74.3 | 73.8 | 65.8 | 63.4 | 73.6 | 63.5 | 72.7 | 64.4 | 73.4 | 72.7 |
| Ti02 | 0.16 | 0.27 | 0.91 | 1.09 | 0.29 | 1.08 | 0.33 | 1.04 | 0.28 | 0.33 |
| Al203 | 14.1 | 14.3 | 16.0 | 16.6 | 14.2 | 16.5 | 14.7 | 16.3 | 14.6 | 14.6 |
| Fe0* | 1.64 | 1.72 | 4.49 | 5.46 | 1.80 | 5.42 | 2.01 | 5.07 | 1.84 | 2.01 |
| MnO | 0.06 | 0.05 | 0.10 | 0.11 | 0.05 | 0.11 | 0.06 | 0.12 | 0.06 | 0.06 |
| MgO | 0.07 | 0.30 | 1.35 | 1.72 | 0.44 | 1.73 | 0.37 | 1.45 | 0.33 | 0.52 |
| Ca0 | 0.70 | 1.41 | 3.52 | 4.29 | 1.65 | 4.21 | 1.52 | 3.80 | 1.52 | 1.84 |
| Na20 | 5.36 | 4.71 | 5.30 | 5.06 | 4.48 | 5.09 | 5.03 | 5.27 | 4.86 | 4.55 |
| K20 | 3.48 | 3.26 | 2.34 | 2.00 | 3.42 | 2.00 | 3.20 | 2.10 | 3.10 | 3.21 |
| P205 | 0.02 | 0.06 | 0.26 | 0.29 | 0.07 | 0.29 | 0.07 | 0.33 | 0.07 | 0.08 |
| TOTAL | 101.56 | 100.63 | 100.80 | 100.60 | 100.82 | 100.83 | 100.88 | 100.46 | 101.11 | 101.68 |
| Rb | 86 | 72 | 46 | 38 | 84 | 38 | 66 | 40 | 62 | 76 |
| Ba | 785 | 818 | 653 | 597 | 834 | 575 | 850 | 608 | 808 | 763 |
| Sr | 50 | 165 | 329 | 393 | 191 | 388 | 181 | 356 | 180 | 217 |
| Cs | 1.35 | 2.00 | 1.27 | 0.65 | 2.45 | 0.61 | 0.99 | 1.16 | 1.39 | 2.11 |
| Sc | 4.52 | 2.73 | 11.45 | 14.50 | 3.40 | 14.47 | 3.60 | 13.23 | 3.18 | 3.78 |
| V | 0 | 13 | 67 | 104 | 14 | 95 | 21 | 53 | 17 | 16 |
| Co | 0.83 | 1.92 | 7.73 | 10.59 | 2.69 | 10.59 | 2.21 | 7.90 | 1.97 | 2.89 |
| Ni | 10 | <20 | <40 | 4 | <30 | 26 | 12 | <40 | <20 | 14 |
| Cr | 2 | 2 | 5 | 4 | 3 | 3 | 2 | <5 | 2 | 3 |
| Zn | 95 | 73 | 140 | 180 | 81 | 122 | 160 | 146 | 78 | 78 |
| La | 28.6 | 19.9 | 20.6 | 20.3 | 20.8 | 19.6 | 21.2 | 20.1 | 20.5 | 19.7 |
| Ce | 54.5 | 43.8 | 47.1 | 43.7 | 45.4 | 45.7 | 47.4 | 47.6 | 46.7 | 44.2 |
| Nd | 24.9 | 15.2 | 20.4 | 21.7 | 15.5 | 22.5 | 16.6 | 22.2 | 16.3 | 16.4 |
| Sm | 4.88 | 2.75 | 4.95 | 5.07 | 3.11 | 5.10 | 3.03 | 5.46 | 3.08 | 3.07 |
| Eu | 0.62 | 0.56 | 1.28 | 1.48 | 0.62 | 1.43 | 0.69 | 1.59 | 0.64 | 0.62 |
| Tb | 0.82 | 0.42 | 0.90 | 0.76 | 0.53 | 0.87 | 0.49 | 1.02 | 0.50 | 0.53 |
| Yb | 3.3 | 1.7 | 2.6 | 2.7 | 1.8 | 2.5 | 2.0 | 2.8 | 1.9 | 1.8 |
| Lu | 0.52 | 0.28 | 0.41 | 0.41 | 0.30 | 0.41 | 0.33 | 0.47 | 0.31 | 0.28 |
| La/Sm | 3.70 | 4.57 | 2.63 | 2.53 | 4.22 | 2.43 | 4.42 | 2.32 | 4.20 | 4.05 |
| ZHREE | 49.26 | 25.70 | 44.22 | 41.72 | 29.14 | 43.11 | 30.17 | 49.59 | 29.28 | 28.52 |
| Eu/Eu* | 0.38 | 0.62 | 0.75 | 0.89 | 0.59 | 0.83 | 0.68 | 0.84 | 0.62 | 0.59 |
| Zr Hf Ga Y Nb Ta Th U | 223 7.6 16 25 18.3 1.04 9.0 2.7 | 197 5.4 11 16 11.9 0.73 6.9 2.4 | 206 5.6 22 26 10.9 0.74 4.7 1.7 | 197 5.1 21 11.7 0.70 3.8 1.9 | 168 5.0 13 18 10.8 0.61 8.8 3.1 | 197 5.4 21 12.0 0.69 3.8 1.4 | 228 6.2 14 17 13.7 0.82 6.6 2.6 | 212 5.7 19 29 12.7 0.72 3.8 1.3 | 176 5.1 17 16 12.5 0.81 6.8 2.5 | 172 4.9 15 18 10.4 0.73 7.6 2.9 |
| Series | TVC | ss | ss | SS | SS | SS | SS | SS | ss | ss |
| Label | + | o | o | o | o | o | o | o | o | o |

Lat 43.95220 44.07730 44.07628 44.07624 44.07363 44.07574 44.07912 44.08730 44.07056 44.06664 -Long 121.6224 121.7651 121.7598 121.7757 121.7769 121.7870 121.7980 121.8187 121.7459 121.7429

3S117 Dome 1.5 mi NW of Edison Bu., 5730' 3S118 Flow 1 mi N. of Moraine Lk., 7380' 3S119 Flow 1 mi NNE of Moraine Lk., 7320' 3S120 Flow 0.3 mi N of Rock Mesa NE, 7000' 3S122 Rock Mesa NE, N. flank of dome 3S123 Flow 0.5 mi NNE of Rock Mesa, 6620'. WDaLa6 3S124 Flow(?) 1 mi NW of Rock Mesa, 6640'. WDaLa6 3S124 Flow(?) 1 mi NW of Rock Mesa, 5980'. WDaLa1 3S126 Dome on SE flank of SS, 6600' RdLa9 3S127 Goose Ck. Holocene dome, E. Flank, 6480'

| Sample | 35128 | 3\$130 | 3s131 | 3s132 | 3\$133 | 3\$134 | 3s135 | 3\$136 | 3\$137 | 35138 |
|--|--|--|--|--|--|--|--|--|--|--|
| Si02 | 56.0 | 62.0 | 55.0 | 61.9 | 58.2 | 65.7 | 62.6 | 63.9 | 64.7 | 62.3 |
| Ti02 | 1.76 | 0.75 | 1.74 | 1.10 | 1.30 | 0.73 | 1.18 | 1.01 | 0.92 | 1.25 |
| Al203 | 16.5 | 16.5 | 16.8 | 17.2 | 17.2 | 16.2 | 16.6 | 16.4 | 16.2 | 16.4 |
| Fe0* | 9.18 | 5.25 | 9.41 | 5.91 | 7.82 | 5.03 | 5.87 | 5.47 | 5.35 | 6.16 |
| MnO | 0.16 | 0.10 | 0.16 | 0.14 | 0.14 | 0.14 | 0.14 | 0.14 | 0.14 | 0.14 |
| MgO | 3.36 | 3.85 | 3.83 | 1.86 | 2.92 | 0.99 | 1.75 | 1.40 | 1.26 | 1.76 |
| Ca0 | 7.03 | 5.72 | 7.44 | 4.59 | 6.11 | 2.72 | 4.30 | 3.66 | 3.23 | 4.33 |
| Na20 | 4.69 | 4.08 | 4.60 | 5.52 | 4.86 | 6.26 | 5.60 | 5.80 | 6.09 | 5.62 |
| K20 | 0.97 | 1.64 | 0.86 | 1.47 | 1.23 | 1.89 | 1.59 | 1.72 | 1.79 | 1.65 |
| P205 | 0.33 | 0.12 | 0.29 | 0.33 | 0.29 | 0.25 | 0.37 | 0.37 | 0.33 | 0.39 |
| T0TAL | 100.70 | 100.62 | 100.81 | 101.44 | 100.80 | 101.27 | 101.25 | 100.92 | 100.48 | 100.53 |
| Rb | 17 | 35 | 14 | 24 | 19 | 31 | 27 | 28 | 27 | 29 |
| Ba | 299 | 456 | 312 | 508 | 436 | 607 | 540 | 583 | 555 | 553 |
| Sr | 440 | 352 | 485 | 483 | 520 | 272 | 412 | 392 | 318 | 404 |
| Cs | 0.23 | 0.69 | 0.43 | 0.68 | 0.53 | 0.85 | 0.76 | 0.95 | 0.87 | 0.86 |
| Sc V Co Ni Cr Zn | 29.59 286 23.73 2 9 | 15.68 109 20.66 49 84 | 30.03 318 27.16 10 21 | 16.01 58 9.76 <40 3 152 | 21.75 184 19.01 3 3 | 12.81 12 3.56 18 <6 208 | 17.14 74 9.04 4 100 | 15.75 50 7.12 <40 3 181 | 14.21 14 4.74 <70 <7 | 17.64 77 9.30 7 <9 |
| La Ce Nd Sm Eu Tb Tb Yb Lu La/Sm ZHREE Eu/Eu* | 13.3 30.3 18.6 4.83 1.58 0.86 3.0 0.41 1.74 45.28 0.95 | 14.1 28.5 15.0 3.15 0.90 0.57 2.0 0.28 2.83 30.32 0.83 | 12.0 27.5 17.8 4.39 1.51 0.72 2.5 0.37 1.73 38.68 1.03 | 18.6 44.1 5.29 1.61 1.04 2.7 0.41 2.22 47.68 0.86 | 15.8 35.5 20.4 4.67 1.45 0.76 2.6 0.36 2.14 39.70 0.93 | 21.3 53.8 28.0 6.75 1.80 1.27 4.2 0.62 1.99 66.22 0.77 | 19.6 41.7 23.9 5.81 1.80 0.91 3.1 0.45 2.13 48.06 0.93 | 22.0 53.2 26.0 6.27 1.82 1.13 3.2 0.49 2.22 54.45 0.84 | 20.5 48.5 27.7 6.55 1.87 1.04 3.7 0.58 1.98 57.70 0.86 | 20.4 45.9 25.6 6.13 1.85 0.94 3.5 0.50 2.10 52.15 0.92 |
| Zr Hf Ga Y Nb Ta Th U | 145 3.6 19 28 9.8 0.56 1.5 0.5 | 110 2.9 13 19 7.4 0.45 3.0 1.1 | 134 3.3 20 25 8.1 0.45 1.1 <2 | 176 4.6 20 27 10.2 0.54 2.3 0.9 | 154 3.7 18 24 8.2 0.48 1.8 0.7 | 241 7.0 21 38 12.7 0.62 2.9 1.3 | 190 5.0 22 31 12.3 0.62 2.6 1.0 | 203 5.8 20 32 11.9 0.59 2.6 1.2 | 225 6.1 23 14.3 0.69 2.7 1.2 | 196 5.3 24 31 11.1 0.58 2.6 1.3 |
| Series | BT | NSMS | NSMS | MS | NSMS | MS | MS | MS | MS | MS |
| Label | + | O | O | X | O | X | X | X | X | X |

Lat 44.06570 44.16157 44.15917 44.13528 44.14319 44.15447 44.15859 44.16345 44.16512 44.17088 -Long 121.7356 121.8175 121.8165 121.8048 121.8190 121.8264 121.8073 121.8079 121.8036 121.8040

3S128 Flow along Fall Ck., 6200'. AnLa7 3S130 Flow 0.5 mi SSW of Obsidian Falls, 6560' 3S131 Flow 0.75 mi SSW Obsidian Falls, 6580', on trail 3S132 Flow 1 mi E. of Linton Sp., 7000' 3S133 Flow on N. Fork of Linton Ck., 6580' 3S134 Lane Mesa, E. margin, 6460' 3S135 Flow 1 mi SSE of Sister Sp., 6980' 3S136 Platy flow 0.5 mi SE of Sister Sp., 6800' 3S137 Flow 0.5 mi ESE of Sister Sp., 7040' 3S138 Flow 0.5 mi E. of Sister Sp., 6880'

| Sample SiO2 TiO2 Al2O3 FeO* MnO MgO CaO Na2O K2O P2O5 TOTAL | 35139 76.1 0.14 13.4 1.16 0.04 0.11 0.97 4.46 3.41 0.04 101.99 | 3s140 73.7 0.26 14.2 1.95 0.09 0.31 1.43 4.42 3.21 0.06 100.90 | 35143 73.3 0.29 14.4 2.07 0.06 0.35 1.56 4.13 3.14 0.07 100.73 | 3s145 71.8 0.35 14.9 2.20 0.07 0.53 2.11 4.61 3.09 0.09 100.67 | 35146 59.9 1.28 16.1 7.29 0.14 3.37 5.33 4.69 1.59 0.29 100.70 | 35147 64.8 0.89 16.3 5.13 0.14 1.19 3.23 6.07 1.78 0.32 101.28 | 3s148 72.5 0.33 14.5 2.06 0.61 1.86 4.77 3.15 0.07 100.85 | 35149 56.7 1.18 17.8 7.31 0.13 4.06 7.16 4.18 1.00 0.25 100.80 | 3s150 65.4 0.85 16.1 4.53 0.09 1.72 4.10 4.73 2.25 0.20 101.24 | 3s151 65.8 0.94 16.0 4.56 0.10 1.31 3.50 5.40 2.31 0.27 101.08 |
|--|--|--|--|---|--|--|--|--|--|--|
| Rb Ba Sr Cs | 79 876 119 1.94 | 70 813 168 1.60 | 69 785 179 1.50 | 73 755 251 1.93 | 30 432 380 0.91 | 29 550 325 0.75 | 71 742 193 2.02 | 19 310 504 0.53 | 48 609 335 1.42 | 47 629 331 1.30 |
| Sc V Co Ni Cr Zn | 1.63 9 0.58 <10 1 64 | 2.66 16 1.96 <10 11 56 | 3.10 19 2.09 <10 11 53 | 3.92 29 3.53 16 7 90 | 19.56 164 19.51 17 43 | 14.25 20 4.53 <70 4 | 4.60 31 3.75 9 9 | 20.17 169 22.45 32 58 | 12.24 87 11.14 26 16 96 | 12.12 57 7.80 6 3 |
| La Ce Nd Sm Eu Tb Tb Yb Lu La/Sm ZHRE E Eu/Eu* | 20.0 44.6 14.9 2.63 0.42 0.39 1.5 0.24 4.80 22.87 0.49 | 19.1 35.9 13.5 2.84 0.54 0.38 1.7 0.27 4.25 24.54 0.59 | 18.9 34.2 14.0 2.74 0.60 0.38 1.8 0.26 4.36 24.71 0.68 | 18.8 43.1 16.7 2.97 0.70 0.54 1.7 0.27 4.00 27.94 0.68 | 16.0 36.0 18.6 5.04 1.43 0.79 2.8 0.40 2.00 42.53 0.86 | 20.6 49.6 26.3 1.90 1.09 3.8 0.58 2.04 59.24 0.88 | 19.2 38.6 17.4 3.06 0.64 0.41 1.9 0.29 3.96 26.75 0.65 | 11.6 25.5 14.2 3.85 1.29 0.57 2.2 0.30 1.90 31.89 1.02 | 17.2 41.7 16.9 4.05 1.11 0.68 2.1 0.34 2.68 34.99 0.81 | 21.2 44.6 21.7 5.23 1.39 0.74 2.9 0.42 2.56 42.56 0.82 |
| Zr Hf Ga Y Nb Ta Th U | 96 3.2 15 10.0 0.74 7.3 2.6 | 167 4.3 16 17 11.5 0.75 6.1 2.5 | 174 4.3 16 18 10.7 0.75 5.7 2.3 | 168 4.6 16 17 10.6 0.68 7.2 2.6 | 182 4.6 16 27 10.5 0.53 2.7 1.1 | 224 6.2 20 34 13.0 0.68 2.7 1.0 | 153 4.5 17 18 8.7 0.67 7.0 2.4 | 143 3.4 17 20 8.1 0.40 1.6 0.7 | 185 5.1 16 21 10.4 0.67 4.7 1.5 | 211 5.6 20 12.9 0.71 4.5 1.8 |
| Series Label | MS X | SS o | SS o | SS o | SS | MS | SS o | ss o | SS o | SS o |

Lat 44.17746 44.12136 44.08801 44.12244 44.11236 44.12094 44.12735 44.09796 44.12165 44.13126 -Long 121.8157 121.7546 121.7726 121.7629 121.7642 121.7941 121.7876 121.7670 121.7595 121.7649

3S139 Obsidian Cliffs, by Prouty Plaque, 6560' 3S140 W. flank of largre dome, easternmost Chambers Lakes, 7600' 3S143 Dome at terminus of Clark Glacier, SS, 8400' 3S145 Holocene dome 0.5 mi NW of Carver Lk., 7500',BT-205 3S146 Flow on W. flank of Prouty Glacier, SS, 8300'. WAnLa2 3S147 E. arm of flow, W. Chambers Lk., 7600'. TS-1603 3S148 Dome W. Chambers Lk, 7600'. TS-1601 3S149 Dike near SS summit, 9300', TS-1712 3S150 Flow at E. Chambers Lk, 7440'. WDaLa2, TS-1716 3S151 Flow on SE flank of SS at 9800', TS-1711

| | | | | | | Misc. | Analyses |
|----------------------------|----------|----------|----------|----------|----------|----------|----------|
| Sample | 3S152 | 3S157 | 3S159 | 3S162 | 3SXBP | 3\$071 | 3\$063 |
| SiOZ | 63.6 | 53.6 | 52.9 | 64.4 | 74.2 | 71.7 | 65.2 |
| Ti02 | 1.08 | 1.07 | 1.62 | 0.72 | 0.16 | 0.31 | 0.66 |
| AL203 | 16.4 | 18.4 | 15.8 | 15.8 | 14.2 | 14.5 | 16.6 |
| FeO* | 5.42 | 7.49 | 9.55 | 5.17 | 1.82 | 3.47 | 4.32 |
| MnO | 0.11 | 0.14 | 0.17 | 0.10 | 0.07 | 0.12 | 0.09 |
| MgO | 1.63 | 5.20 | 5.66 | 1.65 | 0.08 | 0.00 | 1.94 |
| CaO | 4.26 | 9.00 | 8.21 | 3.87 | 0.81 | 0.62 | 4.38 |
| Na20 | 5.13 | 3.77 | 3.60 | 4.90 | 4.46 | 5.68 | 4.82 |
| K20 | 2.05 | 0.62 | 0.71 | 2.01 | 3.47 | 2.79 | 1,92 |
| P205 | 0.30 | 0.18 | 0.28 | 0.28 | 0.03 | 0.03 | 0.16 |
| TOTAL | 100.54 | 99.47 | 98.50 | 98.90 | 99.63 | 100.35 | 101.78 |
| | | | | | | | |
| Rb | 40 | 7 | 13 | 73 | 84 | 54 | 39 |
| ва | 564 | 270 | 204 | 611 | 794 | 750 | 5/5 |
| Sr | 389 | 550 | 400 | 349 | 65 | (6 | 388 |
| CS | 1.10 | | | | 3.10 | 0.40 | 0.90 |
| Sc | 14.86 | | | | 4.29 | 11.54 | 10.47 |
| V | 94 | | | | 0 | 0 | 88 |
| Со | 10.60 | 68 | 74 | 25 | 0.50 | 0.27 | 11.26 |
| Ni | 23 | 72 | 10 | 2 | 11 | | |
| Cr | 2 | 93 | 113 | 7 | 1 | 2 | 23 |
| Zn | | 76 | 94 | 73 | 105 | 130 | 71 |
| | 10.8 | | | | 28.7 | 37 0 | 17.6 |
| Co | 14.0 | | | | 50.7 | 53 7 | 33 7 |
| Nd | 20.5 | | | | 26.5 | 48.0 | 16.2 |
| Sm | 5 13 | | | | 5 05 | 10.62 | 3 53 |
| Fu | 1 52 | | | | 0.83 | 1.81 | 0.99 |
| Th | 0.80 | | | | 0.05 | 1.74 | 0.48 |
| Yh | 2 7 | | | | 4 2 | 6.0 | 1 9 |
| Lu | 0 41 | | | | 0 58 | 0.94 | 0.30 |
| la/Sm | 2 44 | | | | 3 05 | 2.20 | 3.15 |
| <i>SHREE</i> | 42.57 | | | | 57.77 | 94.69 | 28.55 |
| Eu/Eu* | 0.89 | | | | 0.42 | 0.51 | 0.87 |
| , | | | | | | | |
| Zr | 200 | | | | 231 | 336 | 167 |
| Ht | 5.4 | | | | 6.7 | 9.1 | 4.2 |
| Ga | 21 | | | | 20 | 22 | 17 |
| Y | 27 | | | | | 63 | 19 |
| Nb | 11.5 | | | | 16.9 | 19.9 | 10.2 |
| Та | 0.77 | | | | 1.20 | 1.02 | 0.60 |
| Th | 3.9 | | | | 8.1 | 4.2 | 3.6 |
| U | 1.1 | | | | 2.7 | 2.1 | 1.2 |
| Series | SS | NS | MS | SS | TVC | NBRY | SC |
| Lbl | 0 | 0 | x | 0 | + | N | С |
| Lat | 11 17/17 | 44 18055 | 66 13120 | // 138// | 44 04200 | /3 03900 | 43 85802 |
| -long | 121.7700 | 121.7470 | 121.7676 | 121.7100 | 121.3330 | 121,4130 | 121.7840 |

3S152 Flow at Chambers Lk., 7480'. DaLa3, TS-1718 3S157 N. Sister flow, E. flank, 7080' UO AA data, 7/90 3S159 Flow on S. flank of MS at 7160'. UO AA data, 7/90 3S162 Flow 3 mi NE of SS a 5800', SF Squaw Ck. UO AA data, 7/90 3SXBP Bulk Bend Pumice, fresh interiors, quarry ≂1mi SW Bend

Miscellaneous analyses from the Oregon central Cascades:

3S071 Benham Butte; Main dome at parking pad along Deschutes River 3S063 E. flank of Benchmark Butte, near Crane Prarie along Century Drive

| Sample | 35002 | 35006 | 38077 | 38080 | 35087 | 3S088M | 350885 | 3\$141 | 3S142 | 35144 |
|---------------------------------|---------------------------------------|---------------------------------------|-------------------------------------|---------------------------------------|---------------------------------|---|-------------------------------------|--------------------------------------|-------------------------------------|-------------------------------|
| Si02 | 54.9 | 59.3 | 68.2 | 68.7 | 54.3 | 57.8 | 67.4 | 65.7 | 69.9 | 67.1 |
| Ti02 | 1.87 | 1.13 | 0.70 | 0.59 | 1.85 | 1.49 | 0.78 | 0.87 | 0.59 | 0.69 |
| Al203 | 16.8 | 17.1 | 15.9 | 16.3 | 16.5 | 16.7 | 16.0 | 16.0 | 15.9 | 16.3 |
| FeO* | 9.25 | 7.31 | 3.65 | 3.48 | 10.46 | 8.08 | 3.71 | 4.55 | 3.35 | 4.45 |
| MnO | 0.16 | 0.14 | 0.14 | 0.14 | 0.16 | 0.15 | 0.12 | 0.16 | 0.08 | 0.10 |
| Mg0 | 3.20 | 2.79 | 0.61 | 0.67 | 3.28 | 3.04 | 0.96 | 1.19 | 0.48 | 0.37 |
| Ca0 | 7.30 | 6.11 | 1.78 | 2.17 | 7.40 | 6.32 | 2.45 | 2.91 | 1.61 | 2.24 |
| Na20 | 4.61 | 6.23 | 5.50 | 4.28 | 4.33 | 4.83 | 5.98 | 5.50 | 3.04 | 6.16 |
| K20 | 1.15 | 1.40 | 2.75 | 2.83 | 0.97 | 1.25 | 2.34 | 2.12 | 3.82 | 2.23 |
| P205 | 0.54 | 0.25 | 0.18 | 0.14 | 0.59 | 0.32 | 0.19 | 0.25 | 0.11 | 0.18 |
| T0TAL | 100.18 | 100.82 | 100.52 | 98.95 | 100.05 | 100.29 | 99.41 | 100.48 | 98.42 | 100.98 |
| Rb | 17 | 30 | 45 | 64 | 17 | 21 | 45 | 31 | 56 | 44 |
| Ba | 413 | 468 | 946 | 850 | 429 | 367 | 650 | 768 | 740 | 805 |
| Sr | 686 | 485 | 240 | 304 | 574 | 497 | 288 | 316 | 135 | 289 |
| Cs | 0.70 | 0.90 | 1.10 | 1.60 | <.7 | 0.81 | 1.43 | 1.40 | 1.70 | 0.64 |
| Sc V Co Ni Cr Zn | 27.06 188 17.90 0 1 73 | 20.45 192 19.60 8 7 76 | 12.32 7 2.24 8 <3 96 | 10.28 9 1.19 11 <3 109 | 32.96 281 25.90 3 8 | 25.88 211 20.92 3 11 107 | 11.89 18 2.23 8 4 89 | 12.56 38 4.52 6 2 100 | 7.37 28 2.49 11 3 55 | 13.78 5 1.65 7 <6 |
| La | 13.4 | 18.8 | 25.5 | 24.9 | 18.9 | 15.2 | 23.8 | 25.5 | 22.4 | 22.9 |
| Ce | 29.5 | 37.8 | 70.1 | 50.8 | 38.8 | 37.6 | 60.1 | 57.5 | 44.9 | 49.2 |
| Nd | 19.3 | 19.7 | 31.8 | 29.5 | 23.0 | 20.9 | 30.5 | 32.0 | 25.0 | 26.9 |
| Sm | 5.31 | 4.71 | 7.35 | 7.09 | 6.93 | 5.14 | 6.42 | 7.35 | 5.69 | 6.41 |
| Eu | 1.83 | 1.40 | 2.35 | 1.76 | 2.42 | 1.65 | 1.64 | 2.01 | 1.00 | 2.05 |
| Tb | 0.77 | 0.70 | 1.30 | 1.06 | 1.19 | 0.84 | 1.16 | 1.13 | 0.84 | 1.01 |
| Yb | 2.2 | 2.6 | 4.5 | 4.0 | 2.9 | 2.7 | 3.9 | 4.5 | 3.9 | 3.5 |
| Lu | 0.34 | 0.37 | 0.65 | 0.61 | 0.45 | 0.46 | 0.61 | 0.65 | 0.58 | 0.50 |
| La/Sm | 1.59 | 2.52 | 2.19 | 2.22 | 1.72 | 1.87 | 2.34 | 2.19 | 2.49 | 2.26 |
| ZHREE | 37.39 | 38.73 | 69.02 | 60.53 | 53.06 | 44.97 | 62.14 | 65.59 | 54.40 | 53.64 |
| Eu/Eu* | 1.06 | 0.91 | 0.93 | 0.76 | 1.03 | 0.96 | 0.74 | 0.83 | 0.54 | 0.96 |
| Zr | 118 | 163 | 278 | 240 | 136 | 153 | 241 | 269 | 365 | 244 |
| Kf | 2.6 | 4.2 | 7.5 | 5.7 | 3.3 | 4.1 | 6.9 | 7.1 | 7.7 | 7.1 |
| Ga | 19 | 19 | 22 | 21 | 23 | 16 | 22 | 21 | 22 | 23 |
| Y | 28 | 28 | 47 | 44 | 30 | 28 | 35 | 43 | 43 | 31 |
| Nb | 7.7 | 10.9 | 17.8 | 16.1 | 10.6 | 8.6 | 15.8 | 16.9 | 18.3 | 15.5 |
| Ta | 0.36 | 0.57 | 1.13 | 0.80 | 0.48 | 0.57 | 0.83 | 0.89 | 0.85 | 0.81 |
| Th | 1.3 | 2.8 | 5.1 | 4.4 | 1.3 | 4.3 | 4.6 | 2.7 | 4.0 | 4.0 |
| U | <2 | <3 | 0.8 | 1.6 | <3 | <3 | 1.4 | 0.8 | 1.5 | 2.3 |
| Series | DF | DF | DF | DF | DF | DF | DF | DF | DF | DF |
| Label | > | > | > | > | > | > | > | > | > | > |

Appendix D-2: Compositions and locations of Deschutes Formation rocks in the Three Sisters-Bend area.

Lat 44.11051 44.10129 44.13000 44.09803 44.10604 44.11393 44.11393 44.29165 44.28996 44.10994 -Long 121.4869 121.4659 121.4200 121.3407 121.3466 121.3387 121.3387 121.4068 121.4088 121.4638

35002 Upper Pmag-reversed flow, Bull Sp. Inlier 35006 Lower pmag-normal flow, Bull Sp. Inlier 35077 Lapilli from devitrified ignimbrite at S wall of Tumalo Dam 35080 Lapilli from ignimbrite remnant on NNW flank of Awbrey Butte, 3460' 35087 Pmag-normal flow under Bend Pumice, Coates gravel pits, N. Awbrey Bu. 35088M Mafic pumice from ignimbrite at Tumalo Ck-Des.R. junction, 3420' 35088S Silicic pumice from ignimbrite at Tumalo Ck-Des.R. junction, 3420' 35141 Pumice lapilli from Buckhorn Canyon Tuff, Hwy 26, Deep Cyn Grade. HB-14 35142 Flow on E. flank of Bull Sp. Inlier. B-65

| | | | | | | | | | | | | PL | AGIOC | LASE | | | | |
|---------|------|------|------|------|------|------|----|------|--------|------|--------|------|-------|------|------|-------|------|-----------------------------|
| Sample | Rb | Ba | Sr | Cs | Zr | Hf | Nb | Ta | Th | U | Sc | V | Co | Ni | Cr | Zn Ho | l An | Reference |
| XI | 0.08 | | 1.50 | | | | | | | | 0.02 | | 0.05 | 0.04 | | | | Irving, 1978 |
| D-Kyu | 0.46 | 1.43 | 7.0 | 0.78 | | 0.07 | | 0.05 | 0.03 | | 0.15 | | 0.29 | | 0.67 | 0.49 | 49 | Dudas et al., 1971 |
| D-GLPk | 0.06 | 0.79 | 6.0 | 0.03 | | 0.13 | | 0.03 | 0.03 | | 0.03 | | 0.06 | | 0.06 | 0.51 | 41 | Dudas et al., 1971 |
| HB-39 | 0.03 | 0.33 | 2.04 | 0.03 | | | | | | | | | | | | | 70 | Hart & Brooks, 1974 |
| SP-19 | 0.06 | 0.26 | 1.77 | | | | | | | | | | | | | | 79 | Schnitzler & Philpots, 1968 |
| PS-225 | 0.14 | 0.59 | 2.75 | | | | | | | | | | | | | | 46 | Philpots & Schnitzler, 1970 |
| HN-225 | 0.02 | 0.30 | 2.87 | | | | | | | | | | | | | | | Higuchi & Nagasawa, 1969 |
| D-34b | 0.30 | | 1.67 | | | 0.02 | | | 0.01 0 | 0.01 | | | | | | | | Dostal et al., 1983 |
| PS-208 | 0.03 | 0.15 | 1.36 | | | | | | | | | | | | | | 65 | Philpots & Schnitzler, 1970 |
| H-688MS | 0.05 | 0.21 | 1.98 | 0.06 | 0.05 | 0.01 | | 0.01 | | | 0.02 0 | 0.03 | 0.04 | 0.05 | 0.05 | | 61 | Hughes, 1983 |
| PS-204 | 0.19 | 0.24 | 1.66 | | | | | | | | | | | | | | 87 | Philpots & Schnitzler, 1970 |
| LC-11 | | 0.08 | | | | 0.02 | | | | | 0.02 | | | | | 0.13 | 71 | Luhr & Carmichael, 1980 |
| PS-240 | 0.03 | 0.06 | 1.30 | | | | | | | | | | | | | | 97 | Philpots & Schnitzler, 1970 |
| PS-233 | 0.07 | 0.14 | 1.31 | | | | | | | | | | | | | | 93 | Philpots & Schnitzler, 1970 |
| PS-184 | 0.04 | 0.13 | 1.66 | | | | | | | | | | | | | | 80 | Philpots & Schnitzler, 1970 |
| PS-271 | 0.06 | 0.20 | 1.83 | | | | | | | | | | | | | | - 77 | Philpots & Schnitzler, 1970 |
| LC-6 | | 80.0 | | | | 0.01 | | | | | 0.02 | | | | | 0.25 | 58 | Luhr & Carmichael, 1980 |
| LC-7 | | 0.10 | | | | 0.02 | | | | | 0.02 | | | | | 0.16 | 60 | Luhr & Carmichael, 1980 |
| PS-218 | 0.05 | 0.36 | 2.84 | | | | | | | | | | | | | | 51 | Philpots & Schnitzler, 1970 |
| D-SHy | 0.18 | 0.34 | 4.5 | 0.13 | | 0.17 | | 0.08 | 0.07 | | 0.05 | | 0.06 | | 0.06 | 0.26 | 44 | Dudas et al., 1971 |
| BD-PLA | 0.30 | 0.27 | 3.40 | 0.03 | 0.20 | 0.03 | | 0.03 | 0.01 | | 0.01 | | 0.07 | | 0.01 | 0.44 | | Bacon & Druitt, 1988 |
| NC-15 | 0.19 | 1.20 | 10.5 | 0.09 | 0.05 | 0.06 | | 0.03 | 0.05 0 | 0.07 | 0.10 | | | | | 0.44 | 41 | Nash & Crecraft, 1985 |
| NC-12 | 0.09 | 1.20 | 8.7 | 0.10 | 0.05 | 0.05 | | 0.04 | 0.03 (| 0.06 | 0.10 | | | | | 0.40 | 42 | Nash & Crecraft, 1985 |
| NC-6 | 0.07 | 0.41 | 4.0 | 0.16 | 0.08 | 0.14 | | 0.03 | 0.03 (| 0.10 | 0.01 | | | | | 1.70 | 39 | Nash & Crecraft, 1985 |
| BD-PLR | 0.30 | 0.48 | 4.40 | 0.03 | 0.20 | 0.03 | | 0.03 | 0.01 | | 0.01 | | 0.15 | | 0.01 | 0.48 | | Bacon & Druitt, 1988 |
| NS-1P | 0.04 | 0.31 | 4.4 | | | | | | | | | | | | | | | Nagasawa & Schnetzler, 1971 |
| NS-2P | 0.02 | 0.30 | 3.9 | | | | | | | | | | | | | | | Nagasawa & Schnetzler, 1971 |
| D-Tum | 0.07 | 0.92 | 24.3 | 0.10 | | 0.16 | | 0.09 | 0.09 | | 0.17 | | 0.51 | | 0.15 | 0.39 | 23 | Dudas et al., 1971 |

Appendix E-1: Compilation of mineral/matrix distribution coefficients for calc-alkaline and mildly alkaline rocks. Rare earth element distribution coefficients reported in appendix E-2.

| | CLINOPYROXENE | | | | | | | | | | | | | | | | | |
|--------|---------------|------|------|------|----|------|------|----|------|------|------|------|------|------|------|-----|--------|-----------------------------|
| Sample | Rb | Ba | Sr | Cs | ZΓ | Hf | Nb | Ta | Th | U | Sc | V | Со | Nī | Cr | Zn | Hol Er | Reference |
| XI | 0.00 | 0.00 | 0.10 | 0.00 | | | 0.02 | | 0.01 | 0.01 | 2.00 | 1.00 | 1.00 | 5.00 | 2.50 | | | Irving, 1978 |
| PS-193 | 0.02 | 0.01 | 0.15 | | | | | | | | | | | | | | | Philpots & Schnitzler, 1970 |
| SP-25 | 0.02 | 0.02 | 0.11 | | | | | | | | | | | | | | | Schnitzler & Philpots, 1968 |
| HB-39 | 0.00 | 0.00 | 0.08 | 0.00 | | | | | | | | | | | | | | Hart & Brooks, 1974 |
| SP-19 | 0.02 | 0.02 | 0.12 | | | | | | | | | | | | | | | Schnitzler & Philpots, 1968 |
| On-b | | 0.00 | 0.11 | | | | | | 0.01 | 0.02 | 2.9 | | 1.1 | | | | | Onouma et al., 1968 |
| D-34b | | | | | | 0.30 | | | | | 2.9 | 1.6 | 1.2 | 4.2 | 28.0 | 0.3 | | Dostal et al., 1983 |
| PS-204 | 0.00 | 0.00 | 0.08 | 0.00 | | | | | | | | | | | | | | Philpots & Schnitzler, 1970 |
| LC-11 | | 0.04 | | | | 0.26 | | | | | 6.9 | 1.1 | | 5.7 | 42.9 | 9.5 | 43 | Luhr & Carmichael, 1980 |

| Sample PS-271 LC-6 LC-7 BD-CpxA NC-15 NC-12 NC-6 PS-266 BD-CpxR | Rb 0.01 | Ba 0.04 0.05 0.02 0.10 0.13 0.13 | Sr 0.07 0.52 | Cs | Zr | Hf 0.23 0.29 0.46 0.50 0.84 0.91 0.29 | Nb | Ta 0.50 0.13 0.09 0.46 0.50 | Th 0.10 0.12 0.09 0.22 0.10 | U | CI Sc 12.5 12.3 17.0 30.6 30.0 25.1 33.0 | 3.2 2.8 | ROXENE Co 5.5 17.0 | (cor Ni 5.0 6.0 | nt.) Cr 145.0 126.0 30.0 30.0 | Zn Hol 8.9 10.6 2.0 2.0 | En 44 44 | Reference Philpots & Schnitzler, 1970 Luhr & Carmichael, 1980 Luhr & Carmichael, 1980 Bacon & Druitt, 1988 Nash & Crecraft, 1985 Nash & Crecraft, 1985 Nash & Crecraft, 1985 Philpots & Schnitzler, 1970 Bacon & Druitt, 1988 |
|--|----------------------|---|----------------------|----------------------|------|--|----|--|--|------|--|------------|------------------------------------|--------------------------|--|--|----------------|--|
| | | | | | | | | | | | | ORI | HOPYRO | XENE | | | | |
| Sample XI On-c PS-203 | Rb | Ba 0.01 | Sr 0.02 | Cs | Zr | Hf | NÞ | Та | Th | 0.01 | Sc 1.40 1.2 | ¥ 3.40 | Co 1.30 2.1 | Nī 3.00 | Cr 5.00 | Zn Mol | En | Reference Irving, 1978 Onouma et al., 1968 Philpots & Schnitzler, 1970 Philpots & Schnitzler 1970 |
| PS-271 PS-184 LC-6 | 0.02 | 0.01 | 0.02 | | | 0.22 | | | 0.22 | | 3.9 3.7 | 1.5 | | 7.0 | 21.0 | 3.9 | 71 71 | Philpots & Schnitzler, 1970 Luhr & Carmichael, 1980 |
| BD-OpxA NC-15 NC-12 | 0.01 | 0.10 | 0.01 | 0.01 0.40 0.76 | 0.11 | 0.11 0.20 0.17 | | 0.11 0.19 0.12 | 0.14 | 0.16 | 4.3 12.00 10.0 | ,,,,,, | 12.0 | | 10.0 | 3.7 | | Bacon & Druitt, 1988 Nash & Crecraft, 1985 Nash & Crecraft, 1985 |
| BD-OpxR NS-1H NS-2H | 0.01 0.00 0.00 | 0.10 0.00 0.00 | 0.01 0.01 0.05 | 0.01 | 0.11 | 0.11 | | 0.11 | 0.14 | | 7.1 | | 38.0 | | 10.0 | 6.0 | | Bacon & Druitt, 1988 Nagasawa & Schnetzler, 1971 Nagasawa & Schnetzler, 1971 |
| | | | | | | | | | | | | H | ORNBLEI | ΦE | | | | |
| Sample HN-3 | Rb | Ba | Sr | Cs | Zr | Hf | Nb | Ta | Th | U | Sc 2.2 | v | Со | Nī | Cr | Zn | | Reference Higuchi & Nagasawa, 1969 Sebritzion & Philpots, 1968 |
| SP-29 D34-b PS-233 | 0.43 0.33 0.05 | 0.42 | 0.55 | | | 0.50 | | | 0.50 | 0.10 | 3.2 | 3.4 | 2.0 | 6.8 | 12.5 | 0.4 | | Dostal et al., 1983 Philpots & Schnitzler, 1970 |
| LC-6 LC-7 BD-HDA | | 0.17 | | | | 0.34 0.45 0.54 | | 0.59 | 0.18 | | 10.3 9.5 13.0 | 7.0 6.5 | 6.1 | 14.0 14.0 | 51.3 89.8 40.0 | 5.0 8.7 1.6 | | Luhr & Carmichael, 1980 Luhr & Carmichael, 1980 Bacon & Druitt, 1988 |
| BD-HbR NS-1Hb NS-2Hb | 0.01 | 0.30 | 0.02 | | | 0.52 | | 0.43 | 0.16 | | 14.0 | | 37.0 | | 40.0 | 2.3 | | Bacon & Druitt, 1988 Nagasawa & Schnetzler, 1971. Nagasawa & Schnetzler, 1971. |

| | | | | | | | | | | | | | OLIVI | NE | | | | |
|---------|------|------|------|------|------------|------|-----------|------|------|------|-------|------|--------|------|--------|--------|--------|-----------------------------|
| Sample | Rb | Ba | Sr | Cs | Zr | Hf | Nb | Ta | Th | U | Sc | V | Со | Ni | Cr | Zn H | iol Fo | Reference |
| LS-48 | | | | | | | | | | | 0.18 | | 2.75 | | 1.06 | | 88 | Leeman & Scheidegger, 1977 |
| SP-25 | 0.01 | 0.01 | 0.02 | | | | | | | | | | | | | | | Schnitzler & Philpots, 1968 |
| LS-15 | | | | | | | | | | | 0.22 | | 4.19 | | 3.06 | | 72 | Leeman & Scheidegger, 1977 |
| LS-36 | | | | | | | | | | | 0.15 | | 3.24 | 17.6 | 1.23 | | 87 | Leeman & Scheidegger, 1977 |
| 15-44 | | | | | | | | | | | 0.14 | | 3.24 | 16.1 | 2.70 | | 87 | Leeman & Scheidegger, 1977 |
| HB-39 | 0.00 | 0.00 | 0.00 | 0.00 | | | | | | | | | | | | | 86 | Hart & Brooks, 1974 |
| 15-82 | 0.00 | 0100 | | | | | | | | | 0.14 | | 4.11 | 23.5 | 1.65 | | 88 | Leeman & Scheidegger, 1977 |
| 15-5 | | | | | | | | | | | 0.15 | | 3.49 | 19.0 | 2.92 | | 85 | Leeman & Scheidegger, 1977 |
| PS-19 | 0.01 | 0.01 | 0.01 | | | | | | | | | | | | | | | Schnitzler & Philpots, 1968 |
| D-SVG | | | | | | | | | | | | 0.06 | 6.60 | 22.0 | 0.70 | 1.20 | | Dostal et al., 1983 |
| H-688MS | | | | | | | | 0.1 | 0.1 | | 0.39 | 0.08 | 7.70 | 47.2 | 2.5 | | 75 | Hughes, 1983 |
| LC-11 | | 0.01 | | 0.01 | | 0.01 | | | 0.0 | | 0.30 | 0.08 | | 58.0 | 34.0 | 1.5 | 75 | Luhr & Carmichael, 1980 |
| 20 11 | | | | | | | | | | | | | | | | | | |
| | | | | | | | | | | | | | | | | | | |
| | | | | | | | | | | | | | MAGNET | ITE | | | | |
| Sample | Rb | Ba | Sr | Cs | Zr | Hf | NP | Ta | Th | U | Sc | V | Co | NI | Cr | Zn | τUsp | Reference |
| D-34b | | | | | | | | | | | | 26.0 | 7.4 | 29.0 | 153 | 5.1 | 1.4 | Dostal et al., 1985 |
| LC-11 | | 0.40 | | | | 0.62 | | | | | 1./ | 8.7 | | 9.6 | 93 | 5.4 | 01 | Lunr & Carmichael, 1980 |
| LC-6 | | 0.13 | | | | 0.31 | | | | | 3.2 | 38.7 | | 15.0 | 100 | 12.0 | 03 | Lunr & Carmichael, 1900 |
| LC-7 | | 0.20 | | | | 0.37 | | 47.0 | | 0.5 | 2.5 | 34.3 | | 13.0 | 122 | 11.0 | 09 | Lunr & Carmichaet, 1900 |
| NC-15 | | | | | | 2.80 | | 13.0 | 0.2 | 0.5 | 8.0 | | | | 40 | 50.0 | | Nash & Crecratt, 1905 |
| NC-12 | | | | | | 1.70 | | 5.6 | 0.1 | 0.2 | 1.4 | | | | 50.0 | 50.0 | | Nash & Crecratt, 1905 |
| NC-6 | | | | | | 0.90 | | 7.9 | 0.3 | 0.8 | 4.6 | | | | 50.0 | 60.0 | | Nash & Lrecratt, 1965 |
| BD-MgtR | | | | | 0.24 | 0.24 | | 1.2 | | | 5.0 | | 80.0 | | 50.0 | 15.0 | | Bacon & Druitt, 1900 |
| | | | | | | | | | | | | | | | | | | |
| | | | | | | | | | | | | | TLMEN | ITE | | | | |
| Sample | Rb | Ba | Sr | Cs | Zr | Hf | Nb | Ta | Th | U | Sc | V | Co | Ni | Cr | Zn | | Reference |
| NC-6 | | | | | | 3.00 | | 20.0 | 7.3 | 3.0 | 2.3 | | | | 3.0 | 75.0 | | Nash & Crecraft, 1985 |
| XI | | | | | 0.27 | | 0.80 | | | | | 12 | | | 6 | | | Irving, 1978 |
| | | | | | | | | | | | | | | | | | | |
| | | | | | | | | | | | | | | | | | | |
| | | | | | | | | | | | 0.824 | 65 | APATI | TE | 0.2555 | 1224.2 | | |
| Sample | Rb | Ba | Sr | Cs | Zr | Hf | Nb | Та | Th | U | SC | V | Co | Ni | Cr | Zn | | Reference |
| XI | | | 1.00 | | | | | | 1.50 | 0.50 | | | | | | | | Irving, 1978 |
| F-1 | | | | | 0.64 | 0.73 | | | | | | | | | | | | Fujimaki, 1986 |
| | | | | | | | | | | | | | | | | | | |
| | | | | | | | | | | | | | 7100 | 100 | | | | |
| Sampla | Dh | Po | S. | 60 | 7- | H.F | Mb | Te | Th | ш | Sc | V | Co | Mi | Cr | Zn | | Reference |
| sample | KD | 60 | 21 | LS | Z F | 059 | ND | • d | 6.13 | 0 | 30 | | 20 | 10.1 | | Bell I | | Fujimaki, 1986 |
| r-1 | | | | | | 920 | | | | | | | | | | | | rajimakiy 1700 |

Appendix E-2: Published rare earth element phenocryst/matrix distribution coefficients.

| | | | | | | PL | | ASE | | |
|------------------|-------|-------|------|------|------|------|------|-------------|---------|--|
| Sample | La | Ce | Nd | Sm | Eu | Tb | Yb | Lu | An | Reference |
| D-Kyu | 0.28 | 0.16 | 0.16 | 0.12 | 0.75 | 0.12 | 0.07 | 0.06 | 49 | Oudas et al., 1971 |
| SD-10 | 0.49 | 0.40 | 0.29 | 0.11 | 1.78 | 0.15 | 0.09 | 0.07 | 4 | Uudas et al., 1971 Schnitzlan & Dhilpoto 1069 |
| PS-225 | | 0.00 | 0.07 | 0.10 | 0.33 | 0.09 | 0.01 | | 4 | Philpots & Schnitzler 1970 |
| HN-225 | 0.14 | 0.09 | 0.04 | 0.04 | 0.32 | 0.03 | 0.02 | | 40 | Higuchi & Nagasawa, 1969 |
| PS-208 | | 0.02 | 0.02 | 0.02 | 0.23 | 0.01 | 0.03 | 0.04 | 65 | Philpots & Schnitzler, 1970 |
| H-688MS | 0.06 | 0.06 | | 0.03 | 0.30 | 0.03 | | 0.01 | 61 | Hughes, 1983 |
| PS-204 | | 0.28 | 0.20 | 0.17 | 0.44 | 0.20 | 0.30 | 0.24 | 87 | Philpots & Schnitzler, 1970 |
| LC-11 | 0.13 | 0.12 | 0.13 | 0.07 | 0.36 | 0.05 | 0.01 | 0.01 | 71 | Luhr & Carmichael, 1980 |
| PS-240 | | 0.08 | 0.02 | 0.03 | 0.06 | 0.01 | | | 97 | Philpots & Schnitzler, 1970 |
| PS-184 DS-271 | | 0.06 | 0.05 | 0.01 | 0.19 | 0.03 | 0.02 | 0.03 | 80 | Philpots & Schnitzler, 1970 |
| 10-6 | 0 21 | 0.20 | 0.14 | 0.11 | 0.73 | 0.05 | 0.03 | 0.05 | 58 | Lubr & Cosmichael 1980 |
| LC-7 | 0.20 | 0.18 | 0.15 | 0.10 | 0.70 | 0.08 | 0.01 | 0.01 | 60 | Luhr & Carmichael 1980 |
| PS-218 | | 0.24 | 0.17 | 0.13 | 2.11 | 0.09 | 0.08 | 0.06 | 51 | Philpots & Schnitzler, 1970 |
| 0-SHy | 0.49 | 0.22 | 0.14 | 0.13 | 0.97 | 0.18 | 0.11 | 0.10 | 44 | Oudas et al., 1971 |
| AB-BP1 | | 0.18 | 0.17 | 0.10 | 0.80 | 0.07 | 0.06 | 0.05 | | Arth & Barker, 1976 |
| BO-PLA | 0.18 | 0.12 | 0.09 | 0.06 | 0.75 | 0.15 | 0.10 | 0.10 | | Bacon & Oruitt, 1988 |
| NS-6P | 0.70 | 0.07 | 0.12 | 0.08 | 0.96 | 0.40 | 0.30 | 0.29 | | Nagasawa & Schnetzler, 1971. |
| NC-12 | 0.39 | 0.27 | 0.18 | 0.16 | 3.40 | 0.10 | 0.07 | 0.07 | 41 | Nash & Crecraft, 1985 |
| NC-6 | 0.30 | 0.20 | 0.20 | 0.10 | 3.30 | 0.11 | 0.00 | 0.05 | 44 | Nash & Crecraft, 1985 |
| RÚ-PIR | 0.30 | 0.27 | 0.24 | 0.17 | 2 00 | 0.15 | 0.10 | 0.09 | 23 | Racon & Orwitt 1988 |
| NS-1P | 0.50 | 0.27 | 0.19 | 0.13 | 2.35 | 0.06 | 0.04 | 0.04 | | Nagasawa & Schnetzier 1971 |
| NS-2P | | 0.11 | 0.06 | 0.05 | 0.82 | 0.03 | 0.02 | 0.02 | | Nagasawa & Schnetzler, 1971 |
| 0-Tum | 0.30 | 0.22 | 0.19 | 0.17 | 4.20 | 0.21 | 0.11 | 0.11 | 23 | Oudas et al., 1971 |
| NS-8P | | 0.35 | 0.24 | 0.15 | 2.81 | 0.08 | 0.07 | 0.07 | | Nagasawa & Schnetzler, 1971 |
| NS-4P | | 0.28 | 0.29 | 0.14 | 2.49 | 0.08 | 0.56 | 0.54 | | Nagasawa & Schnetzler, 1971 |
| | | | | | | CL 1 | | WENE | | |
| Sample | La | Ce | Nd | S | Eu | Tb | Yb | Lu | Nol Er | Reference |
| XI | 0.15 | 0.20 | 0.30 | 0.50 | 0.40 | 0.65 | 0.60 | 0.60 | | Irving, 1978 |
| PS-193 | | 0.10 | 0.18 | 0.26 | 0.26 | 0.30 | 0.23 | | | Philpots & Schnitzler, 1970 |
| SP-25 | | 0.18 | 0.32 | 0.43 | 0.48 | 0.57 | 0.48 | | | Schnitzler & Philpots, 1968 |
| SP-19 | 0.00 | 0.12 | 0.26 | 0.38 | 0.39 | 0.48 | 0.43 | | | Schnitzler & Philpots, 1968 |
| 0n-b | 0.08 | 0.07 | 0.38 | 0.74 | 0.75 | 0.97 | 1.01 | 0.95 | | Onouma et al., 1968 |
| 0-340 PS-20/ | 0.14 | 0.10 | 0.07 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | | Dellasta et al., 1985 |
| LC-11 | 0.14 | 0.20 | 0.43 | 0.78 | 0.72 | 1 11 | 0.07 | 0.07 | 47 | Lubr & Carmichael 1980 |
| PS-271 | | 0.51 | 0.65 | 0.95 | 0.68 | 1.40 | 1.30 | 0.00 | 4. | Philpots & Schnitzler, 1970 |
| LC-6 | 0.14 | 0.24 | 0.64 | 0.96 | 1.07 | 1.48 | 0.97 | 0.95 | 44 | Luhr & Carmichael, 1980 |
| LC-7 | 0.16 | 0.36 | 0.68 | 1.26 | 1.19 | 1.66 | 1.33 | 1.38 | 44 | Luhr & Carmichael, 1980 |
| BO-CpxA | 0.28 | 0.47 | 0.86 | 1.60 | 1.10 | 2.70 | 2.00 | 2.00 | | Bacon & Oruitt, 1988 |
| NS-6Au | 0.7/ | 0.36 | 0.94 | 1.50 | 1.10 | 2.60 | 2.01 | 1.81 | | Nagasawa & Schnetzler, 1971. |
| NC-15 | 0.76 | 1.20 | 2.10 | 3.30 | 2.20 | 4.40 | 3.50 | 3.30 | | Nash & Crecraft, 1985 |
| NC-6 | 1 50 | 2 50 | 4 50 | 6 00 | 2.20 | 10 3 | 0 50 | 8 80 | | Nash & Crecraft 1985 |
| PS-266 | 1.50 | 0.65 | 1.28 | 1.81 | 2.01 | 1.30 | 1.14 | 1.28 | | Philpots & Schnitzler 1970 |
| B0-CpxR | 0.52 | 0.84 | 1.40 | 2.90 | 2.50 | 3.80 | 3.00 | 3.00 | | Bacon & Oruitt, 1988 |
| | | | | | | | | | | |
| Sample | La | Ce | Nd | S | Eu | Th | Yb | JXENE Lu | Nol F | Reference |
| XI | 0.005 | 0.010 | | 0.02 | 0.02 | 0.03 | 0.15 | | THOL LI | Irving, 1978 |
| On-c | | 0.00 | | 0.01 | 0.02 | 0.05 | 0.11 | 0.11 | | Onouma et al., 1968 |
| PS-203 | | 0.02 | 0.05 | 0.10 | 0.07 | 0.19 | 0.40 | 0.45 | | Philpots & Schnitzler, 1970 |
| PS-271 | | 0.04 | 0.06 | 0.10 | 0.08 | 0.23 | 0.67 | 0.84 | | Philpots & Schnitzler, 1970 |
| PS-184 | 0.70 | 0.03 | 0.03 | 0.05 | 0.05 | 0.09 | 0.24 | 0.40 | - | Philpots & Schnitzler, 1970 |
| LU-D | 0.50 | 0.33 | 0.45 | 0.43 | 0.42 | 0.48 | 0.56 | 0.68 | 7 | Luhr & Carmichael, 1980 |
| NS-6H | 0.20 | 0.31 | 0.47 | 0.10 | 0.34 | 0.09 | 0.77 | 0.71 | | Nadasawa & Schootzlan 1071 |
| NC-15 | 0.74 | 0.87 | 1,20 | 1.40 | 0.54 | 1.70 | 2.10 | 2 10 | | Nash & Crecraft 1085 |
| NS-3H | | 0.26 | 0.34 | 0.38 | 0.27 | 0.55 | 0.99 | 0.92 | | Nagasawa & Schnetzler, 1971. |
| NC-12 | 0.65 | 0.79 | 1.10 | 1.40 | 0.56 | 1.60 | 1.90 | 1.90 | | Nash & Crecraft, 1985 |
| B0-OpxR | 0.40 | 0.46 | 0.60 | 0.78 | 0.61 | 0.85 | 0.91 | 0.90 | | Bacon & Oruitt, 1988 |
| NS-1H | | 0.17 | 0.26 | 0.34 | 0.21 | 0.51 | 0.78 | 0.78 | | Nagasawa & Schnetzler, 1971 |
| NS-4H | | 0.10 | 0.15 | 0.21 | 0.09 | 0.30 | 0.93 | 1.14 | | Nagasawa & Schnetzler, 1971 |

| | | | | | | 11 | ORNBLE | NDE |
|--------|------|------|------|------|------|------|--------|------|
| Sample | La | Се | Nd | S | Eu | Tb | Yb | Lu |
| HN-3 | 0.17 | 0.34 | | 0.93 | 1.11 | 1.02 | 0.98 | 0.82 |
| SP-29 | | 0.12 | 0.16 | 0.24 | 0.26 | 0.30 | 0.23 | |
| D34-b | 0.25 | 0.32 | | 1.40 | 1.20 | 1.30 | 1.20 | 1.10 |
| PS-233 | | 0.09 | 0.19 | 0.34 | 0.36 | 0.58 | 0.46 | 0.44 |
| LC-6 | 0.14 | 0.33 | 0.79 | 1.31 | 1.66 | 1.93 | 1.18 | 1.07 |
| LC-7 | 0.19 | 0.32 | 0.73 | 1.51 | 1.79 | 1.91 | 1.43 | 1.41 |
| AB-BP1 | | 0.90 | 2.80 | 3.99 | 3.44 | 5.80 | 4.90 | 4.51 |
| BO-HbA | 0.31 | 0.53 | 1.20 | 2.00 | 1.90 | 2.00 | 2.10 | 2.00 |
| NS-3Hb | | 1.38 | 4.03 | 7.10 | 5.00 | 12.5 | 9.0 | 6.3 |
| BO-HbR | 0.36 | 0.68 | 1.60 | 2.30 | 3.20 | 2.40 | 1.80 | 1.80 |
| NS-1Hb | | 1.42 | 4.25 | 8.20 | 5.90 | 13.5 | 8.7 | 5.9 |
| NS-2Hb | | 0.43 | 1.06 | 1.61 | 1.39 | 2.3 | 1.9 | 1.8 |
| NS-8Hb | | 1.77 | 4.49 | 8.00 | 4.53 | 13.0 | 7.5 | |

| | | | | | | | | OLIVI | IE . | |
|----|---------|-------|-------|-------|-------|-------|-------|-------|-------|--|
| \$ | Sample | La | Ce | Nd | Sm | Eu | Tb | Yb | Lu | |
| L | .s-48 | | | | 0.009 | | | | | |
| S | SP-25 | | 0.009 | 0.007 | 0.006 | 0.006 | 0.009 | | | |
| L | S-15 | | | | 0.009 | | | | | |
| L | S-36 | | | | 0.003 | | | | | |
| L | S-44 | | | | 0.009 | | | | | |
| L | .s-82 | | | | 0.009 | | | | | |
| L | .s-5 | | | | 0.015 | | | | | |
| F | PS-19 | | 0.009 | 0.010 | 0.011 | 0.010 | 0.013 | 0.023 | | |
| ŀ | 1N-4 | 0.003 | | | 0.003 | 0.005 | 0.005 | 0.009 | 0.014 | |
| ŀ | 1-688MS | 0.008 | | | 0.012 | 0.008 | | 0.070 | 0.300 | |
| L | .C-11 | | | | 0.010 | | | 0.030 | | |

| | | | | | | 1 | AGNET | ITE |
|---------|------|------|------|------|------|------|-------|------|
| Sample | La | Се | Nd | S | Eu | Tb | Yb | Lu |
| LC-11 | 0.45 | 0.42 | 0.43 | 0.45 | 0.41 | 0.48 | 0.47 | 0.44 |
| LC-6 | 0.37 | 0.35 | 0.55 | 0.53 | 0.42 | 0.67 | 0.29 | 0.32 |
| LC-7 | 0.38 | 0.31 | 0.42 | 0.54 | 0.35 | 0.53 | 0.38 | 0.41 |
| NC-15 | 0.89 | 1.10 | 1.60 | 1.90 | 0.25 | 1.90 | 1.10 | 1.10 |
| NC-12 | 0.74 | 0.99 | 1.30 | 1.70 | 0.32 | 1.70 | 0.89 | 0.84 |
| NC-6 | 1.70 | 2.40 | 3.00 | 3.50 | 1.20 | 3.50 | 1.60 | 1.20 |
| B0-MgtR | 0.66 | 0.71 | 0.93 | 1.20 | 0.91 | 1.30 | 0.11 | 0.30 |

| | | | | | | | I LMEN I | TE |
|--------|------|------|------|------|------|------|----------|------|
| Sample | La | Се | Nd | Sm | Eu | Tb | Yb | Lu |
| NC-6 | 6.20 | 6.80 | 6.30 | 5.50 | 1.00 | 4.80 | 3.10 | 2.10 |

| | APATITE | | | | | | | | |
|--------|---------|------|-------|-------|------|------|------|------|-----------------------------|
| Sample | La | Се | Nd | S | Eu | Tb | Yb | Lu | Reference |
| XI | 8.00 | 9.00 | 11.00 | 11.00 | 10.0 | 8.00 | 4.00 | 3.50 | Irving, 1978 |
| E-1 | 14.5 | 21.1 | 32.8 | 46.0 | 25.5 | 38.0 | 15.4 | 13.8 | Fujimaki, 1986 |
| N-6A | | 18.0 | 27.4 | 29.3 | 20.5 | 25.0 | 13.1 | 11.2 | Nagasawa, 1970 |
| NS-2A | | 16.6 | 21.0 | 20.7 | 14.5 | 18.0 | 9.4 | 7.9 | Nagasawa & Schnetzler, 1971 |
| N-1A | | 52.5 | 81.1 | 89.8 | 50.2 | 75.0 | 37.0 | 30.0 | Nagasawa, 1970 |
| N-4A | | 33.6 | 61.3 | 69.6 | 23.8 | 60.0 | 23.2 | 20.1 | Nagasawa, 1970 |

| | | | | | | | ZIRCO | M | | |
|--------|------|------|------|------|------|----|-------|-----|-------------|------|
| Sample | La | Ce | Nd | Sm | Eu | Tb | Yb | Lu | Reference | |
| F-1 | 3.11 | 3.49 | 3.80 | 4.72 | 4.23 | 35 | 254 | 333 | Fujimaki, 1 | 1986 |
| N-1A | | 3.00 | 2.43 | 3.70 | 5.22 | 40 | 299 | 366 | Nagasawa, ' | 1970 |
| N-4A | | 2.29 | 1.97 | 2.58 | 1.07 | 20 | 242 | 281 | Nagasawa, 1 | 1970 |

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